



the Extended Baryon Oscillation Spectroscopic Survey

Kyle Dawson University of Utah

on Behalf of the eBOSS Collaboration

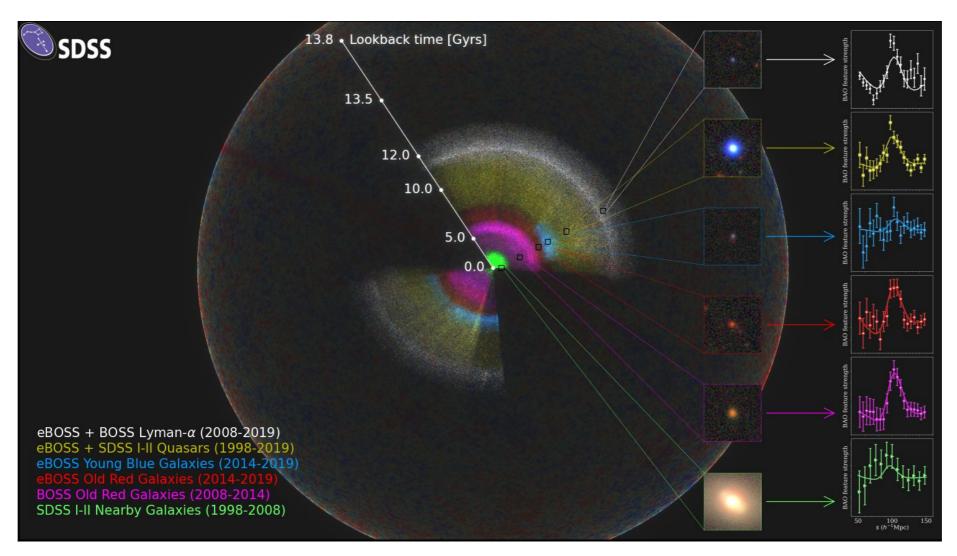






Outline

- Background and Survey Overview (completed March 1, 2019!)
- Measurements of BAO and RSD
- Cosmology Interpretation

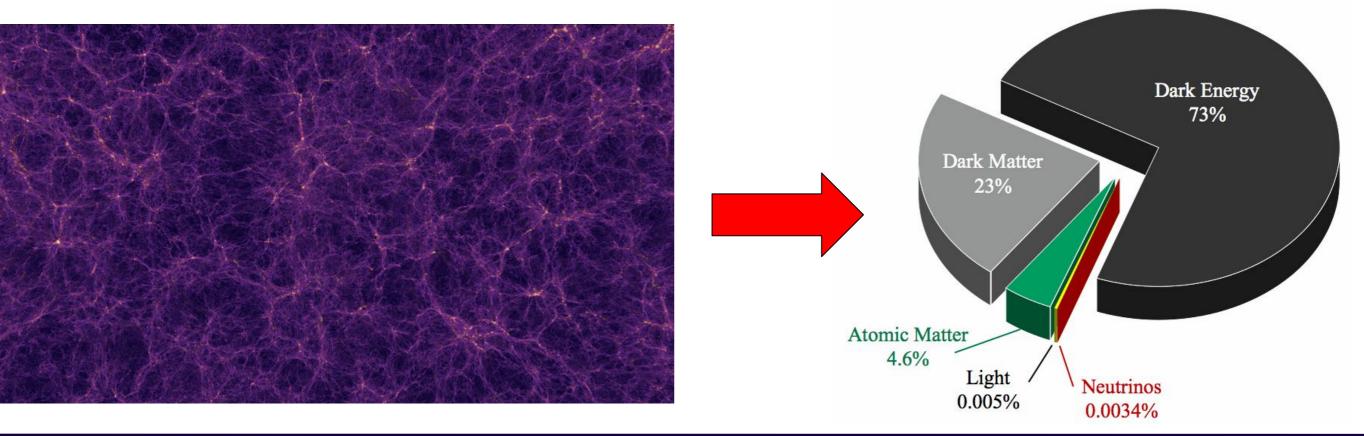




Cosmological Background



- Evolving distribution of matter in Universe
 - Cosmic expansion and growth of structure
- Derived Measurements: H(z), $D_M(z)$, $f\sigma_8(z)$
 - Physics of dark energy
 - Composition of the Universe
 - Neutrino mass, Inflation, Laws of gravity





- Friedmann Equation: $H^2(a) = \frac{8\pi G}{3}\rho(a) \frac{kc^2}{a^2}$
 - Energy Components: dark (dm), baryonic (b), and total matter (m), neutrinos, photons, dark energy, and curvature

$$\begin{split} \Omega_x &= \frac{\rho_x}{\rho_{\rm crit}} = \frac{8\pi G}{3H^2} \rho_x \quad \Omega_k(a) = -\frac{kc^2}{a^2 H^2(a)} \\ \text{Dark Energy equation of state} \\ w(a) &= \begin{cases} -1 \\ w \\ w_0 + w_a(1-a) \end{cases} \\ \frac{\rho_{\rm DE}(a)}{\rho_{\rm DE,0}} &= \begin{cases} 1 \\ a^{-3(1+w)} \\ a^{-3(1+w_0+w_a)} \exp[-3w_a(1-a)] \end{cases} \end{split}$$



BAO measure angular diameter distance and H(z)

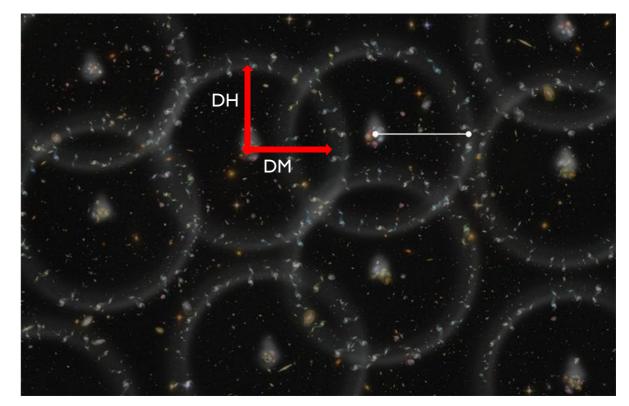
$$H(z) = c\Delta z/r_d$$

$$D_H(z) = \frac{c}{H(z)}$$

$$D_C(z) = \frac{c}{H_0} \int_0^z dz' \frac{H_0}{H(z')}$$

$$D_M(z) = \frac{c}{H_0} S_k \left(\frac{D_C(z)}{c/H_0}\right)$$

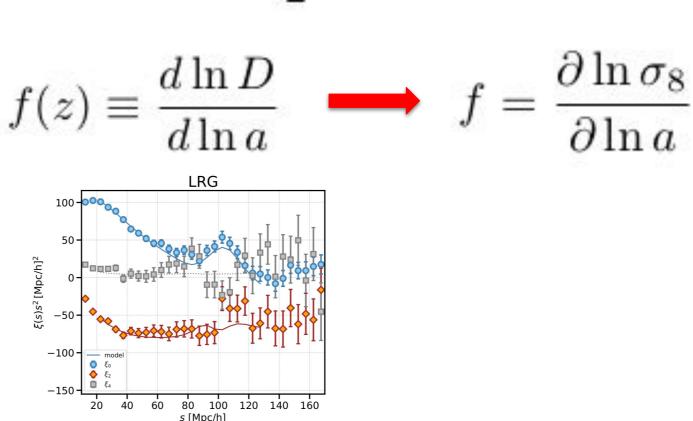
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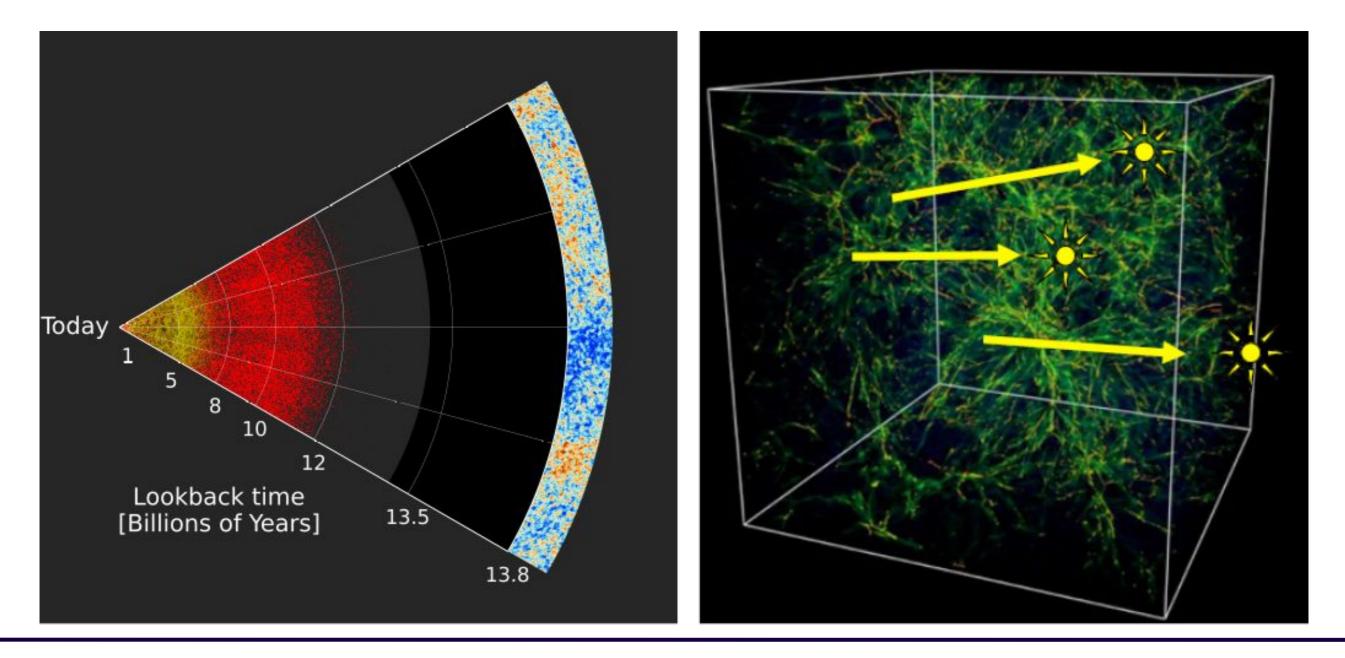
- Scale-independent growth factor: $\delta(\mathbf{x}, t) = D(t)\delta(\mathbf{x}, t_0)$
- Linear growth equation: $\ddot{D}+2H(z)\dot{D}-\frac{3}{2}\Omega_mH_0^2(1+z)^3D=0$
- Linear Growth Rate:







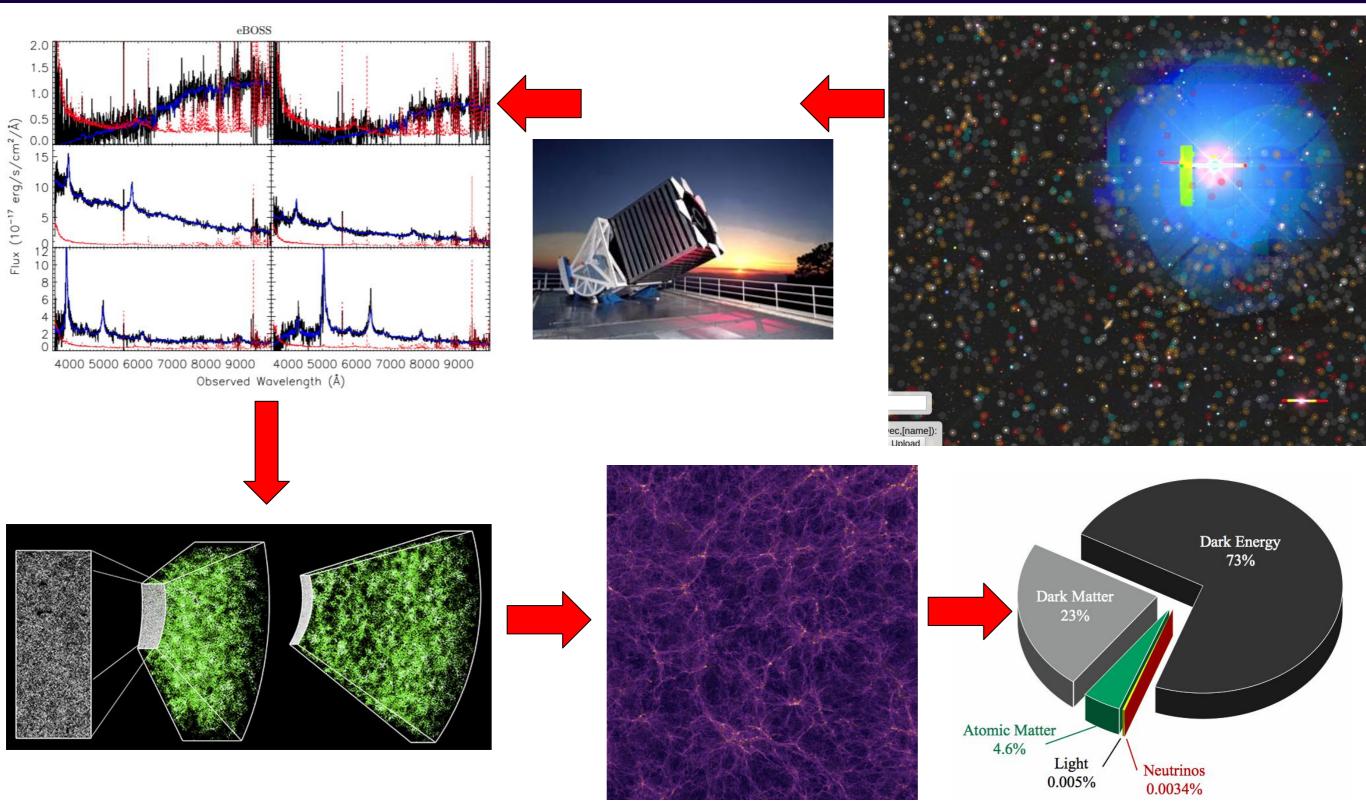
Direct tracers Galaxies and quasars (z<2.1) Absorption in quasar spectra by foreground Lyman-alpha forest (z>2.1)





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Cosmology with Spectroscopy





Survey Overview

BOSS: Dawson, Schlegel, et al., 2013, "The Baryon Oscillation Spectroscopic Survey of SDSS-III"

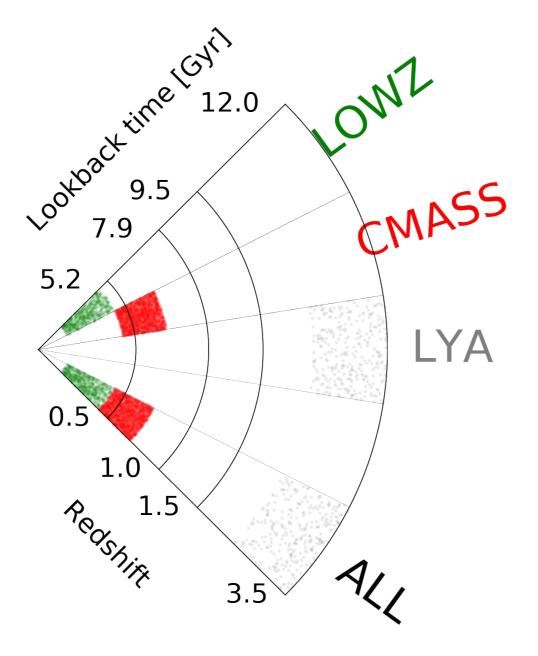
eBOSS: Dawson, Kneib, Percival, et al., 2016, "The SDSS-IV Extended Baryon Oscillation Spectroscopic Survey: Overview and Early Data"



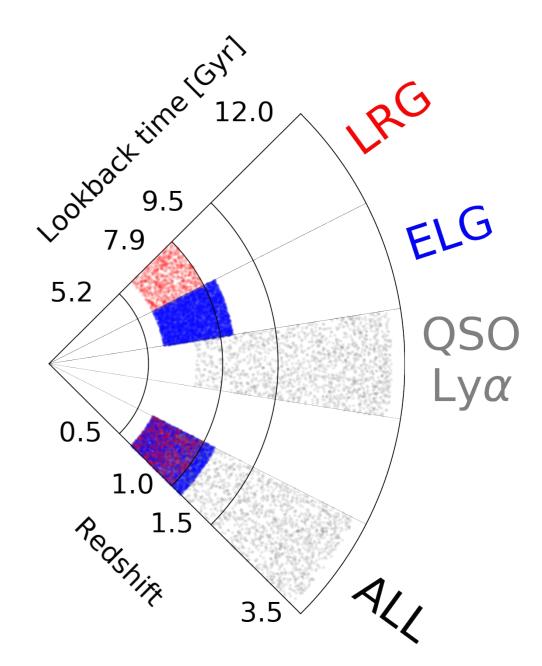
- Target selection: well-understood selection functions near imaging limits
- Mountain Operations: <1% downtime, near-optimal efficiency
- Uniformity: tuned spectroscopic exposure times to real-time data quality
- Data reduction: extractions of spectra to S/N<1 with negligible spurious signal
- **Redshift classification:** >90% efficiency, even for faintest targets
- Catalog creation: imaging/spectroscopic systematics sub-dominant



BOSS (2009-2014)



eBOSS (2014-2019)





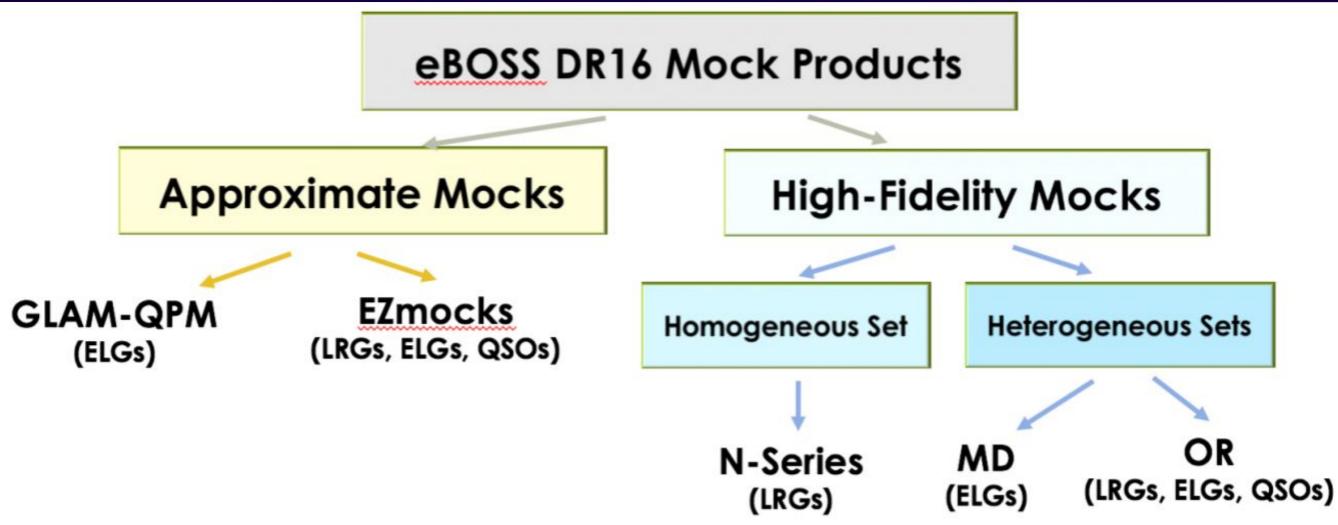
Local MP4



BAO and **RSD** Measurements

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Each synthetic product serves a different science target

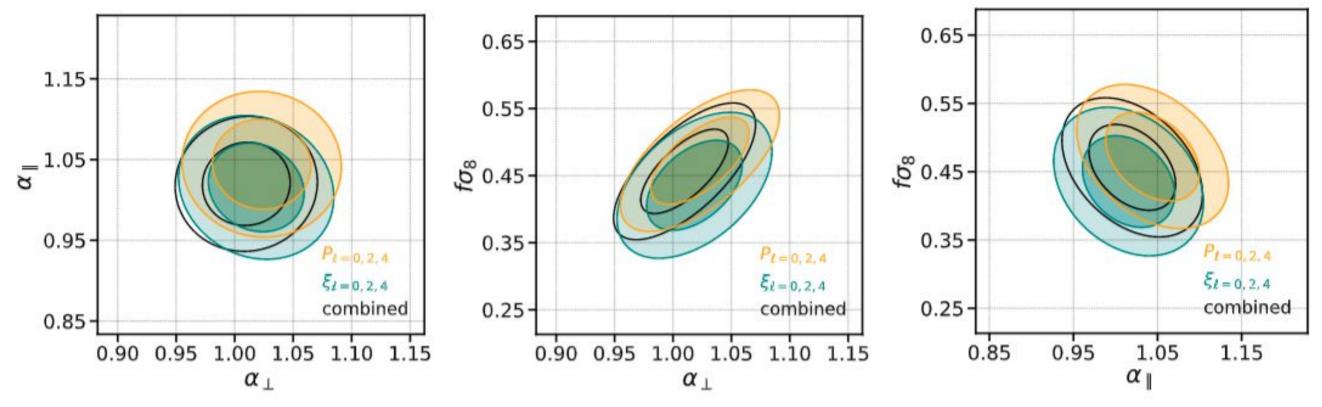
Approximate Mocks: covariances, observational systematics High-Fidelity Mocks: theoretical systematics, analysis pipeline validation, performance & accuracy, analysis biases



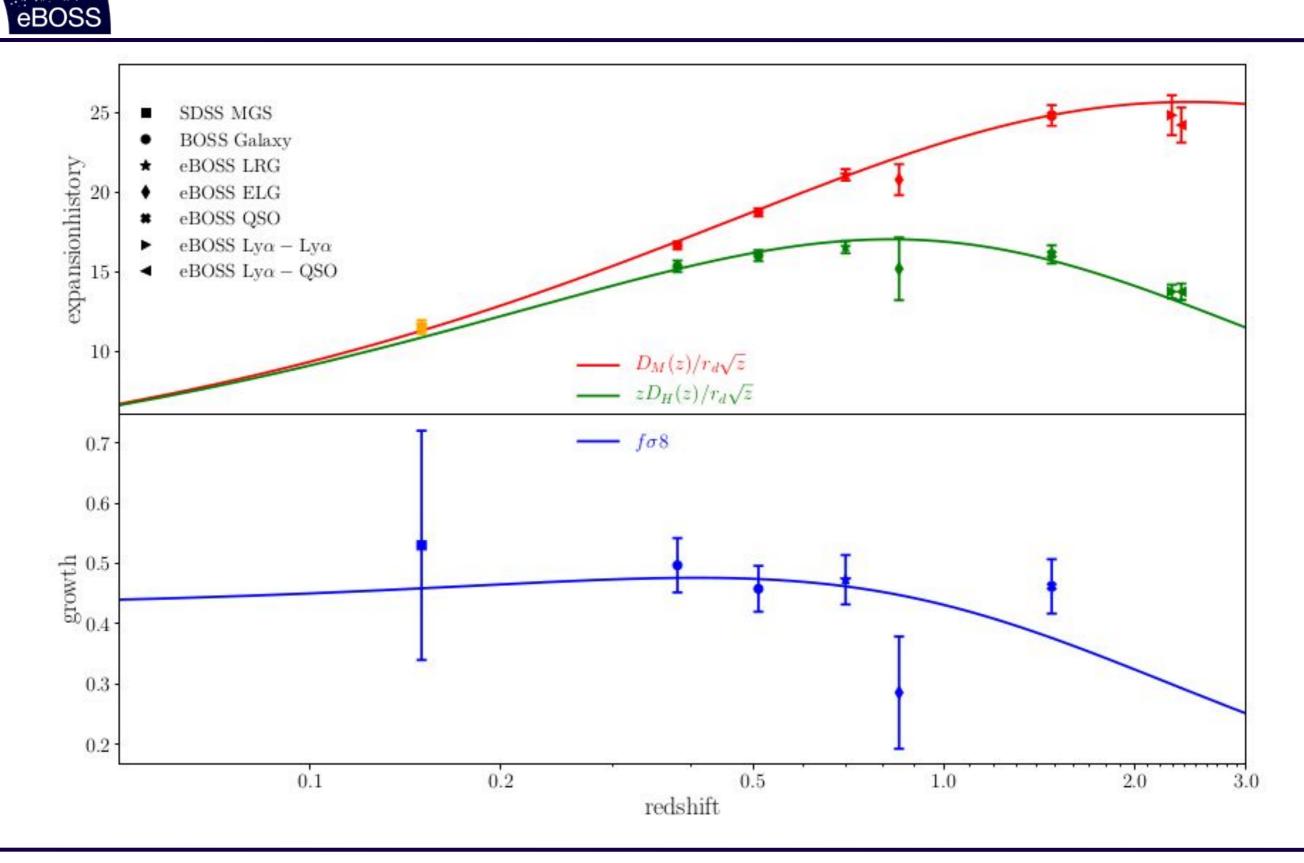
- Modeling, observational artifacts, and fiducial cosmology
- Larger affect on fsigma8 estimates than on BAO estimates
 - ~0.5 sigma on LRG and ELG measurements
 - ~0.3 sigma on quasar measurements
 - Added in quadrature to statistical errors
- No additional increase in Lyman-alpha BAO studies: systematics determined to be sub-dominant



- BAO measured w/reconstruction in LRG and ELG
- BAO/RSD measured in full shape with LRG/ELG/QSO
- Configuration Space and Fourier Space Measurements for each tracer
- Combine BAO/RSD results using mock-calibrated covariance matrices



All BAO and RSD Measurements

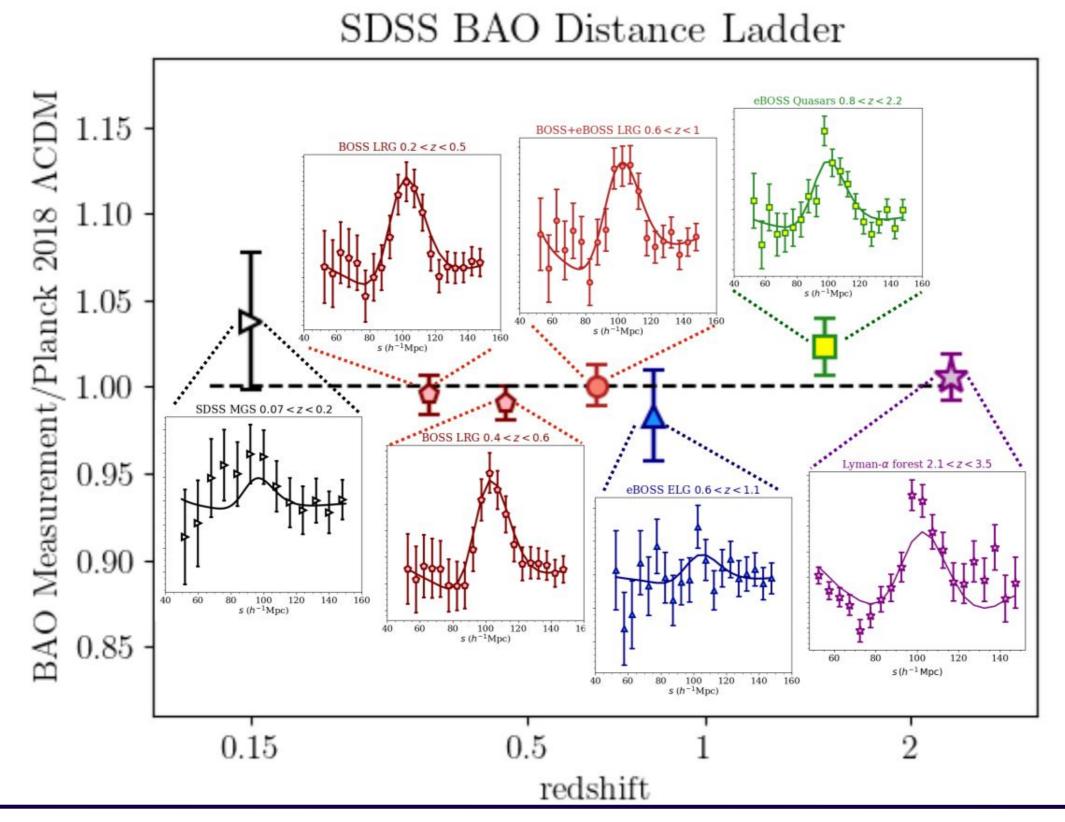


LBL Research Progress Meeting – July 9, 2020 18

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Final Results (Primary Science Drivers)





- Aggregate precision of the expansion history measurements is 0.70% at redshifts z < 1
- Aggregate precision of the expansion history measurements is 1.19% at redshifts z > 1
- Aggregate precision of the growth measurements (fsigma8) is
 4.78% over the redshift interval 0 < z < 1.5.

https://test.sdss.org/final-bao-and-rsd-measurements/



Cosmology Interpretation

eBOSS Collaboration, "The Completed SDSS-IV extended Baryon Oscillation Spectroscopic Survey: Cosmological Implications from two Decades of Spectroscopic Surveys at the Apache Point observatory", to appear on arXiv July 20, 2020

Special thanks to Eva-Maria Mueller, Andreu Font-Ribera, Anze Slosar, and Zheng Zheng











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Extensions to LCDM

- -	Parameter	Definition			
	Ω_m	density parameter of matter			
	Ω_c	density parameter of cold dark matter			
	Ω_b	density parameter of baryons			
	Ω_{Λ}	density parameter of cosmological constant			
	$\Omega_{\rm DE}$	density parameter of dark energy			
0	Ω_k	curvature parameter			
	$\omega_c = \Omega_c h^2$	physical density parameter of cold dark matter			
	$\omega_b = \Omega_b h^2$	physical density parameter of baryons			
	H_0	current expansion rate (Hubble constant)			
	h	$H_0/100 \rm km s^{-1} Mpc^{-1}$			
	θ_{MC}	approximate angular scale of sound horizon (CosmoMC)			
	A_s	power of the primordial curvature perturbations at $k = 0.05 \text{Mpc}^{-1}$			
	σ_8	amplitude of matter fluctuation on $8h^{-1}$ Mpc comoving scale			
	n_s	power-law index of the scalar spectrum			
	$ au^{-}$	Thomson scattering optical depth due to reionization			
	$N_{\rm eff}$	effective number of neutrino-like relativistic degrees of freedom			
w	$w(w_0)$	dark energy equation of state, $w = p_{\rm DE}/\rho_{\rm DE}$ ($c = 1$ units)			
w _a	wa	time derivative of dark energy equation of state parameter (eq.6)			
nu	$\sum m_{\nu}$	sum of neutrino masses			

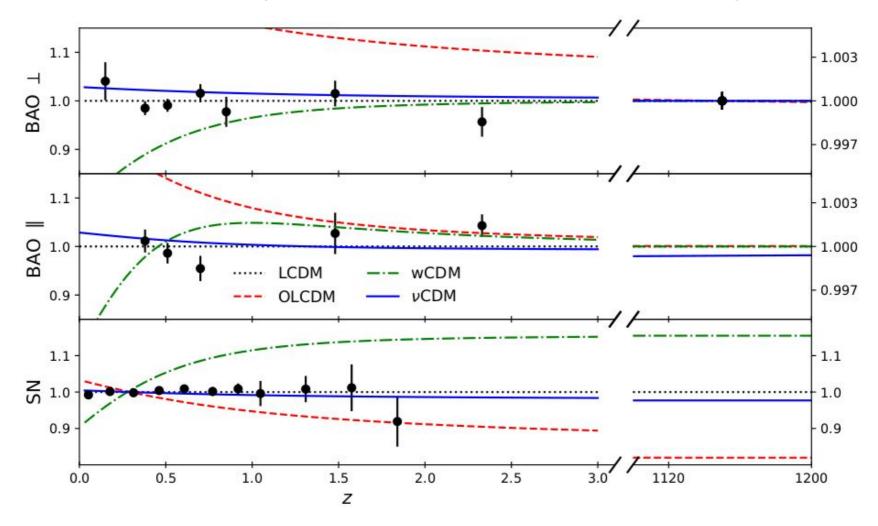


Legacy of BOSS/eBOSS

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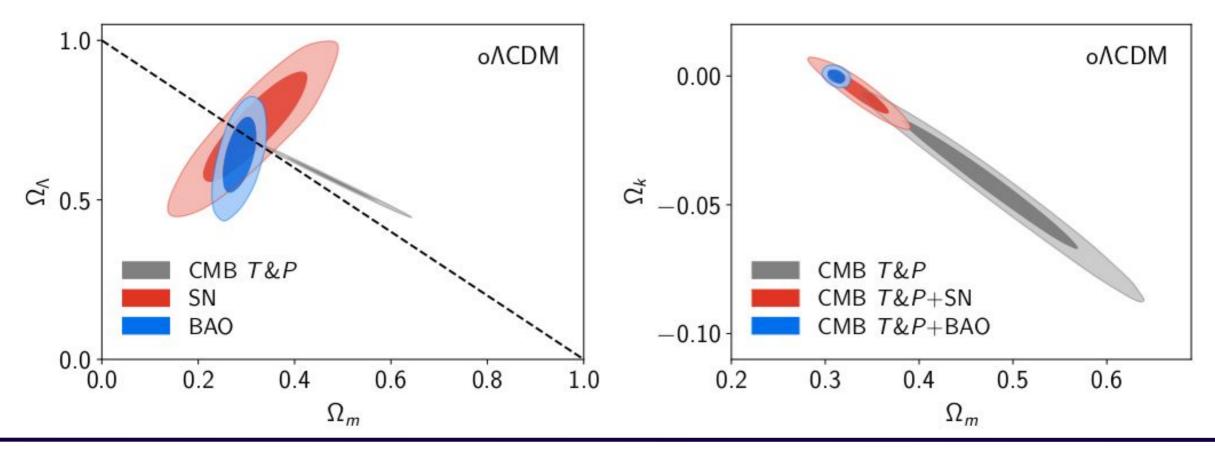
- BAO-only from SDSS/BOSS/eBOSS
- Pantheon sample of SNe (Scolnic et al, 2018)
- Planck 2018 results (Planck Collaboration, 2018)





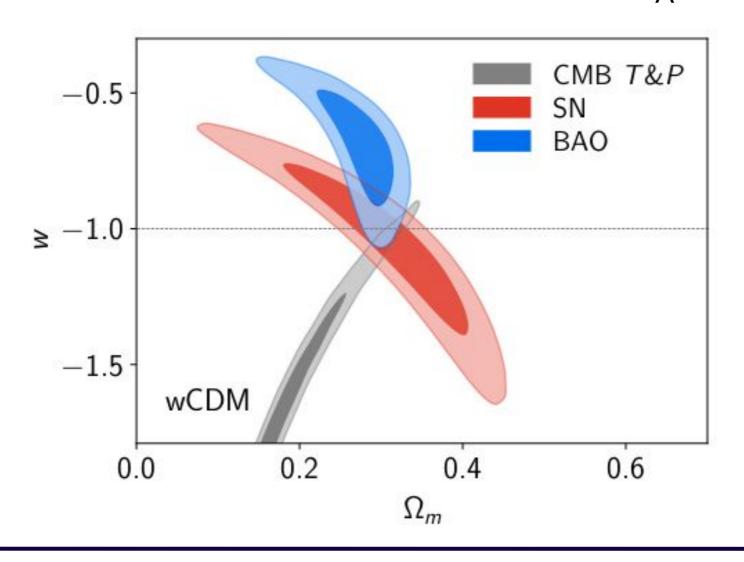
oLCDM Cosmology

- BAO measurements alone \rightarrow constraint on the dark energy density with a ~8 sigma confidence detection.
- Combined with Planck temperature and polarization data \rightarrow order of magnitude improvement on curvature constraints
- Strong evidence for a nearly flat geometry \rightarrow roughly one order of magnitude within the fundamental limit set by cosmic variance.





- BAO-only offer tighter constraints than SNe-only
- Degeneracies well-aligned for SNe+CMB
 - SNe+CMB vs BAO+CMB: I.2X better in Ω_{Λ} ; I.5X in w





One Parameter Extensions

		Ω_{DE}	$H_0[{\rm km/s/Mpc}]$	Ω_k	w	$\Sigma m_{\nu} [\mathrm{eV}]$			
	CMB T&P	0.6836 ± 0.0084	67.29 ± 0.61	—		—			
ACDM	CMB T&P + BAO	0.6881 ± 0.0059	67.61 ± 0.44	-					
ACDM	CMB T&P + SN	0.6856 ± 0.0078	67.43 ± 0.57	_	<u> </u>	_			
	CMB T&P + BAO + SN	0.6891 ± 0.0057	67.68 ± 0.42	-	-	-			
8	CMB T&P	$0.561^{+0.050}_{-0.041}$	$54.5^{+3.3}_{-3.9}$	$-0.044^{+0.019}_{-0.014}$					
oΛCDM	CMB T&P + BAO	0.6882 ± 0.0060	67.59 ± 0.61	-0.0001 ± 0.0018		-			
ONODW	CMB T&P + SN	0.670 ± 0.017	65.2 ± 2.2	$-0.0061^{+0.0062}_{-0.0054}$		-			
	CMB T&P + BAO + SN	0.6891 ± 0.0057	67.67 ± 0.60	-0.0001 ± 0.0018					
	CMB T&P	$0.801^{+0.057}_{-0.022}$	> 82.3	_	$-1.58^{+0.16}_{-0.35}$	—			
wCDM	CMB T&P + BAO	0.694 ± 0.012	$68.4^{+1.4}_{-1.5}$	-	$-1.034_{-0.053}^{+0.061}$	—			
webm	CMB T&P + SN	0.692 ± 0.010	68.3 ± 1.1	—	-1.035 ± 0.037	-			
	CMB T&P + BAO + SN	0.6929 ± 0.0075	68.21 ± 0.82	—	-1.026 ± 0.033	_			
	CMB T&P	$0.680^{+0.016}_{-0.0087}$	$\begin{array}{r} 67.0^{+1.2}_{-0.67} \\ 67.70^{+0.53}_{-0.48} \end{array}$	_	<u> </u>	< 0.268			
$\nu \Lambda \text{CDM}$	CMB T&P + BAO	$0.6890^{+0.0069}_{-0.0061}$	$67.70_{-0.48}^{+0.53}$			< 0.134			
PRODM	CMB T&P + SN	$0.686^{+0.011}_{-0.0083}$	$67.47_{-0.65}^{+0.83}$	-	—	< 0.174			
	CMB T&P + BAO + SN	0.6898 ± 0.0061	67.76 ± 0.47	—		< 0.125			
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- BAO in isolation \rightarrow relative measures of expansion history
- BAO in combination with early Universe physics \rightarrow calibrated ruler
- Calibrated ruler \rightarrow absolute expansion rates, including current expansion rate, H₀
 - Standard physics (e.g. 3 neutrino species) \rightarrow baryon density, matter density, and CMB temperature offer calibration of standard ruler
 - CMB or Big Bang Nucleosynthesis can offer baryon density
 - Matter density from CMB or distance probes

$$H(z) = c\Delta z/r_d$$

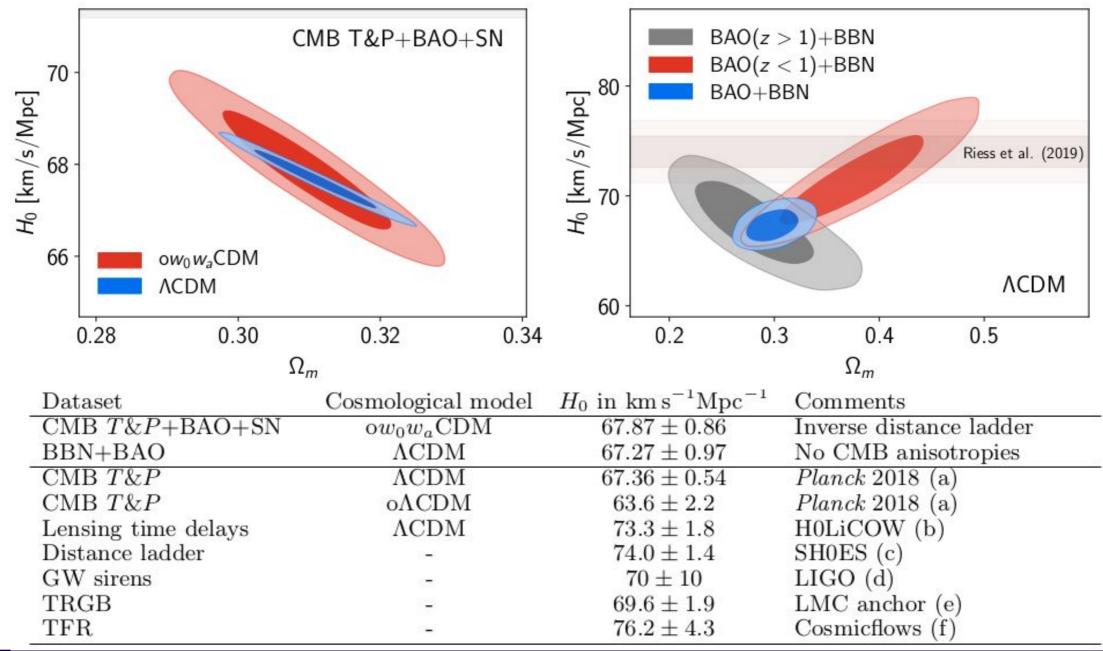
$$D_H(z) = \frac{c}{H(z)}$$

$$r_d = \int_{z_d}^{\infty} \frac{c_s(z)}{H(z)} dz$$



New Physics in Expansion History?

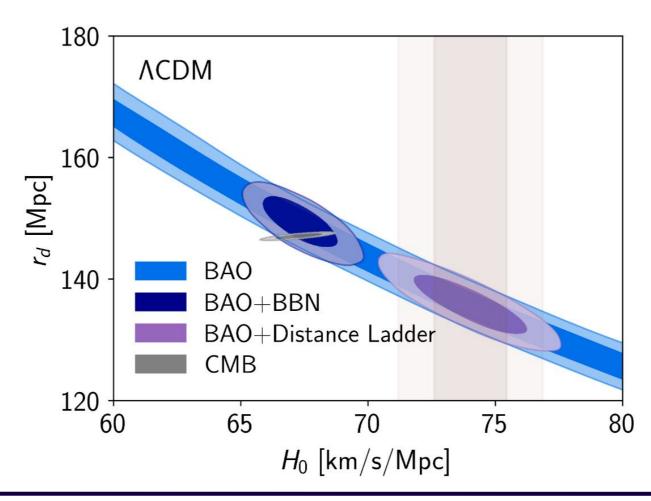
- BAO \rightarrow insensitive to the strict cosmological priors in CMB-only estimates.
- BAO \rightarrow insensitive to CMB anisotropies if using LCDM and BBN



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- H0 from BAO $\rightarrow \sim 10\%$ smaller than those from the Cepheid distance ladder and strong-lensing time delays.
- `H0 tension' can not be restricted to systematic errors in Planck or to the strict assumptions of the LCDM model \rightarrow new physics?

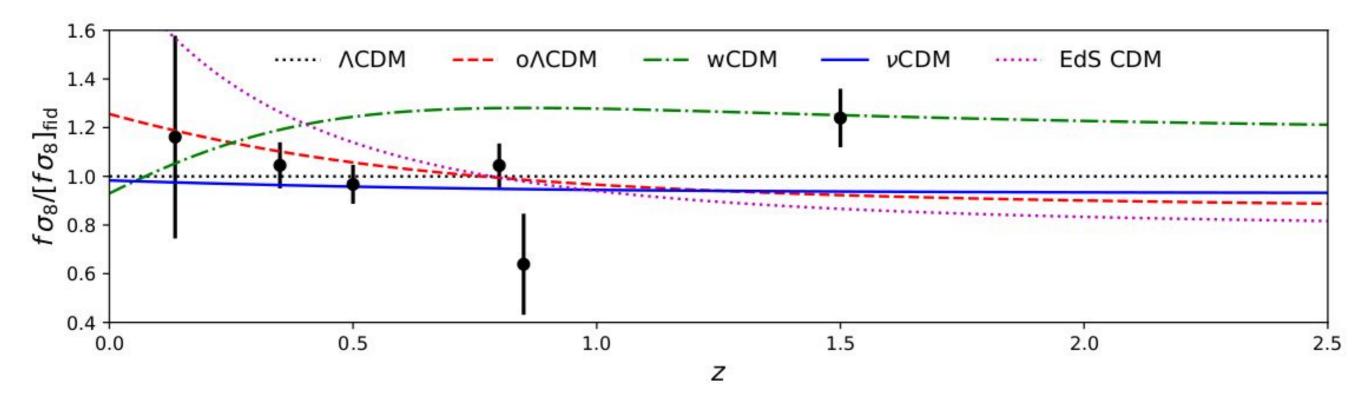




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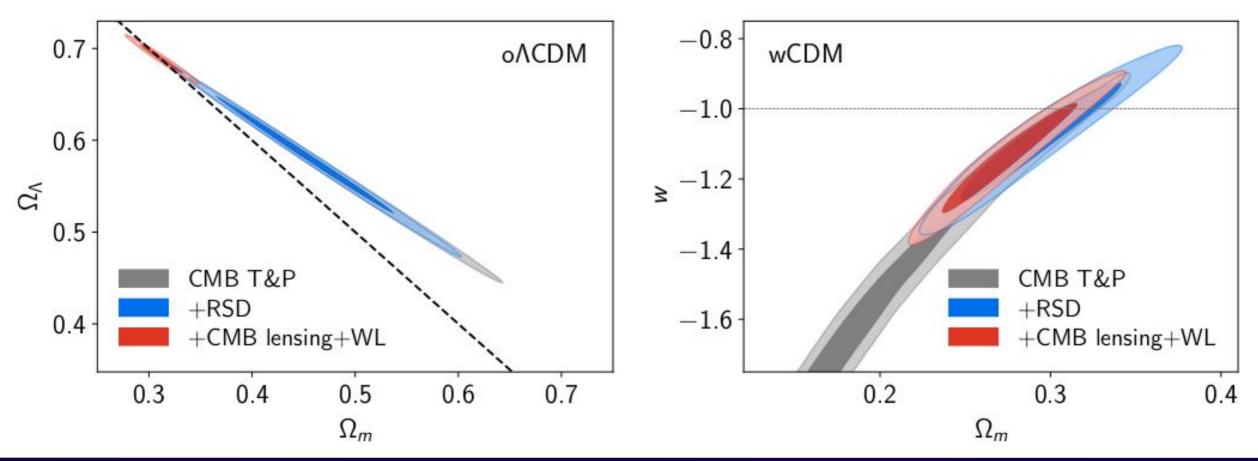


- RSD-only from SDSS/BOSS/eBOSS
- Weak Lensing (WL) from DES (Troxel et al, 2018) and Planck
- Planck temperature and polarization (Planck Collaboration, 2018)



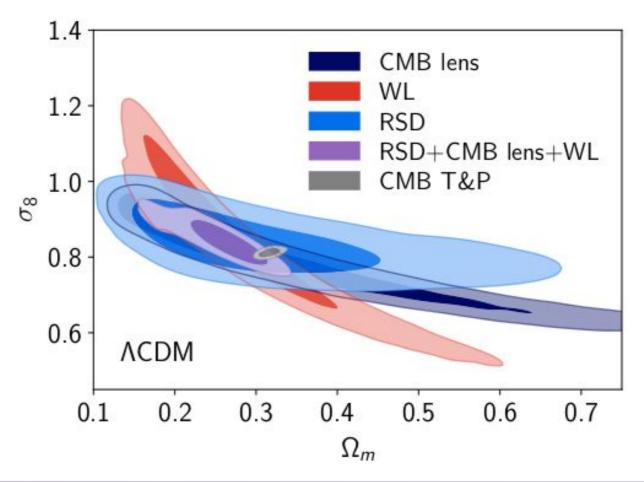


- Growth measurements \rightarrow factor two to three improvements in precision compared to CMB temperature and polarization data alone.
- Weak lensing data instill a preference for a flat geometry
- RSD instill a preference for a cosmological constant.





- Analogy to H0 test: local fluctuation amplitude vs CMB fluctuation amplitude
- RSD + WL \rightarrow current amplitude of matter fluctuations.
- LCDM predictions and GR \rightarrow consistent picture of structure growth





- Differential measurements of growth to test modified gravity
- Allow linear perturbations to Poisson equation
 - RSD measurements probe the gravitational response of matter
 - WL measurements probe that of photons

$$\begin{split} k^{2}\Psi &= -4\pi Ga^{2}(1+\mu(a))\rho\delta \\ k^{2}(\Psi+\Phi) &= -8\pi Ga^{2}(1+\Sigma(a))\rho\delta \\ \mu(z) &= \mu_{0}\frac{\Omega_{\Lambda}(z)}{\Omega_{\Lambda}}, \quad \Sigma(z) &= \Sigma_{0}\frac{\Omega_{\Lambda}(z)}{\Omega_{\Lambda}} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ 0.0 \\ g \end{array} \\ & \begin{array}{c} 0.5 \\ g \end{array} \\ & \begin{array}{c} 0.0 \\ g \end{array} \\ \\ & \begin{array}{c} 0.0 \\ g \end{array} \\ \\ \\ \end{array}$$



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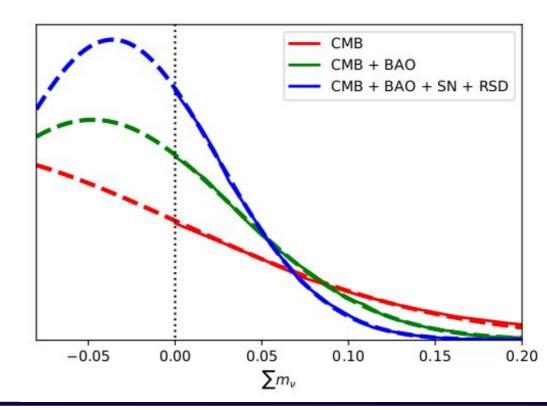
- BAO + RSD from SDSS/BOSS/eBOSS
- Pantheon sample of SNe (Scolnic et al, 2018)
- Weak Lensing (WL), galaxy-galaxy lensing, clustering from DES (Abbott et al, 2018)
- Planck temperature, polarization, and lensing (Planck Collaboration, 2018)



- Planck + Pantheon SNe Ia + DES + BAO + RSD data \rightarrow tightest constraints to date
 - Uncertainty $\rightarrow \sim$ lower bound of 60 meV allowed by neutrino oscillations
 - Relative to Planck-only \rightarrow the largest improvement in precision comes from the addition of the SDSS BAO measurements
 - RSD improve the precision by another 30%.

 $\sum m_{\nu} > 0.0588 \,\mathrm{eV}$ $\sum m_{\nu} > 0.0995 \,\mathrm{eV}$ inverted hierarchy.

normal hierarchy,





Neutrino Mass

Constraints on neutrino masses and relative probabilities of neutrino models.

Model	95% upper limit [eV]	$P_{\rm inv}/P_{\rm norm}$	$P_{\rm unphy}$	RMS of Gaussian fit [eV]
Planck	0.264	0.64	0.45	-0.144 ± 0.148
Planck + BAO	0.131	0.37	0.65	-0.048 ± 0.081
Planck + BAO + RSD	0.100	0.22	0.77	-0.037 ± 0.062
Planck + BAO + RSD (wACDM)	0.127	0.36	0.70	-0.150 ± 0.104
Planck + SN	0.175	0.50	0.57	-0.110 ± 0.114
Planck + BAO + SN	0.126	0.34	0.67	-0.041 ± 0.076
Planck + BAO + RSD + SN	0.099	0.22	0.78	-0.036 ± 0.060
Planck + BAO + RSD + SN (wACDM)	0.120	0.32	0.71	-0.084 ± 0.083
Planck + BAO + RSD + DES	0.132	0.37	0.64	-0.039 ± 0.079
Planck + BAO + RSD + DES (wACDM)	0.172	0.50	0.58	-0.181 ± 0.134
Planck + BAO + RSD + SN + DES	0.114	0.29	0.72	-0.036 ± 0.068
Planck + BAO + RSD + SN + DES (wACDM)	0.140	0.40	0.66	-0.100 ± 0.096

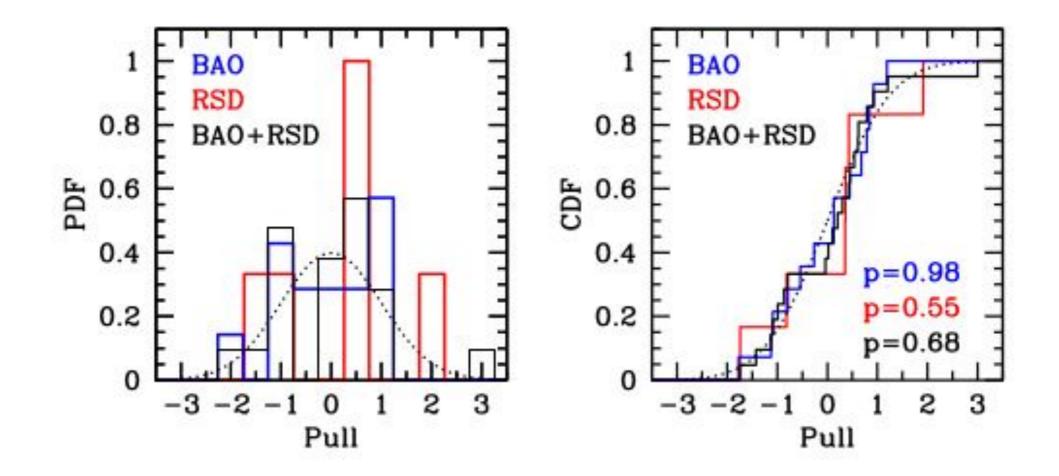


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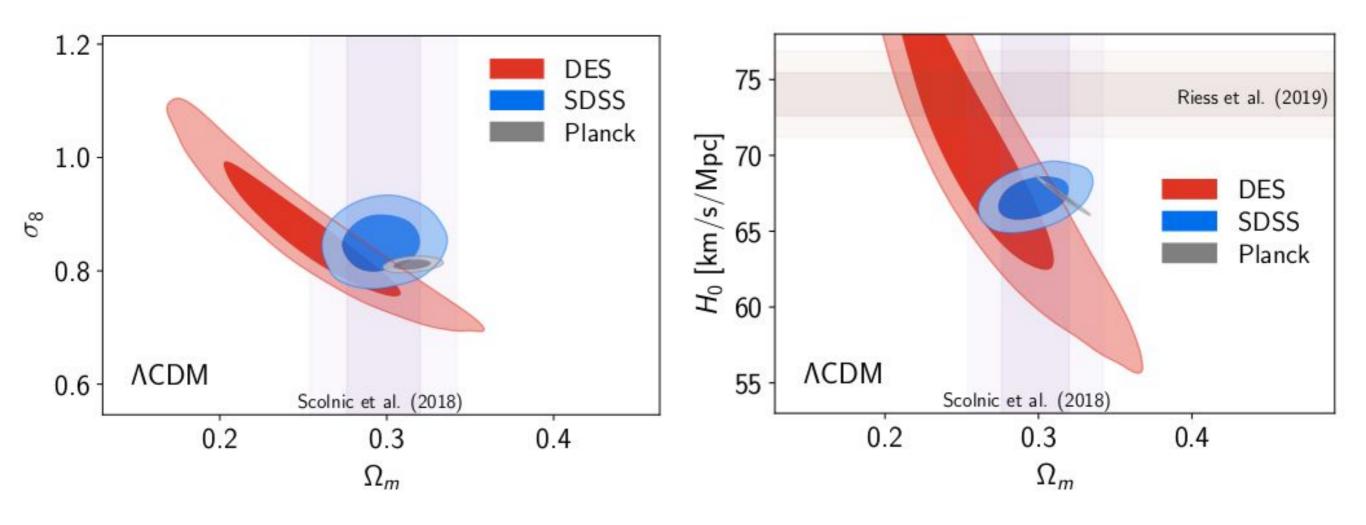


• Pull statistics \rightarrow the full suite of BAO+RSD measurements are fully consistent with the preferred LCDM model





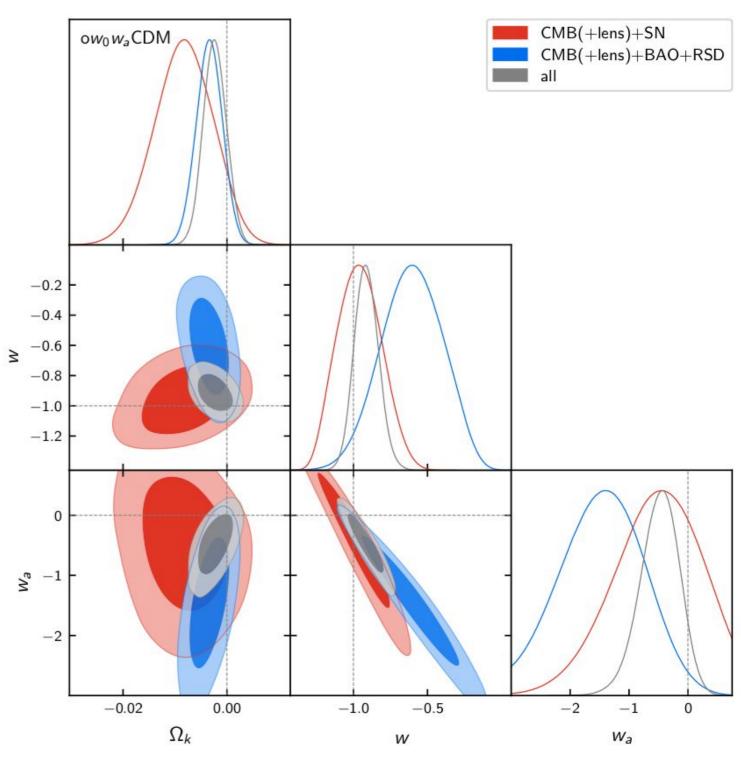
- BAO+RSD consistent with Planck and DES under LCDM
 - CMB and DES: sigma8 tension, Omega_m tension, or no tension?





Constraints from ow0waCDM

- Complementarity of BAO and SNe la → tight constraints of curvature and the dark energy equation of state
- Dark Energy Task Force
 Figure of Merit of 103
- FoM=150 predicted by DETF at conclusion of Stage-III





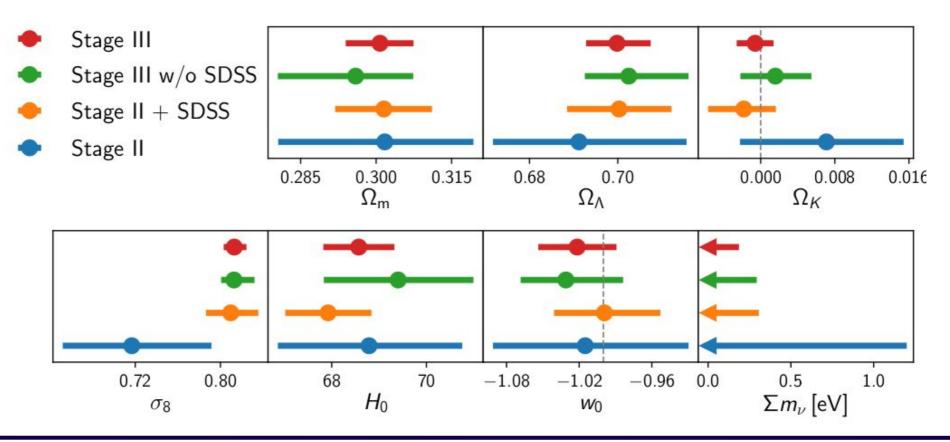
- ~ 1% precision estimates on the dark energy density, H0, and amplitude of matter fluctuation regardless of cosmological model
- Little degradation in curvature precision with increasing parameters
- Little degradation in neutrino mass with increasing parameters
- $w_p(z=0.36) = -1.013 + 0.029$ in w0waCDM, little degradation with w_a

	Ω_{Λ}	H_0	σ_8	Ω_K	w_0	w_a	$\Sigma m_{ u} [\mathrm{eV}]$
ACDM	0.6959 ± 0.0047	68.19 ± 0.36	0.8073 ± 0.0056	-	—	-	-
$o\Lambda CDM$	0.6958 ± 0.0048	68.21 ± 0.55	0.8076 ± 0.0065	0.0001 ± 0.0017		1000	in the second
wCDM	0.6992 ± 0.0066	68.64 ± 0.73	0.8128 ± 0.0092	() —+)	-1.020 ± 0.027		
owCDM	0.6997 ± 0.0069	68.59 ± 0.73	0.8127 ± 0.0091	-0.0004 ± 0.0019	-1.023 ± 0.030		100
$w_0 w_a \operatorname{CDM}$	0.6971 ± 0.0069	68.47 ± 0.74	0.8139 ± 0.0093		-0.939 ± 0.073	$-0.31^{+0.28}_{-0.24}$	<u> </u>
$ow_0 w_a CDM$	0.6988 ± 0.0072	68.20 ± 0.81	0.8140 ± 0.0093	-0.0023 ± 0.0022	-0.912 ± 0.081	$-0.48\substack{+0.36\\-0.30}$	
$m_{\nu}\Lambda \text{CDM}$	0.6975 ± 0.0053	68.34 ± 0.43	$0.8115^{+0.0092}_{-0.0068}$		-	1.200	< 0.111(95%)
m_{ν} wCDM	0.6993 ± 0.0067	68.65 ± 0.73	$0.813^{+0.011}_{-0.0098}$	—	$-1.019^{+0.034}_{-0.029}$	—	< 0.161(95%)



A Decade of Progress (nuow0CDM)

- Stage-II (2010): WMAP + JLA SNe + SDSS/2dFGRS BAO
- SDSS: 50X decrease in curvature/H0/sigma8/w0/neutrino mass posterior volume relative to Stage-II
 - largest improvements in curvature, H0, and neutrino mass precision
 - Stage-II + SDSS: H0 = 67.60 +/- 0.92 km/s/Mpc
- Planck+Pantheon+DES: additional 20X improvent \rightarrow average 4X per parameter





Summary

- **BOSS**/eBOSS
 - Conclusion of Stage-III Dark Energy surveys with spectroscopy
 - BAO measurements over 11 Gyr & RSD measurements to z<1.5

Cosmology

- BAO complement SNe, but higher precision in isolation
- BAO allow robust estimates of H0 not possible otherwise
- RSD complement WL \rightarrow favor LCDM model with Planck and support General Relativity
- SDSS largest role in advancing neutrino mass constraints: I-sigma uncertainty now comparable to minimum allowed mass
- LCDM model is preferred by all data: SDSS leads the way in improving precision of late-time cosmological model since 2010

Special thanks to our partners: DOE Office of Science, SDSS partnering institutions, Alfred P. Sloan Foundation, Utah Center for High Performance Computing (CHPC), and the National Energy Research Scientific Computing Center (NERSC).