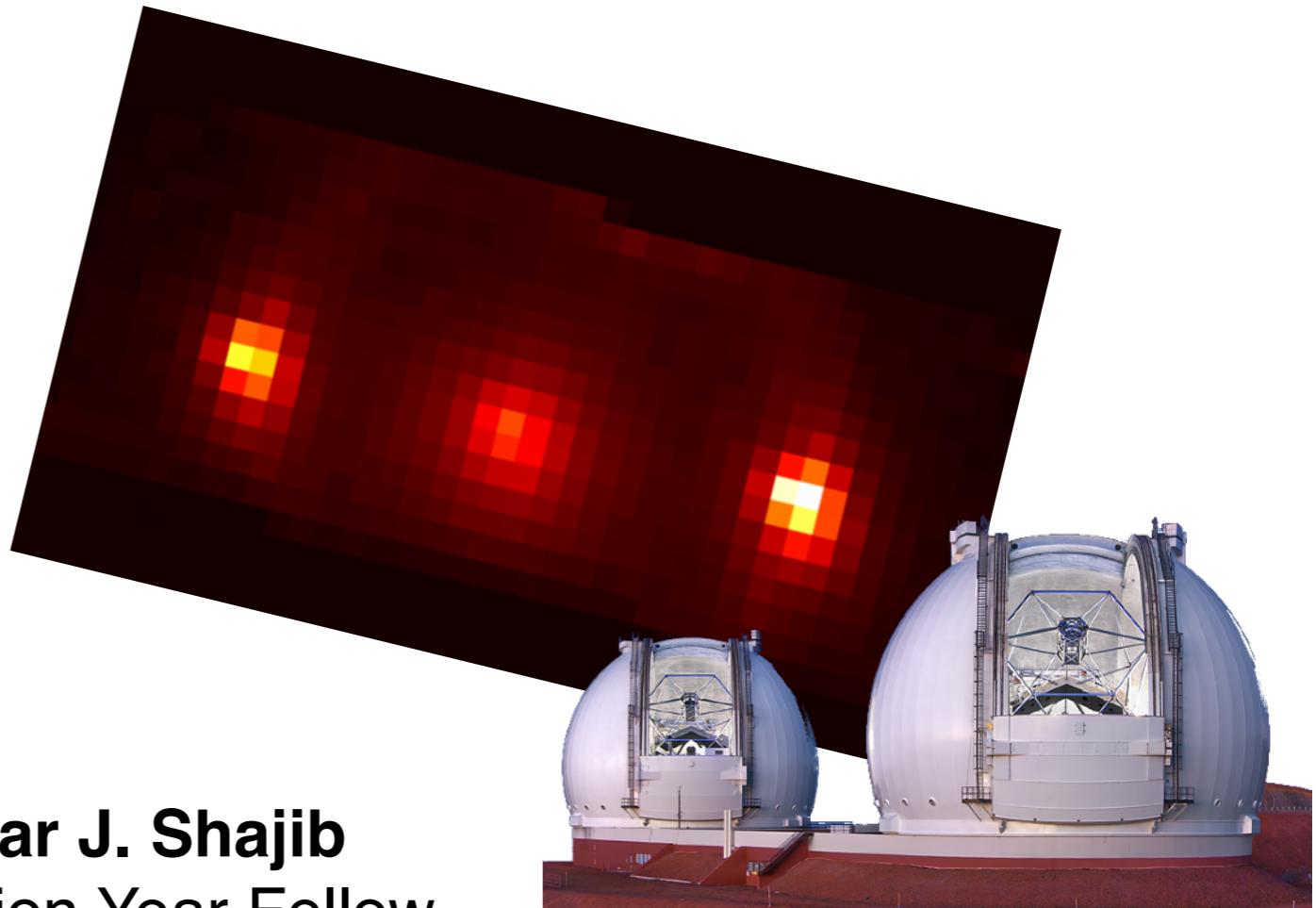


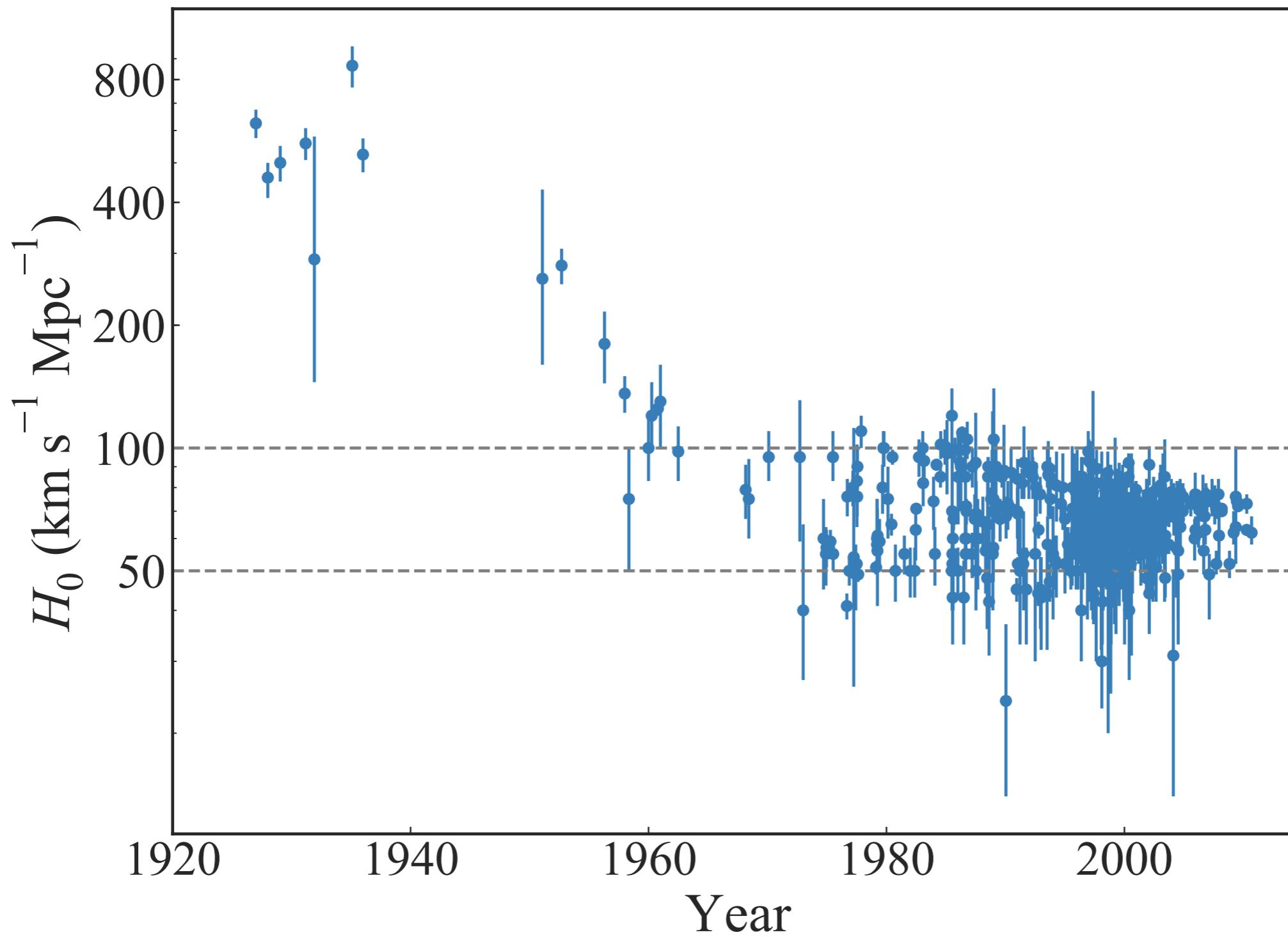
The Hubble tension and strong-lensing time-delays: Hint of new physics?



Anowar J. Shajib
Dissertation Year Fellow
University of California, Los Angeles

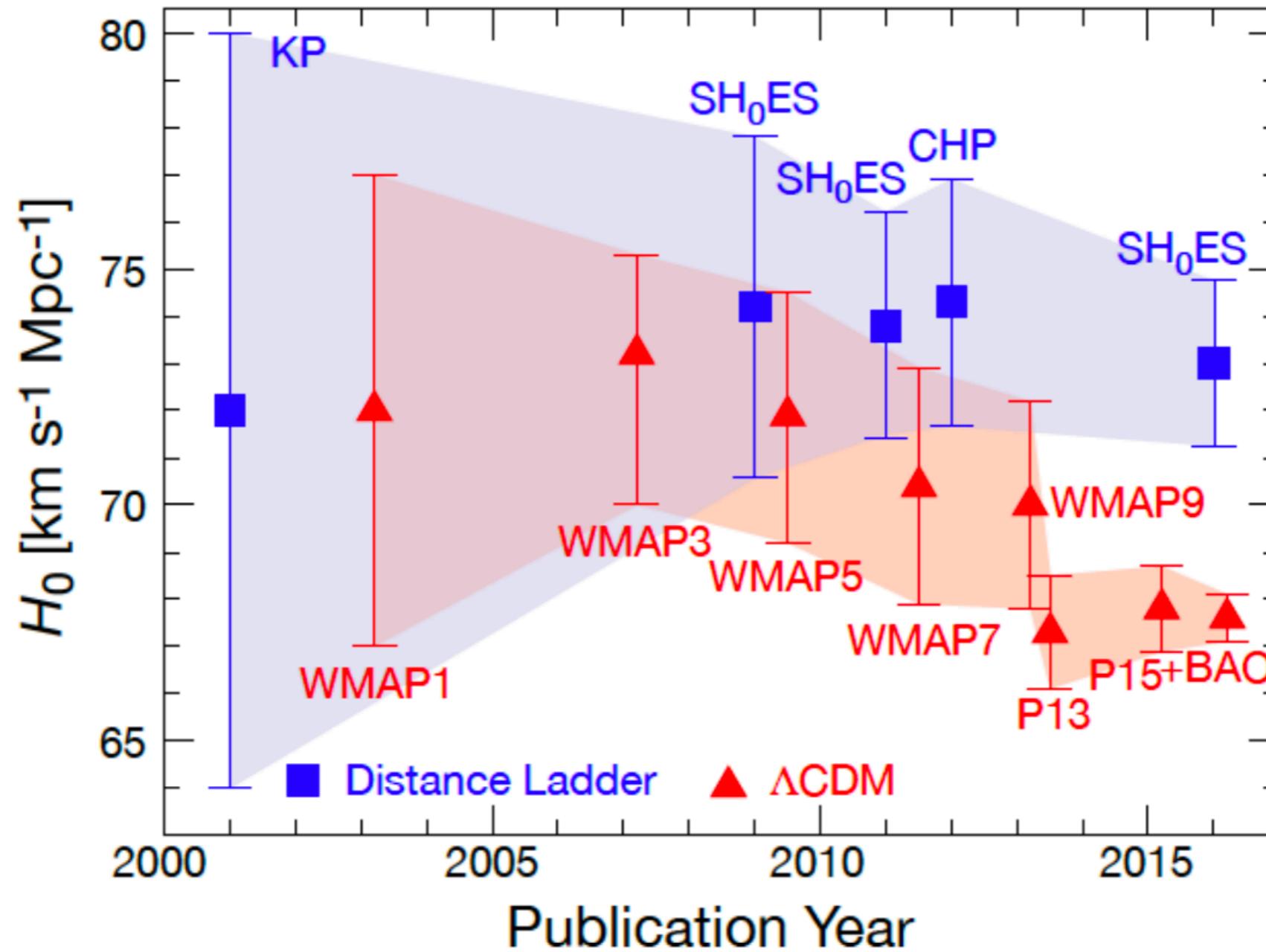
Advisor: Tommaso Treu
Collaborations: STRIDES and H0LiCOW

Hubble constant through the years



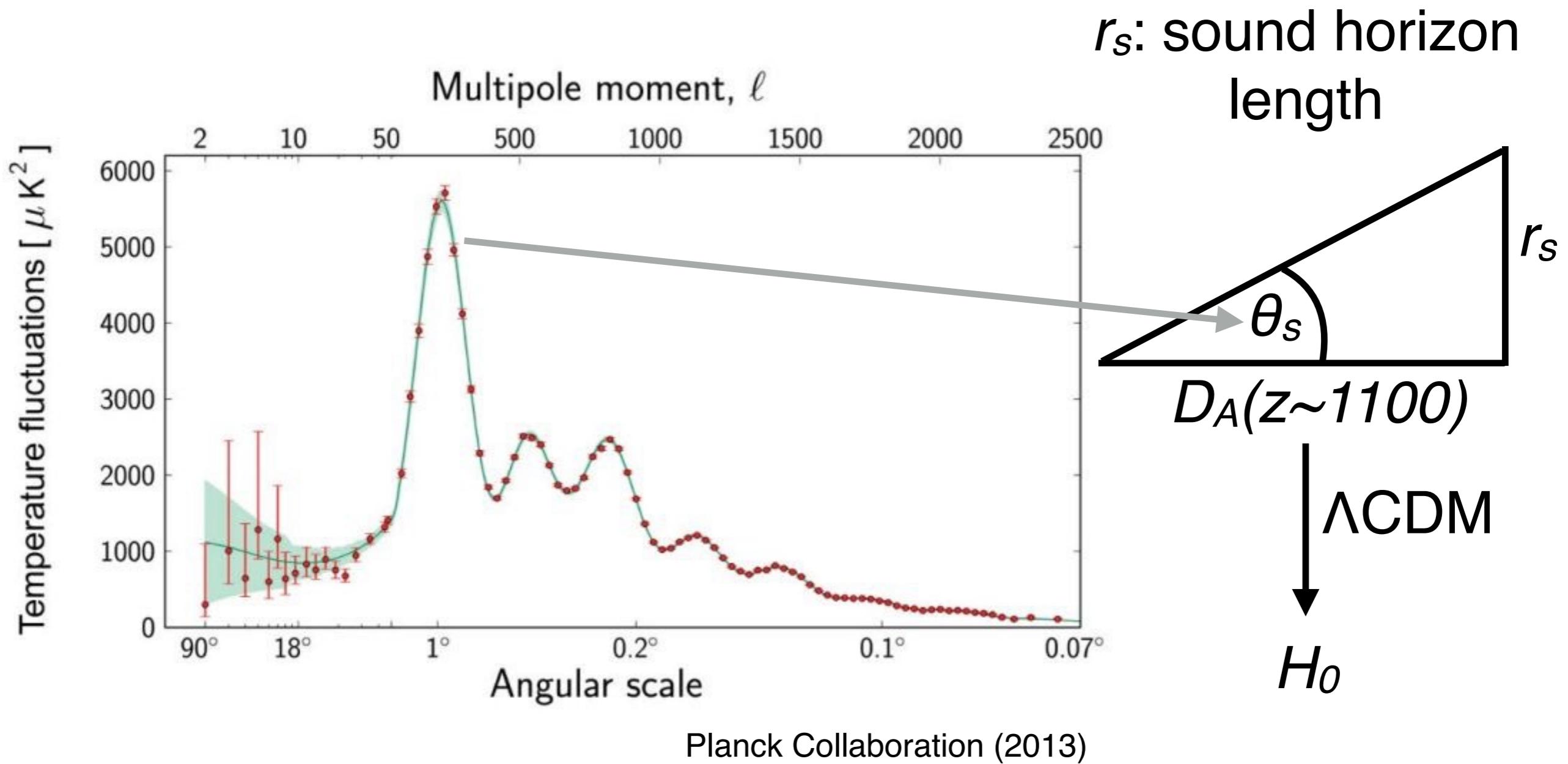
Collected by John Huchra

Recent debate over Hubble constant

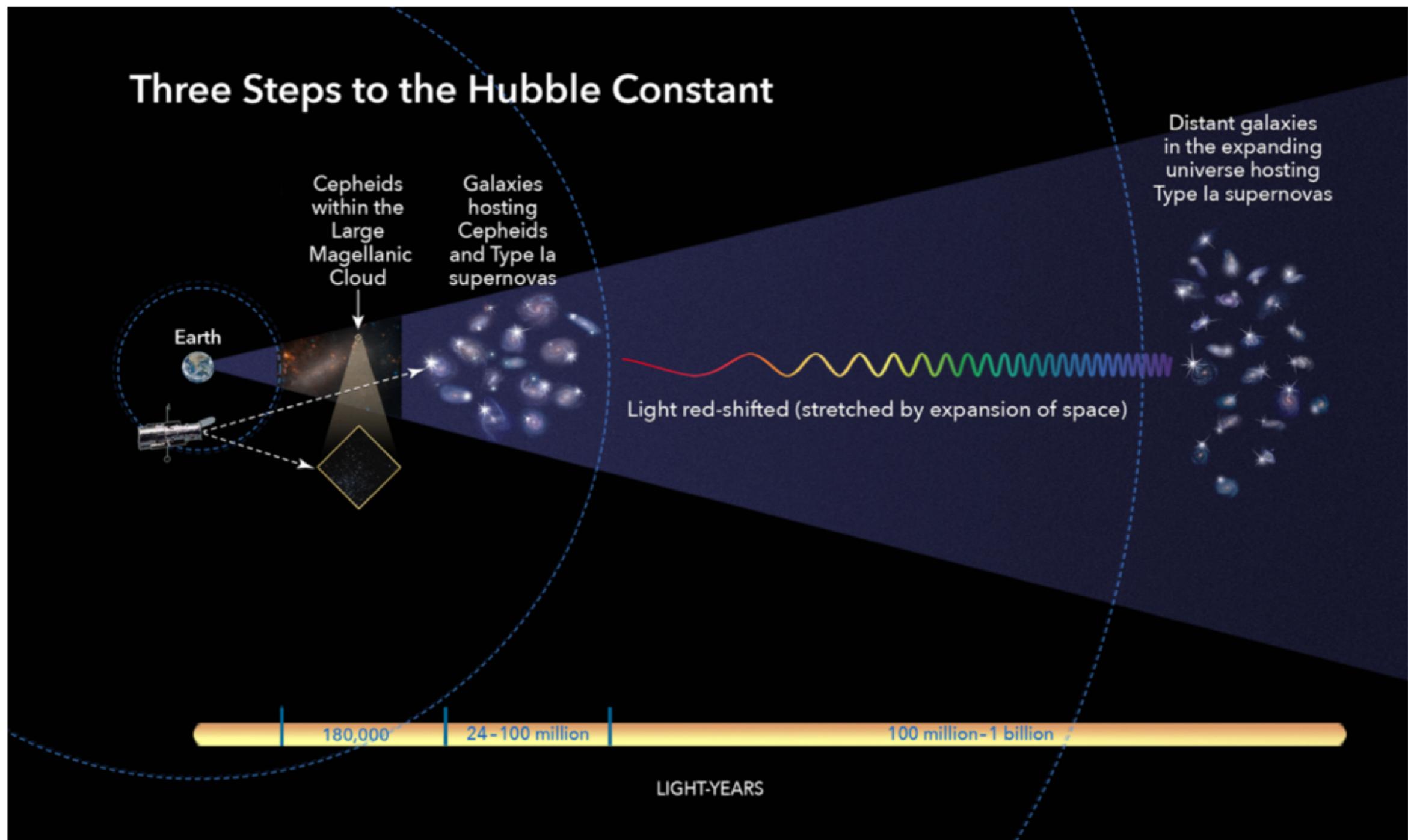


Freedman (2017)

Early-Universe H_0 measurement: CMB



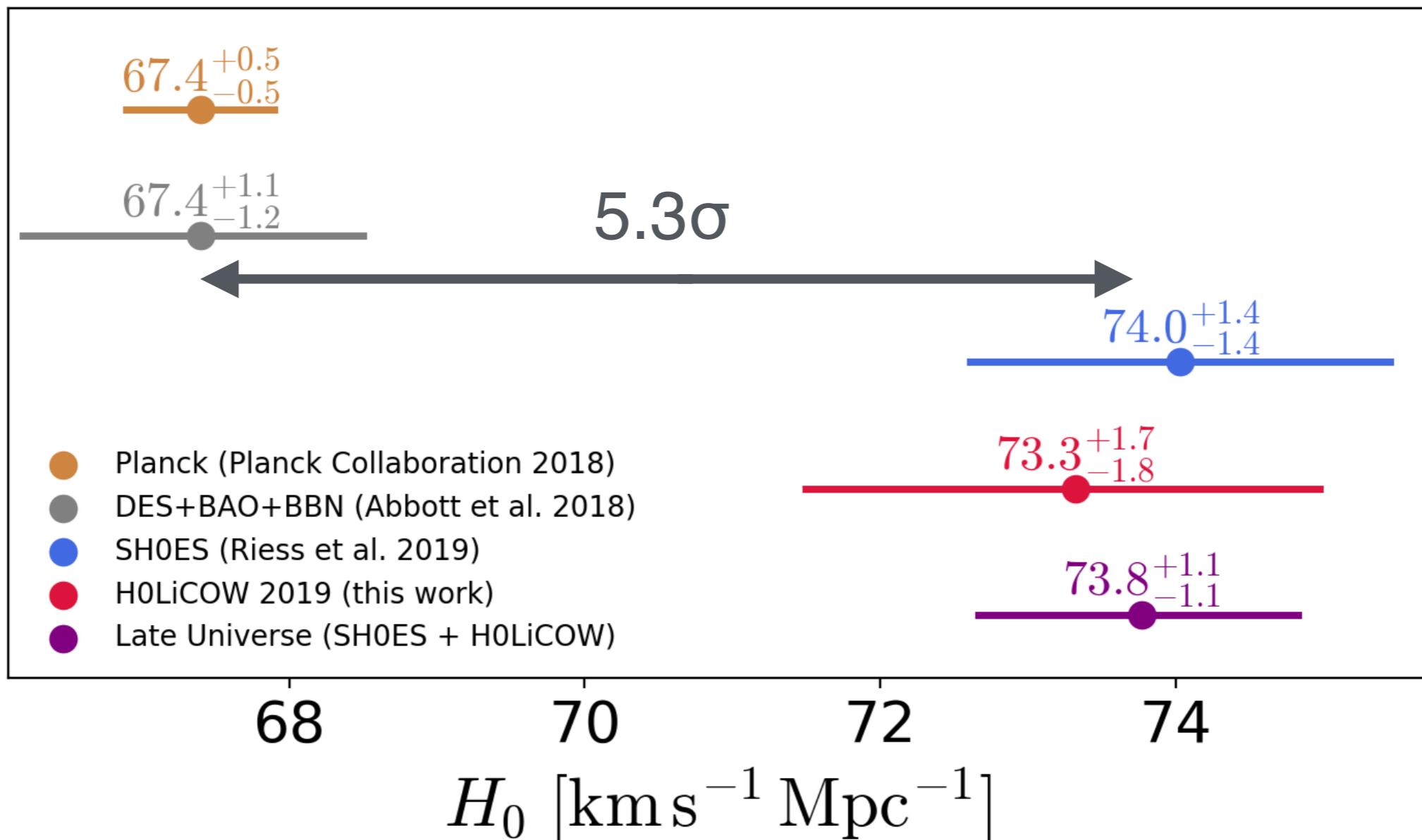
Late-Universe H_0 measurement: cosmic distance ladder



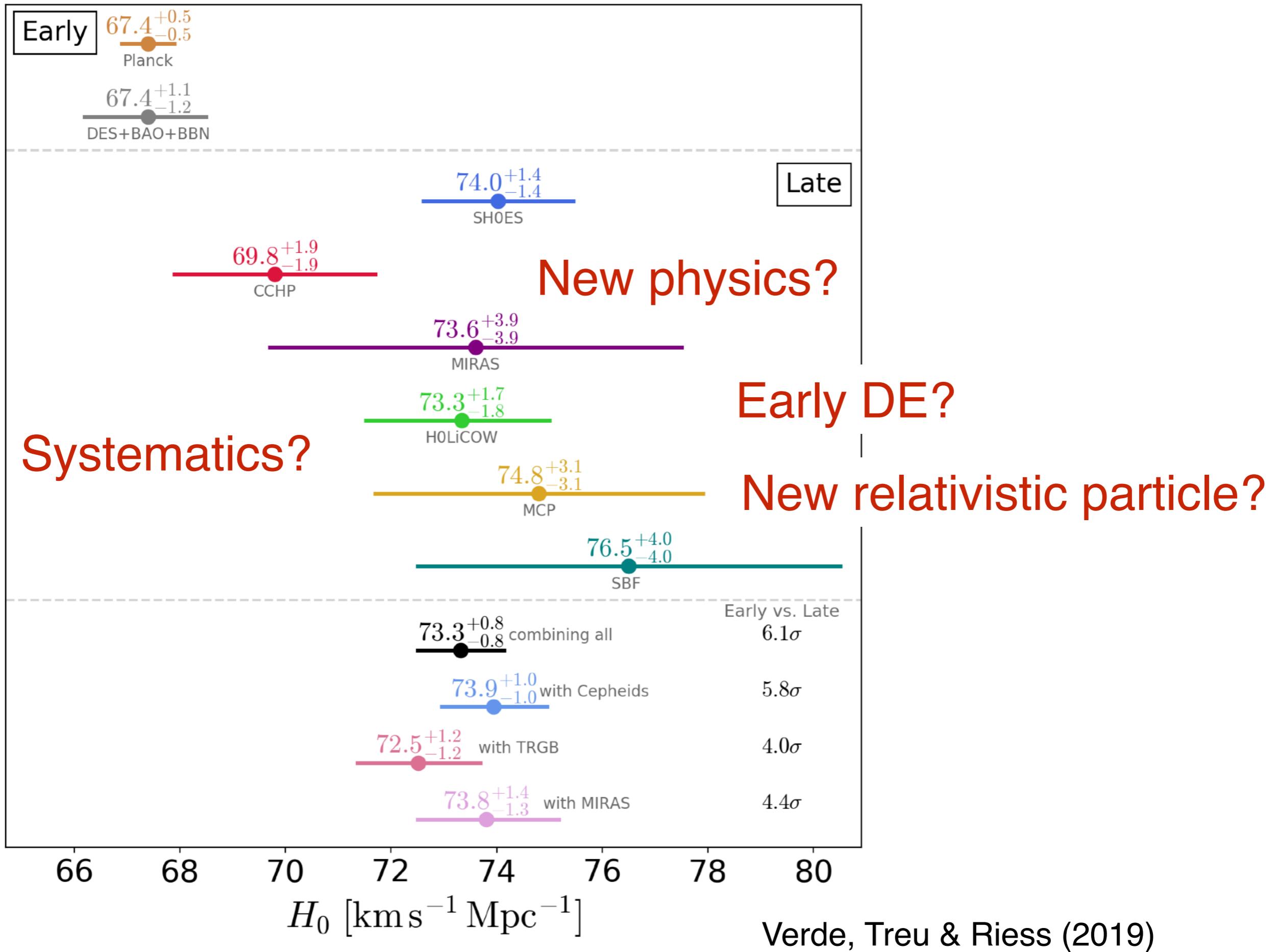
Riess et al. (2019), Image: <https://www.photonicsspectra-digital.com>

The “tension” becomes 5.3σ !

flat Λ CDM



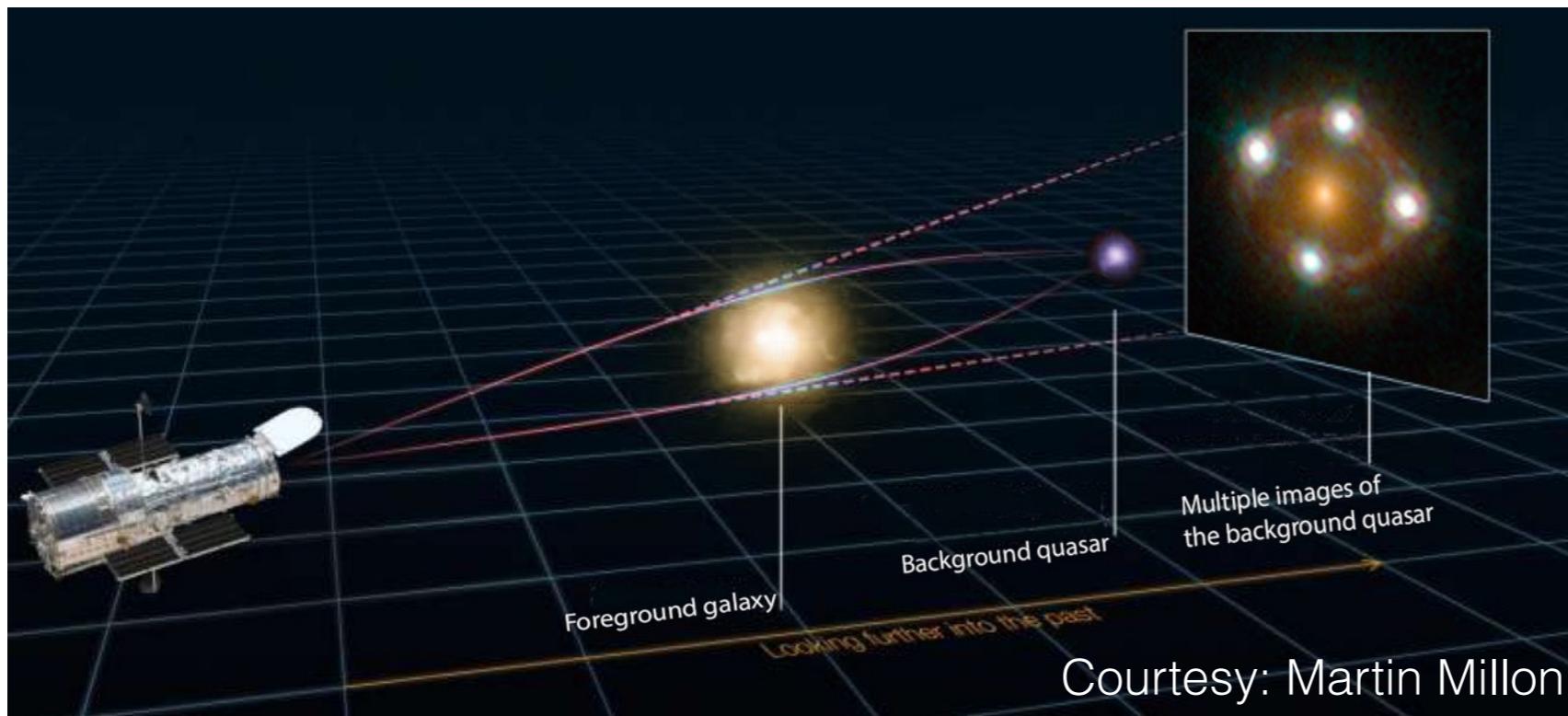
flat – Λ CDM



Time delay cosmography

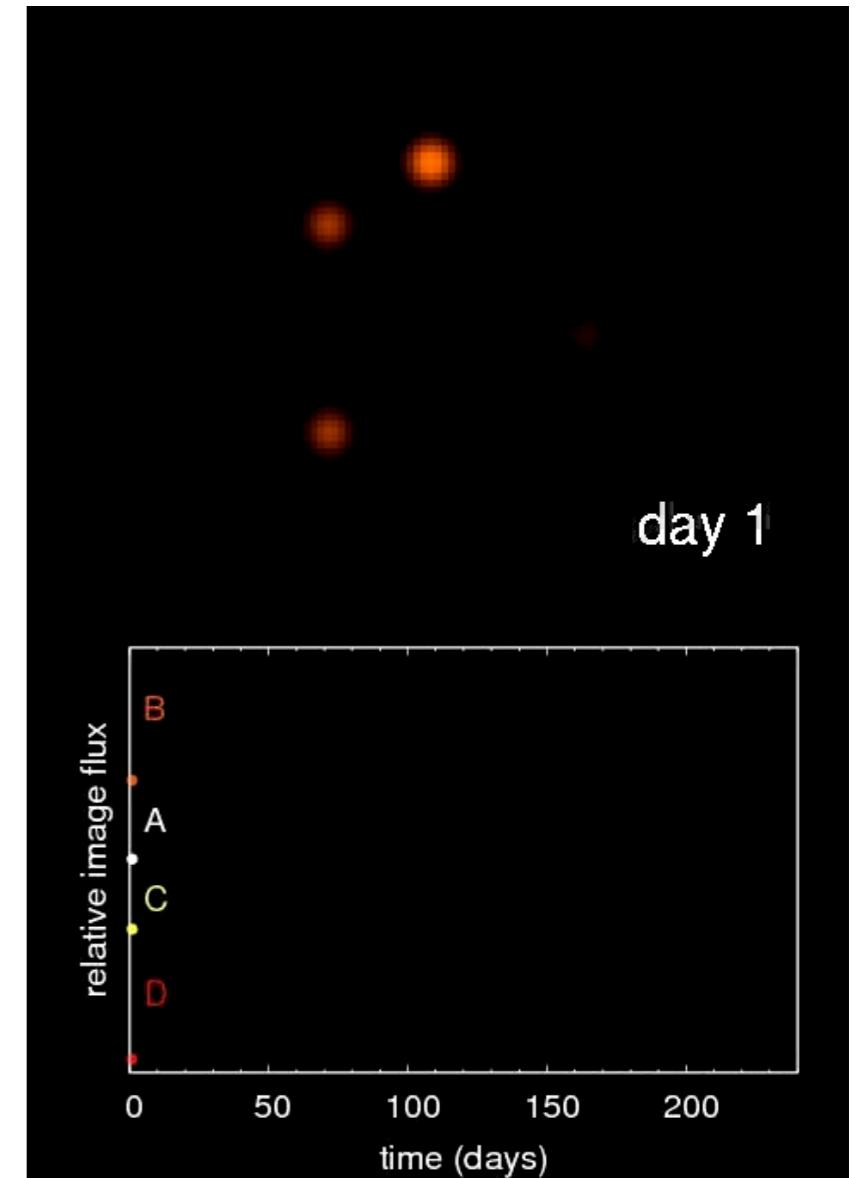
- **Past:** Introduction and recent results
- **Present:** Current works in progress
- **Future:** Further improvements and forecasts

Time-delay Cosmography



- Time-delay distance

$$D_{\Delta t} \propto \frac{\Delta t}{\Delta \Psi} \propto \frac{D_d D_s}{D_{ds}} \propto \frac{1}{H_0}$$

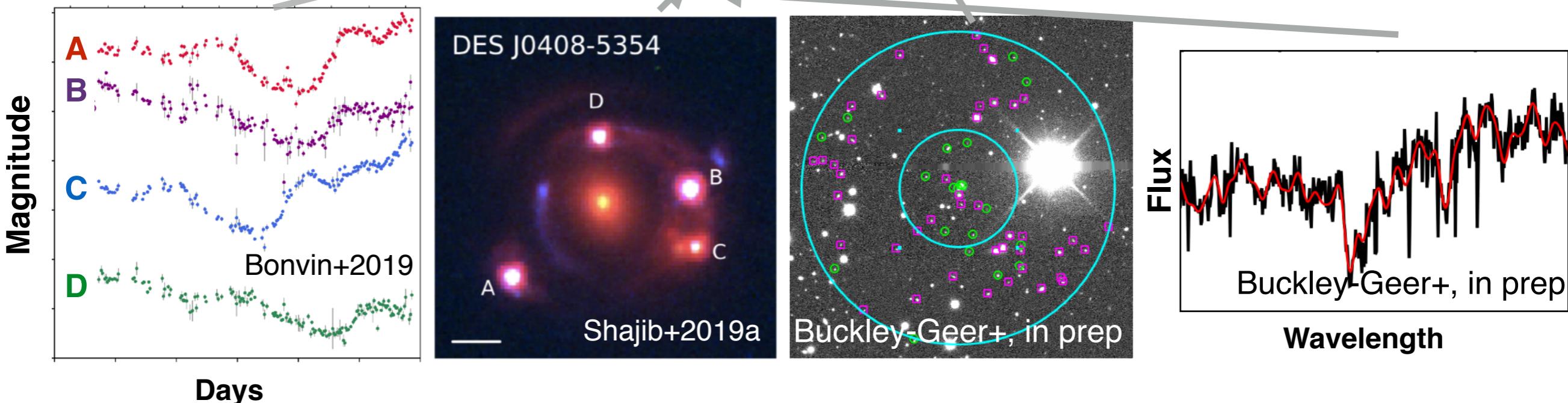


Courtesy: Fred Courbin

Necessary data for time-delay distance measurement

Time delay distance:

$$D_{\Delta t} = \frac{c\Delta t}{\Delta\Psi} \frac{1}{1 - \kappa_{\text{ext}}}$$

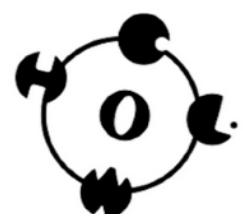


- Time delay measurement

- High resolution imaging of the lens

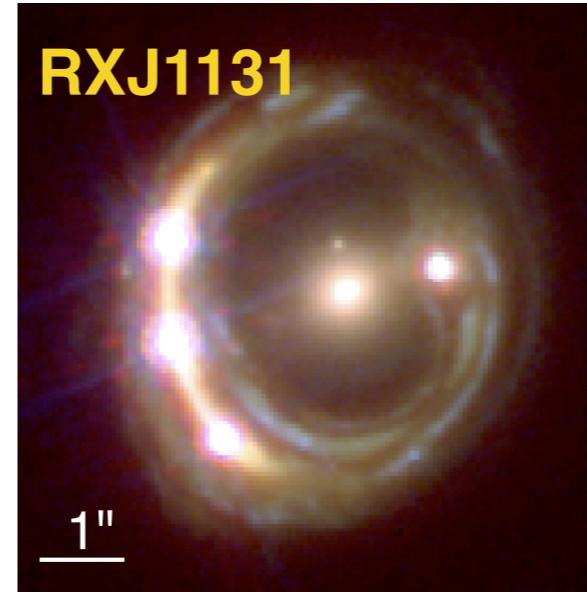
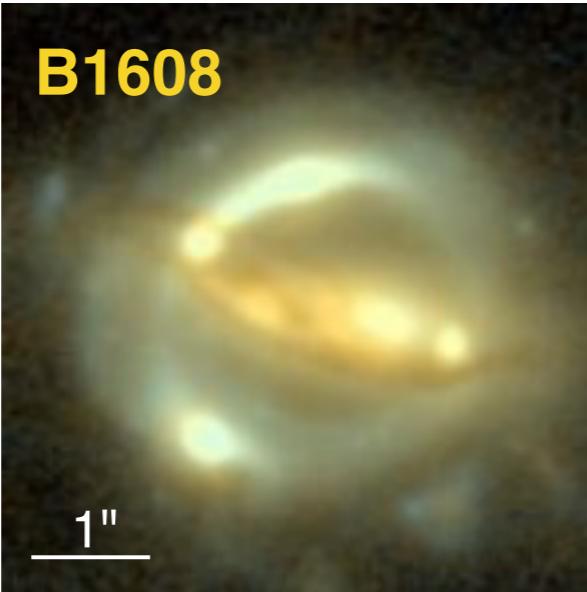
- Estimate of line-of-sight effects

- Kinematics



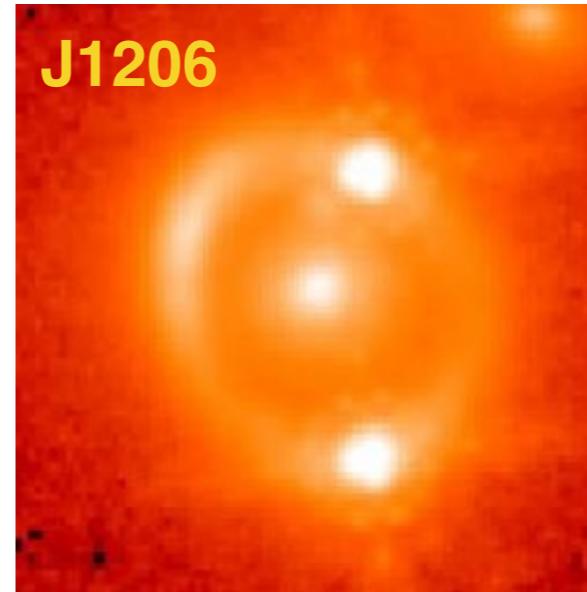
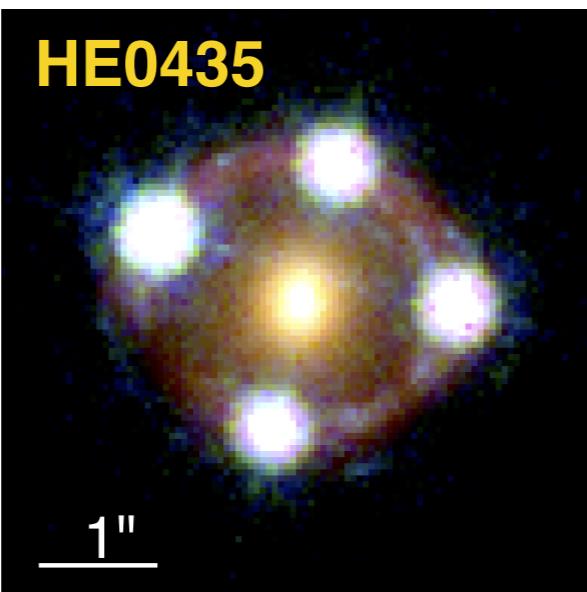
H0LiCOW sample of 6 time-delay lenses

Suyu et al. (2010)



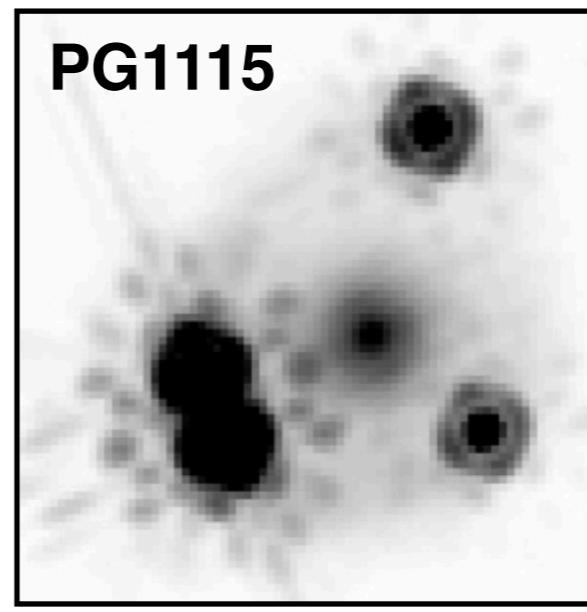
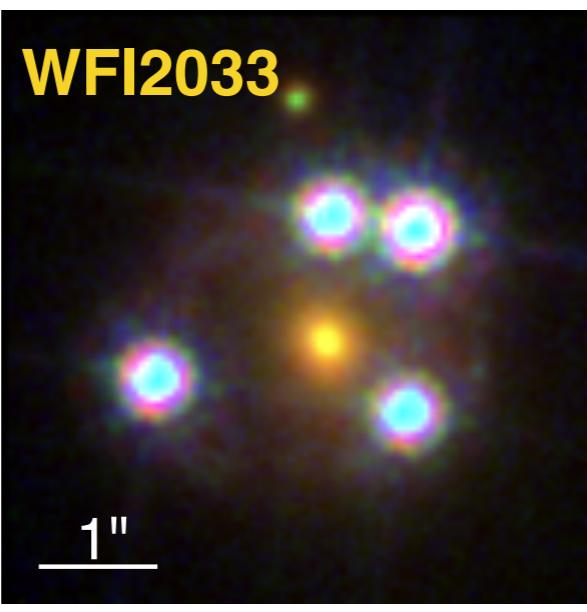
Suyu et al. (2014)

Wong et al. (2017)



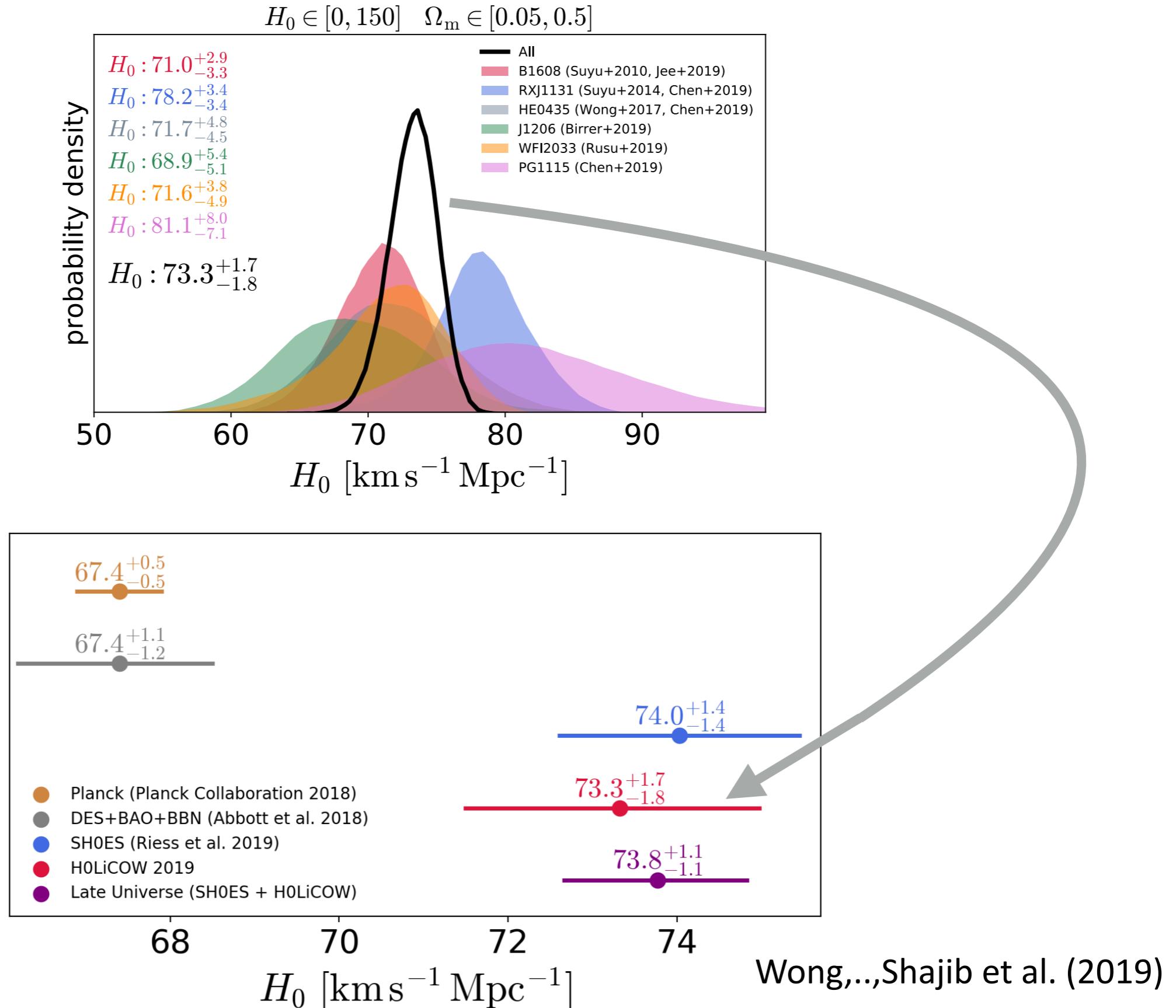
Birrer,..,Shajib et al.
(2019)

Rusu,..,Shajib et al.
(2019)



Chen,..,Shajib et al.
(2019)

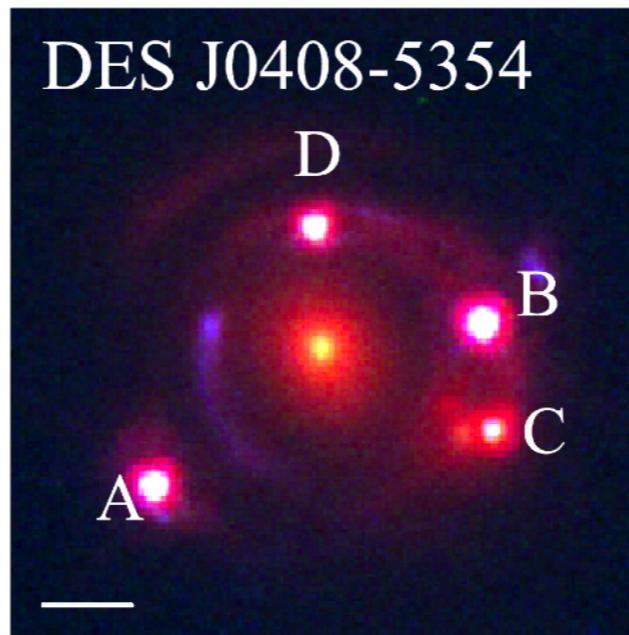
Combining “blind” measurements from H0LiCOW: 2.4% precision in H_0



Time delay cosmography

- **Past:** Introduction and recent results
- **Present:** Current works in progress
- **Future:** Further improvements and forecasts

One new time-delay lens from the STRIDES collaboration



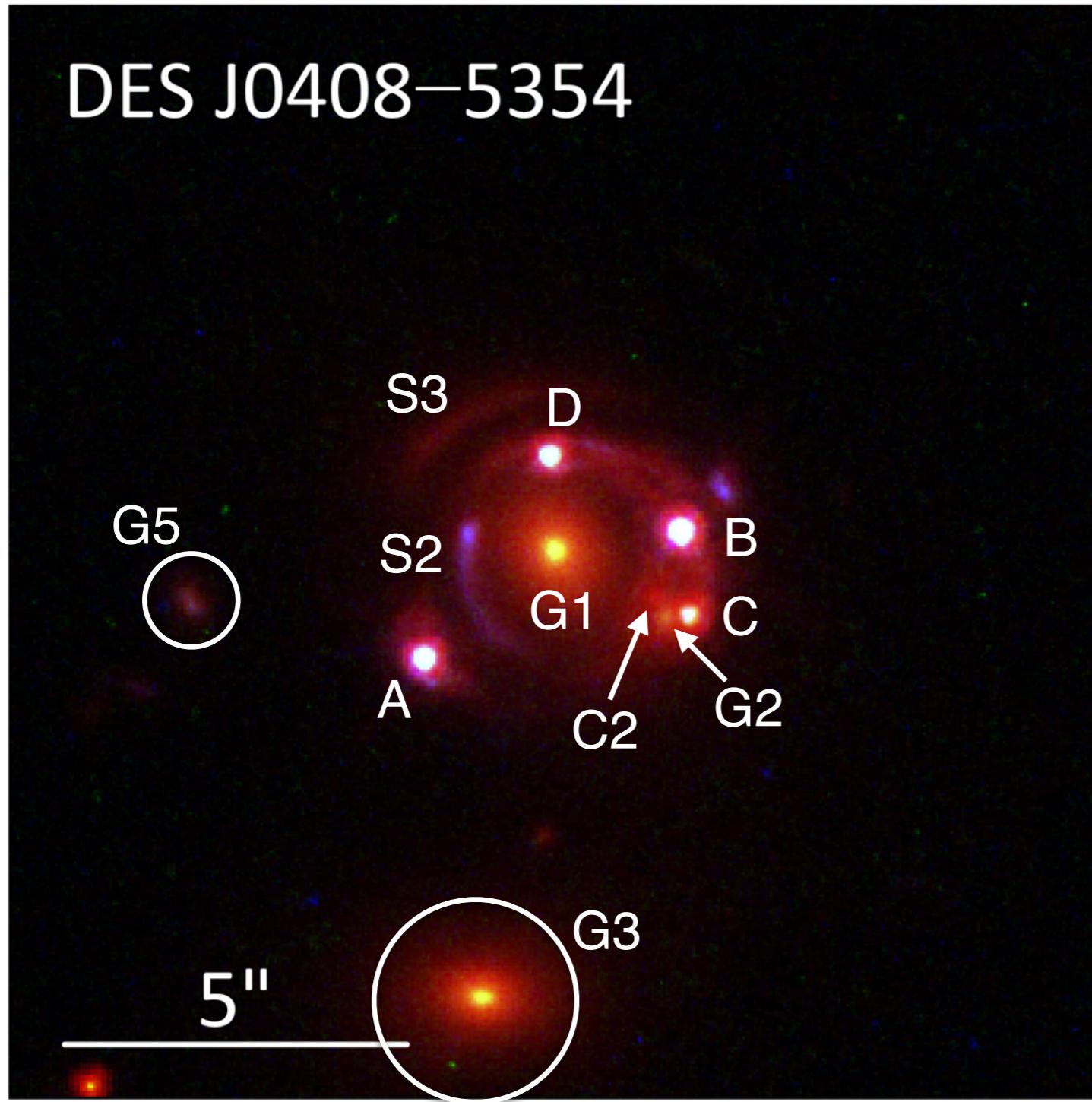
Independent analysis by **2** teams to check for systematics:

- Shajib et al. (UCLA) using “Lenstronomy”
- Yildirim et al. (MPA) using “GLEE”

Each lens requires coordinated effort
by a large team.

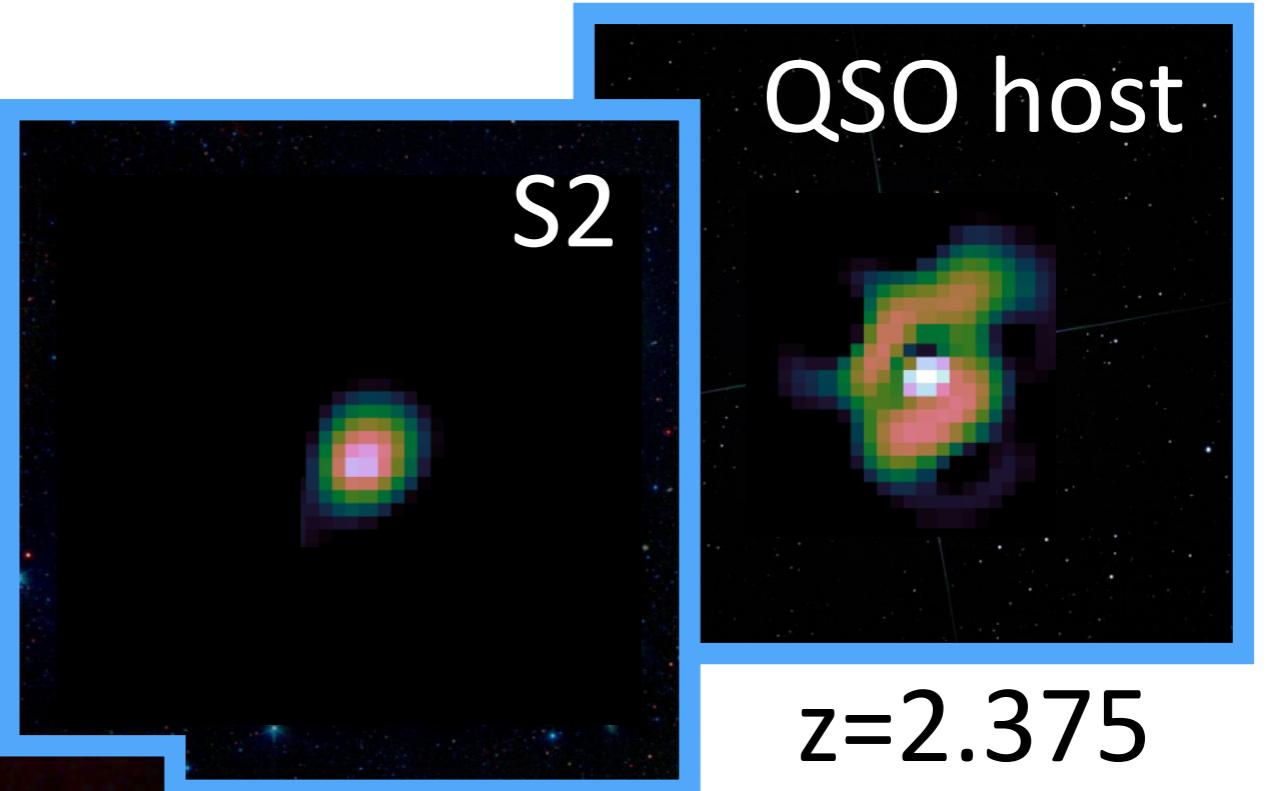
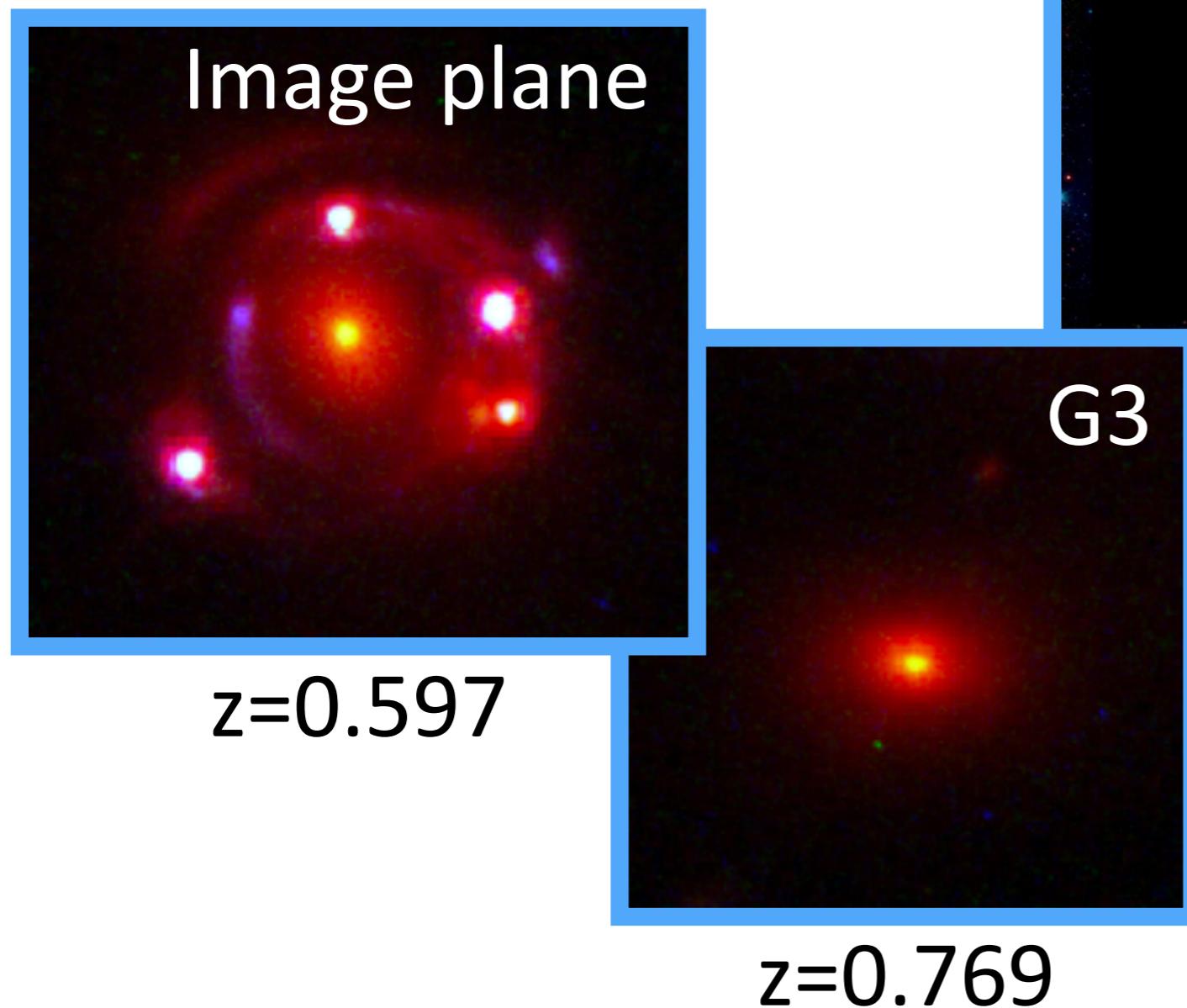
- Lenstronomy lead developer: **S. Birrer**, KIPAC, Stanford
- Time-delays from COSMOGRAIL: **Courbin** et al. (2018)
- External convergence estimate: **E. Buckley-Geer** (Fermilab) and **C. E. Rusu** (NAOJ)
- Kinematics of the deflector: **H. Lin** (Fermilab)
- Nearby galaxy group detection: **J. Poh**, PhD student, (UChicago), **H. Lin** (Fermilab)

The most complex lensed quasar to-date



- Nearby satellite (G2)
- Additional image (C2)
- Multiple sources (S2, S3)
- Line-of-sight perturbers (G3, G5)

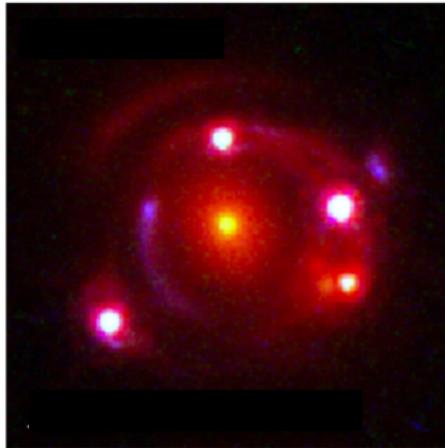
Multi-lens-plane lens modelling



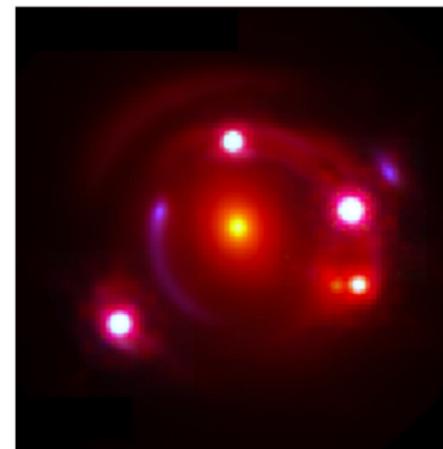
$z=2.375$
 $z=2.228$

- Triple-lens-plane
- Joint lens and source plane for S2

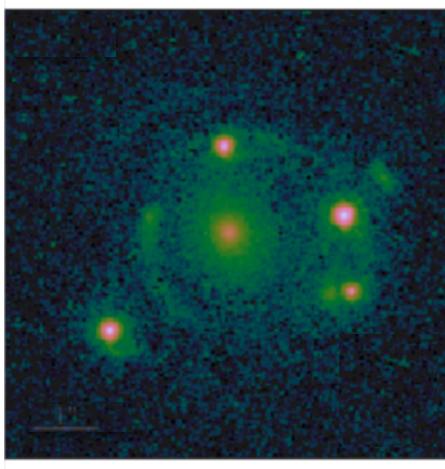
Behind the scene of lens modeling



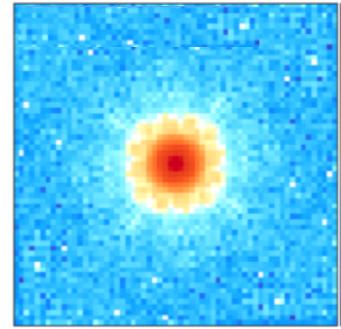
Data: 3-band



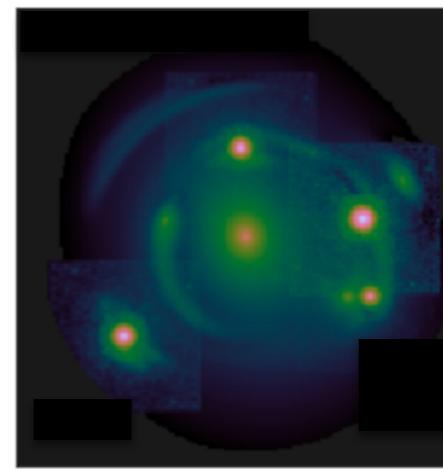
Reconstructed:
multi-band



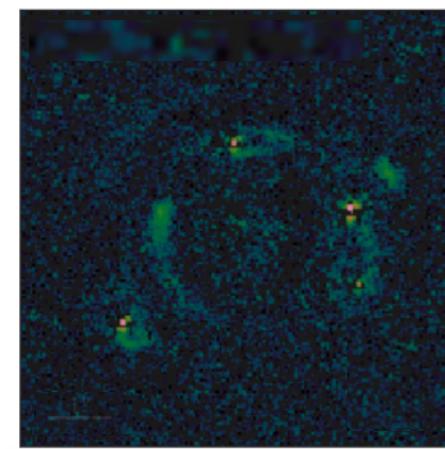
Data: single band



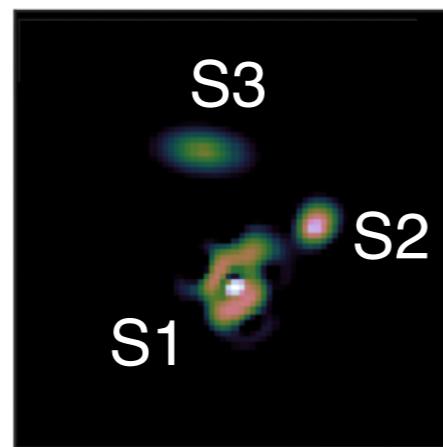
PSF



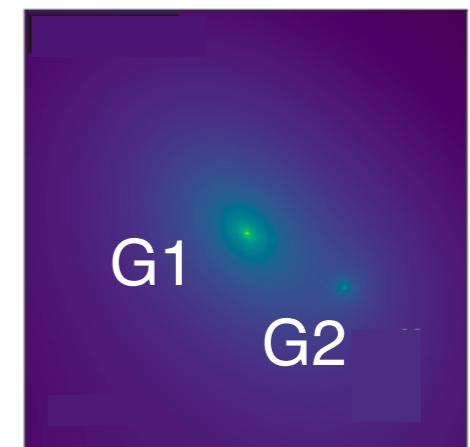
Reconstructed:
single band



Lensed
host
galaxy

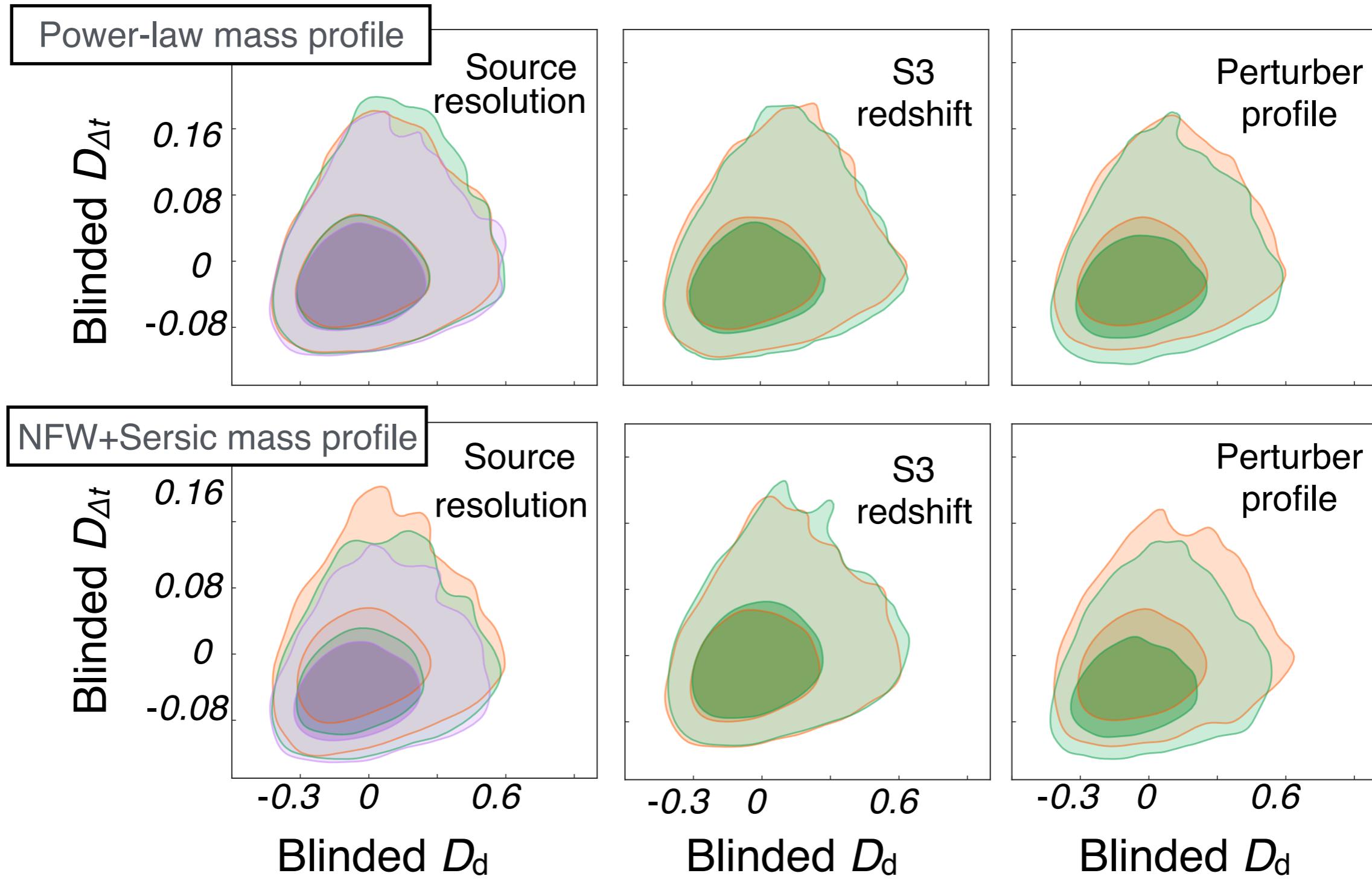


Recons-
tructed
source

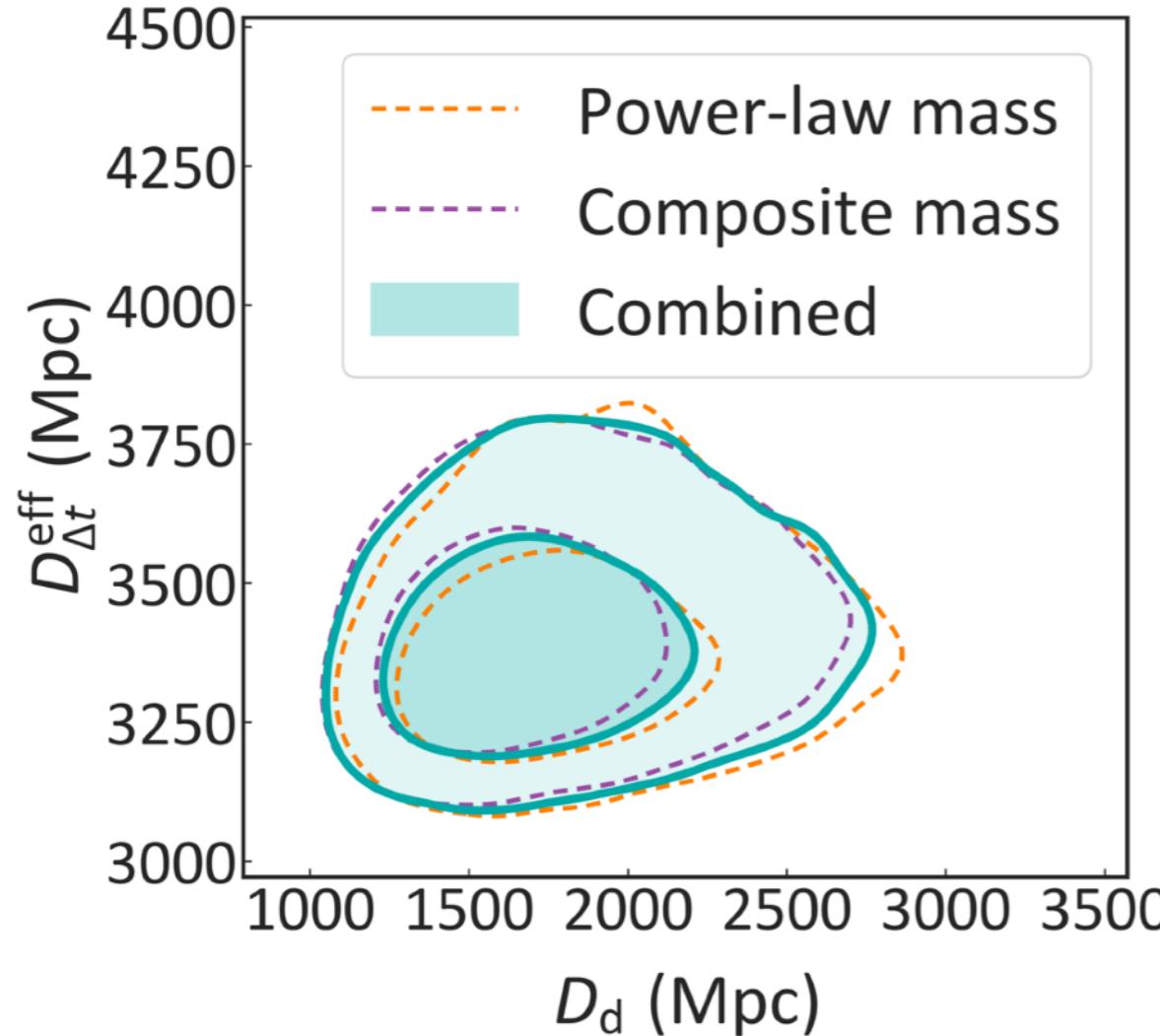


Mass

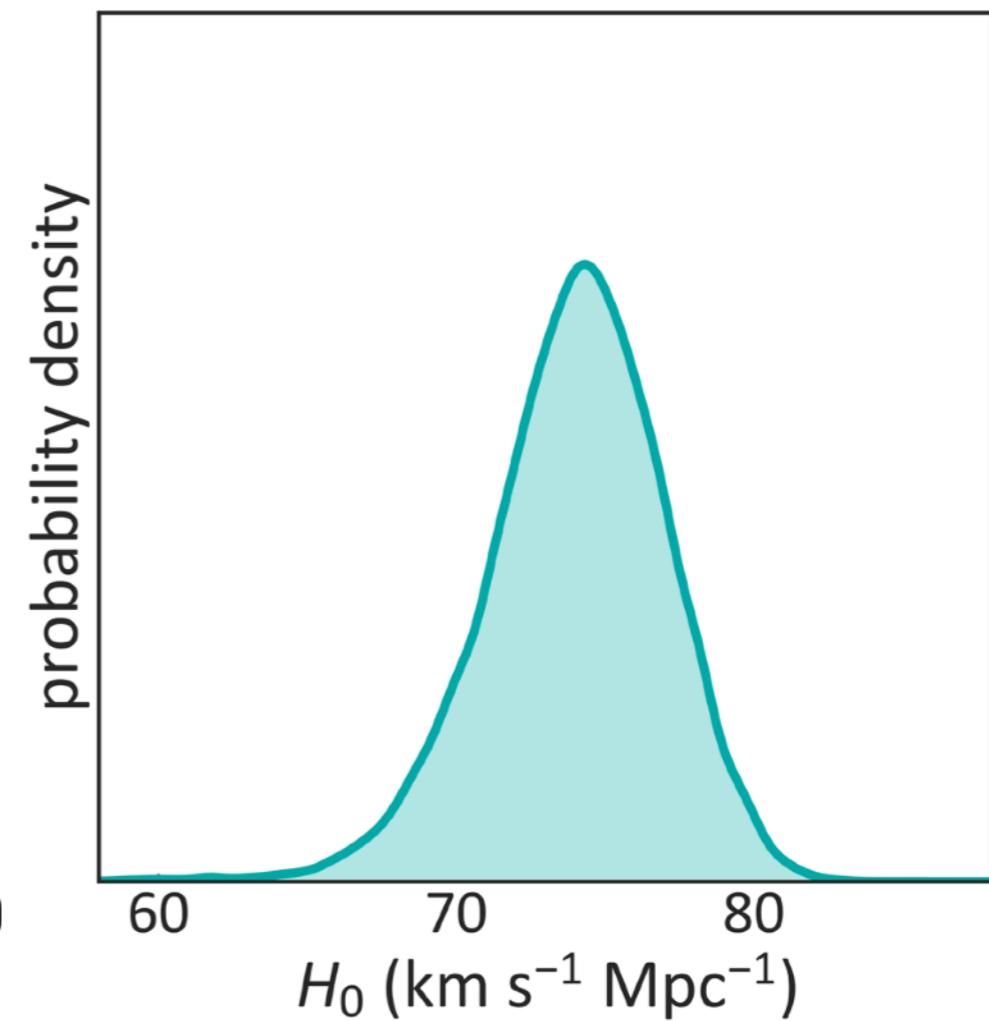
Marginalizing systematics in lens modelling



H_0 from DES J0408-5354

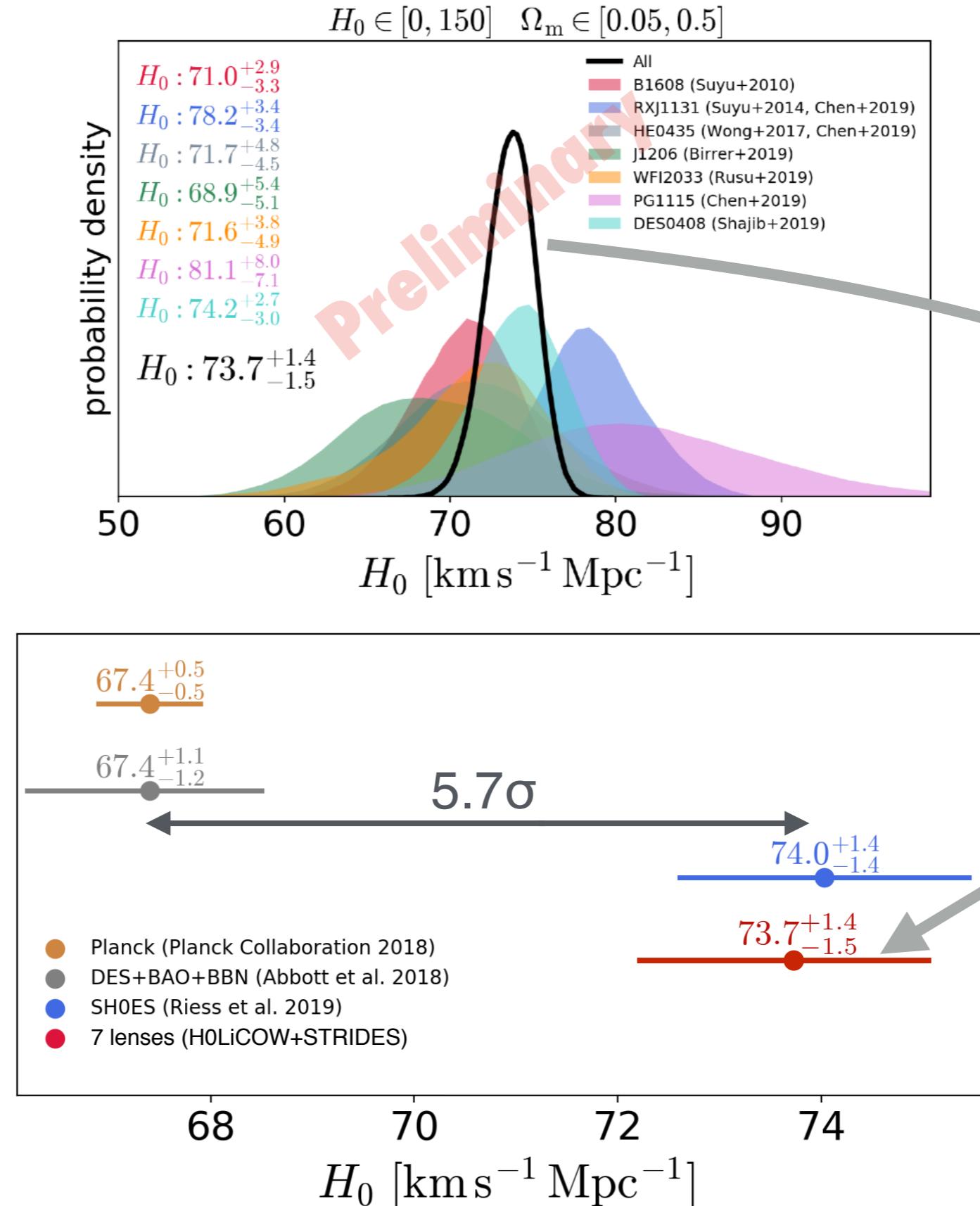


$$H_0 = 74.2^{+2.7}_{-3.0} \text{ km s}^{-1} \text{ Mpc}^{-1}$$



- 3.9% precision from a single lens, highest to-date
- Consistent with the previous sample of 6-lenses

Preliminary combination of 7 lenses: 2% precision in H_0



Time delay cosmography

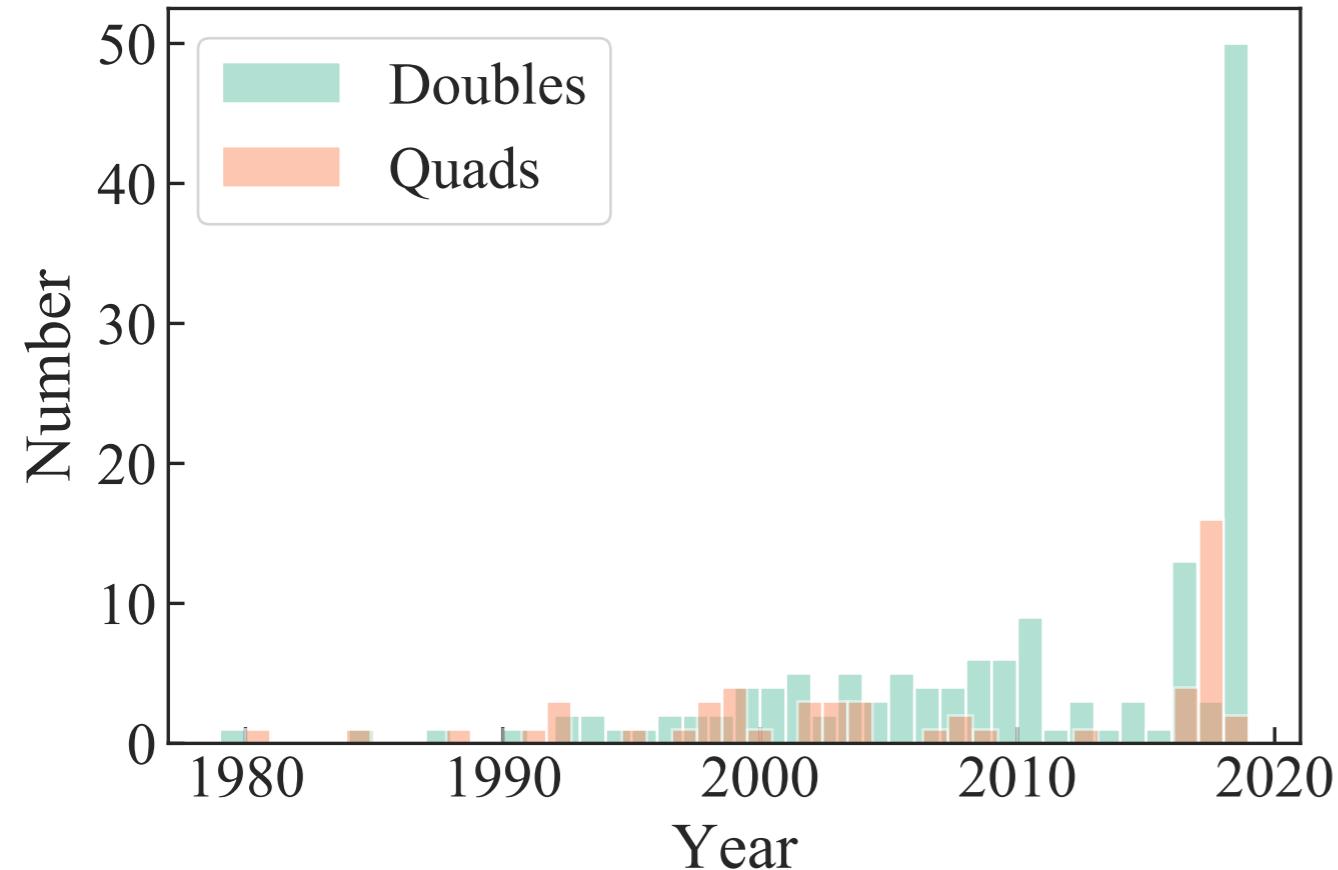
- **Past:** Introduction and recent results
- **Present:** Current works in progress
- **Future:** Further improvements and forecasts

Future goal is 1% H_0 measurement.

- Two Ways to improve precision
 - Increase sample size
 - Improve precision per system
- Thoroughly investigate to find unknown systematics

Approach 1: Increasing sample size

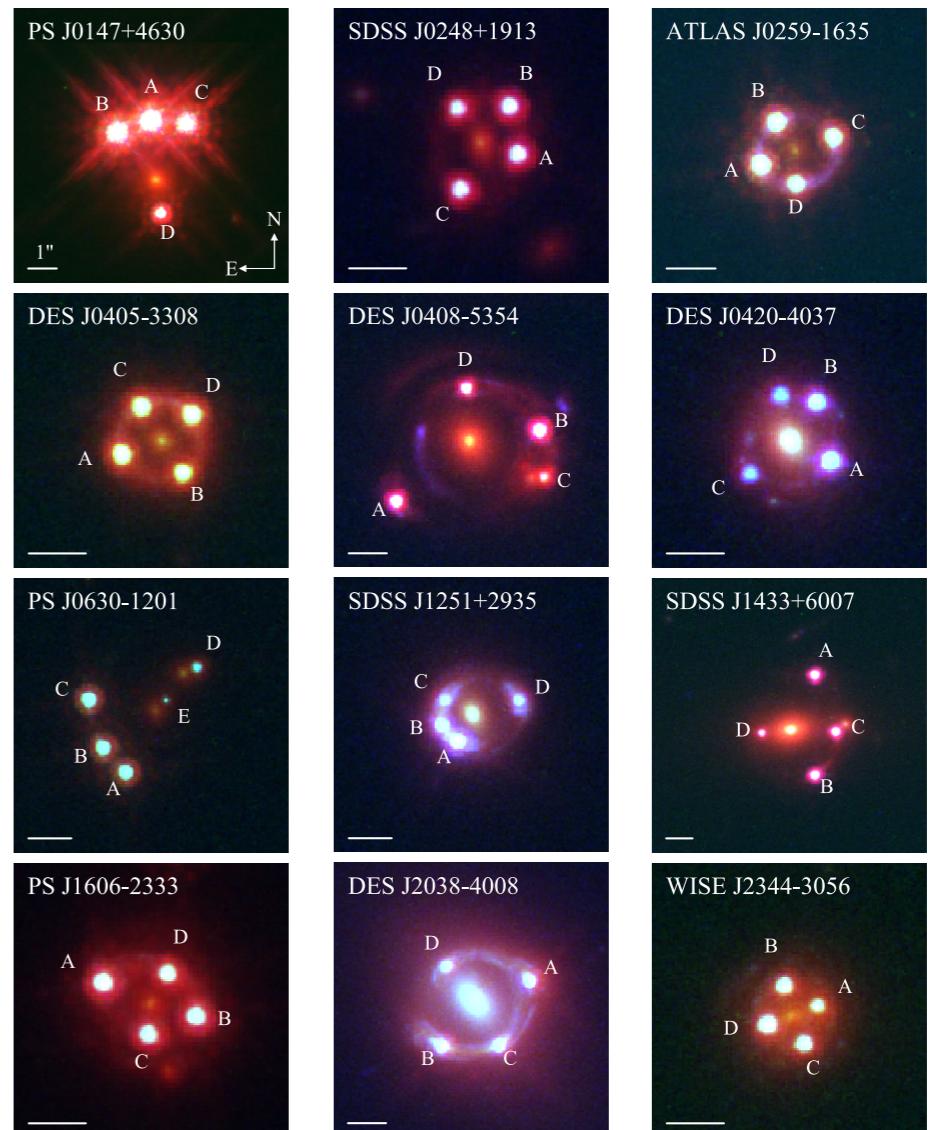
Lens discoveries



Data courtesy: Lens DB by Cameron Lemon

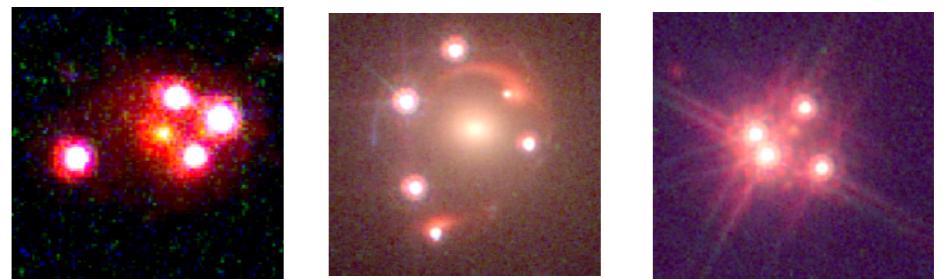
We have already discovered enough quasars to reach 1% in H_0 .

From HST Cycle 25



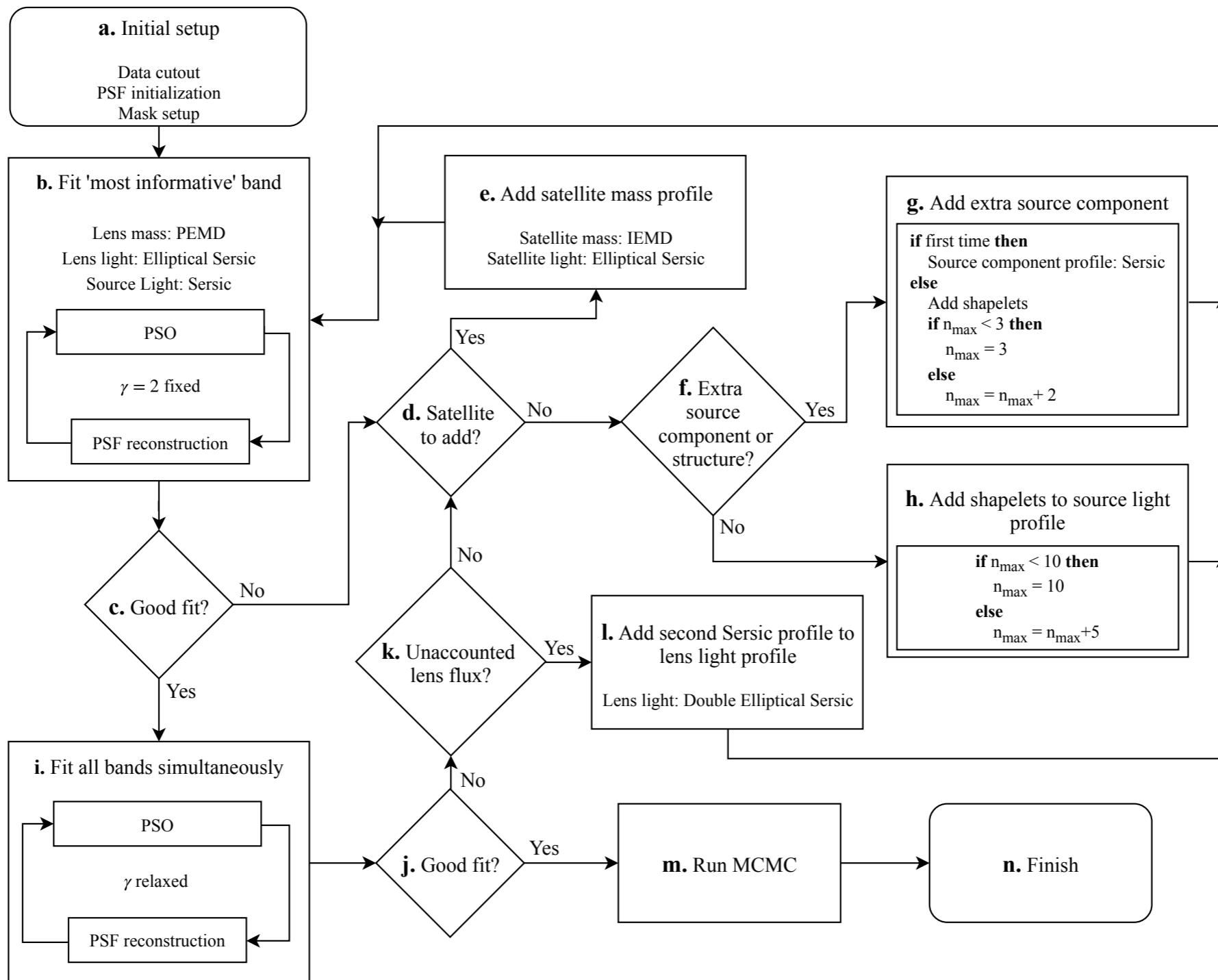
Shajib et al. 2019a

From HST Cycle 26

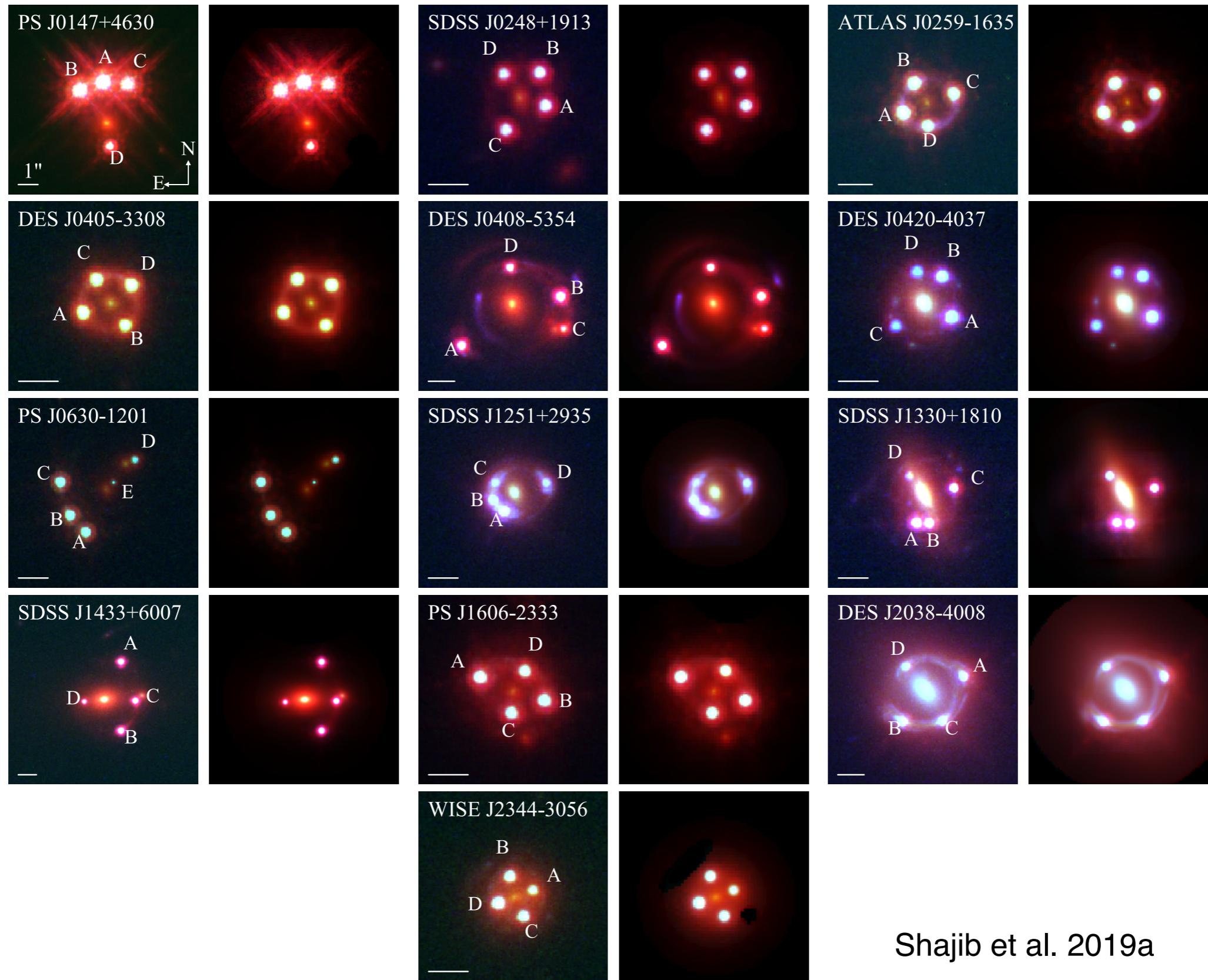


... and 15 more to be observed.

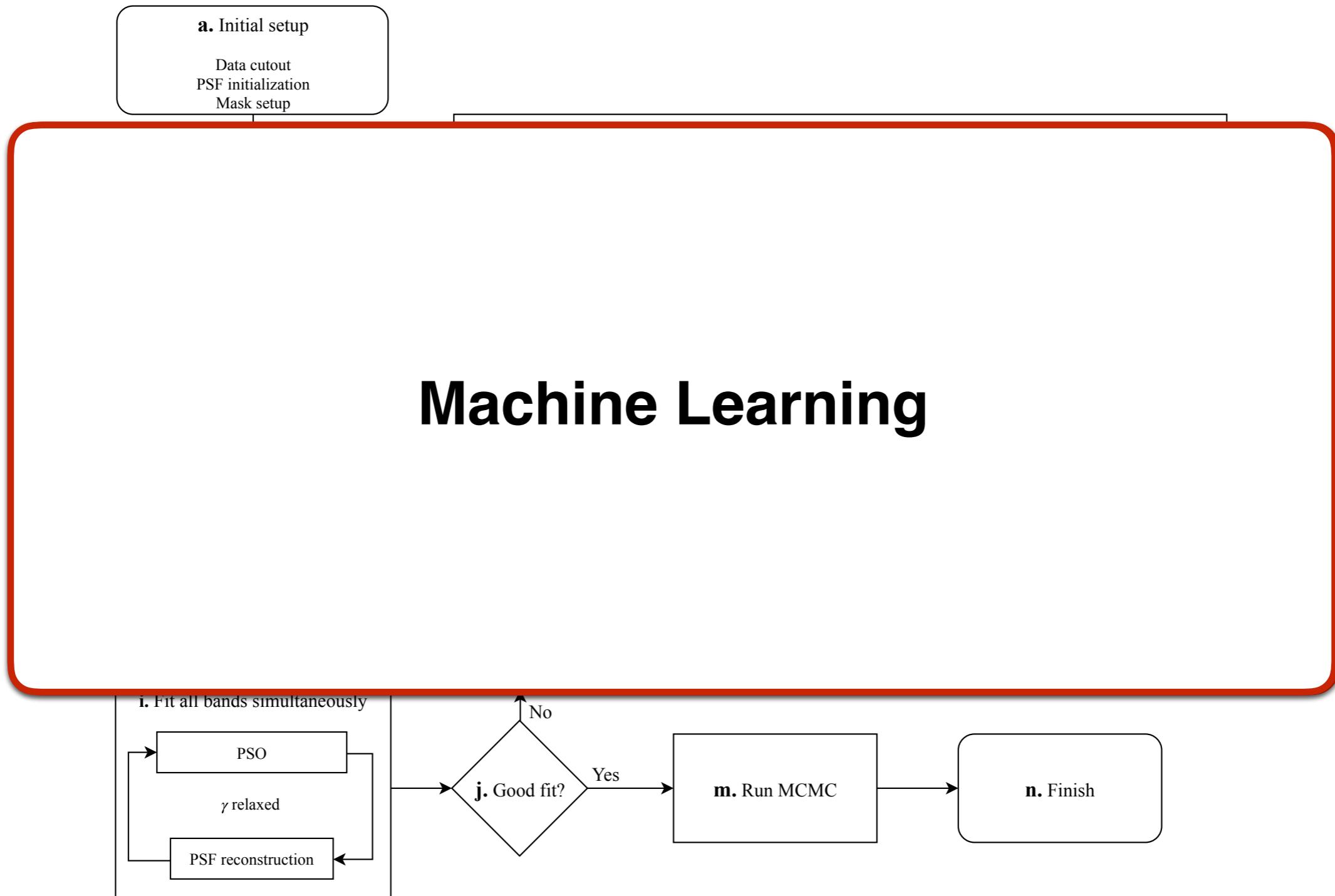
Approach 1: Automating the lens modeling



Approach 1: Automated lens models of 13 lenses



Approach 1: Automating the lens modelling

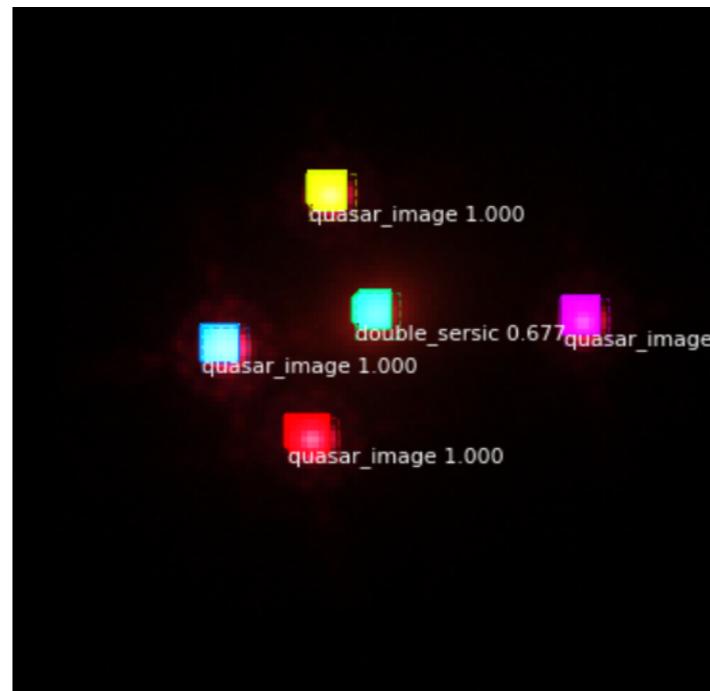
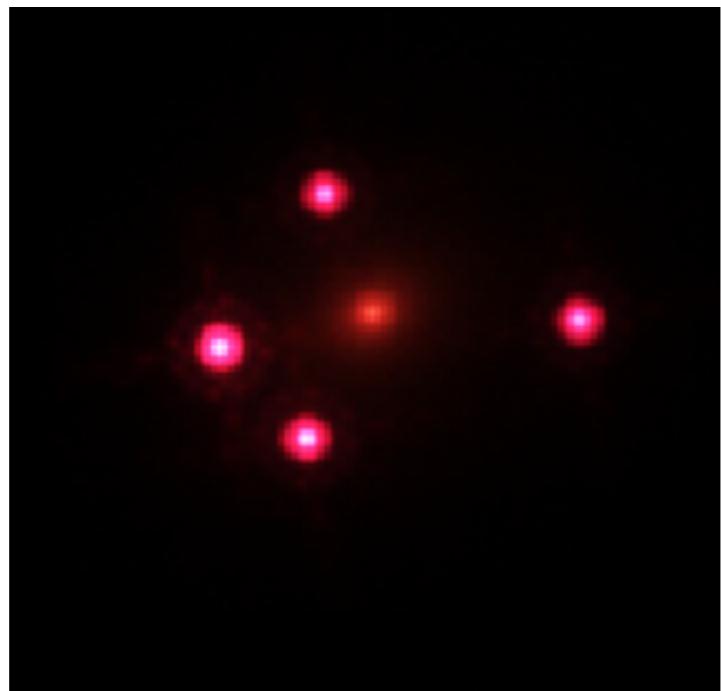


Approach 1: Future direction in automated lens modelling

- **Machine learning** for initializing lens models
 - Work of Vedant Sahu, UCLA undergraduate



Image
from
validation
set

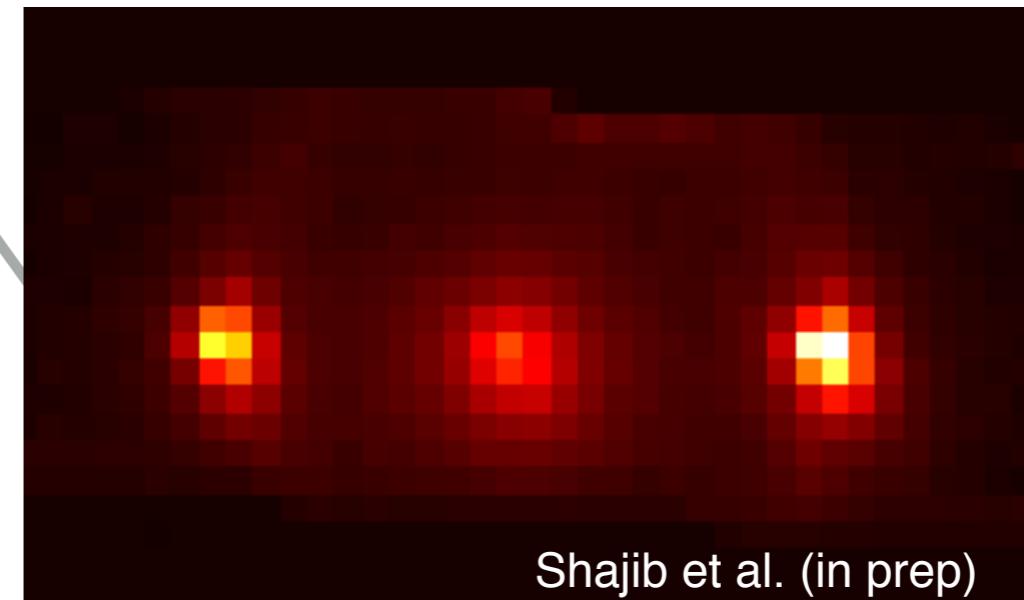
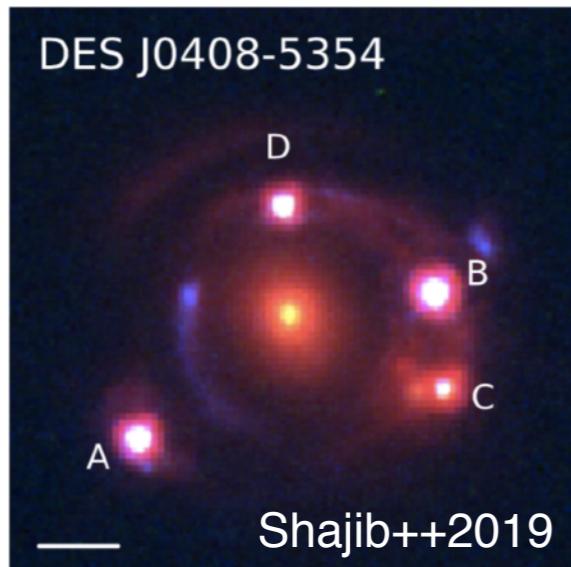


Detected
quasar images
and lensing
galaxy

Approach 2: Improving Precision Per System

Time delay distance:

$$D_{\Delta t} = \frac{c\Delta t}{\Delta\Psi} \frac{1}{1 - \kappa_{\text{ext}}}$$

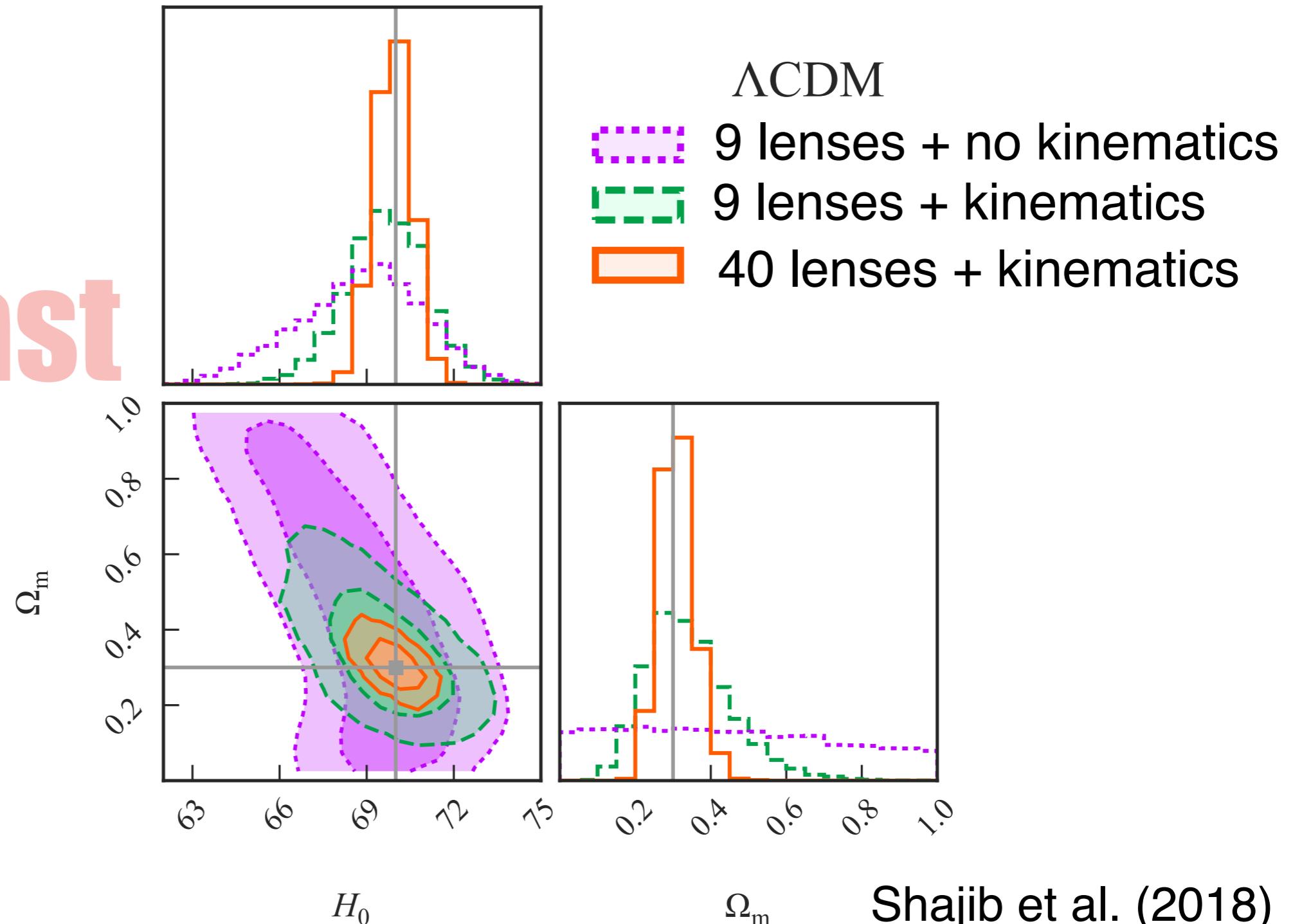


Spatially resolved kinematics

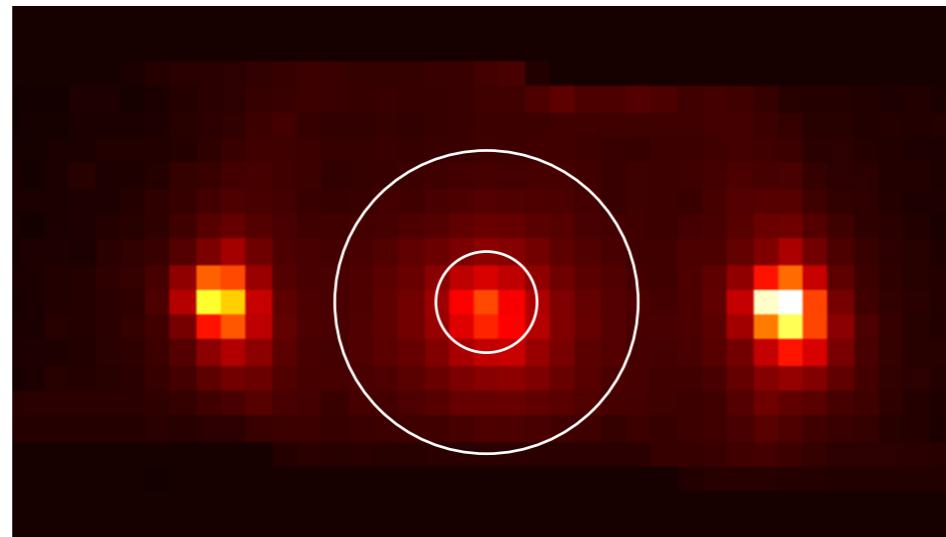
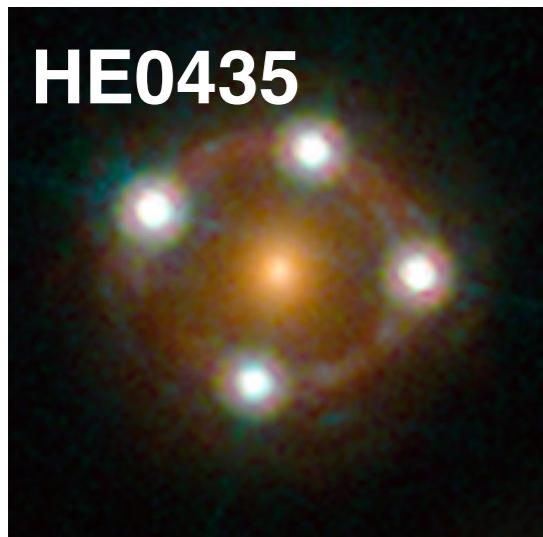
Spatially resolved kinematics improves precision on the mass profile slope.

Approach 2: Spatially resolved kinematics helps determine H_0 1% from a sample of 40 lenses.

Forecast

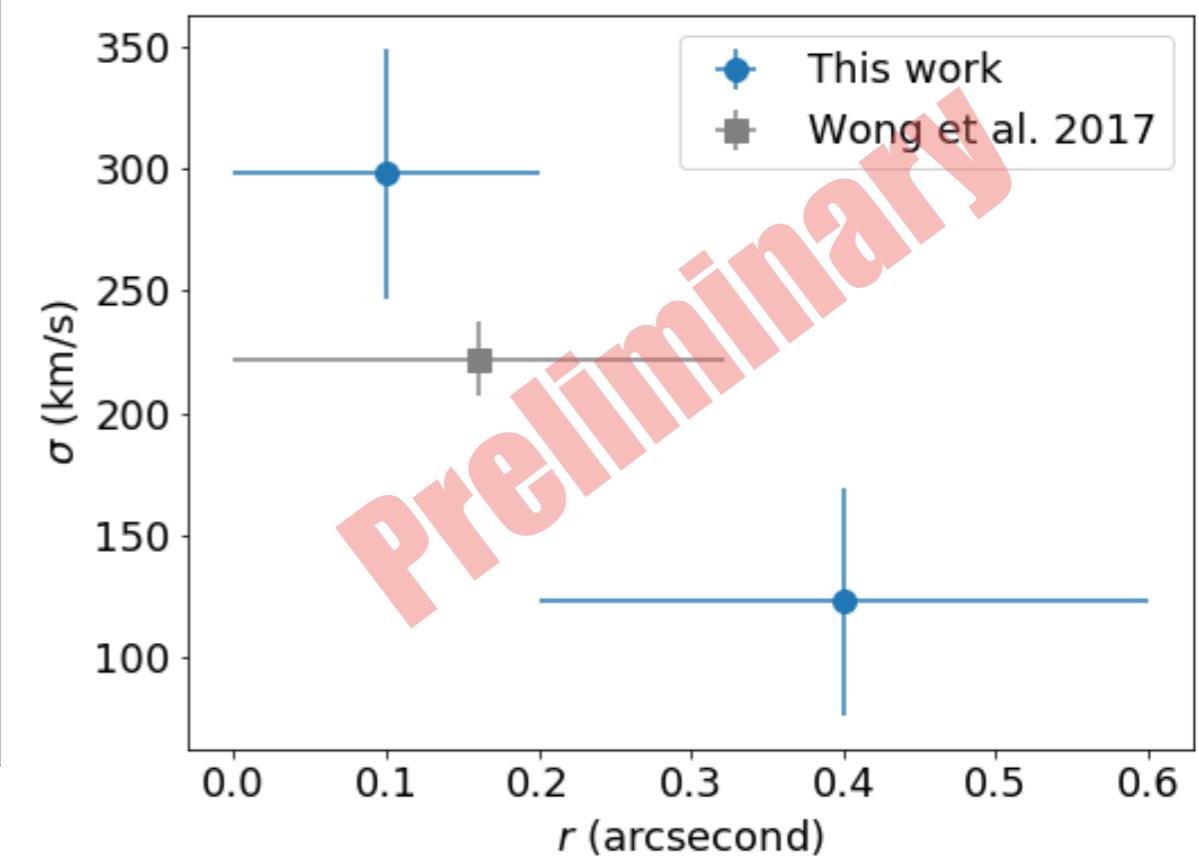
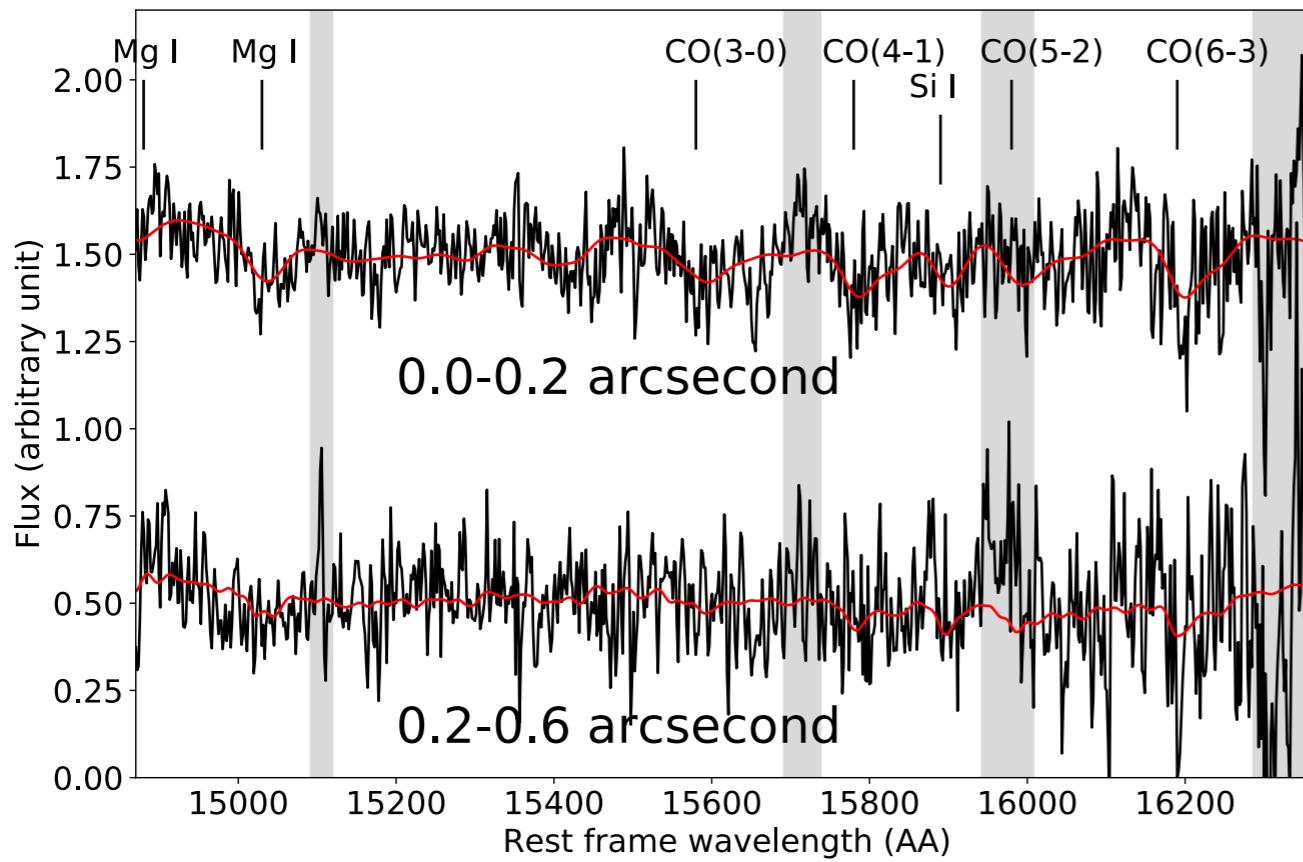


Approach 2: Stellar kinematics from Keck/OSIRIS



Data

Integration time: 4 hours
Target: 8 hours



Tackling systematic: community data challenge

- To check for **bias** from different codes and modelers
- Simulated data with known, but hidden H_0 (Ding, Treu, Shajib et al.)
- 3 different levels with increasing complexity to understand the source of systematics, if any.
- Challenge finished recently, result to be announced soon

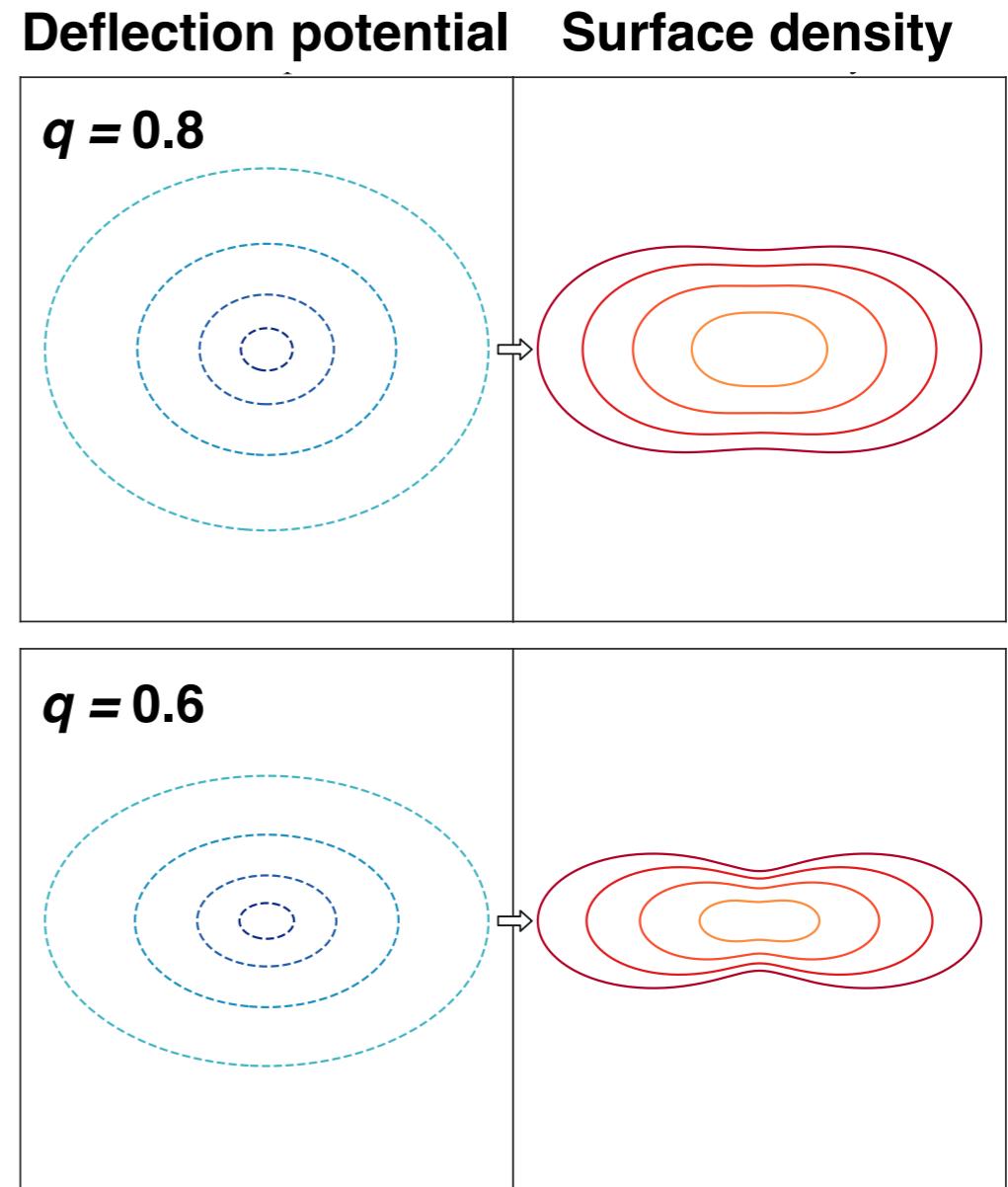
Tackling systematic: exploring more mass models with a novel method

- Elliptical mass profiles are analytically difficult for lensing.

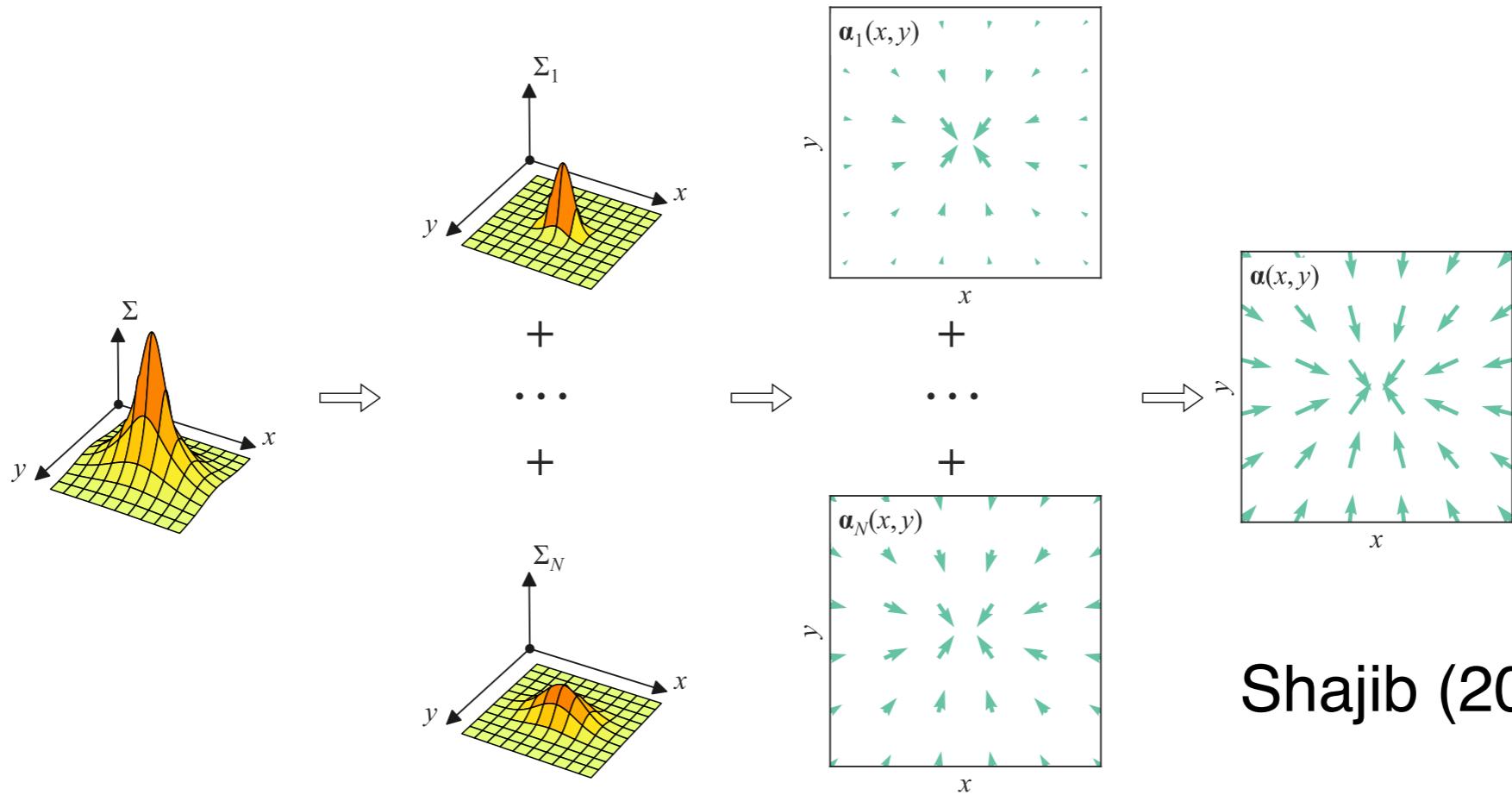
$$\text{deflection angle} = \int \text{surface density}$$

$$\text{potential} = \int \text{deflection angle}$$

- No general solution for three decades.



Tackling systematic: exploring more mass models with a novel method



Shajib (2019b)

- General and efficient
- Readily pluggable to Jeans anisotropic modeling of kinematics → unified lensing and kinematic analysis

Summary

- 7th time-delay lens gives most precise H_0 measurement at 3.9%
- Preliminary combination increases tension with early-Universe probes to $\sim 5.7\sigma$
- Future directions:
 - Automated lens modelling for large samples
 - Spatially resolved kinematics to improve precision per lens
 - 1% H_0 measurement forecasted from ~ 40 lenses