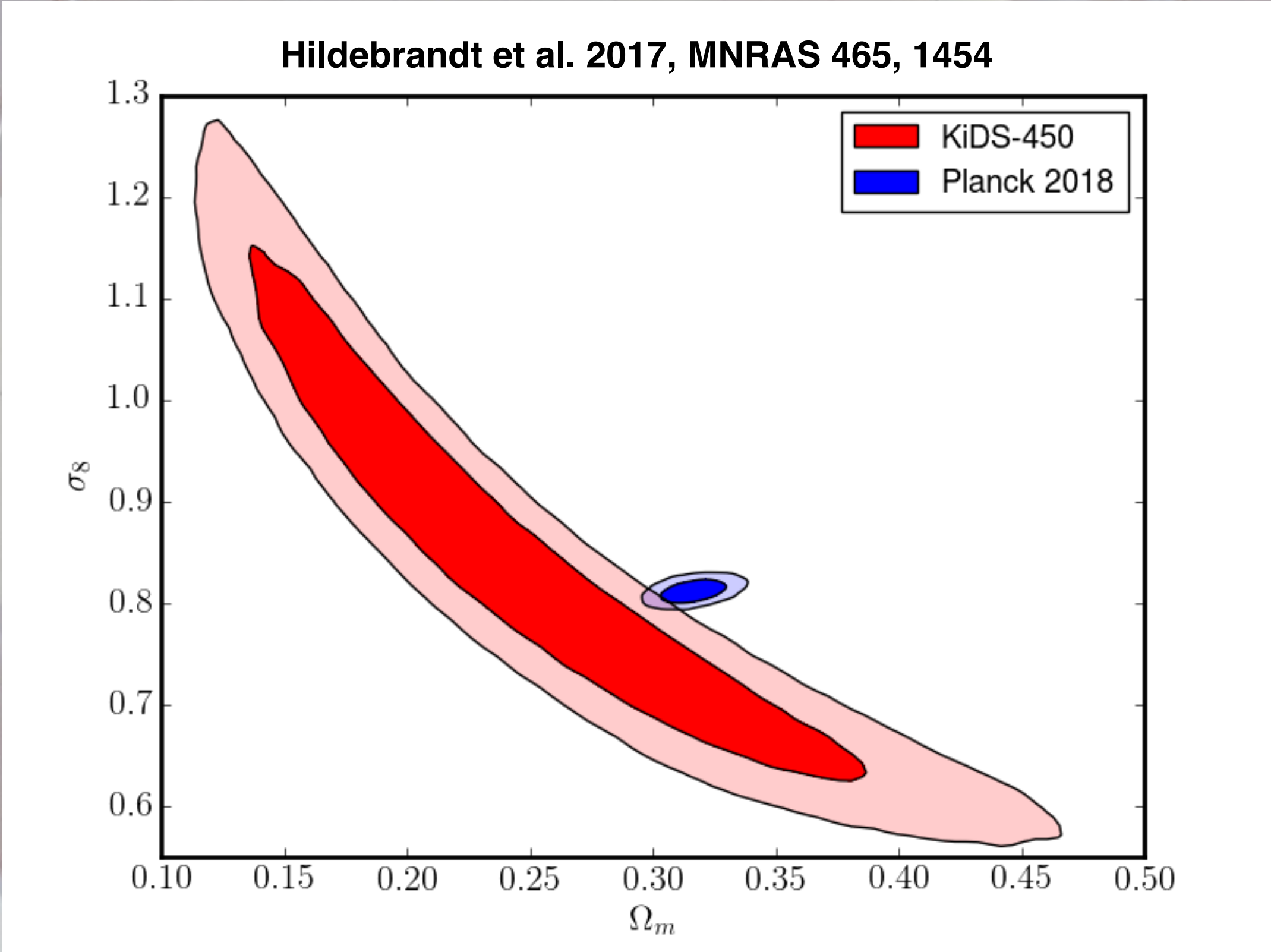


Latest KiDS cosmic shear results: **KiDS-VIKING-450**

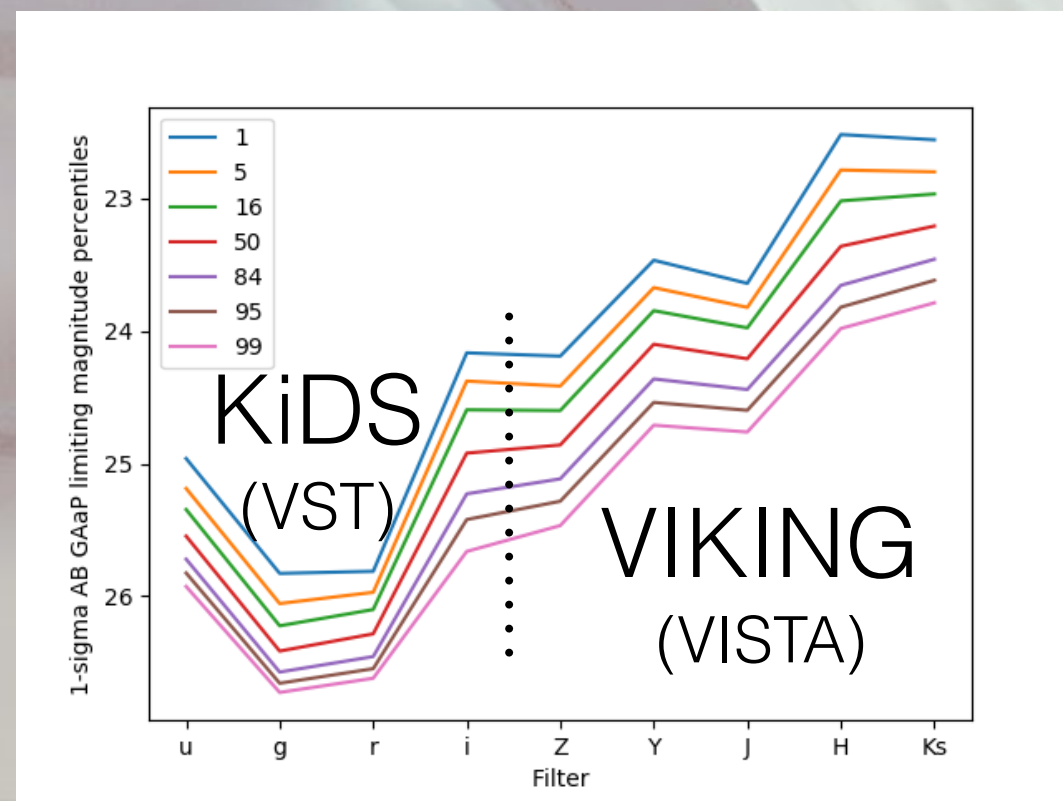
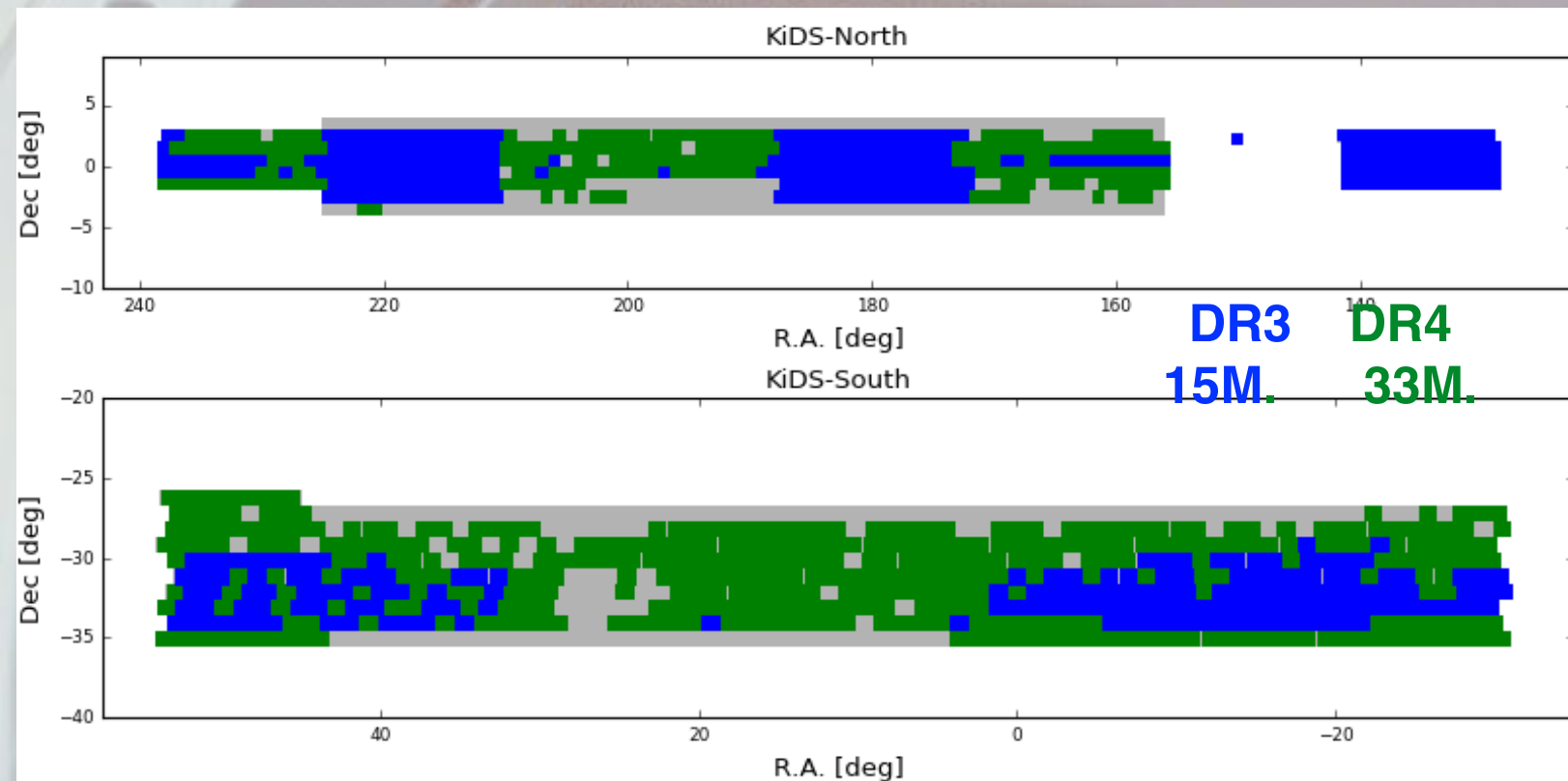
Konrad Kuijken
Leiden Observatory

Hildebrandt, Köhlinger, van den Busch, Heymans, Joachimi, **Arun Kannawadi**, Wright, + many more KiDS

**Alex Amon, Joachim Harnois-Déraps, Harry Johnston,
Naomie Robertson, Cristóbal Sifón**



KiDS The Kilo-Degree Survey: basic stats



Final area **1350 deg²**

Lensing images (r):

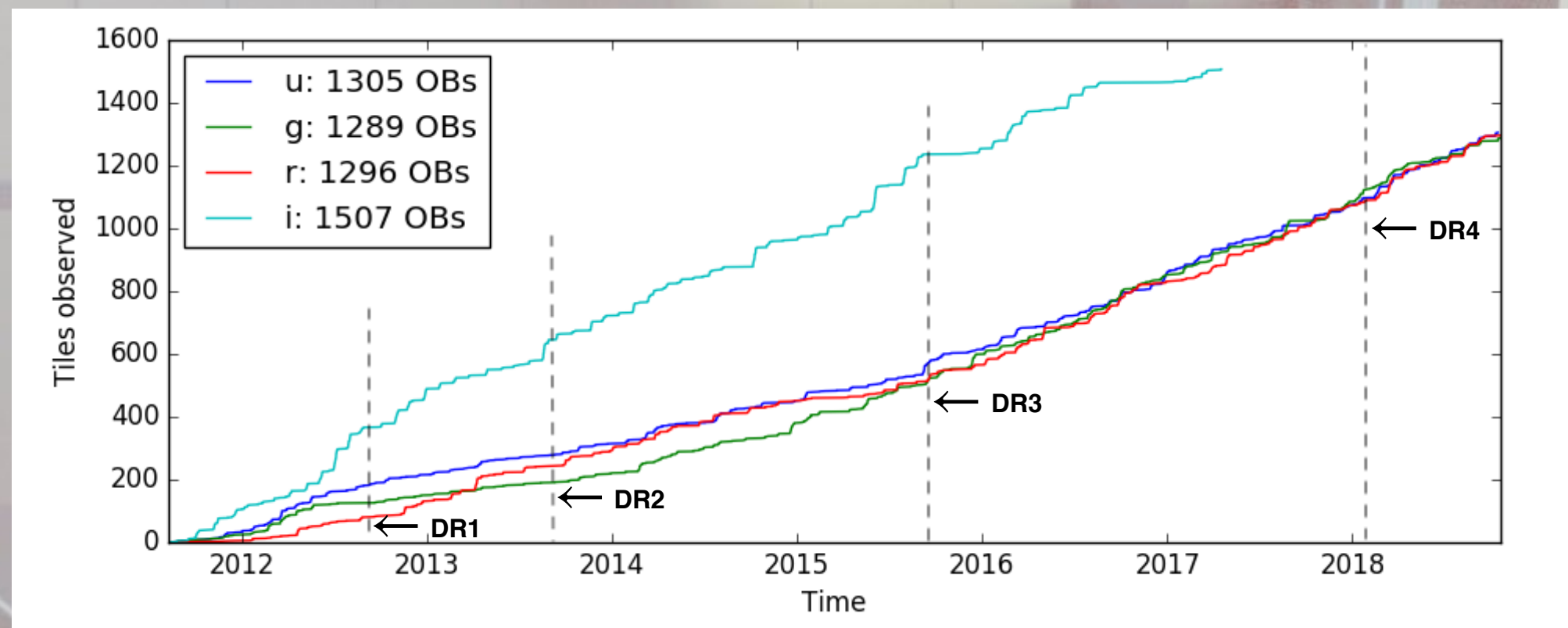
Seeing **0.7''**

Depth **25^m** (5 σ AB)

PSF Ellipticity 0.03

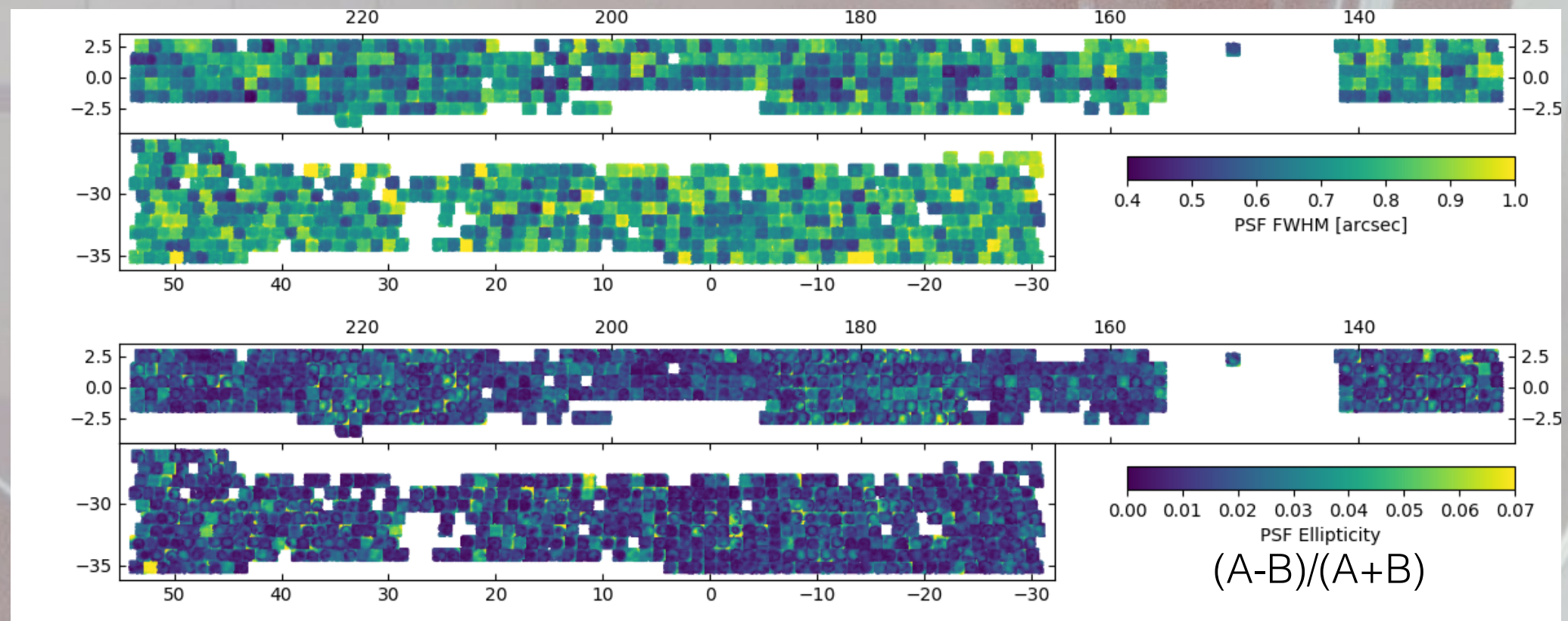
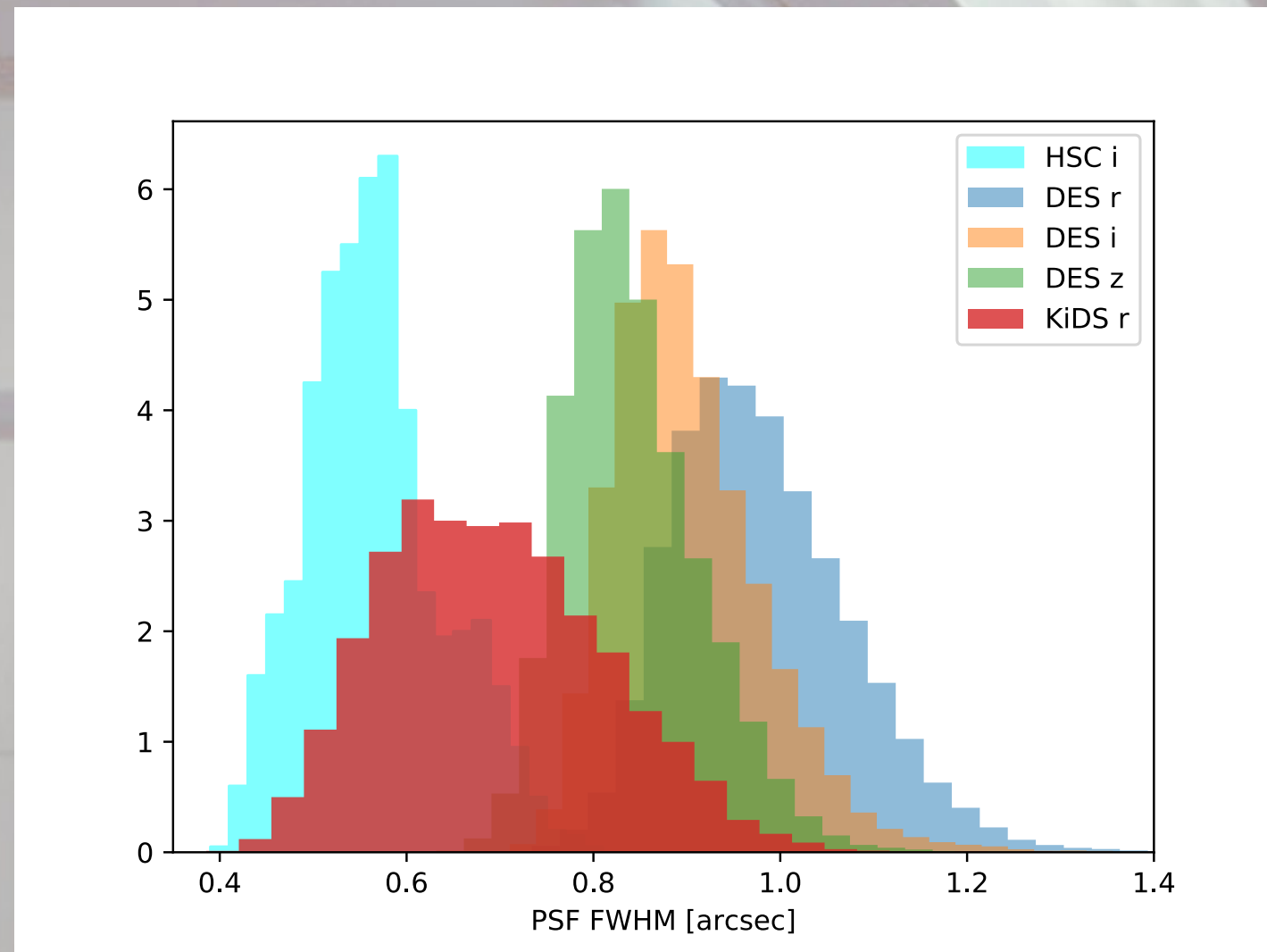
Colours:

ugriZYJHK_s



KiDS Image quality

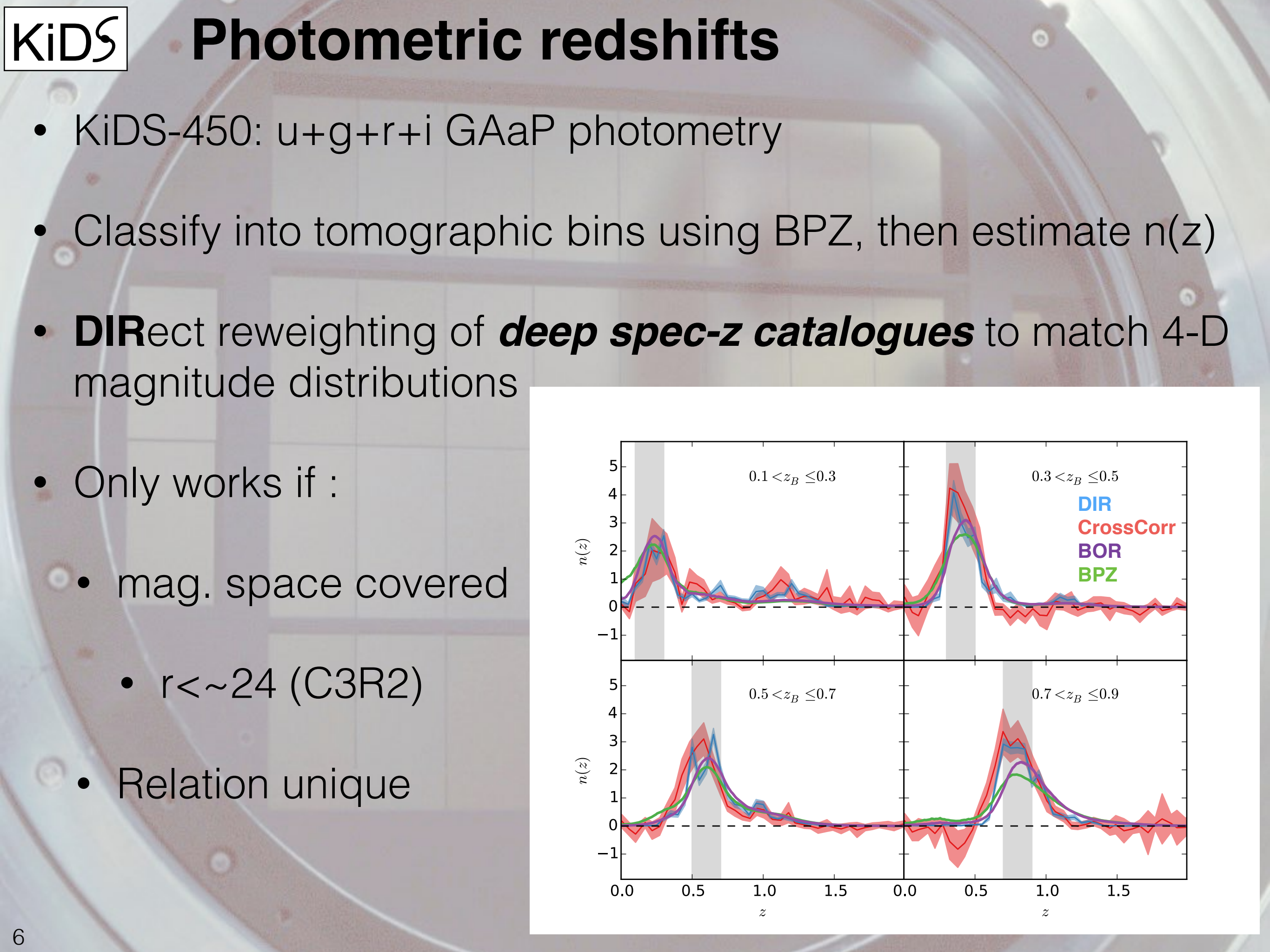
- Median 0.7"
- PSF ellipticity low
- Image scale constant
- Benign PSF patterns



- ***Lensfit***
 - forward modelling code,
 - simultaneous fit to all sub-exposures with individual PSFs,
 - likelihood function for each source ellipticity.
- Self-calibration (response of best-fit model to shear)
- Calibration verified with **realistic image simulations**
- Conservative estimate of systematic shear calibration error: 2%.

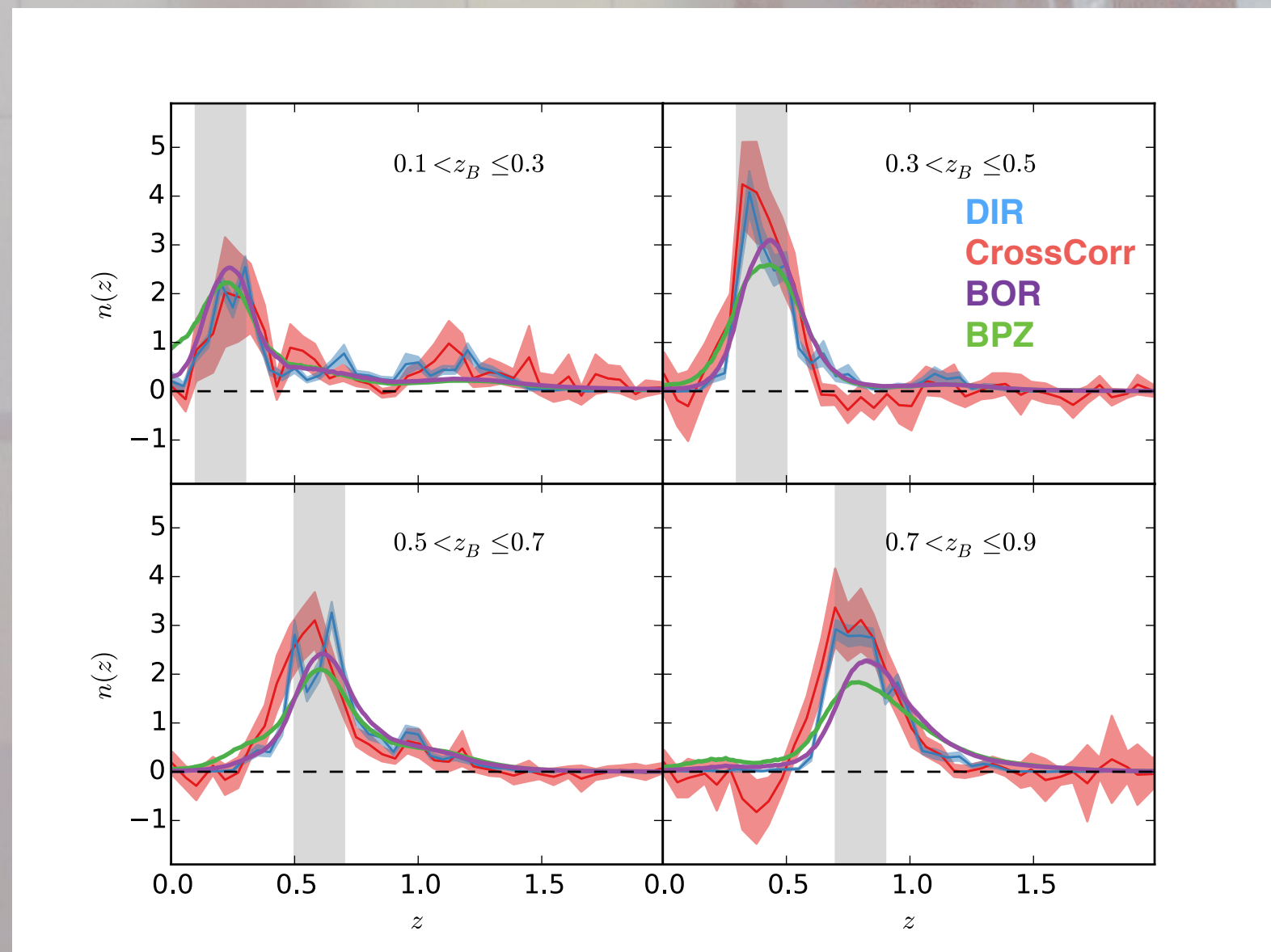


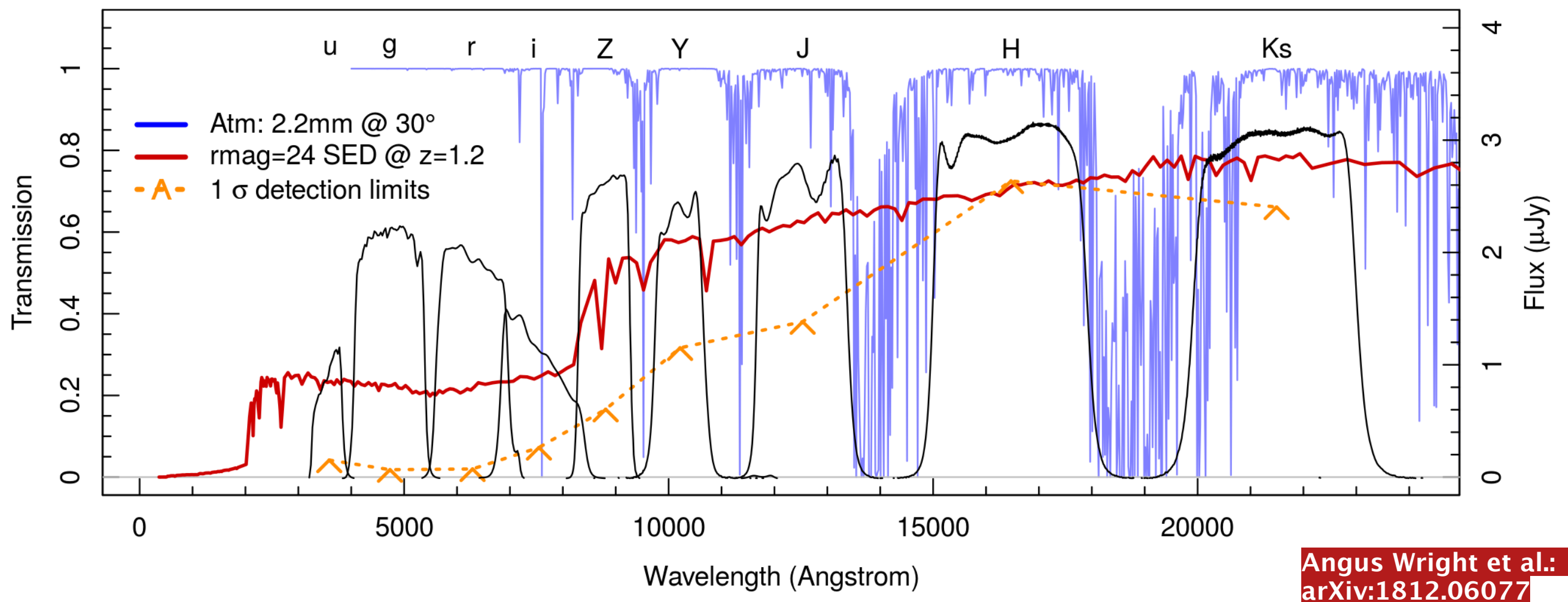
Arun Kannawadi et al.:
arXiv:1812.03983



Photometric redshifts

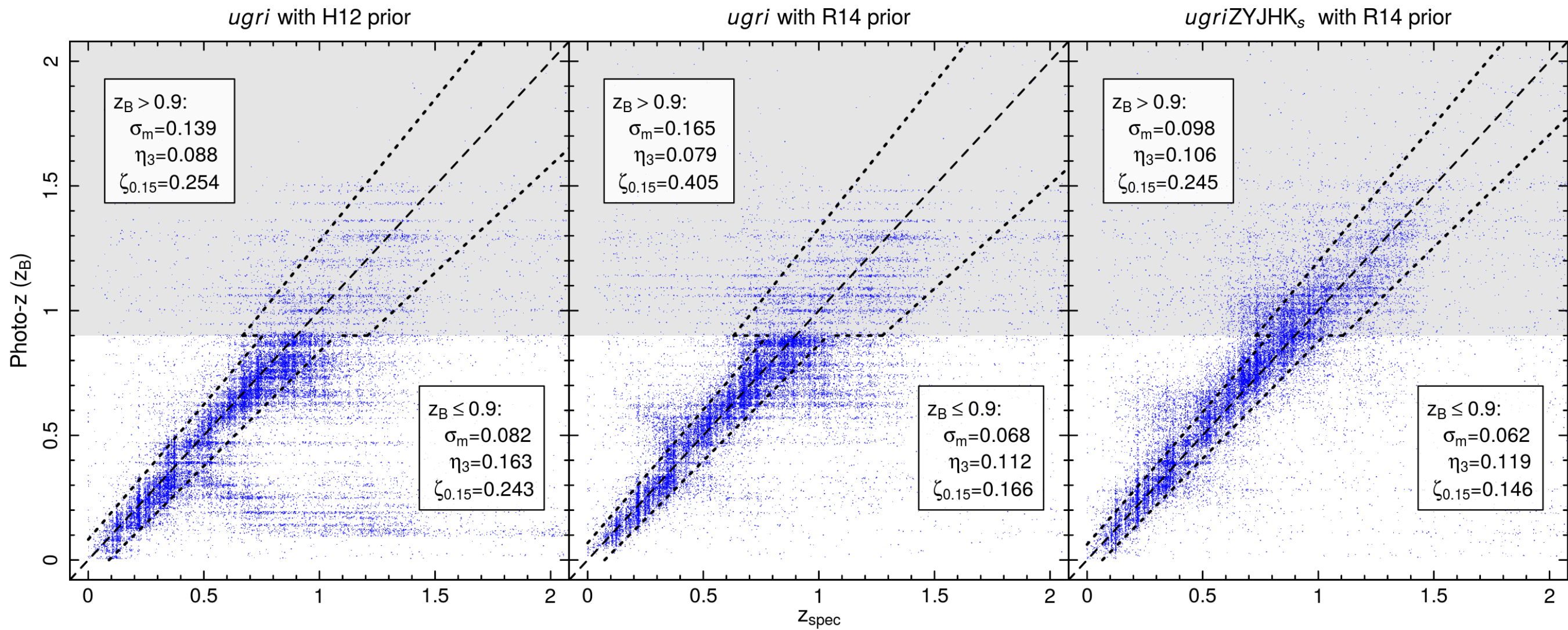
- KiDS-450: u+g+r+i GAaP photometry
- Classify into tomographic bins using BPZ, then estimate $n(z)$
- **DIR**ect reweighting of **deep spec-z catalogues** to match 4-D magnitude distributions
- Only works if :
 - mag. space covered
 - $r < \sim 24$ (C3R2)
- Relation unique

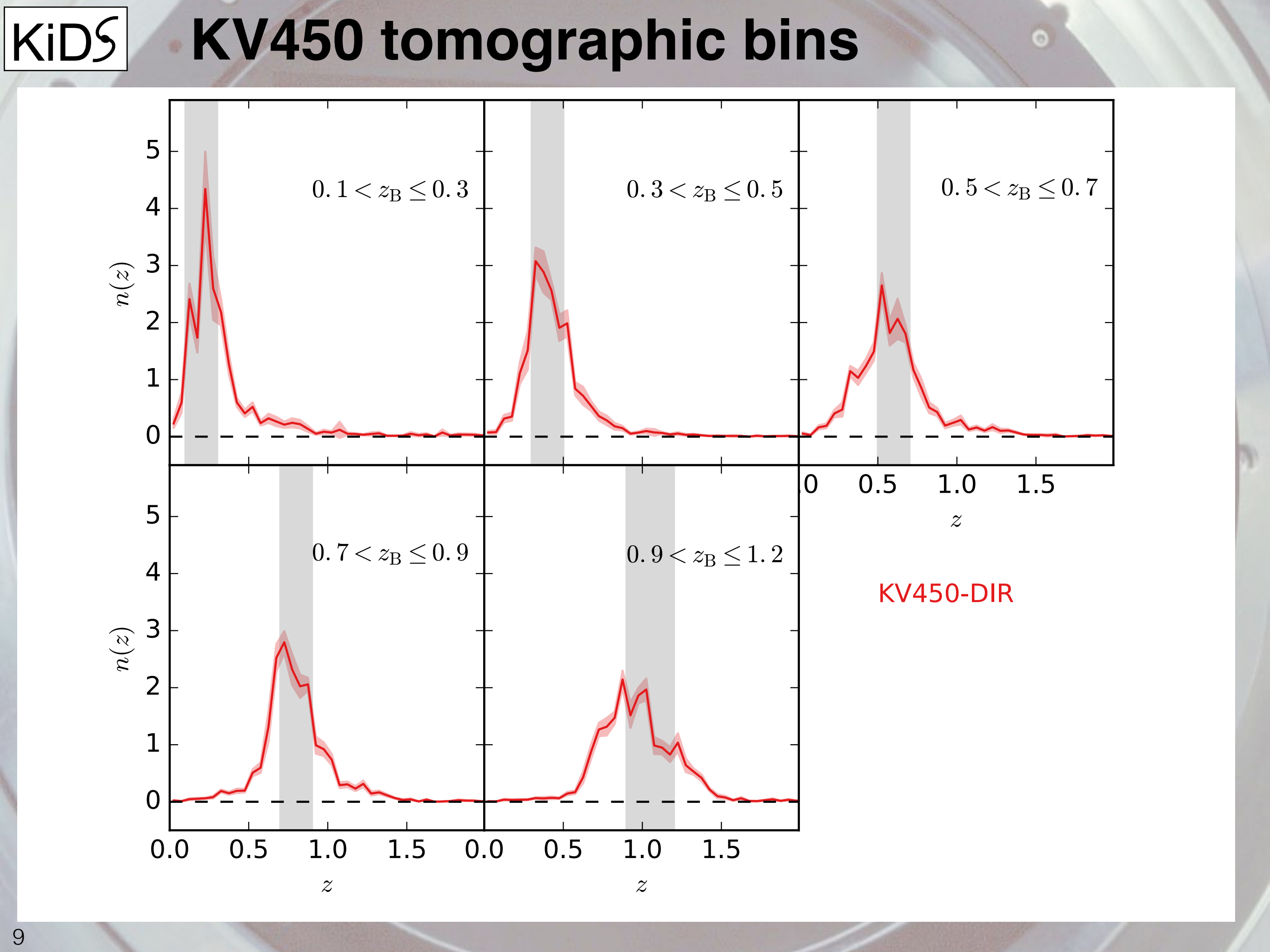




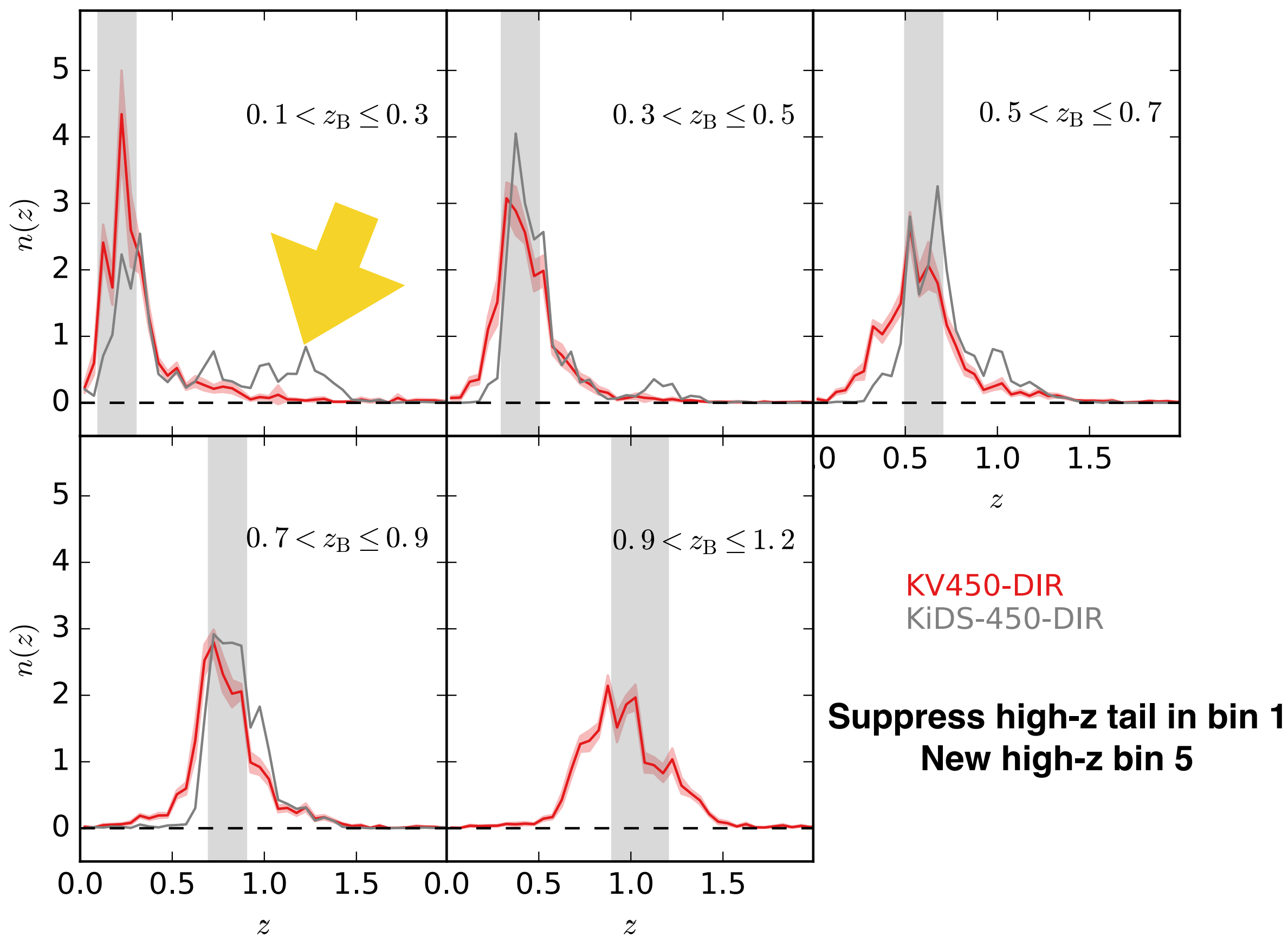
- Include 5-band VIKING survey ZYJHK_s (depth 23-21, 5σ)

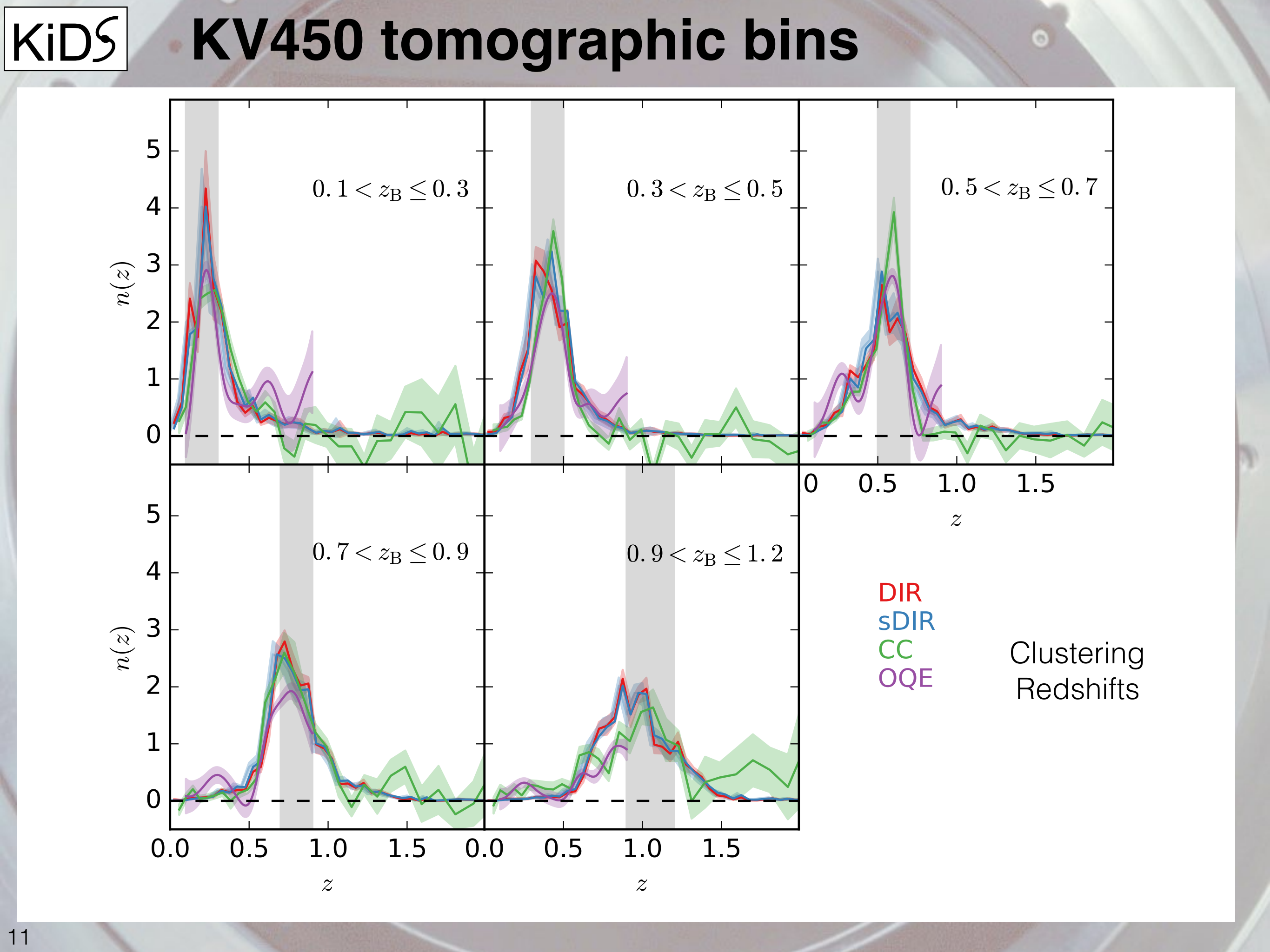
Benefits of adding near-IR





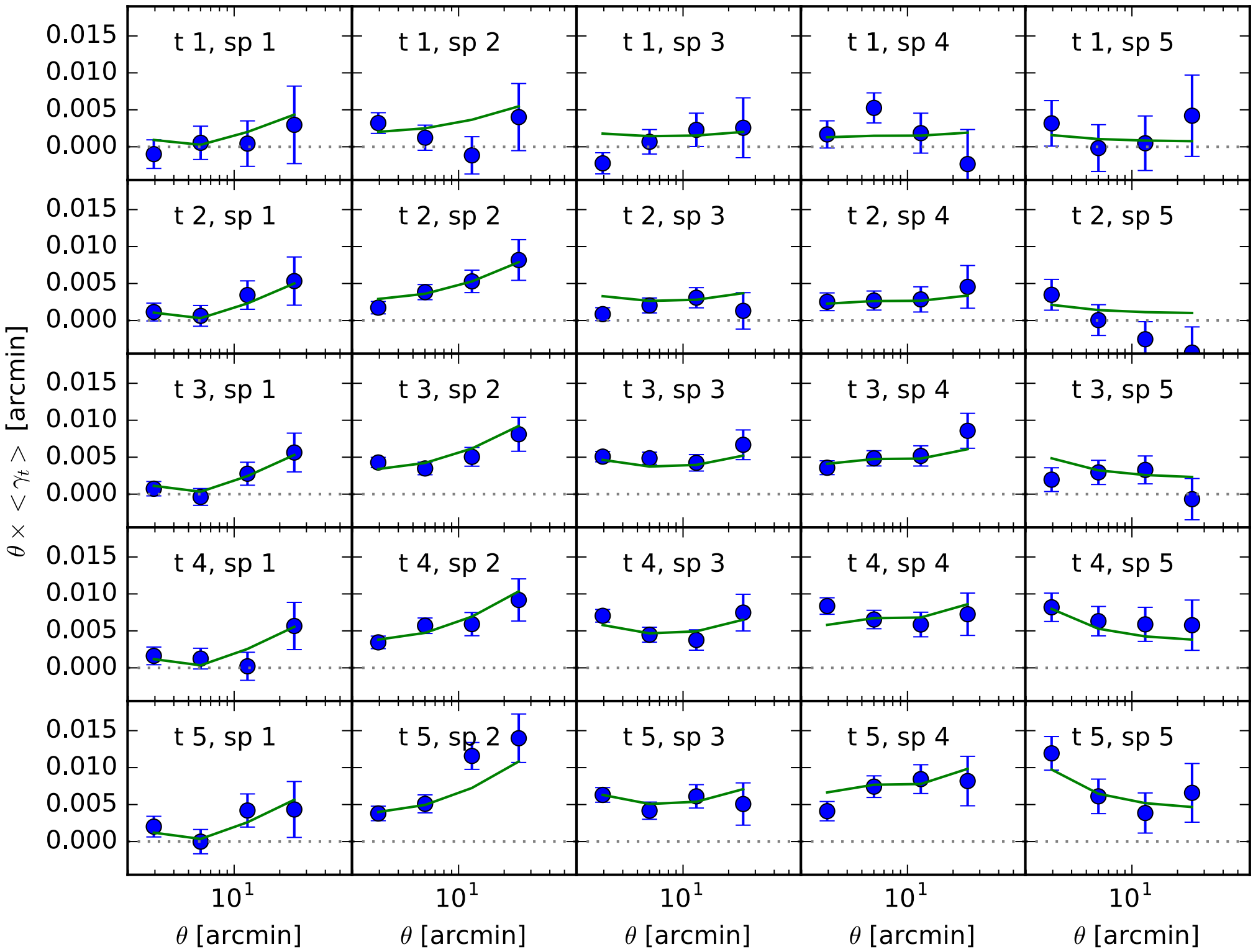
KV450 tomographic bins

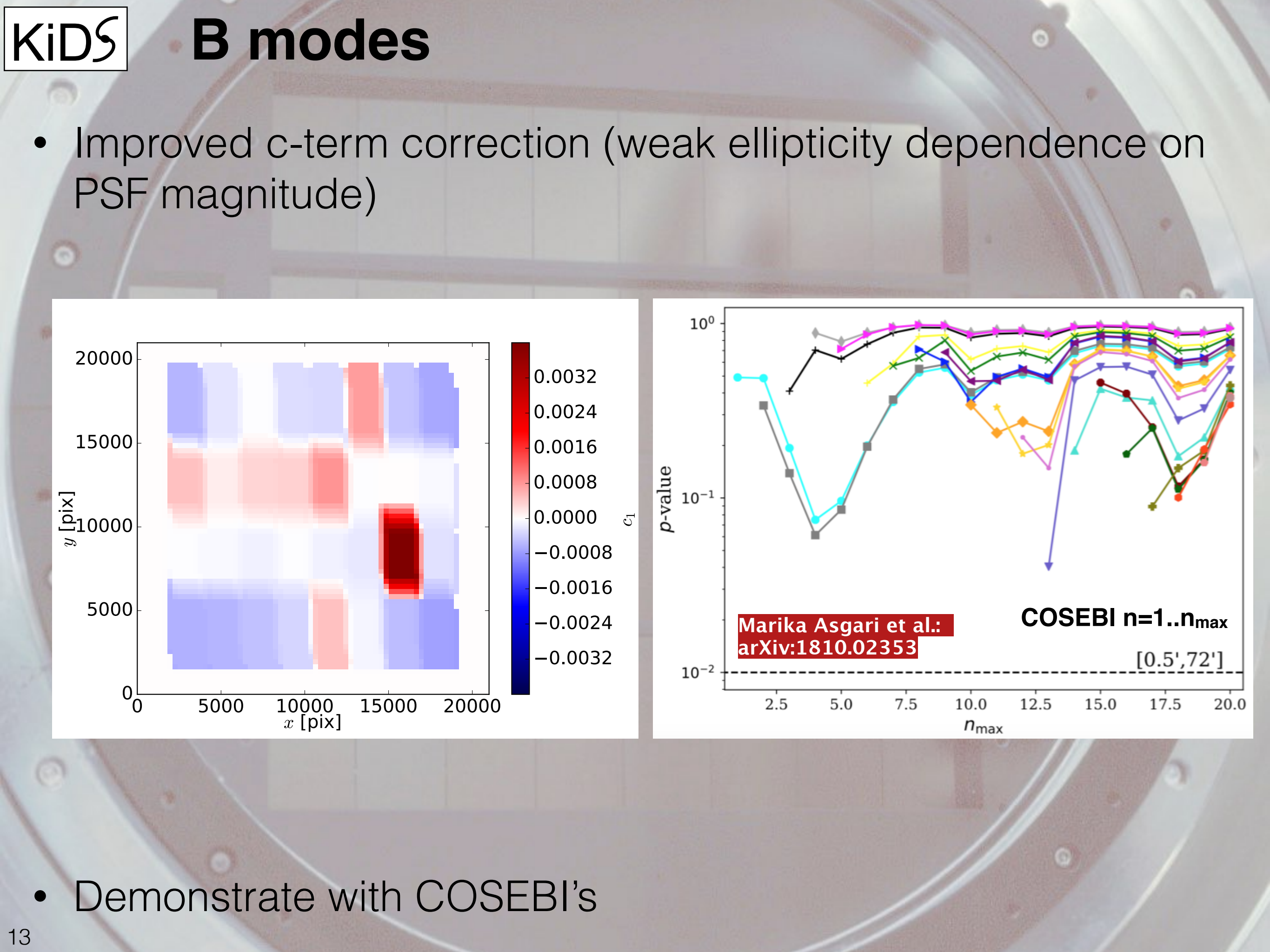




Shear ratio test

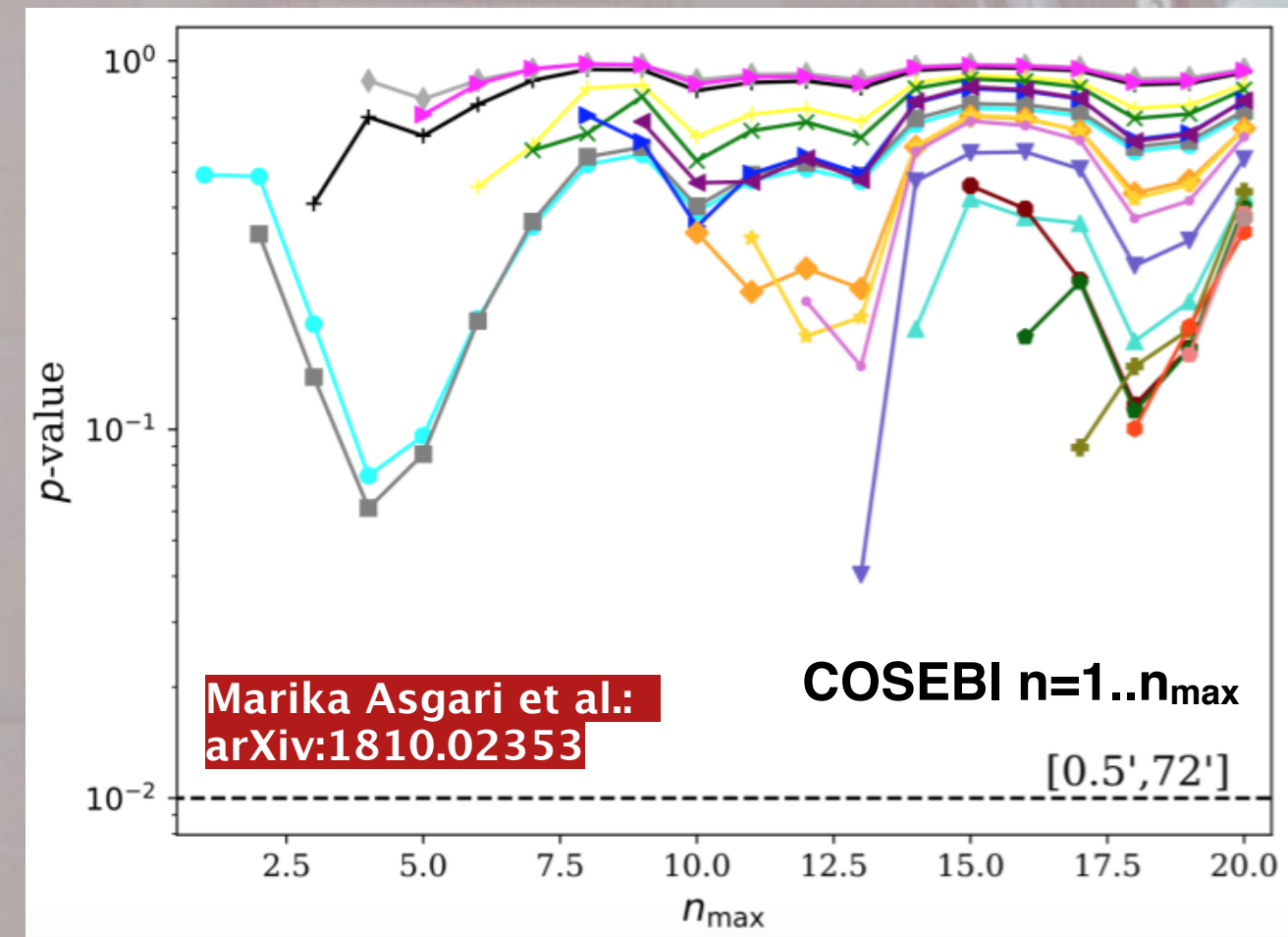
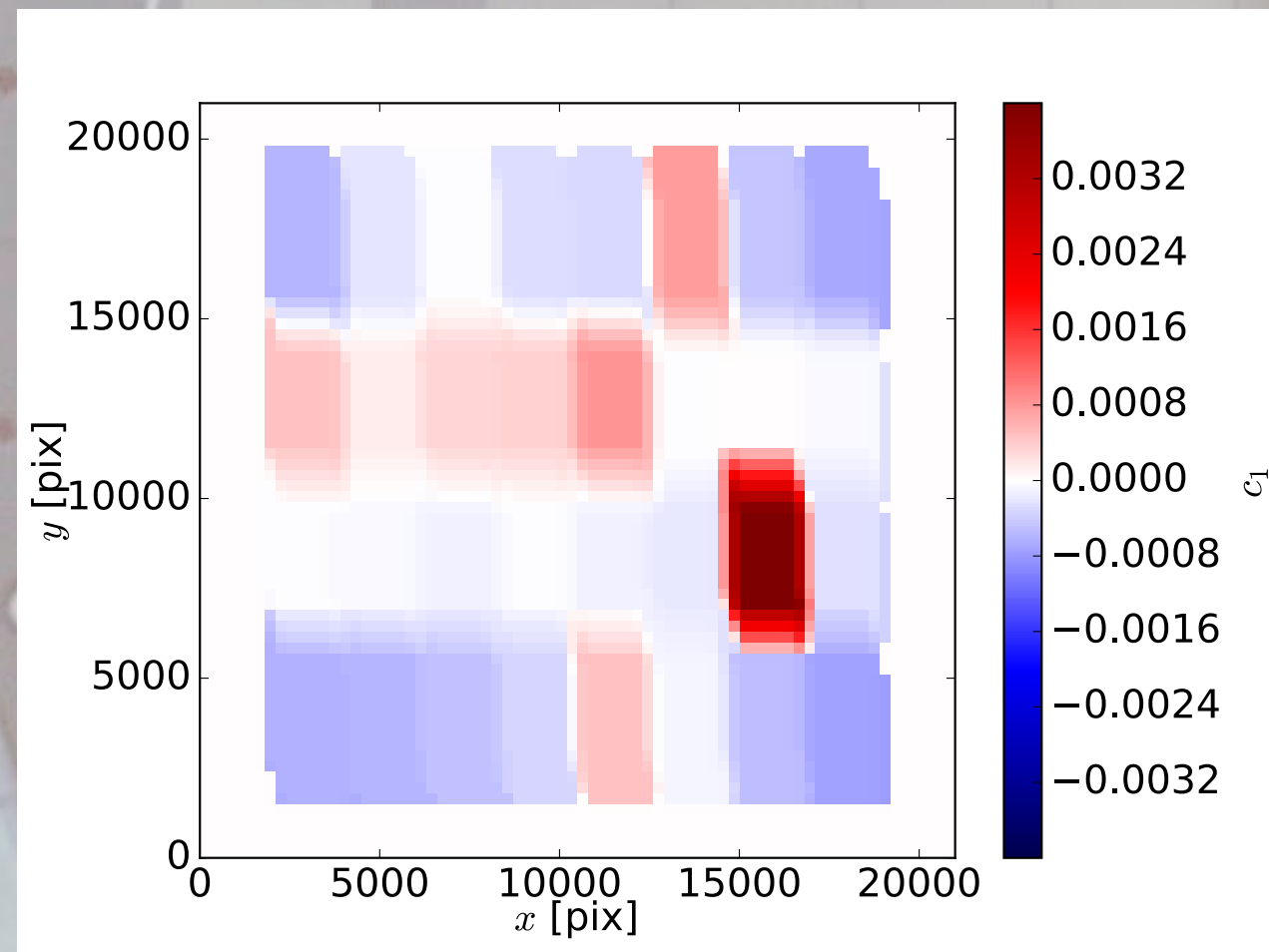
GAMA lenses, KV450 sources, $\chi^2/\text{dof}=1.05$, $p=35.3\%$



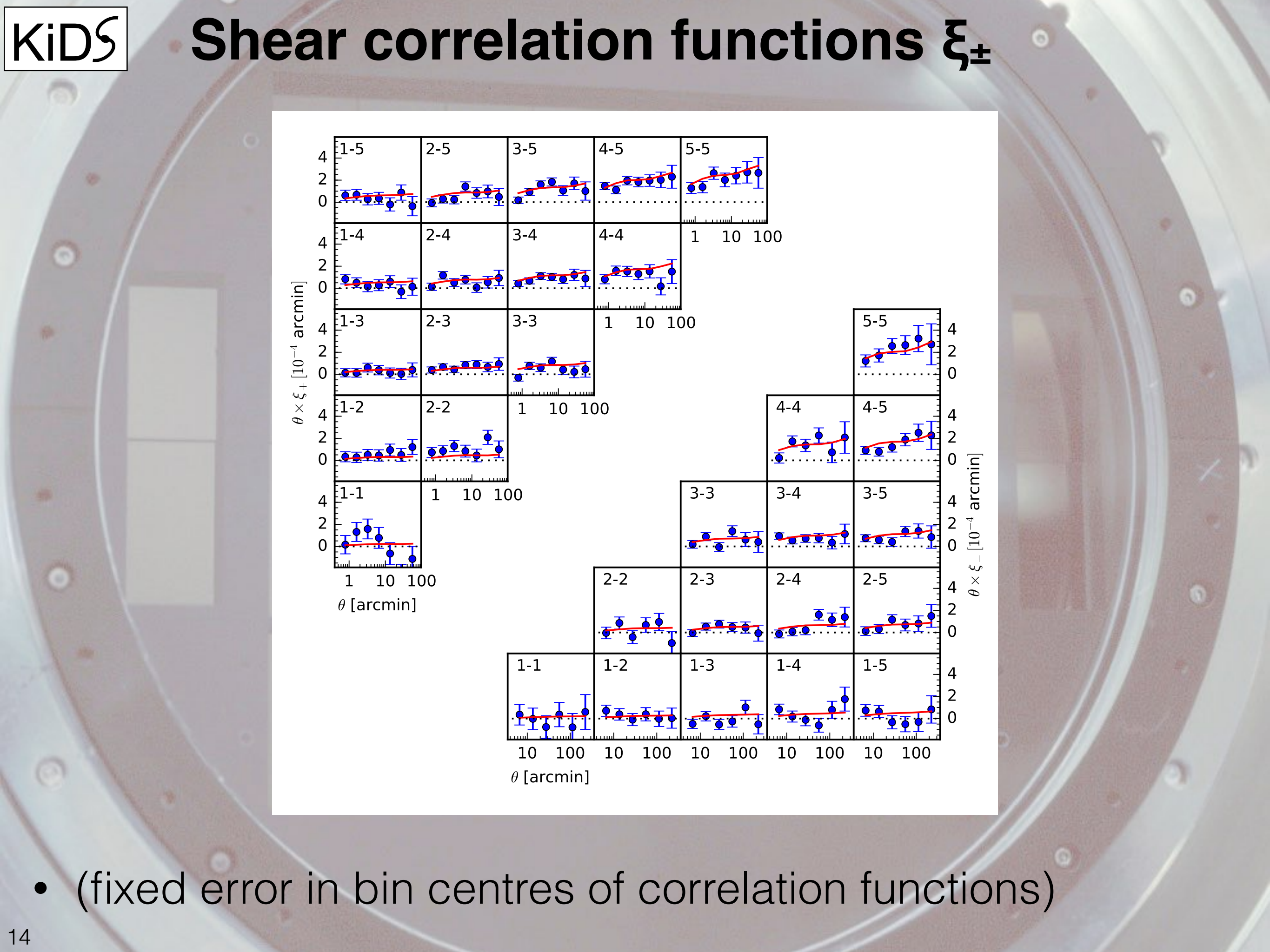


B modes

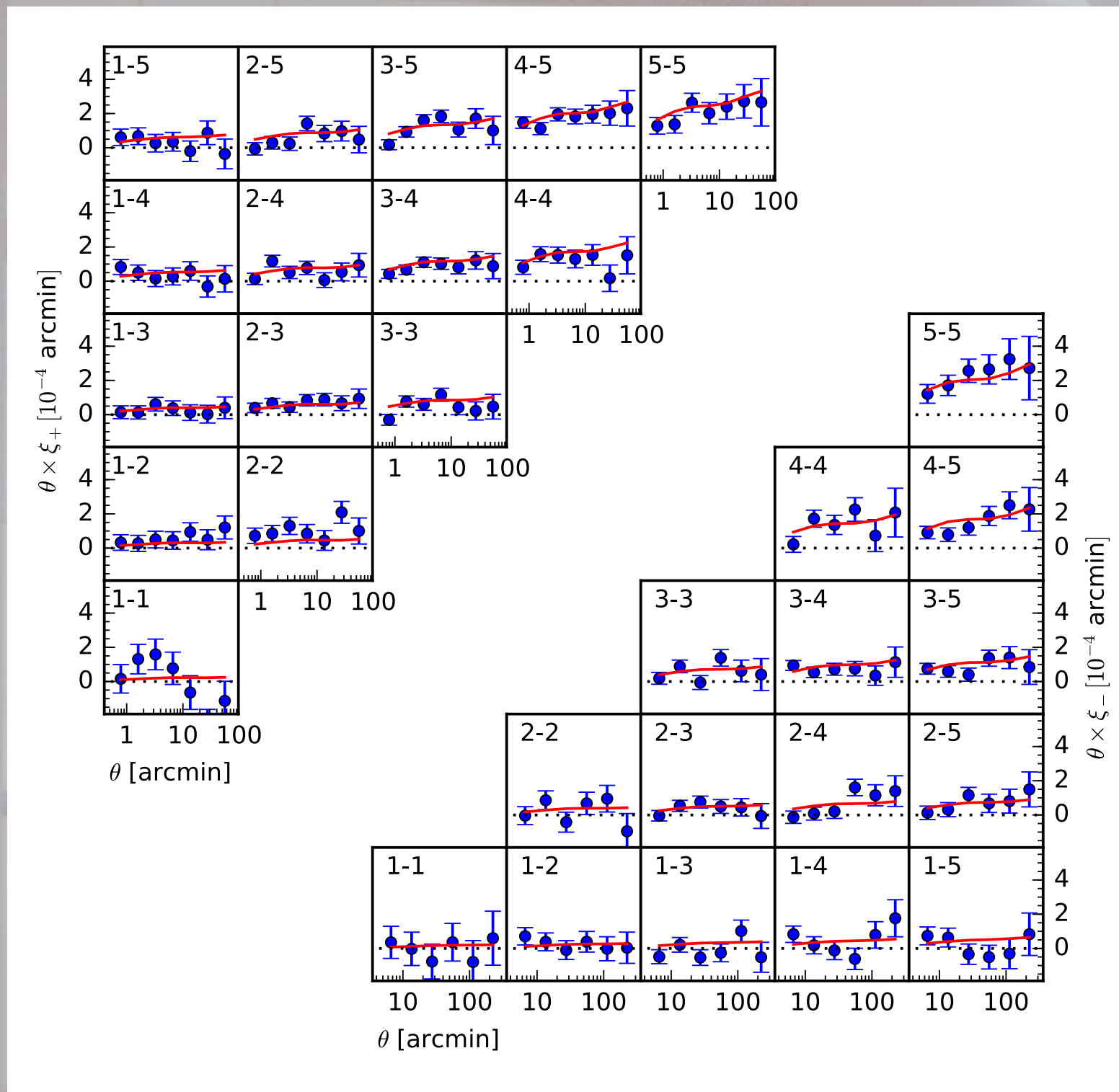
- Improved c-term correction (weak ellipticity dependence on PSF magnitude)



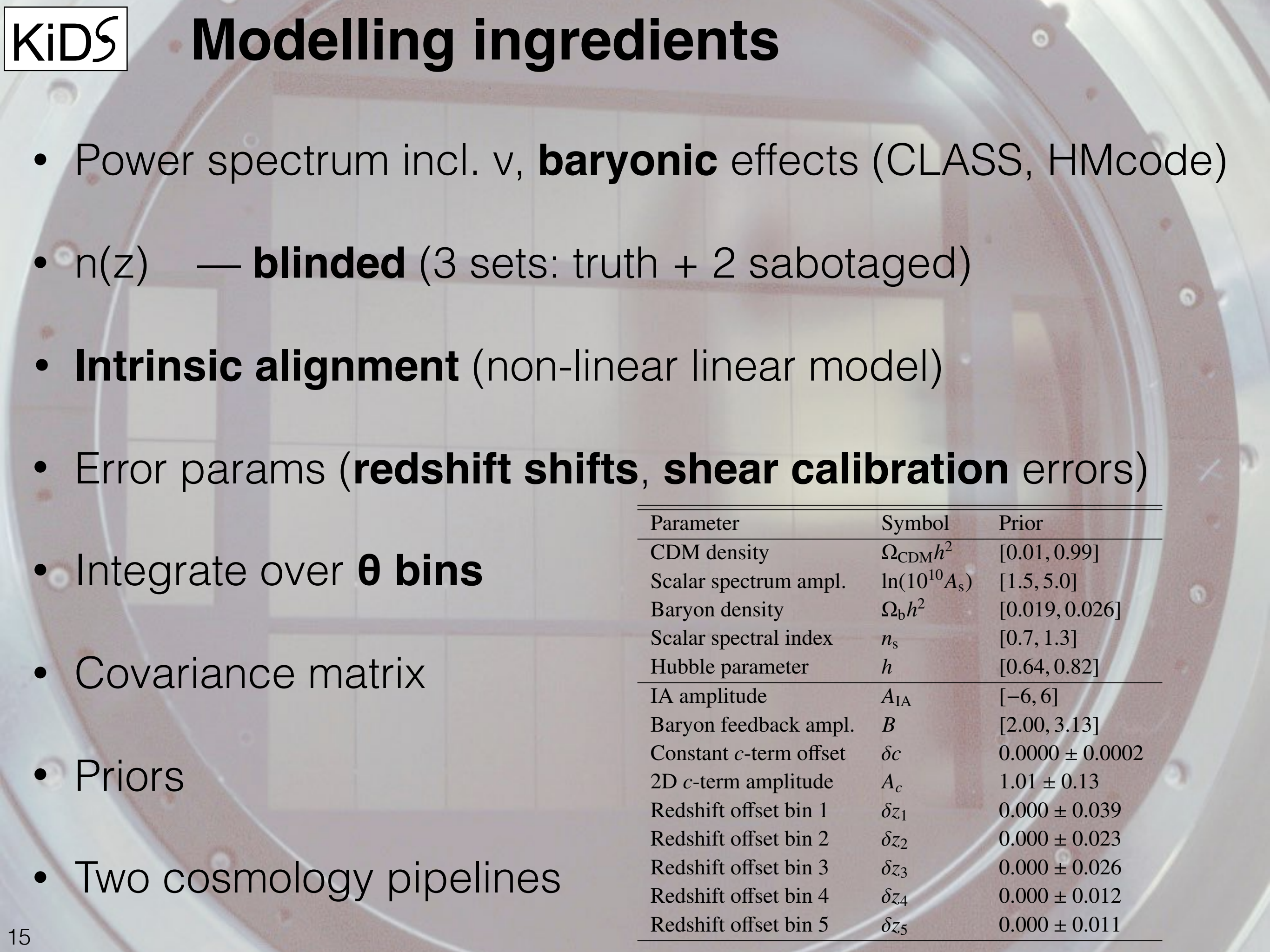
- Demonstrate with COSEBI's



Shear correlation functions ξ_{\pm}



- (fixed error in bin centres of correlation functions)

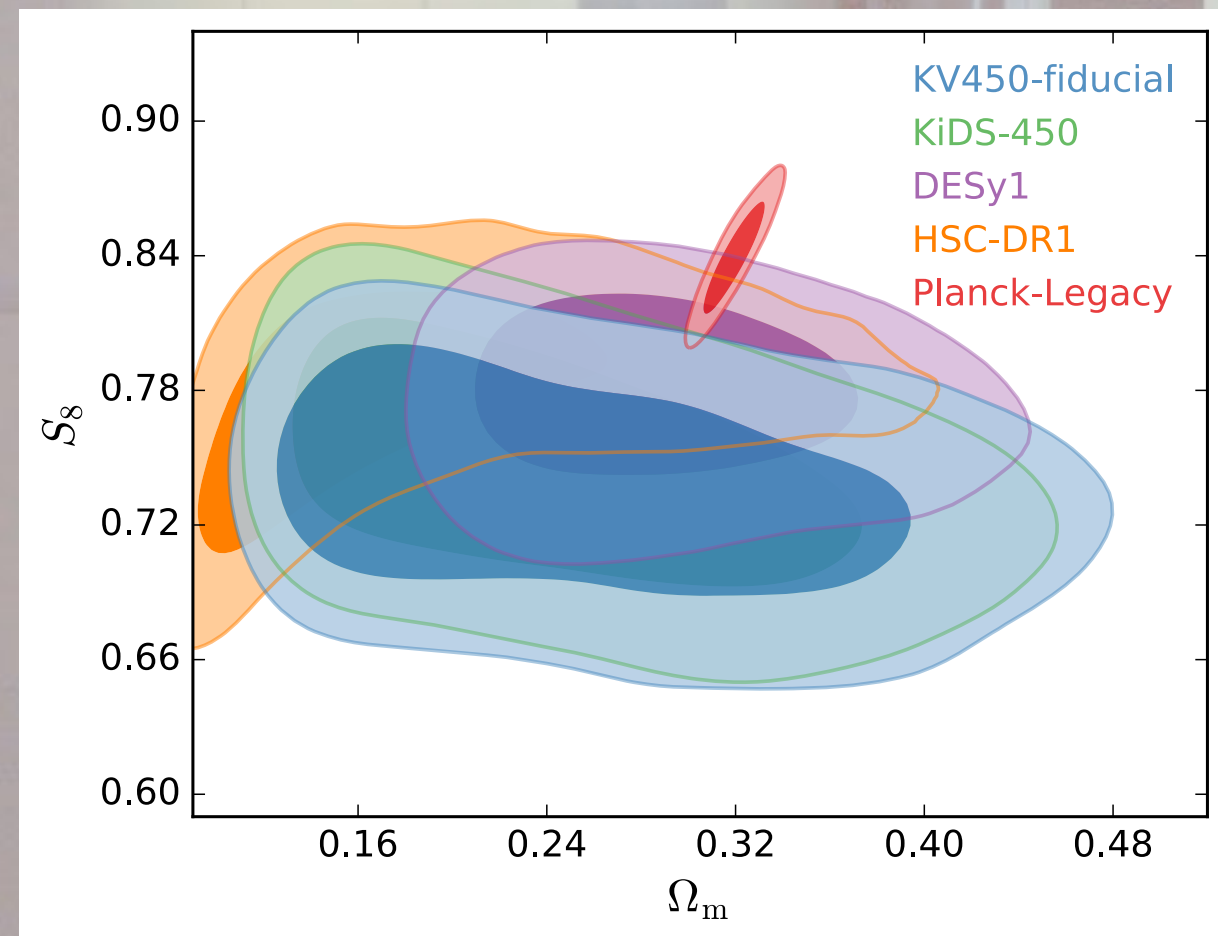
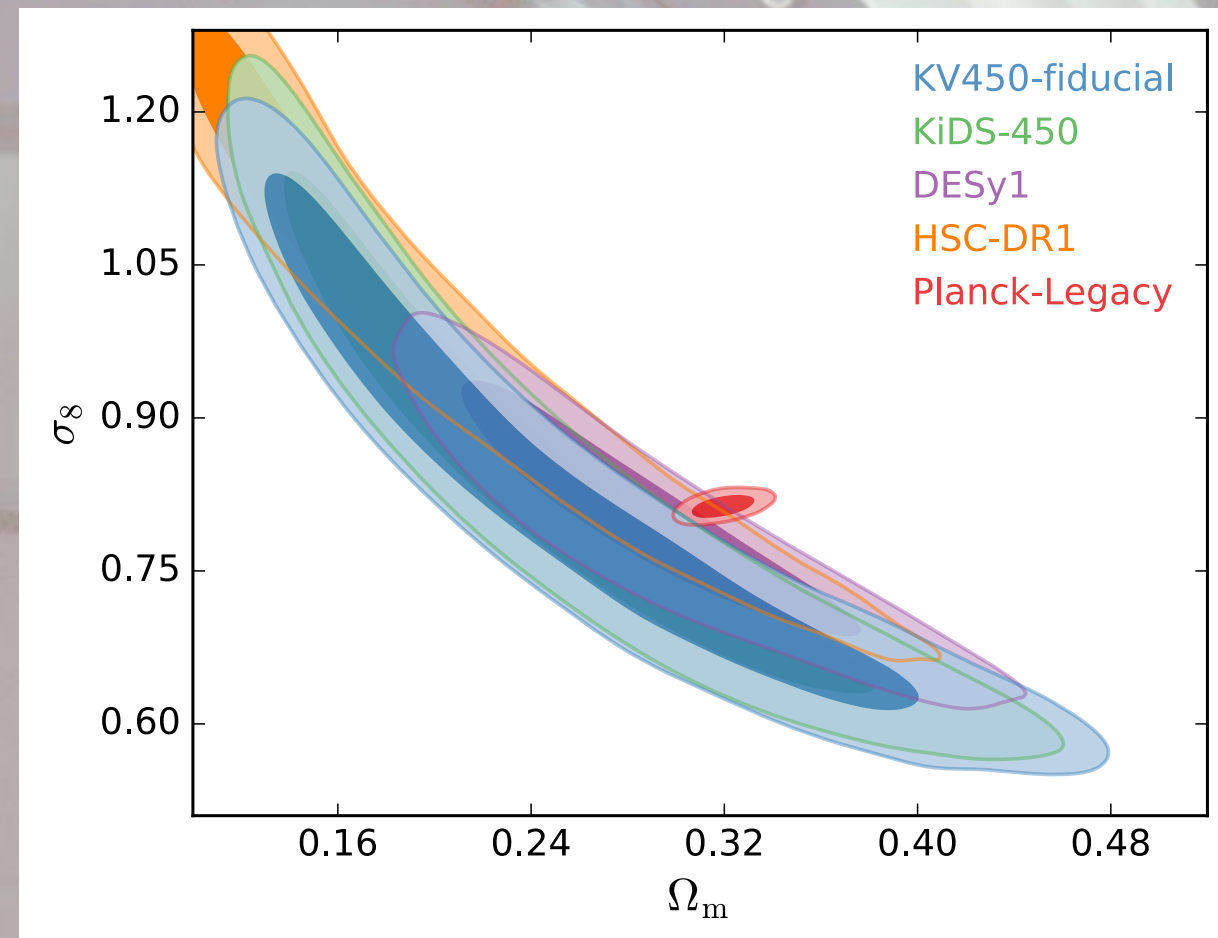


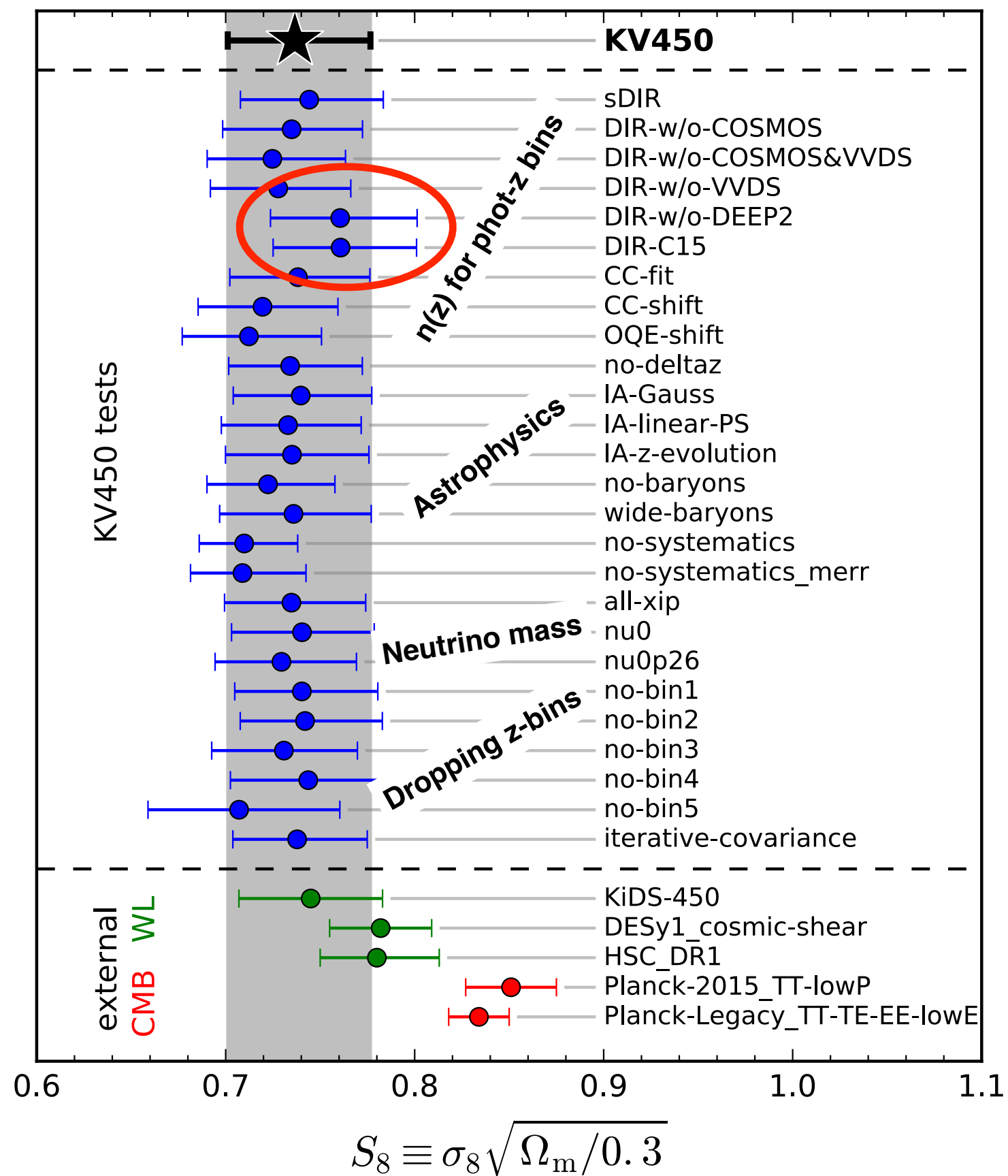
Modelling ingredients

- Power spectrum incl. v, **baryonic** effects (CLASS, HMcode)
- $n(z)$ — **blinded** (3 sets: truth + 2 sabotaged)
- **Intrinsic alignment** (non-linear linear model)
- Error params (**redshift shifts**, **shear calibration** errors)
- Integrate over **θ bins**
- Covariance matrix
- Priors
- Two cosmology pipelines

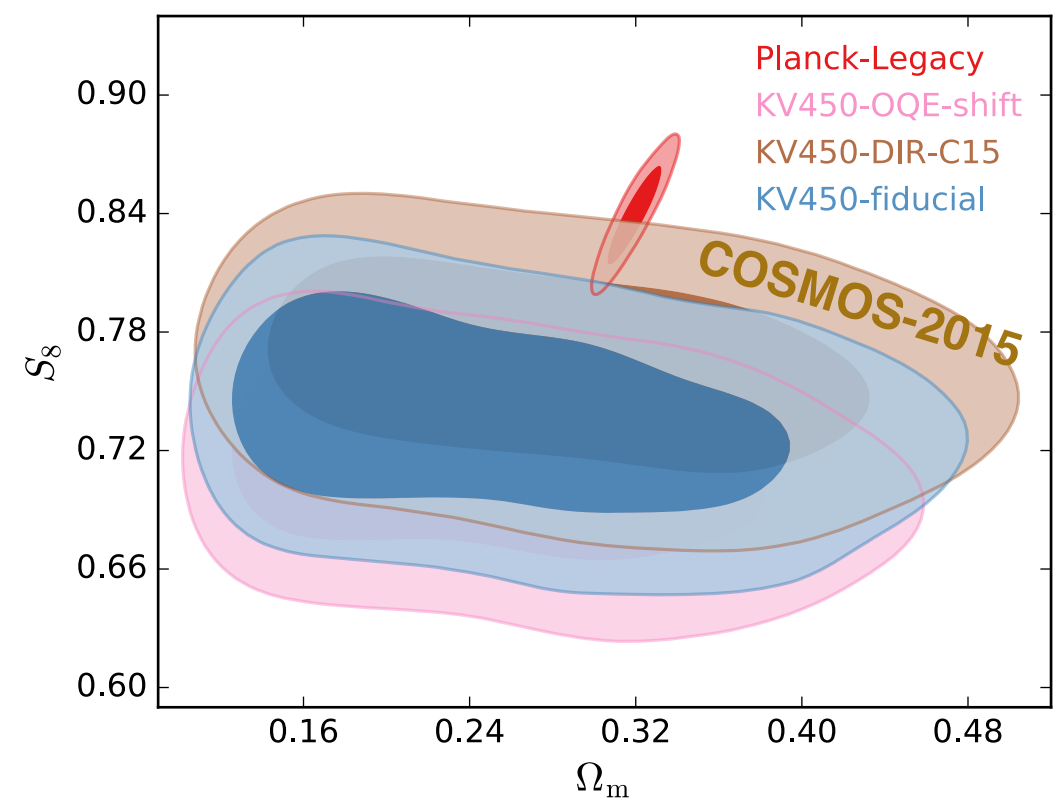
Parameter	Symbol	Prior
CDM density	$\Omega_{\text{CDM}}h^2$	[0.01, 0.99]
Scalar spectrum ampl.	$\ln(10^{10}A_s)$	[1.5, 5.0]
Baryon density	$\Omega_b h^2$	[0.019, 0.026]
Scalar spectral index	n_s	[0.7, 1.3]
Hubble parameter	h	[0.64, 0.82]
IA amplitude	A_{IA}	[-6, 6]
Baryon feedback ampl.	B	[2.00, 3.13]
Constant c -term offset	δc	0.0000 ± 0.0002
2D c -term amplitude	A_c	1.01 ± 0.13
Redshift offset bin 1	δz_1	0.000 ± 0.039
Redshift offset bin 2	δz_2	0.000 ± 0.023
Redshift offset bin 3	δz_3	0.000 ± 0.026
Redshift offset bin 4	δz_4	0.000 ± 0.012
Redshift offset bin 5	δz_5	0.000 ± 0.011

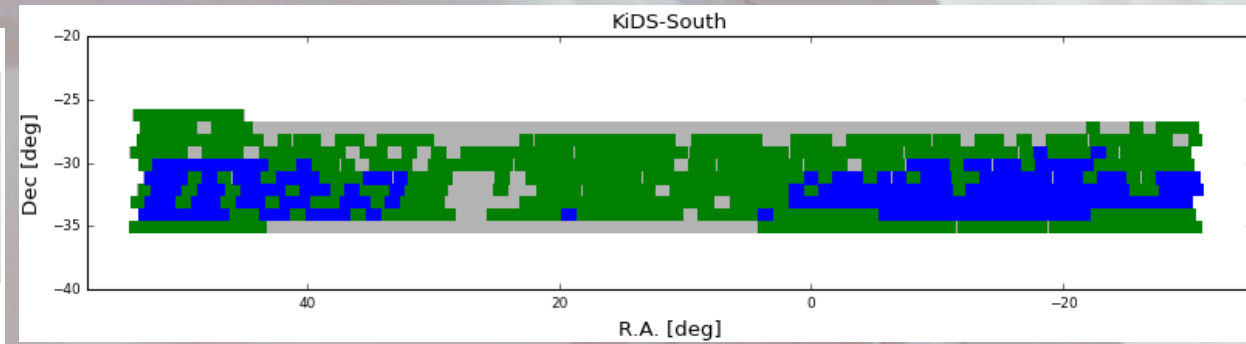
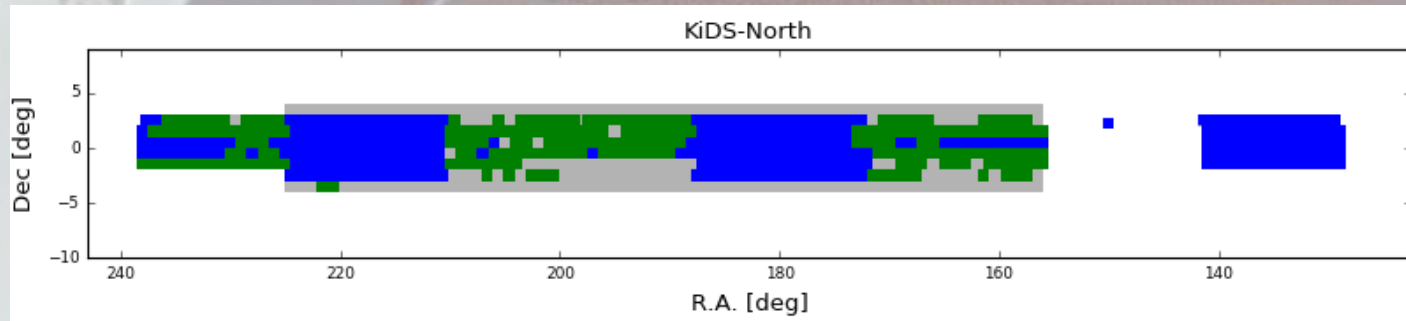
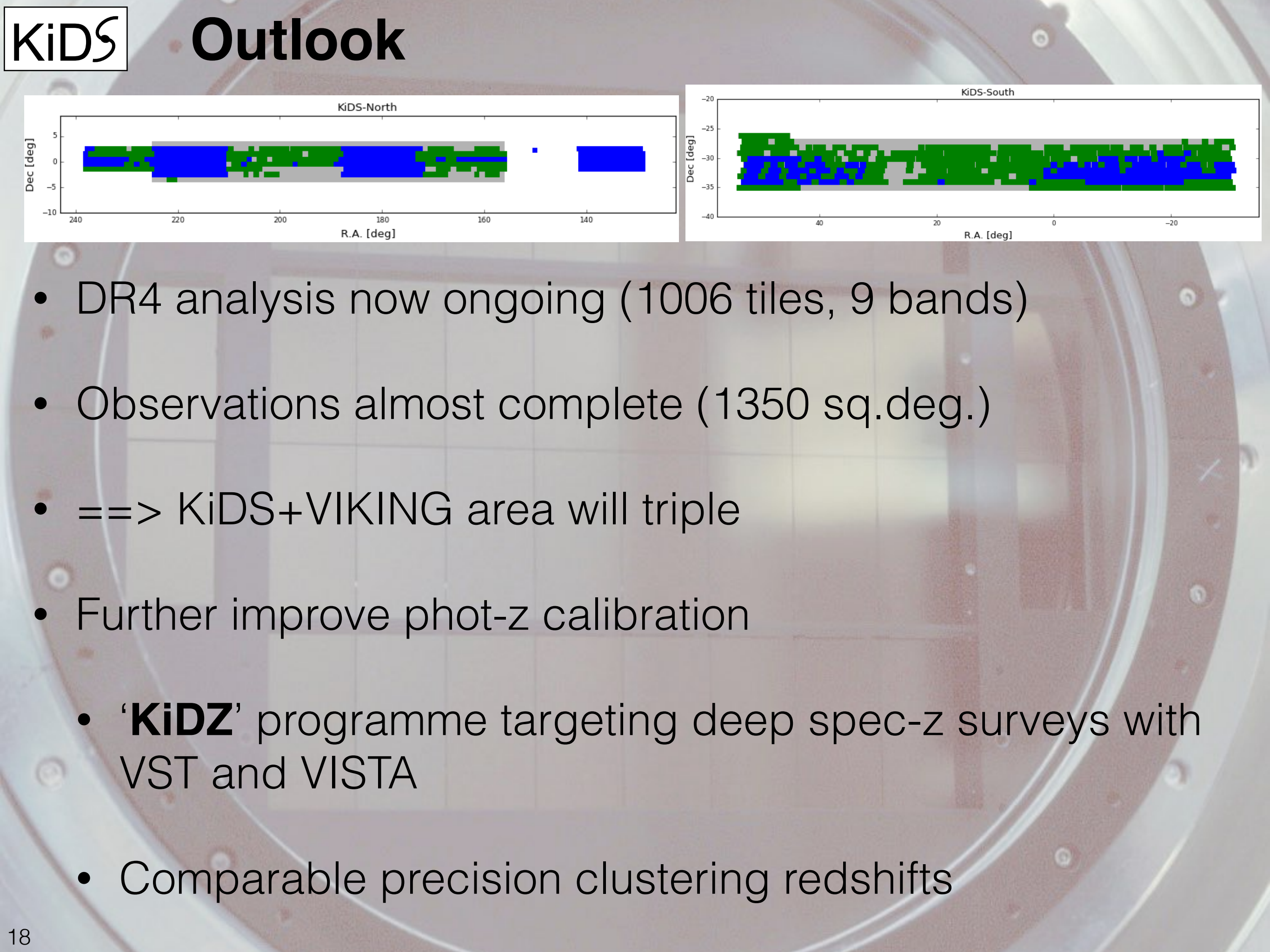
- Despite the changes, KV450 constraints very close to KiDS-450.
- Min. $\chi^2_{181} = 180.6$
- 2.3σ 'tension' with Planck
- $S_8 = 0.737^{+0.040}_{-0.036}$



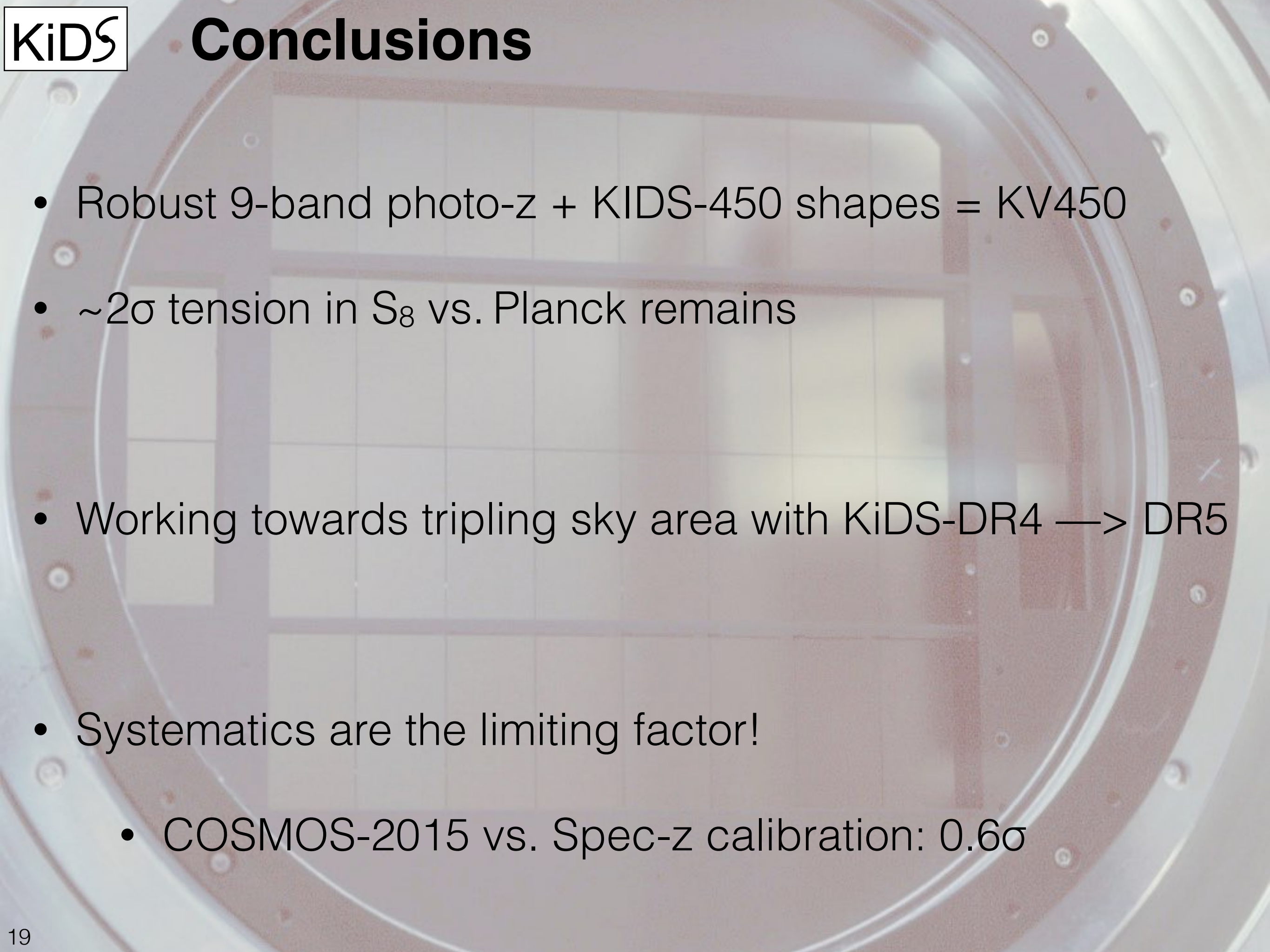


Deep spec-z crucial



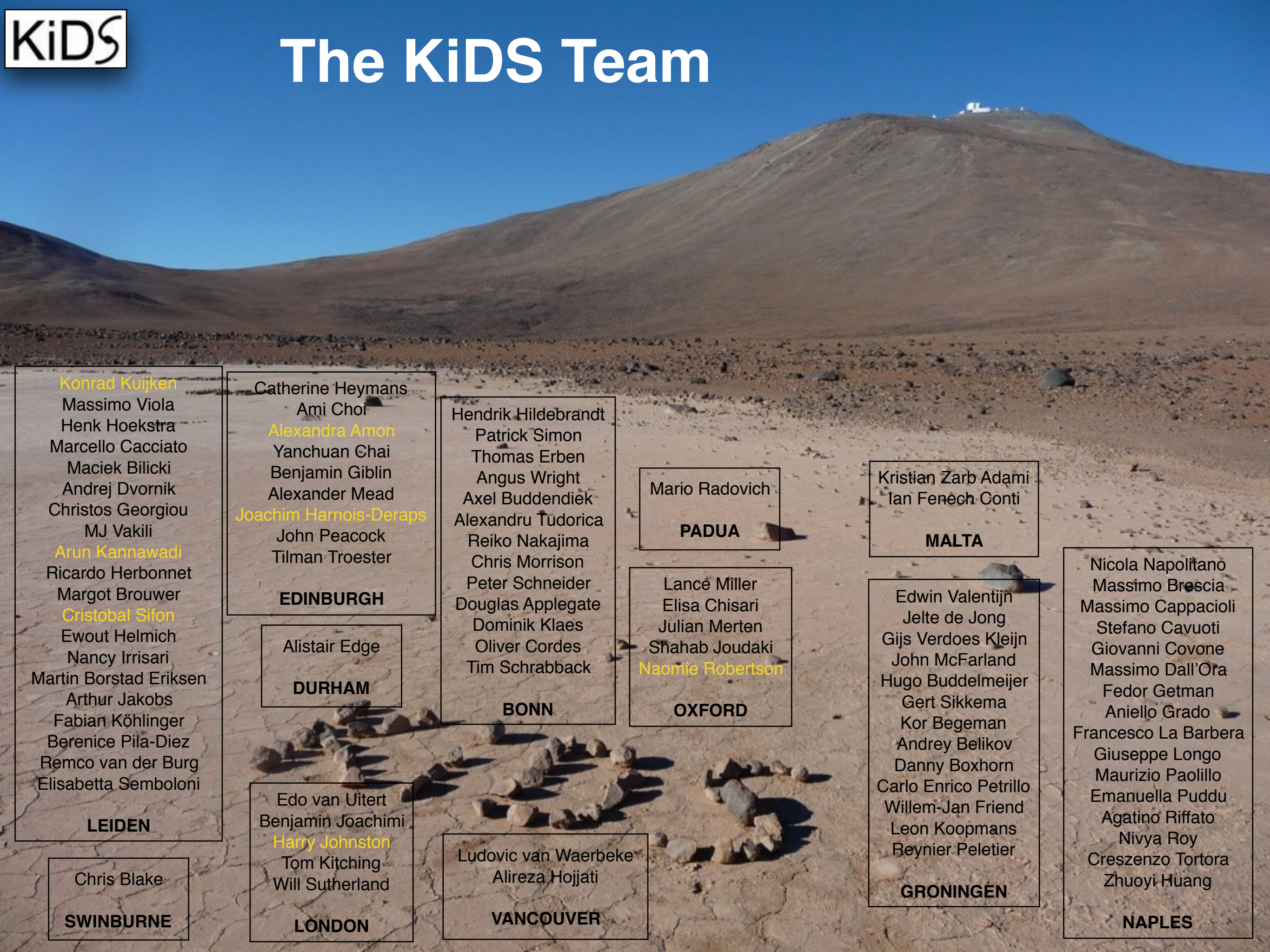


- DR4 analysis now ongoing (1006 tiles, 9 bands)
- Observations almost complete (1350 sq.deg.)
- ==> KiDS+VIKING area will triple
- Further improve phot-z calibration
 - ‘**KiDZ**’ programme targeting deep spec-z surveys with VST and VISTA
- Comparable precision clustering redshifts



Conclusions

- Robust 9-band photo-z + KiDS-450 shapes = KV450
- $\sim 2\sigma$ tension in S_8 vs. Planck remains
- Working towards tripling sky area with KiDS-DR4 \rightarrow DR5
- Systematics are the limiting factor!
 - COSMOS-2015 vs. Spec-z calibration: 0.6σ



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