

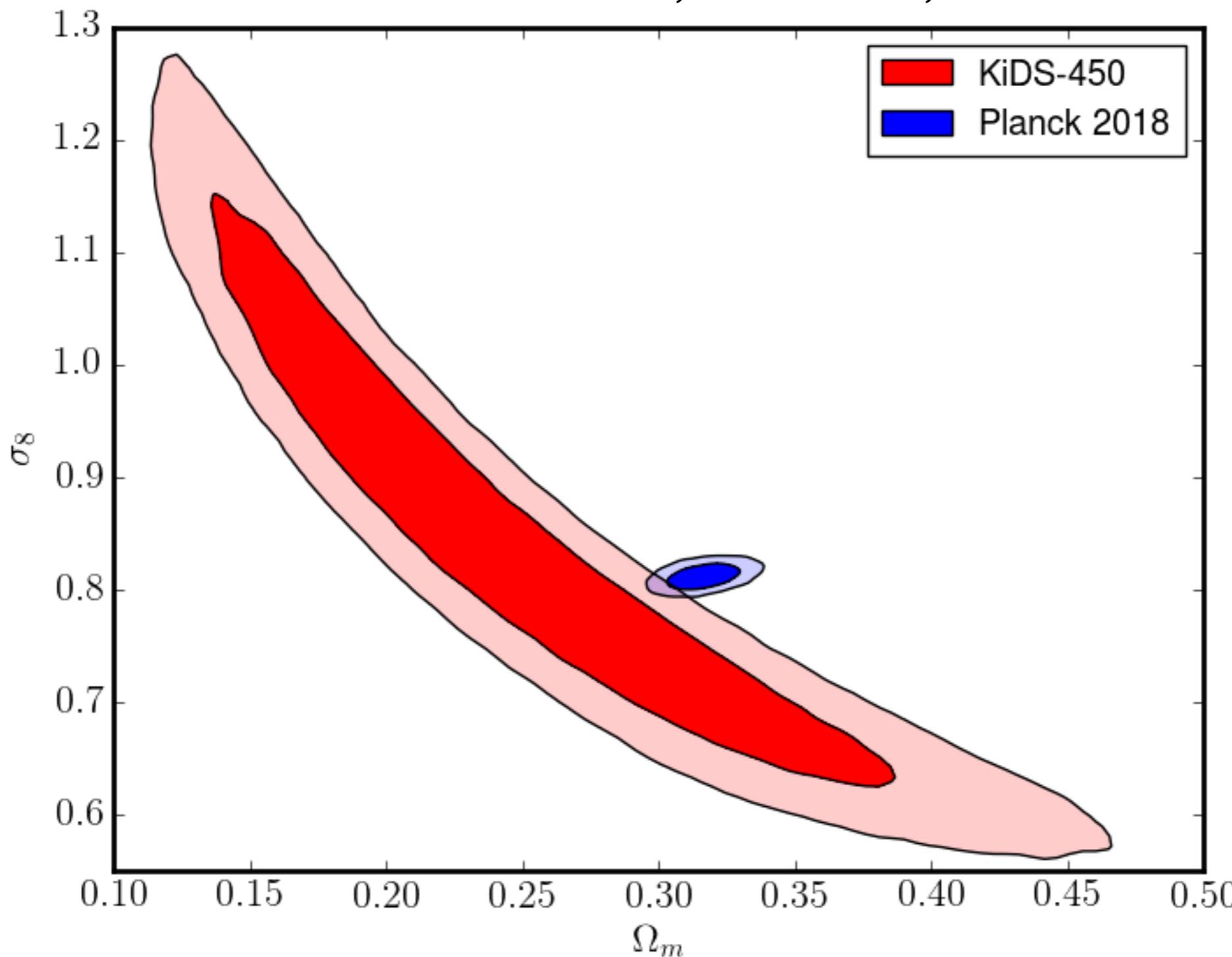
Latest KiDS cosmic shear results: **KiDS-VIKING-450**

Konrad Kuijken
Leiden Observatory

Hildebandt, Köhlinger, van den Busch, Heymans, Joachimi, Arun
Kannawadi, Wright, + many more KiDS

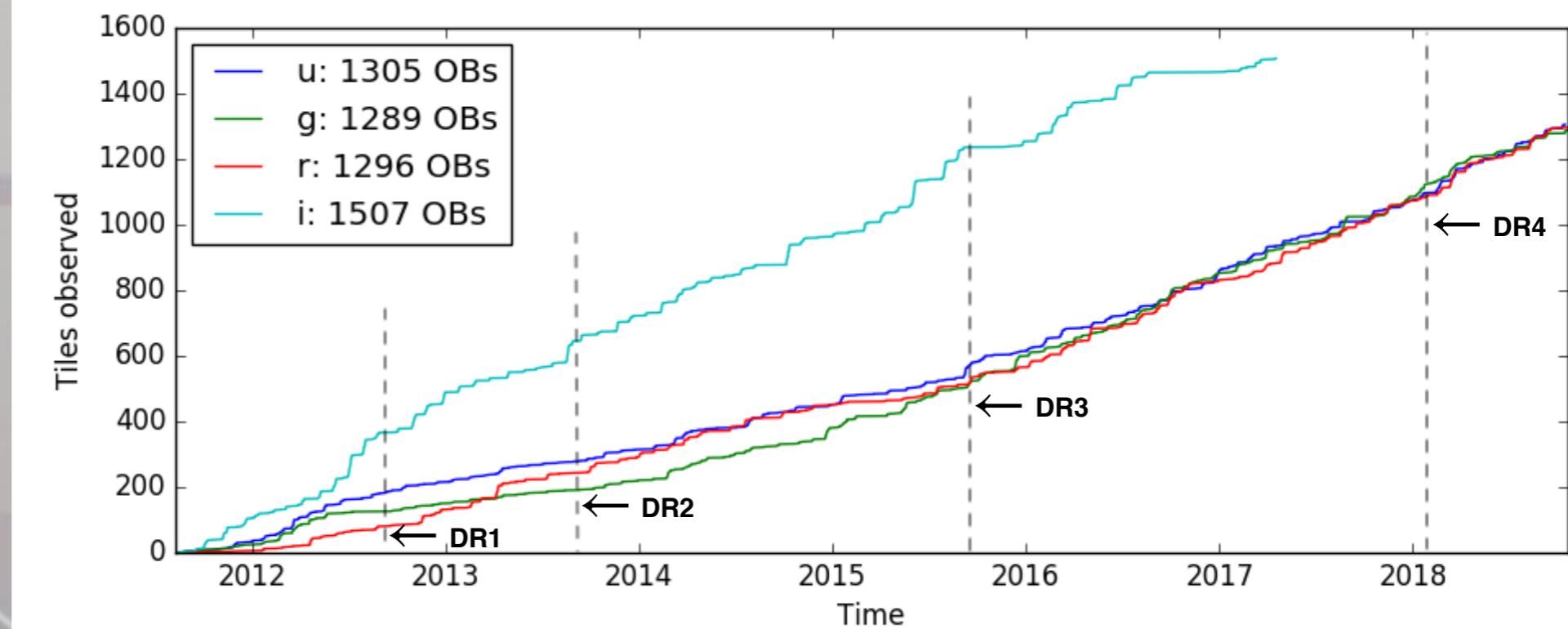
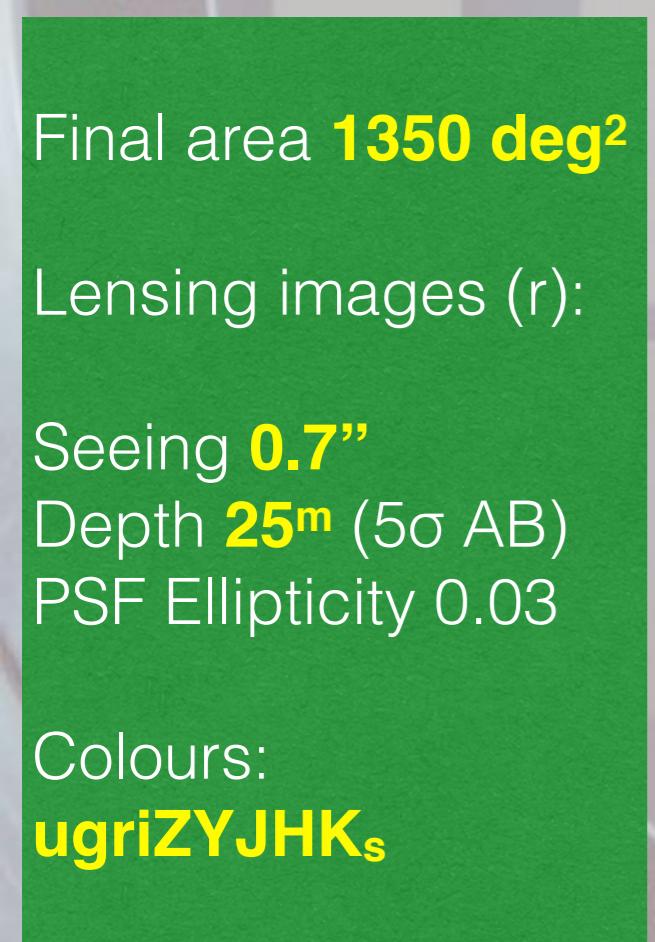
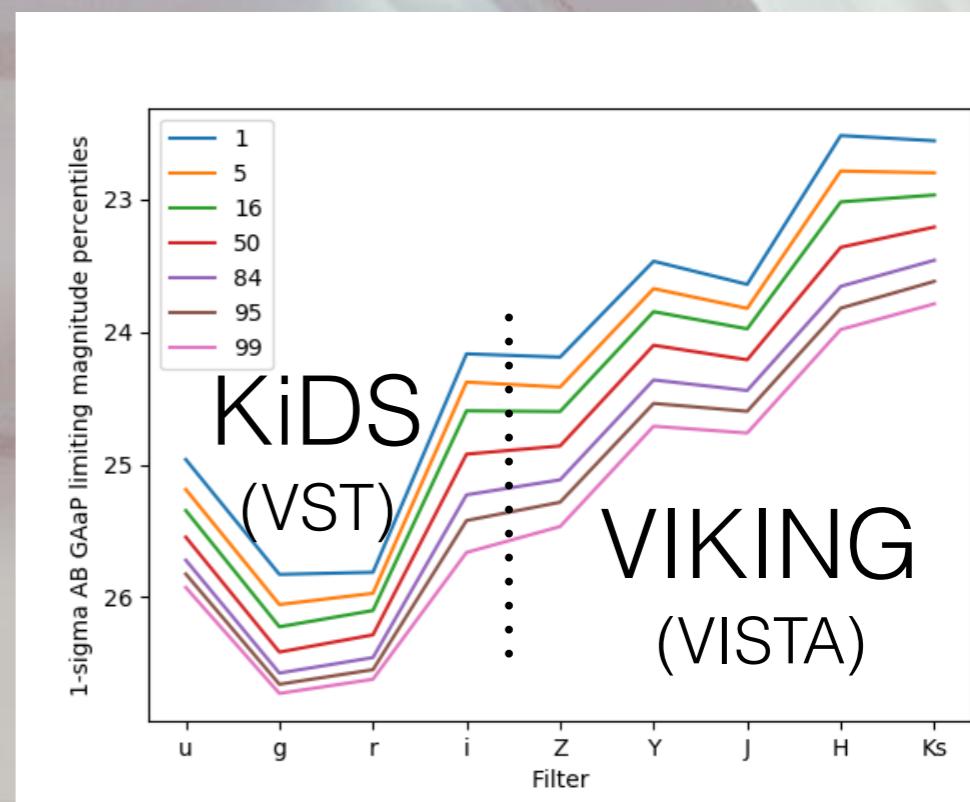
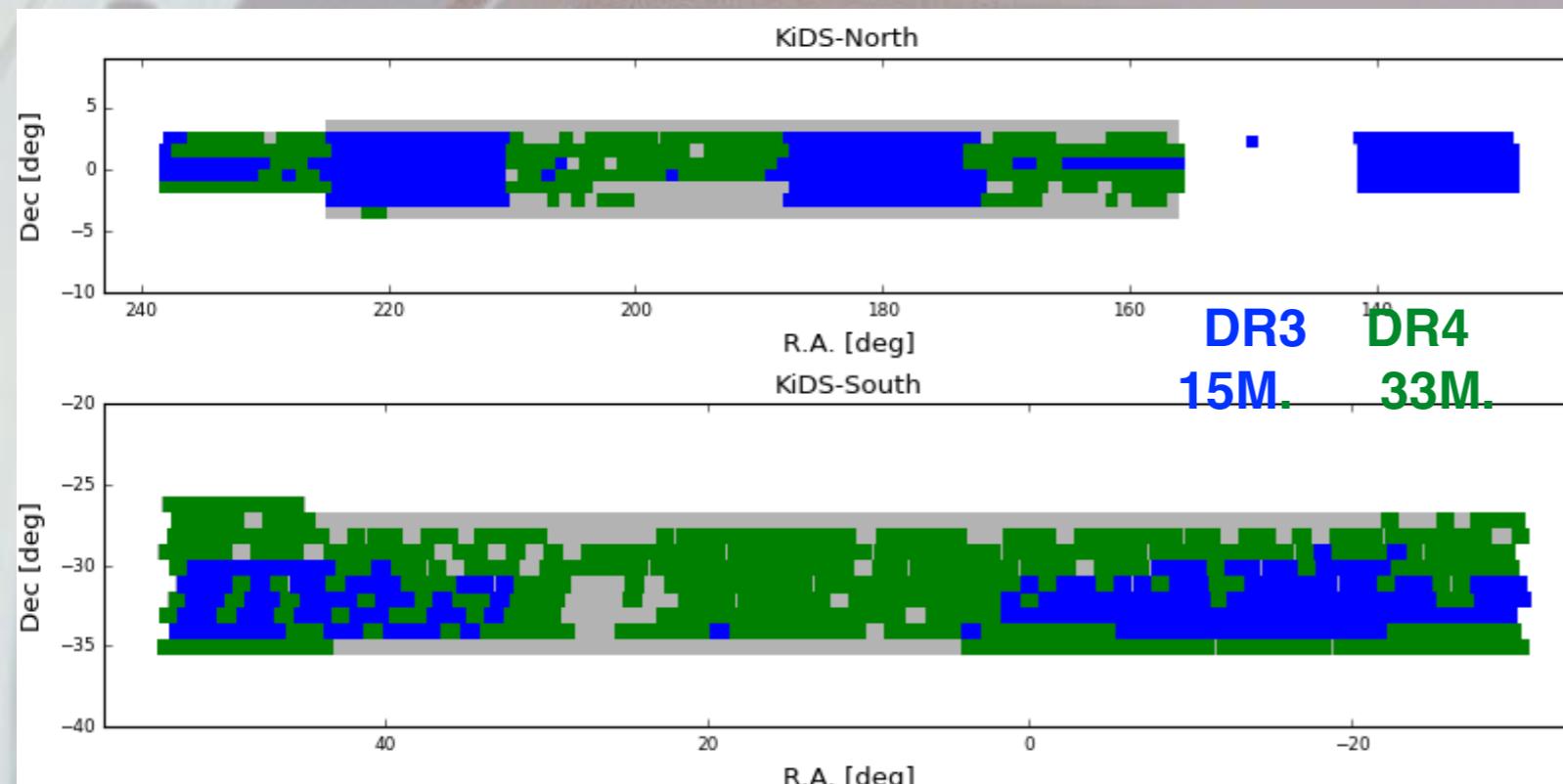
Alex Amon, Joachim Harnois-Déraps, Harry Johnston,
Naomie Robertson, Cristóbal Sifón

Hildebrandt et al. 2017, MNRAS 465, 1454

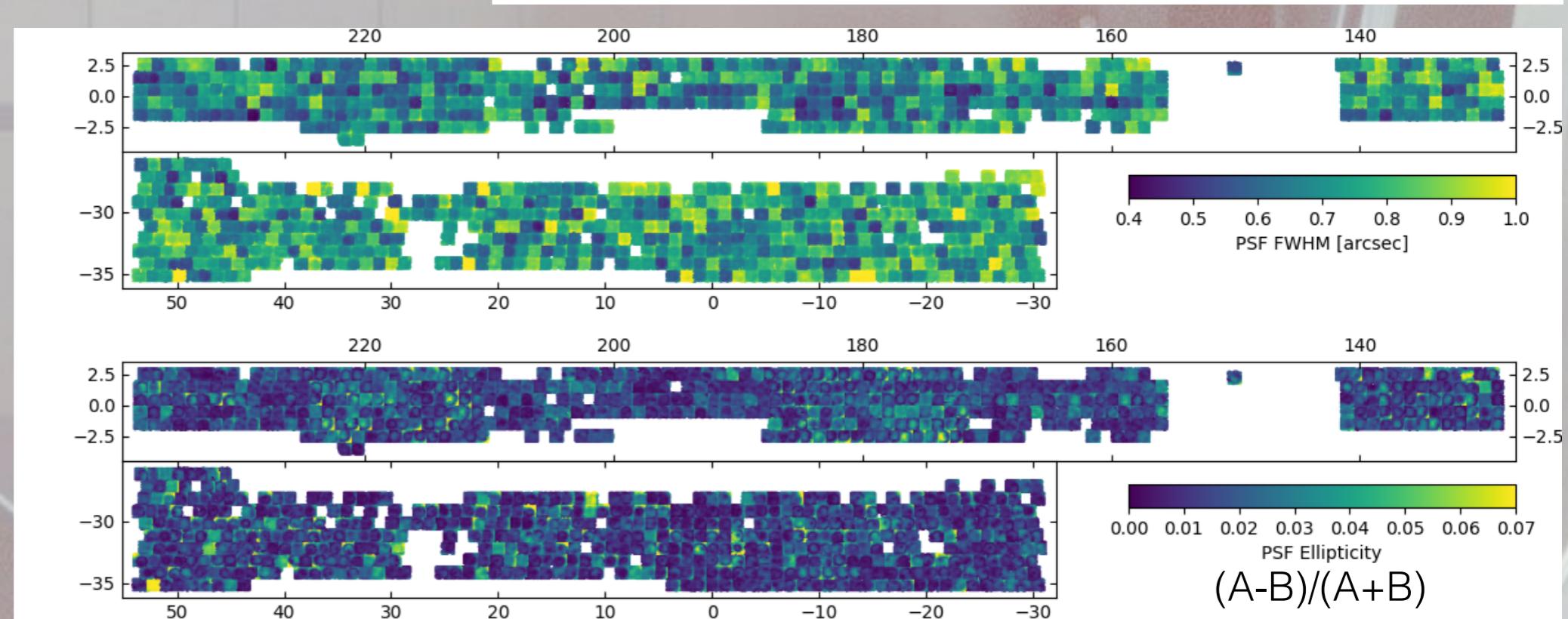
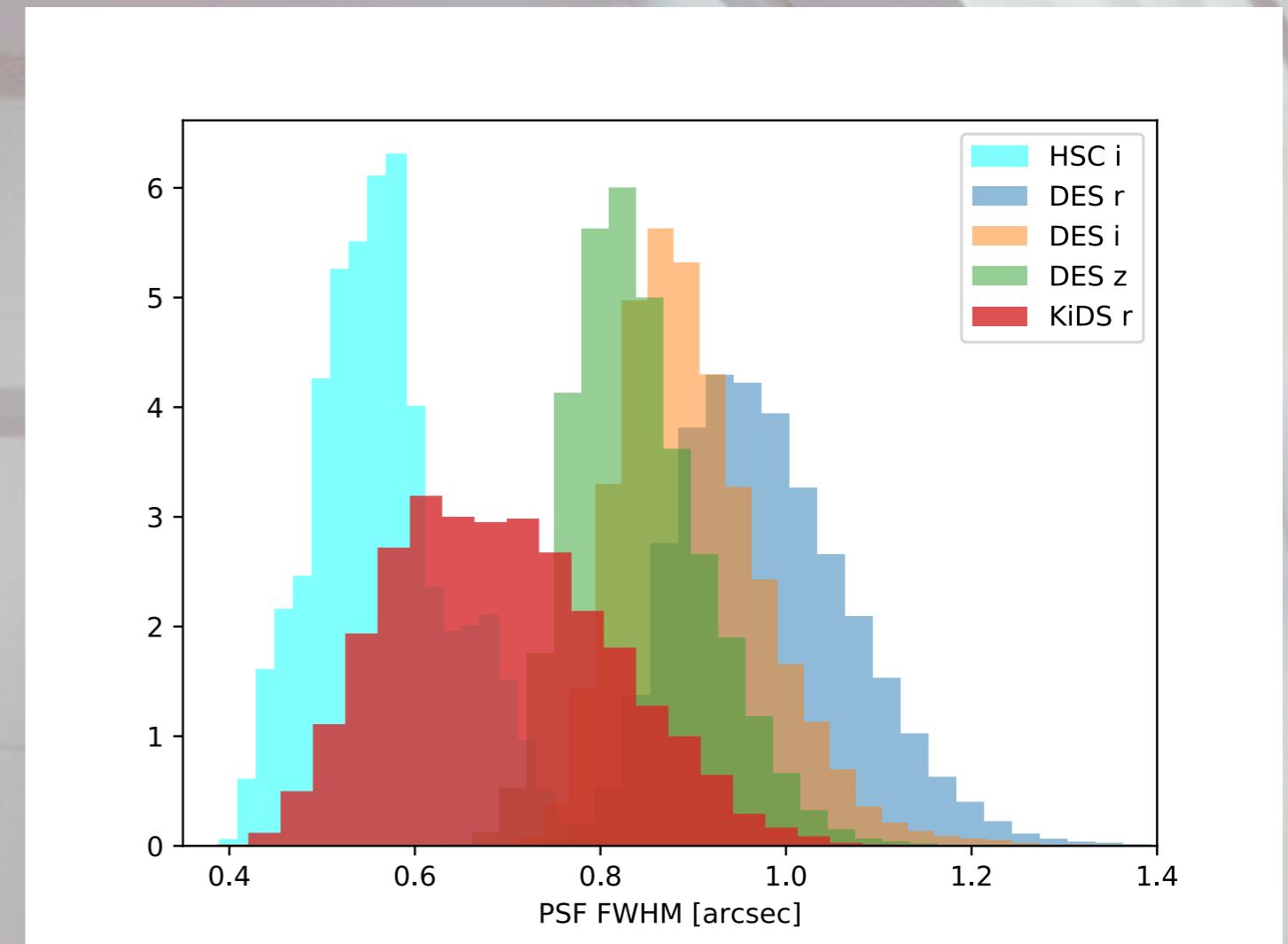


KiDS

The Kilo-Degree Survey: basic stats



- Median 0.7"
- PSF ellipticity low
- Image scale constant
- Benign PSF patterns

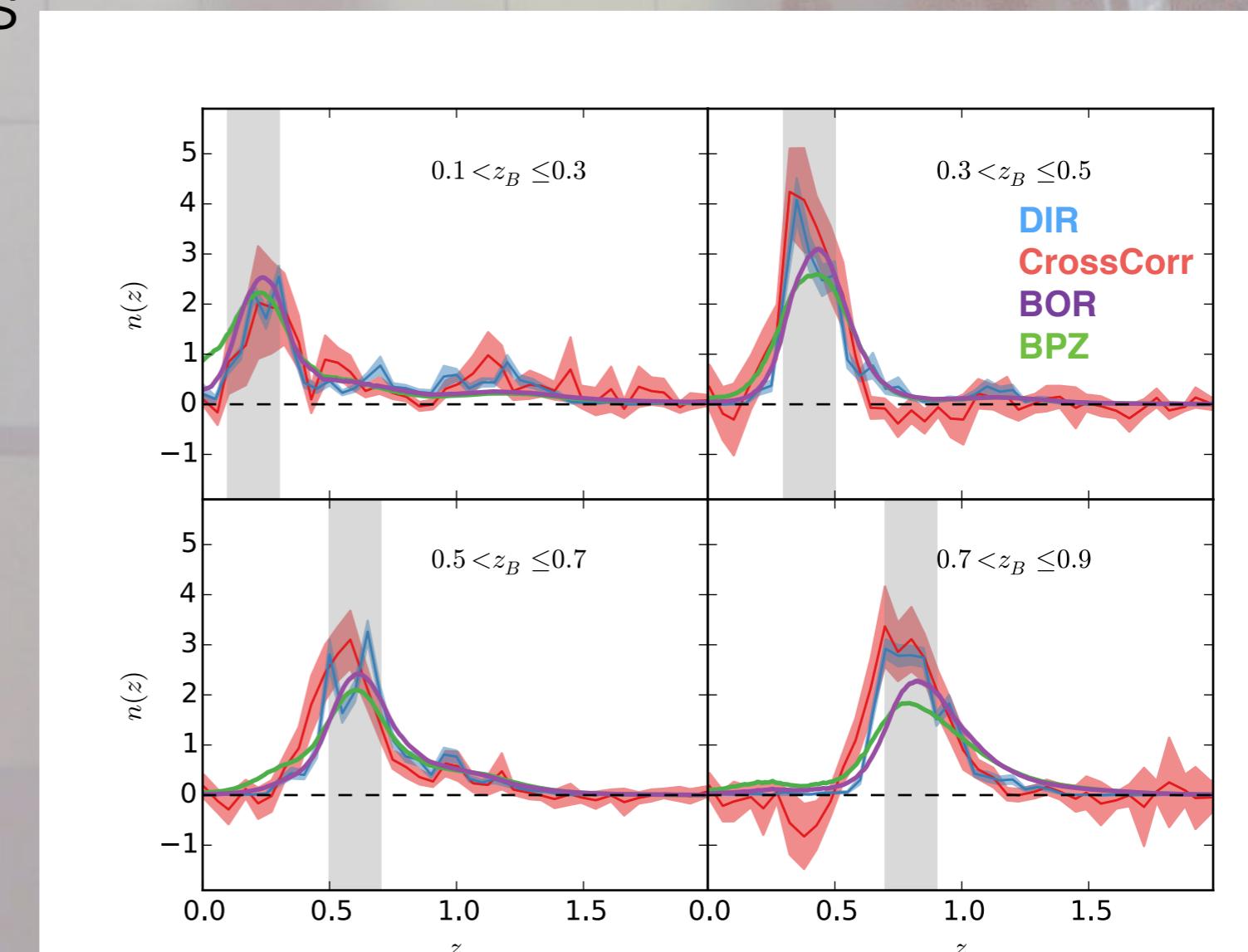


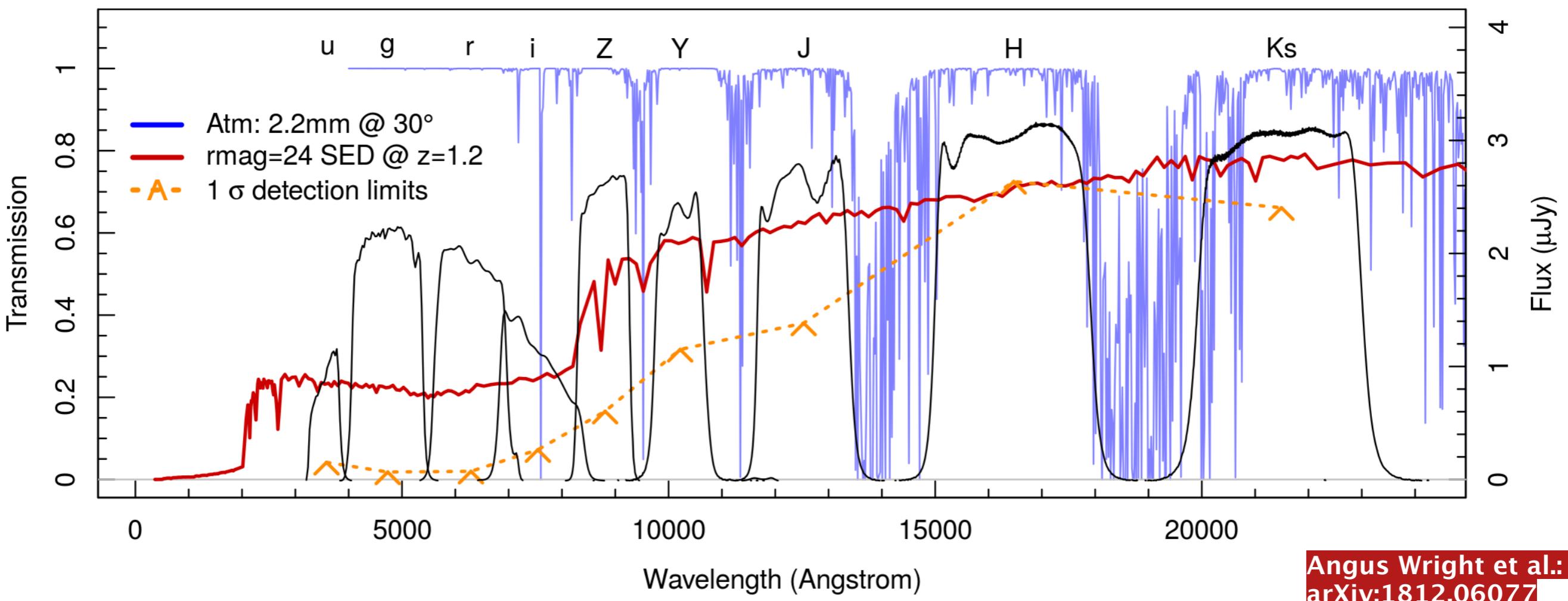
- ***Lensfit***
 - forward modelling code,
 - simultaneous fit to all sub-exposures with individual PSFs,
 - likelihood function for each source ellipticity.
- Self-calibration (response of best-fit model to shear)
- Calibration verified with **realistic image simulations**
- Conservative estimate of systematic shear calibration error: 2%.

Arun Kannawadi et al.:
arXiv:1812.03983



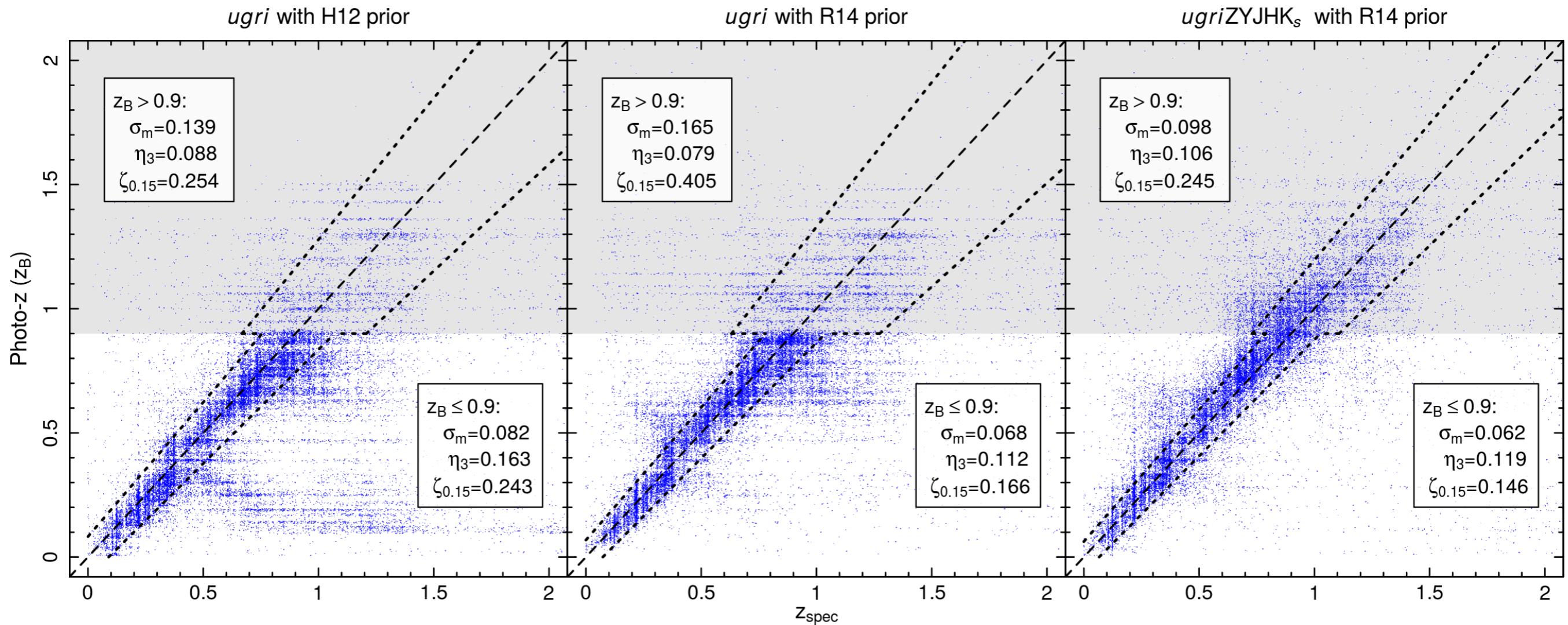
- KiDS-450: u+g+r+i GAaP photometry
- Classify into tomographic bins using BPZ, then estimate $n(z)$
- **DIR**ect reweighting of ***deep spec-z catalogues*** to match 4-D magnitude distributions
- Only works if :
 - mag. space covered
 - $r < \sim 24$ (C3R2)
 - Relation unique



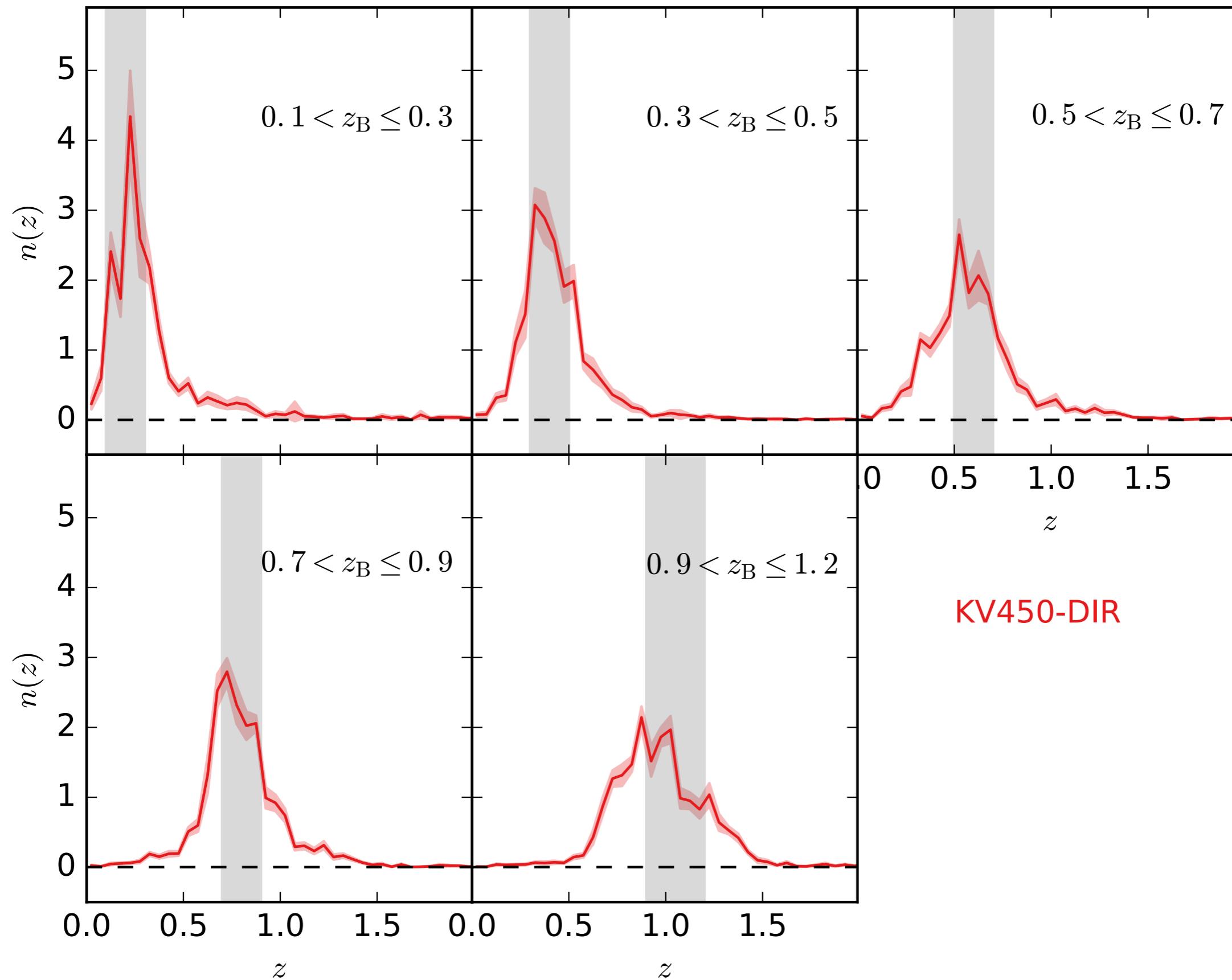


- Include 5-band VIKING survey ZYJHK_s (depth 23-21, 5 σ)

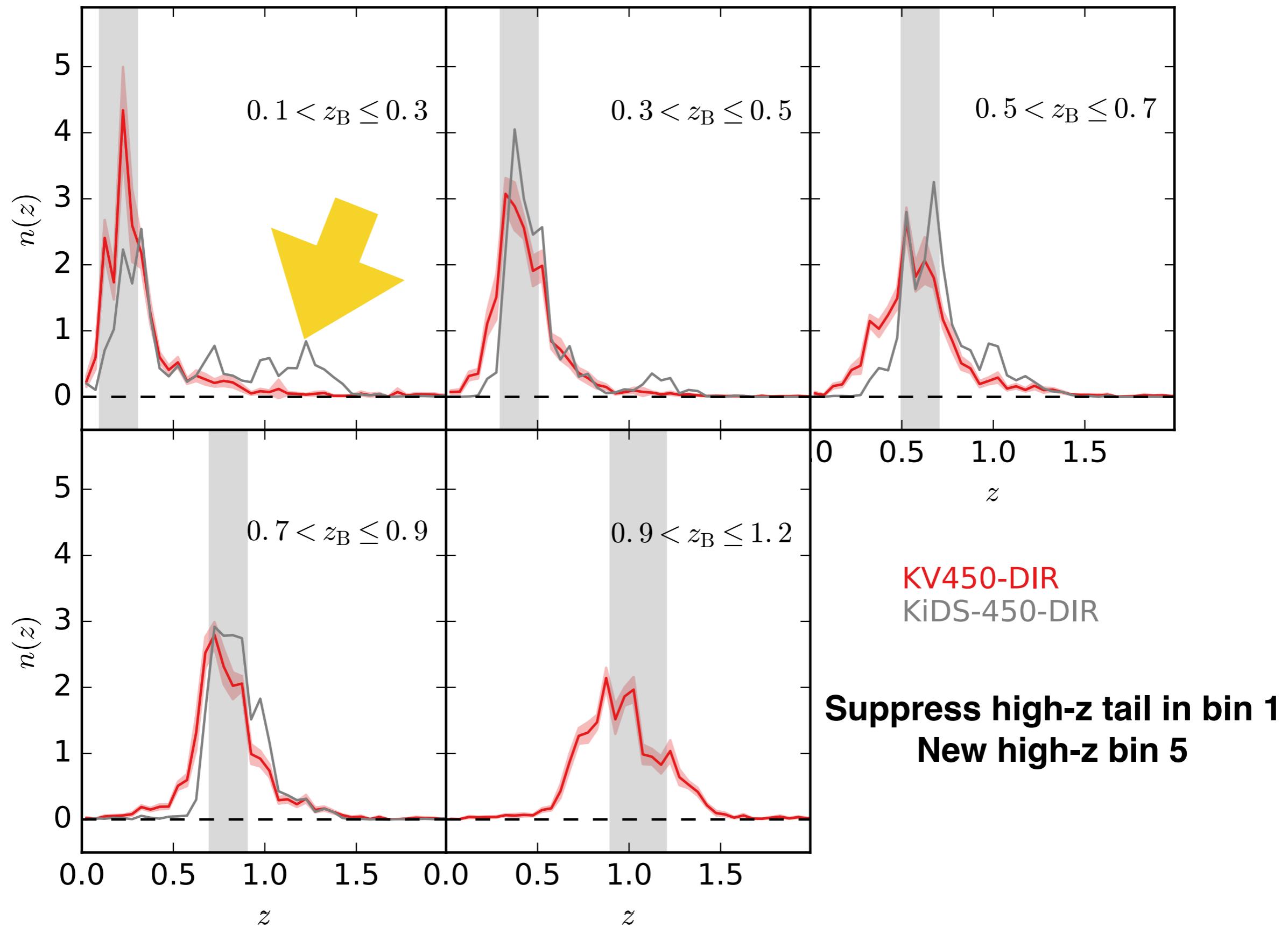
Benefits of adding near-IR



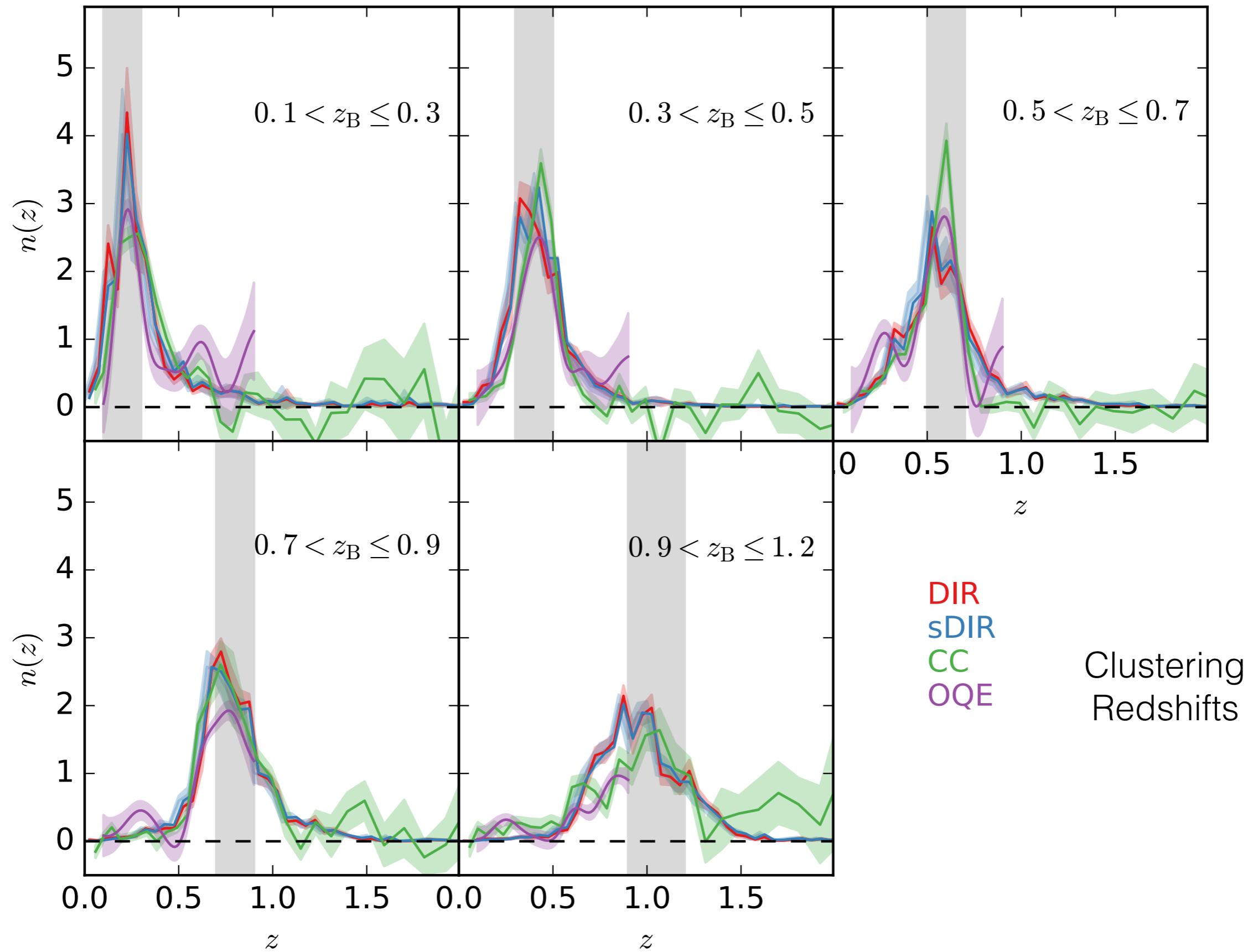
KV450 tomographic bins



KV450 tomographic bins

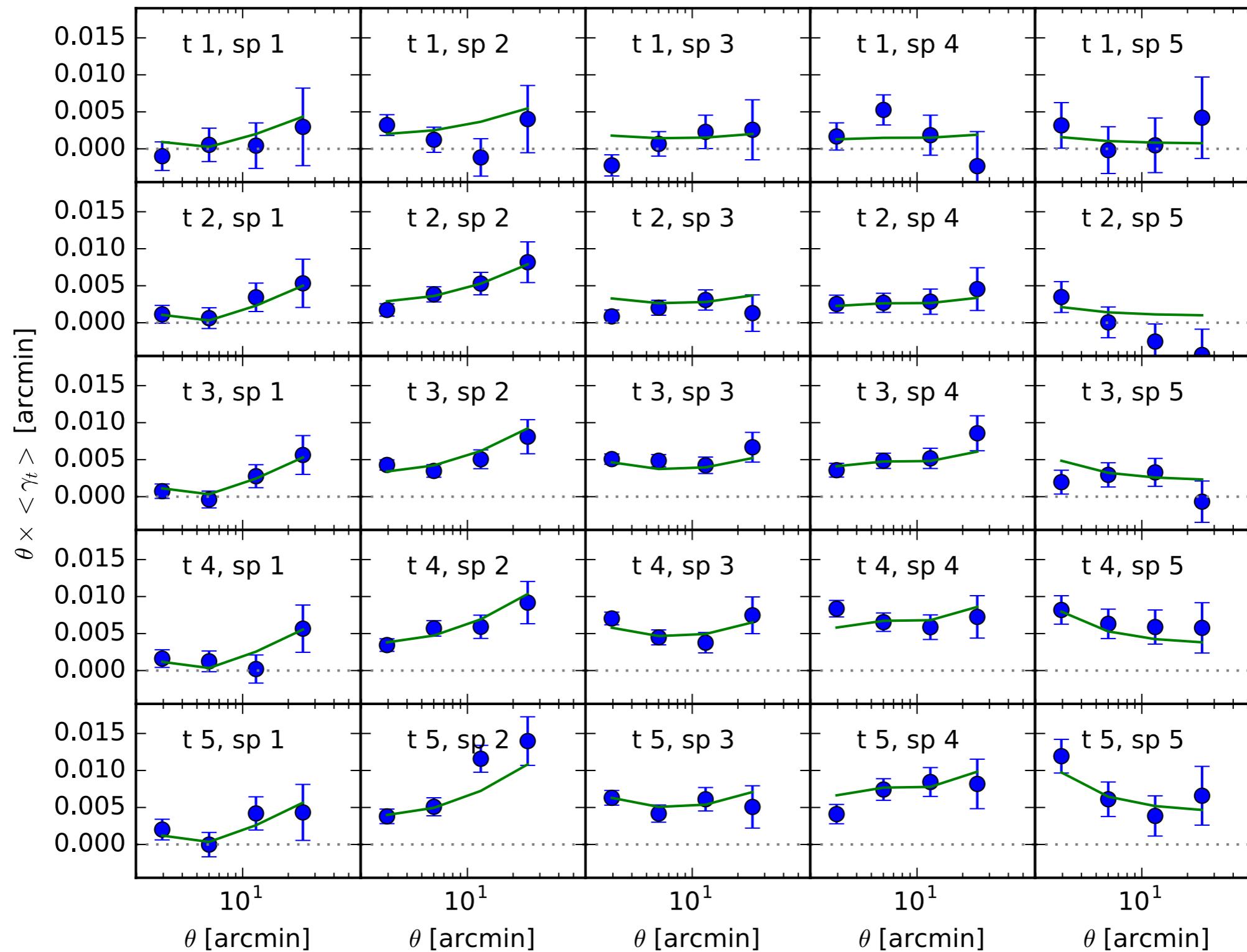


KV450 tomographic bins

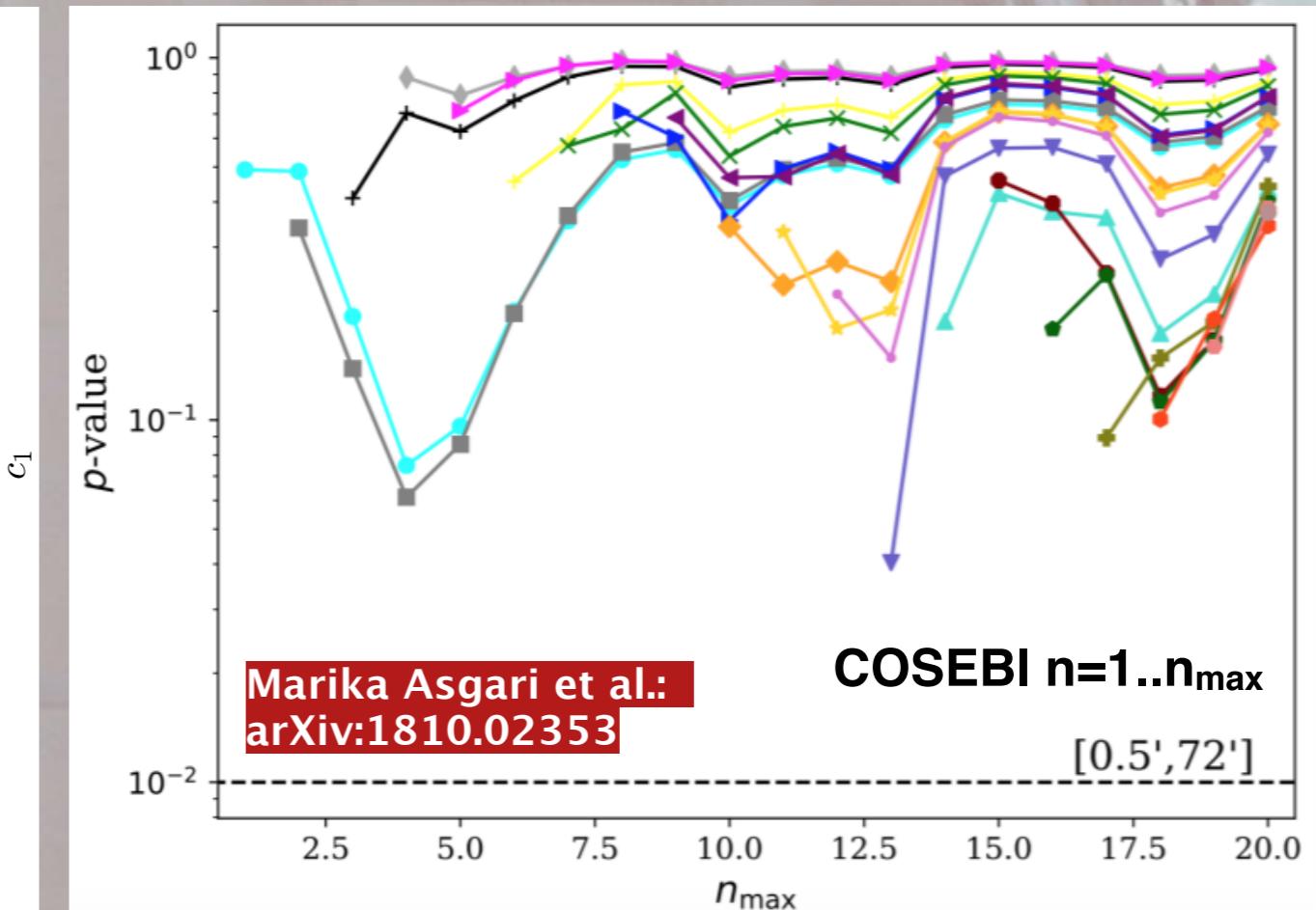
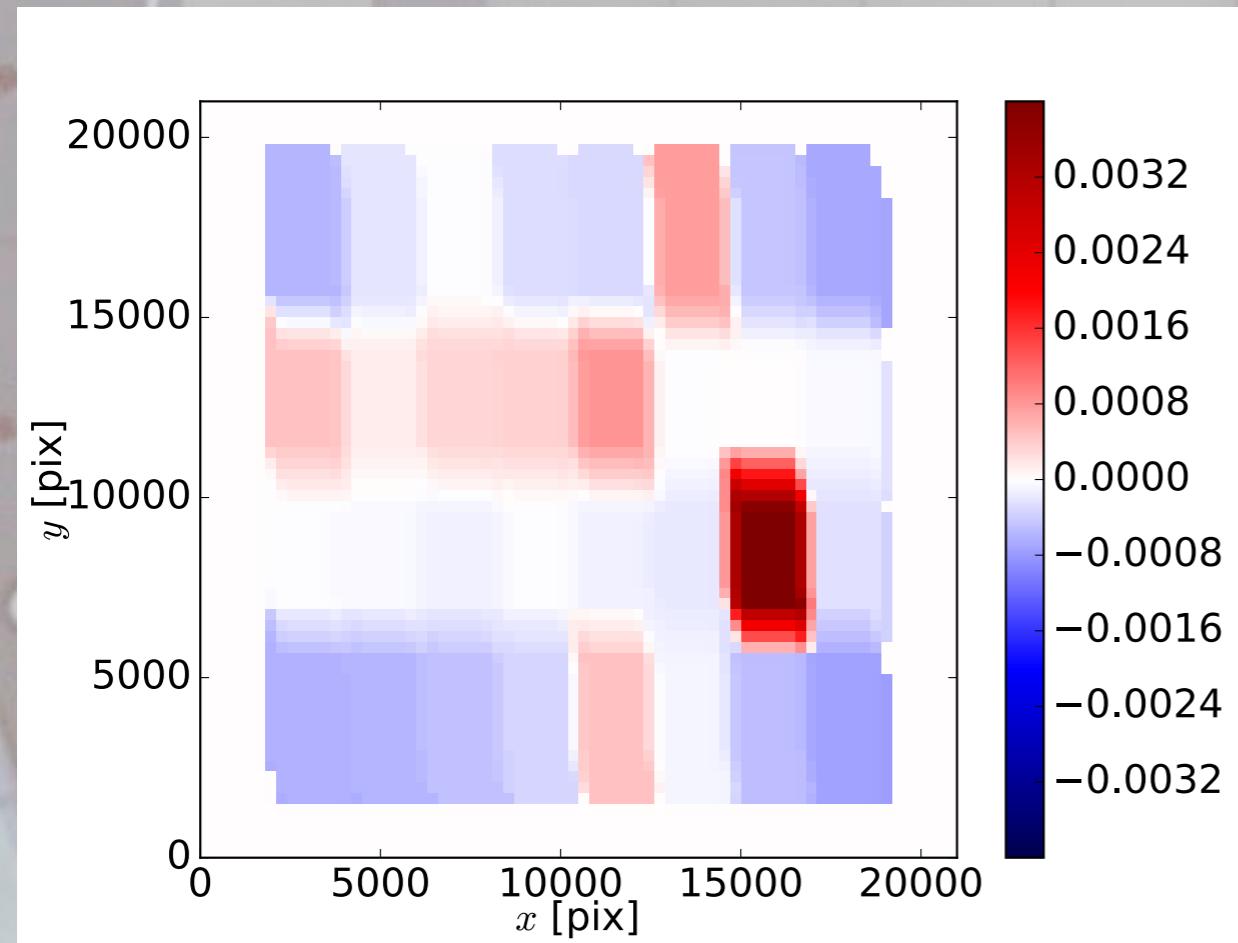


Shear ratio test

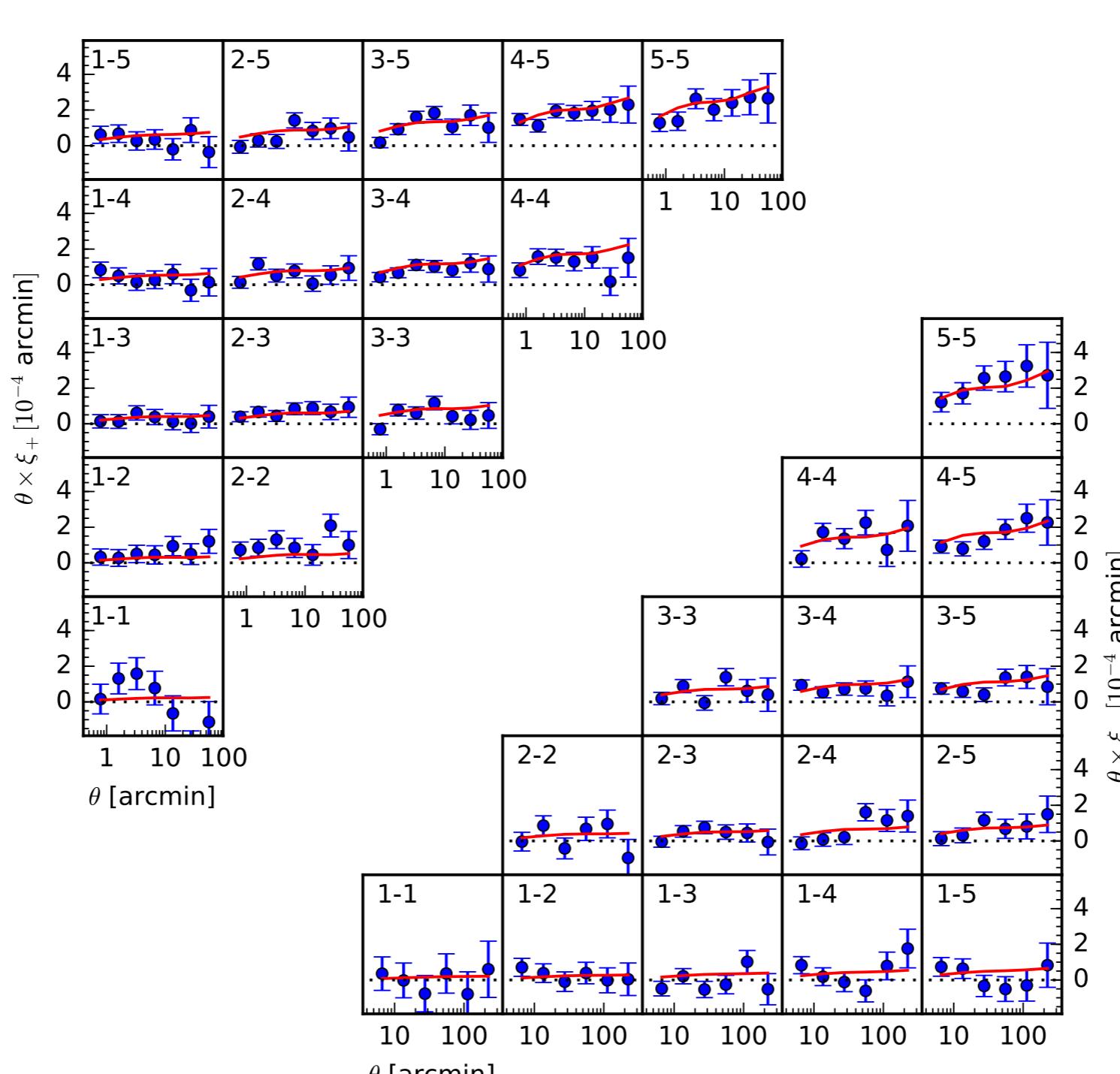
GAMA lenses, KV450 sources, $\chi^2/\text{dof}=1.05$, $p=35.3\%$



- Improved c-term correction (weak ellipticity dependence on PSF magnitude)



- Demonstrate with COSEBI's

Shear correlation functions ξ_{\pm} 

- (fixed error in bin centres of correlation functions)

Modelling ingredients

- Power spectrum incl. v , **baryonic** effects (CLASS, HMcode)
- $n(z)$ — **blinded** (3 sets: truth + 2 sabotaged)
- **Intrinsic alignment** (non-linear linear model)
- Error params (**redshift shifts, shear calibration** errors)

- Integrate over **θ bins**

- Covariance matrix

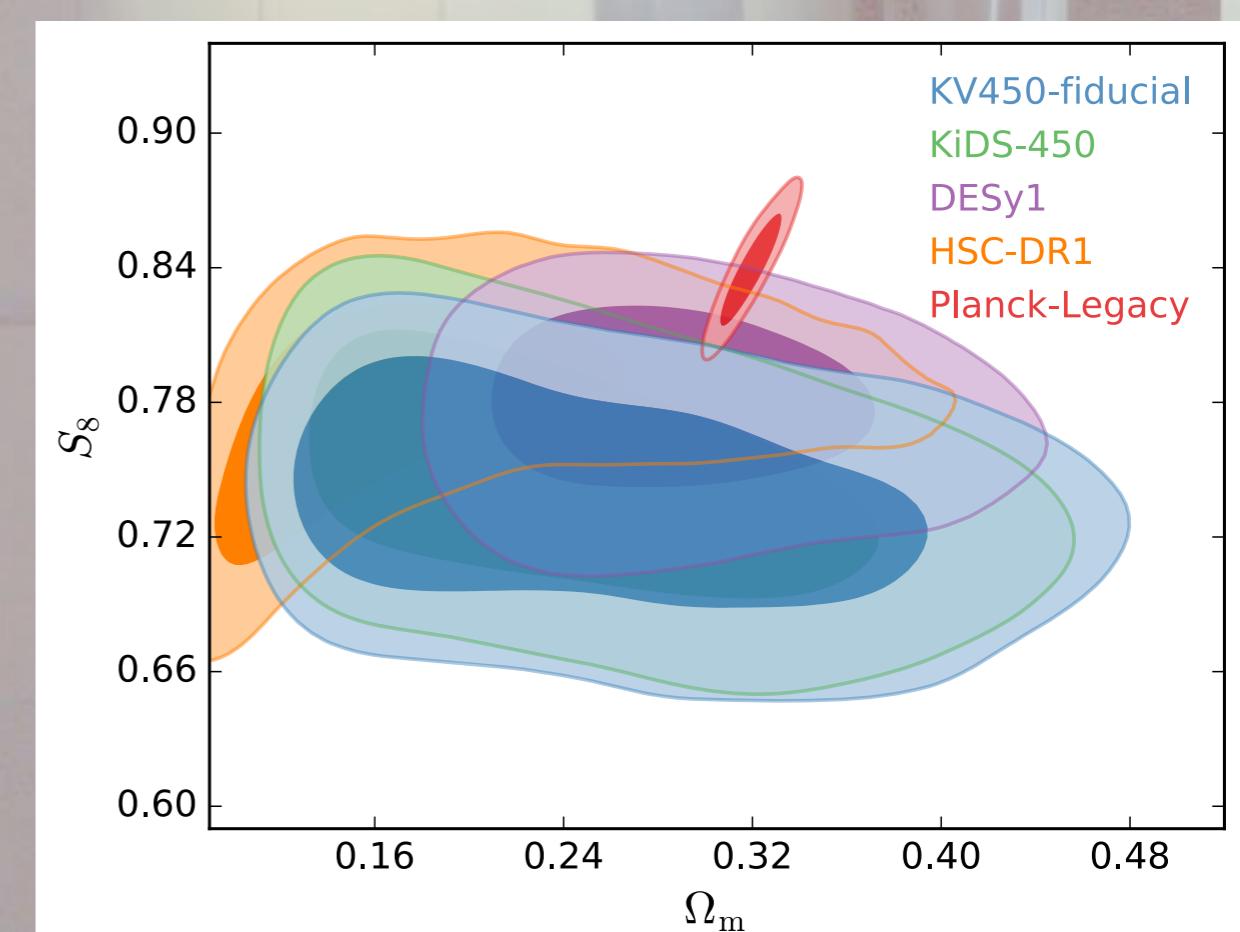
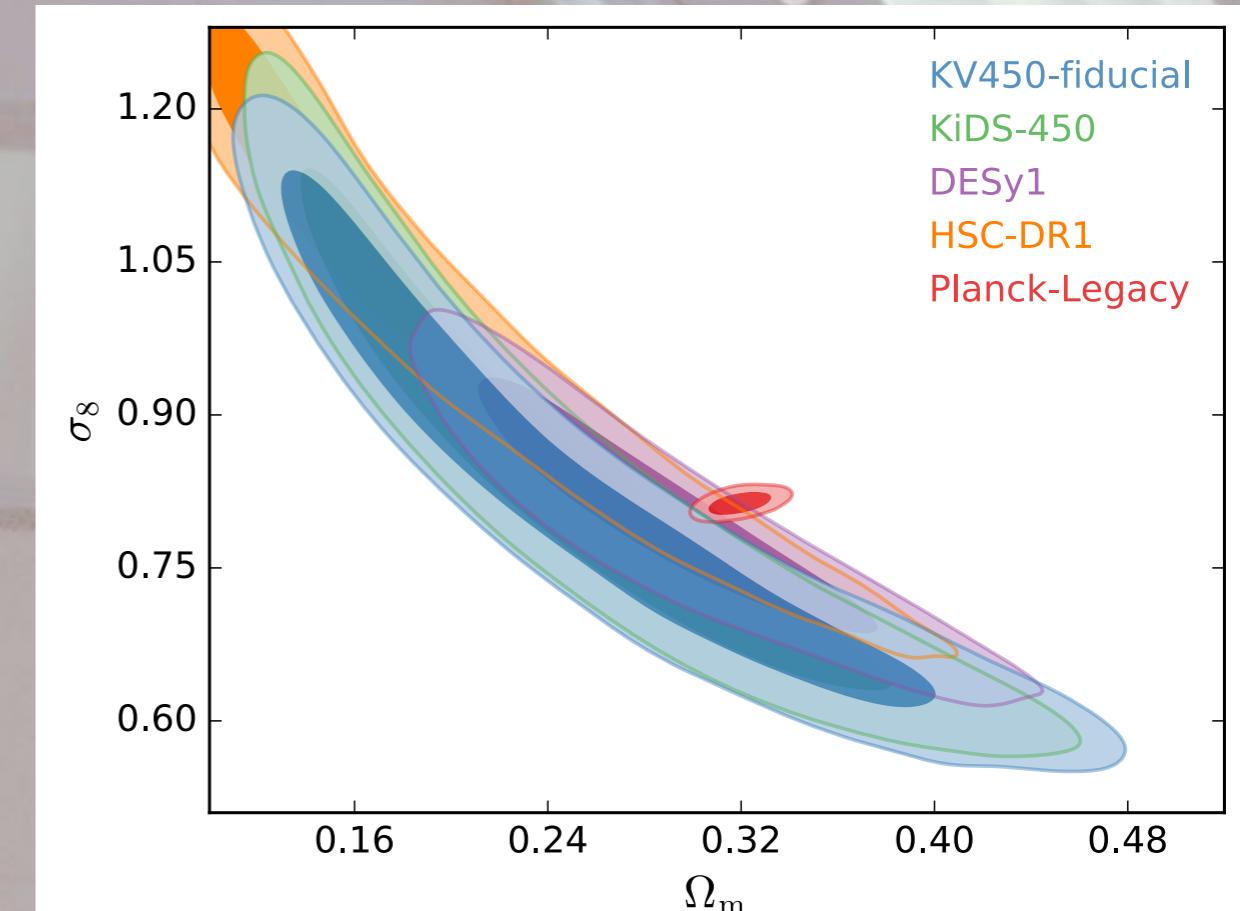
- Priors

- Two cosmology pipelines

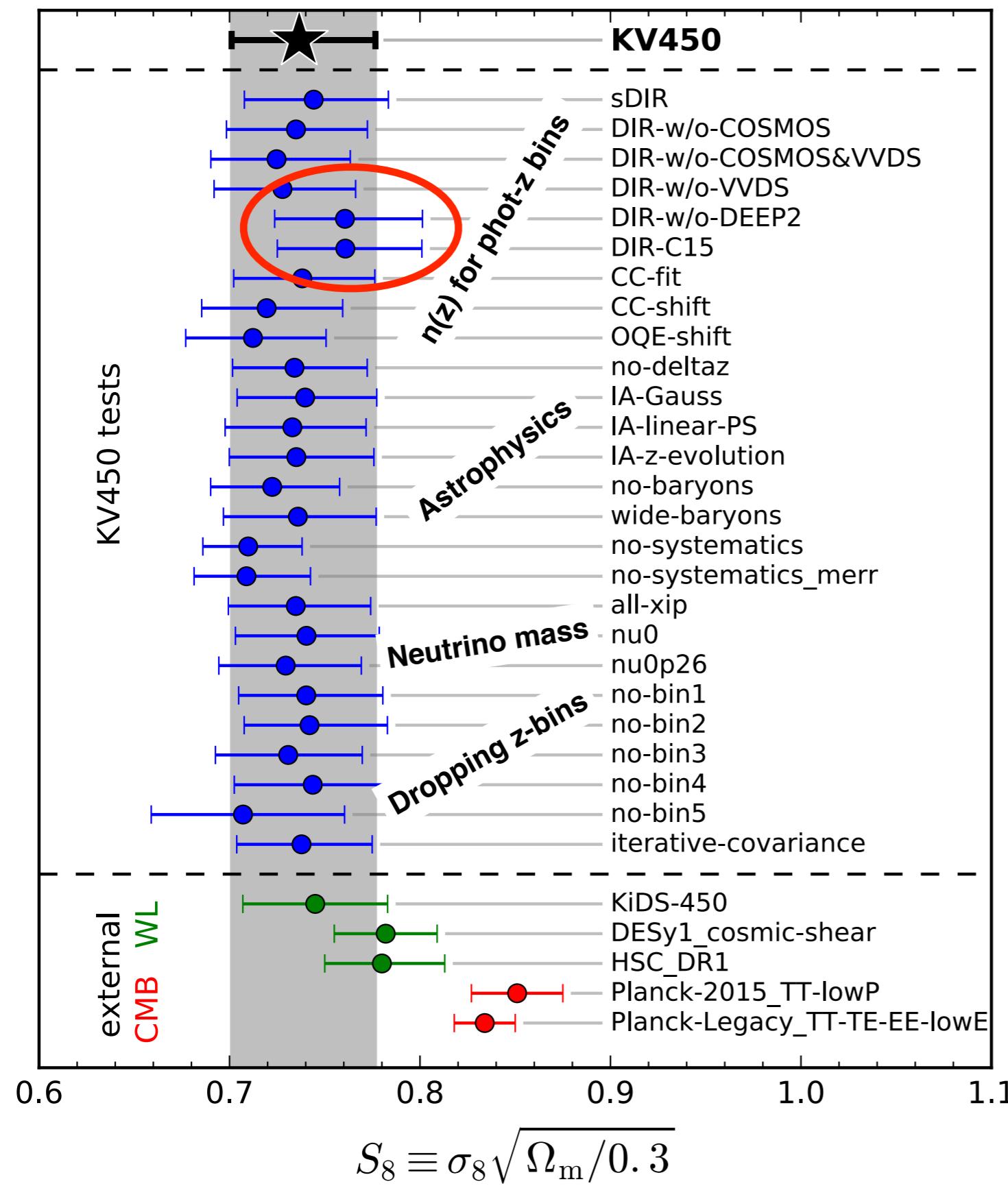
Parameter	Symbol	Prior
CDM density	$\Omega_{\text{CDM}}h^2$	[0.01, 0.99]
Scalar spectrum ampl.	$\ln(10^{10}A_s)$	[1.5, 5.0]
Baryon density	$\Omega_b h^2$	[0.019, 0.026]
Scalar spectral index	n_s	[0.7, 1.3]
Hubble parameter	h	[0.64, 0.82]
IA amplitude	A_{IA}	[-6, 6]
Baryon feedback ampl.	B	[2.00, 3.13]
Constant c -term offset	δc	0.0000 ± 0.0002
2D c -term amplitude	A_c	1.01 ± 0.13
Redshift offset bin 1	δz_1	0.000 ± 0.039
Redshift offset bin 2	δz_2	0.000 ± 0.023
Redshift offset bin 3	δz_3	0.000 ± 0.026
Redshift offset bin 4	δz_4	0.000 ± 0.012
Redshift offset bin 5	δz_5	0.000 ± 0.011

Results

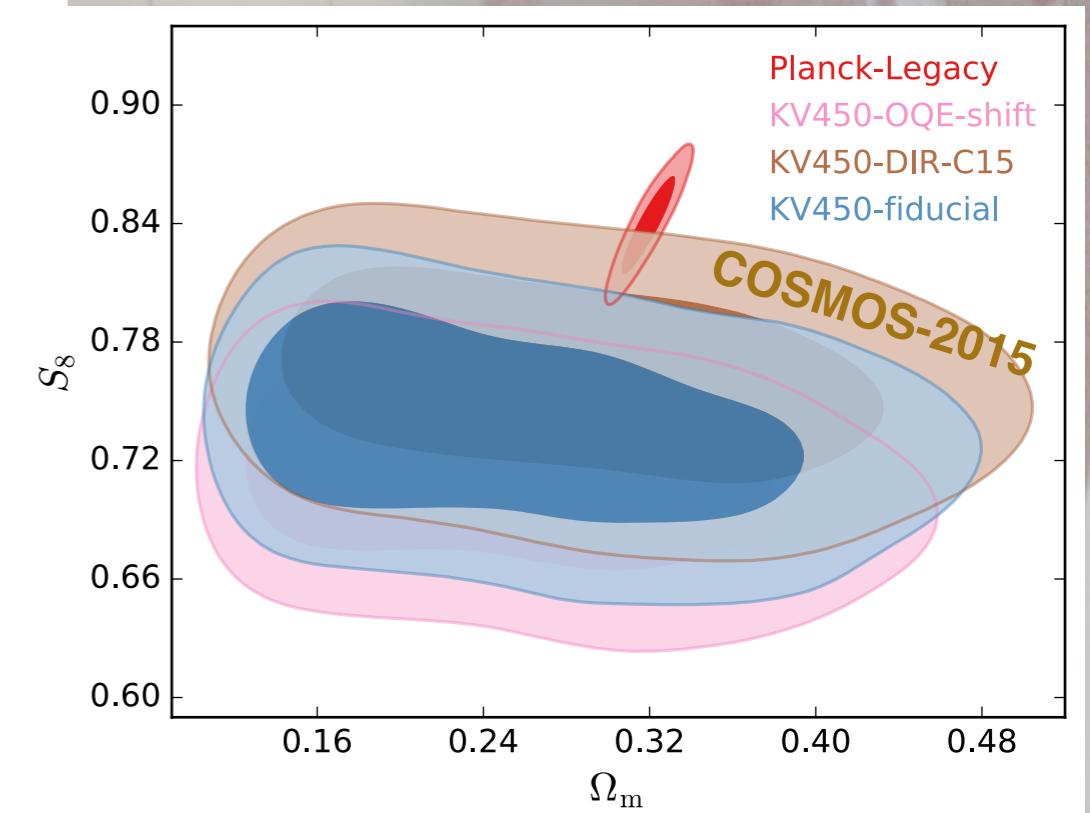
- Despite the changes, KV450 constraints very close to KiDS-450.
- Min. $\chi^2_{181} = 180.6$
- 2.3σ ‘tension’ with Planck
- $S_8 = 0.737^{+0.040}_{-0.036}$

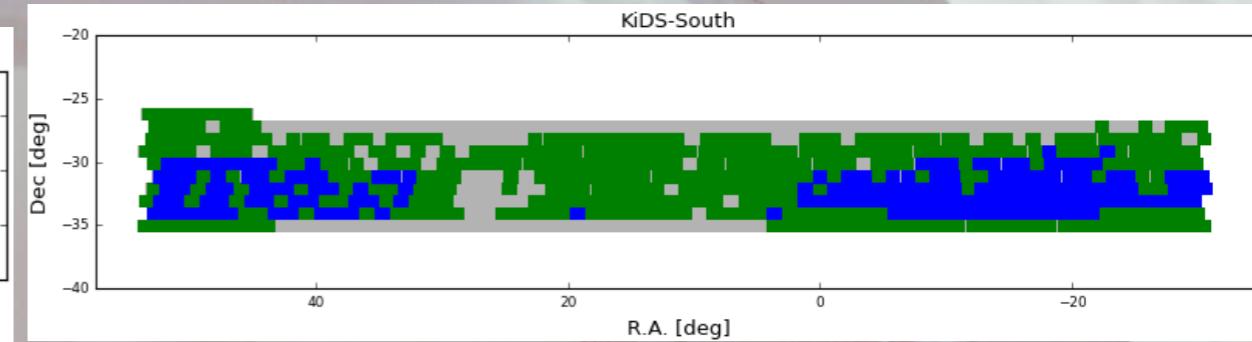
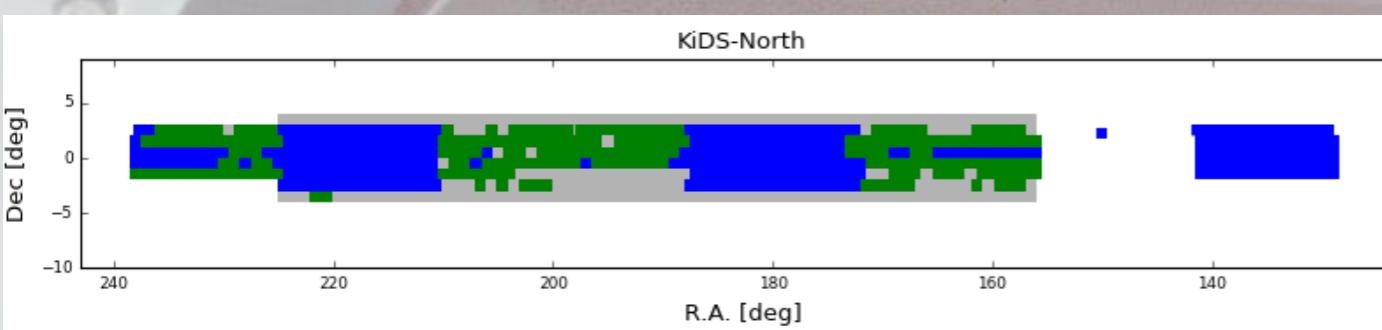


Results: systematic tests



Deep spec-z
crucial





- DR4 analysis now ongoing (1006 tiles, 9 bands)
- Observations almost complete (1350 sq.deg.)
- ==> KiDS+VIKING area will triple
- Further improve phot-z calibration
 - ‘**KiDZ**’ programme targeting deep spec-z surveys with VST and VISTA
 - Comparable precision clustering redshifts

Conclusions

- Robust 9-band photo-z + KiDS-450 shapes = KV450
- $\sim 2\sigma$ tension in S_8 vs. Planck remains
- Working towards tripling sky area with KiDS-DR4 \longrightarrow DR5
- Systematics are the limiting factor!
 - COSMOS-2015 vs. Spec-z calibration: 0.6σ

The KiDS Team

