Cryogenic Dark Matter Search

Current Status & Future Plans



Todd Doughty Cosmology in Northern California Lawrence Berkeley National Laboratory October 22, 2010

Motivation for WIMP Search

 Significant Portion of Mass Density of Universe



 Convergence of Particle Physics & Cosmology

 – Relic DM Density → σ_{DM} ≈ σ_{WF}

Three Types of Searches



Production (LHC)



Indirect Detection (Fermi, Pamela)



Direct Detection (CDMS, XENON)

Direct Detection Methods

DAMA/Nal, ZEPLIN I, CLEAN, KIMS











Transition Edge Sensor



Background Discrimination

- Each Event
- Multiple Dimensions
 Ionization Yield



Background Discrimination

- Each Event
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 Ionization Yield





Background Discrimination

- Each Event
- Multiple Dimensions
 Ionization Yield
 Phonon Timing







CDMS II: Results



• Final Results – 2009

- 2 events in NR Band
- Estimate Leakage:
 0.9 ± .3 events
- 23% chance both events not Wimps
- Analysis can't be significant evidence for Wimps, but neither event can be rejected
- Z. Ahmed et al., *Science* Vol. 327, no. 5973 (2010)

Primary background
 – Surface Events



Primary background

Surface Events

New Detector - IZIP

Interdigitated Electrodes

Side 1 Charge Electrode +2V



Side 2 Charge Electrode -2V















Primary background

Surface Events

New Detector - IZIP

Interdigitated Electrodes





 Primary background

 Surface Events

 New Detector - IZIP

 Interdigitated Electrodes

- 1000:1 rejection

Future Plans





SuperCDMS Collaboration

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