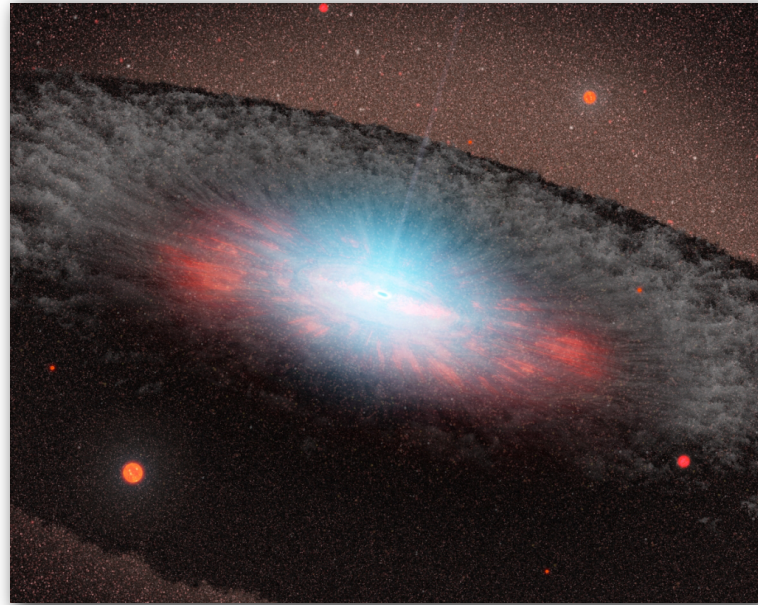


Black hole growth in the local universe



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Why do galaxy and black hole growth trace each other?

Which galaxies host accreting black holes?

Do the AGN interact with their host galaxies?

Data



SDSS DR7 - photometry & spectra for galaxies & AGN



Galaxy Zoo 1 & 2 - detailed visual morphologies for
~1 million SDSS galaxies (publicly available soon!)
Over 230,000 members of the public involved

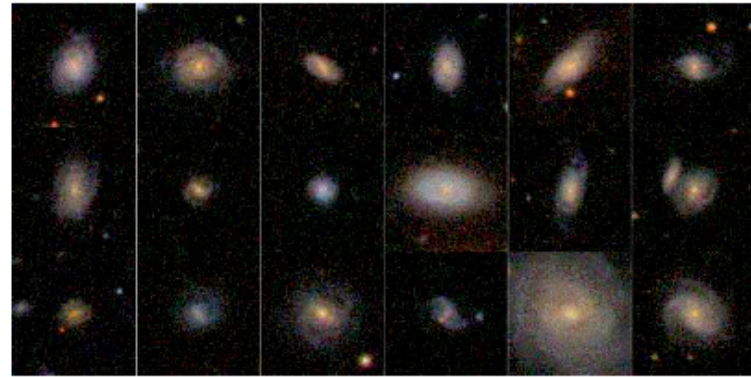


Galaxy Zoo Morphologies

Early-type



Face-on late-type



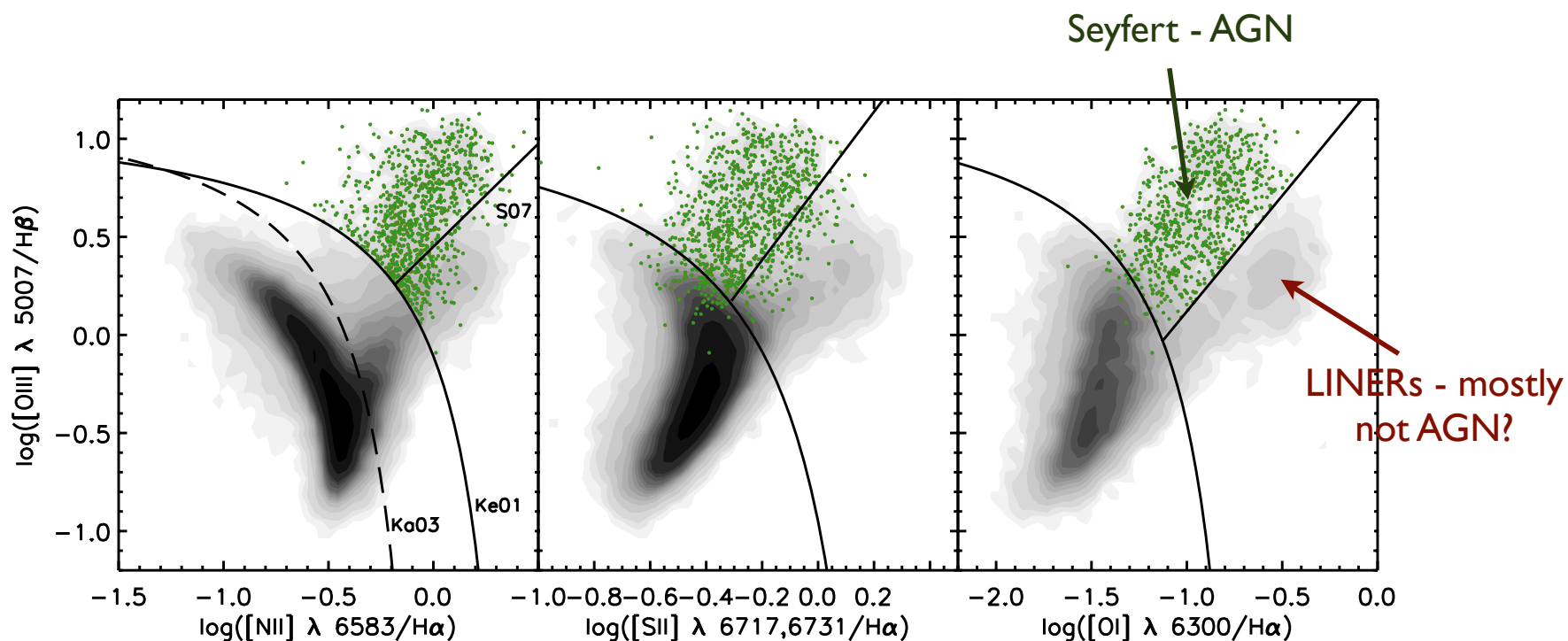
Lintott+08

At ~ 20 classifications per objects, the classifications from citizen scientists are as good as those from professionals.

Avoids biases that plague automated methods that use colour, spectral information or structural parameters as a proxy for morphology.



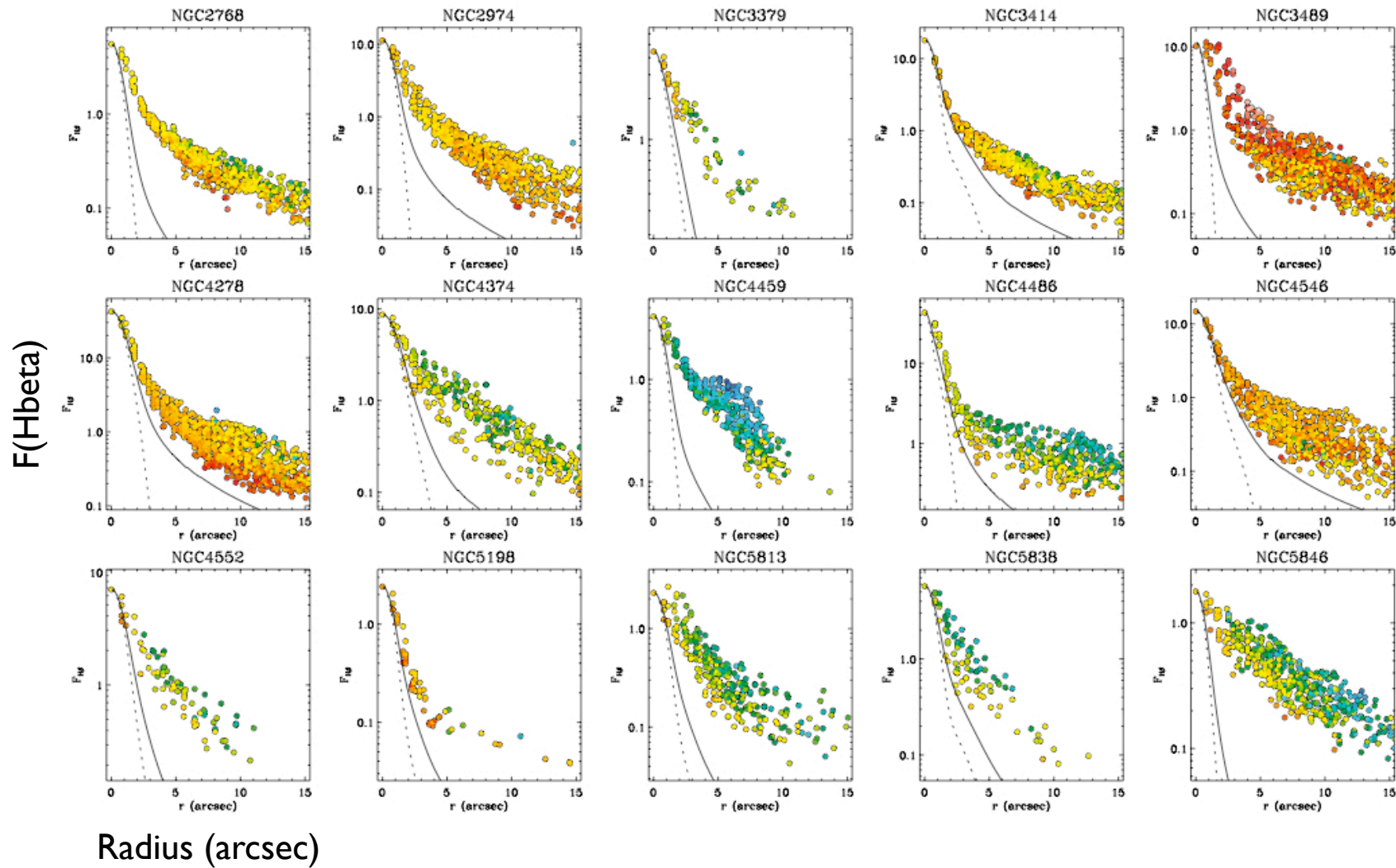
AGN Selection via emission line diagnostics



Emission line ratio diagrams (Kewley+01,+06, Kauffmann+03, Schawinski+07)



Extended LINER emission in SAURON early-type galaxies

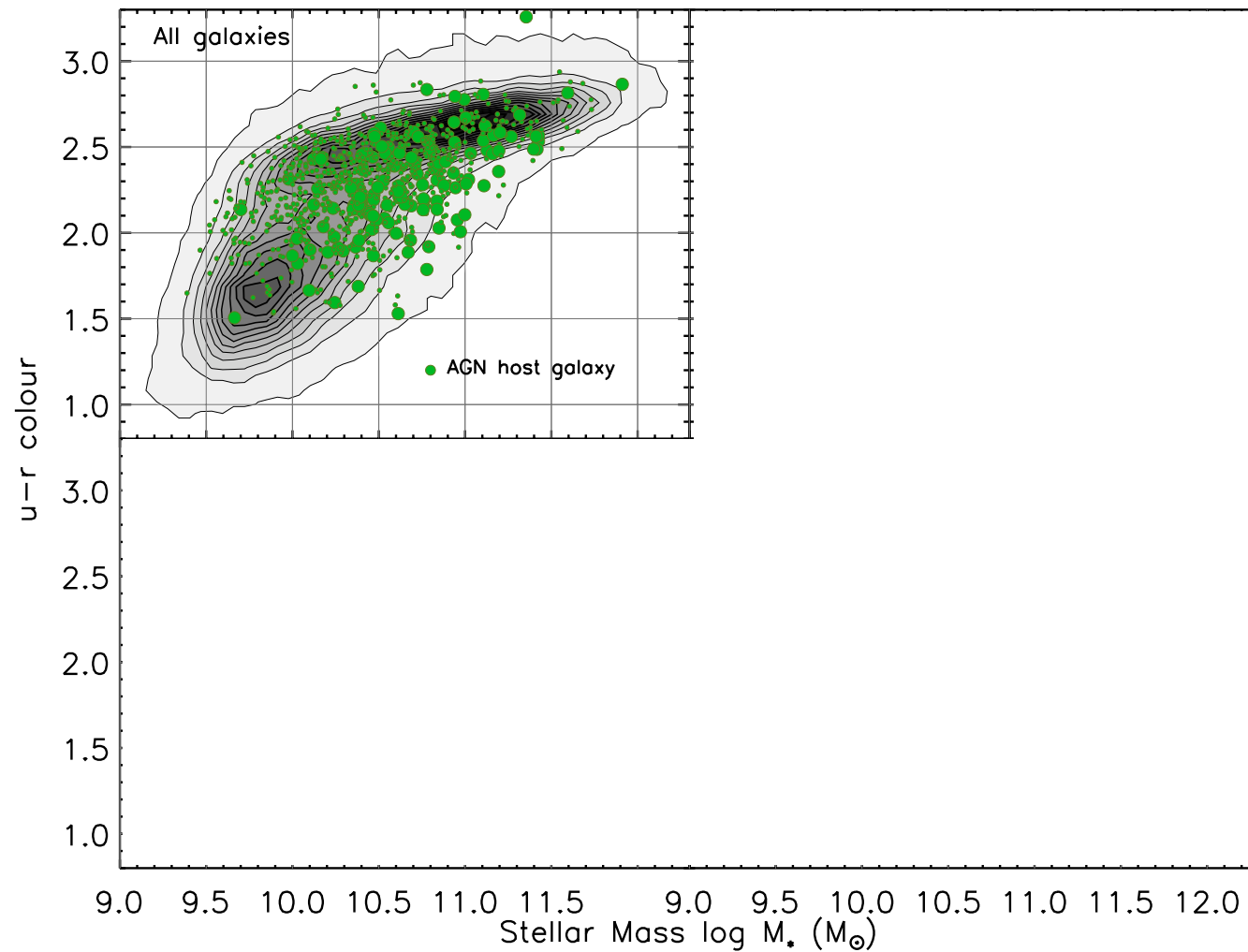


Sarzi, Shields, Schawinski+10



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AGN host galaxies split by morphology

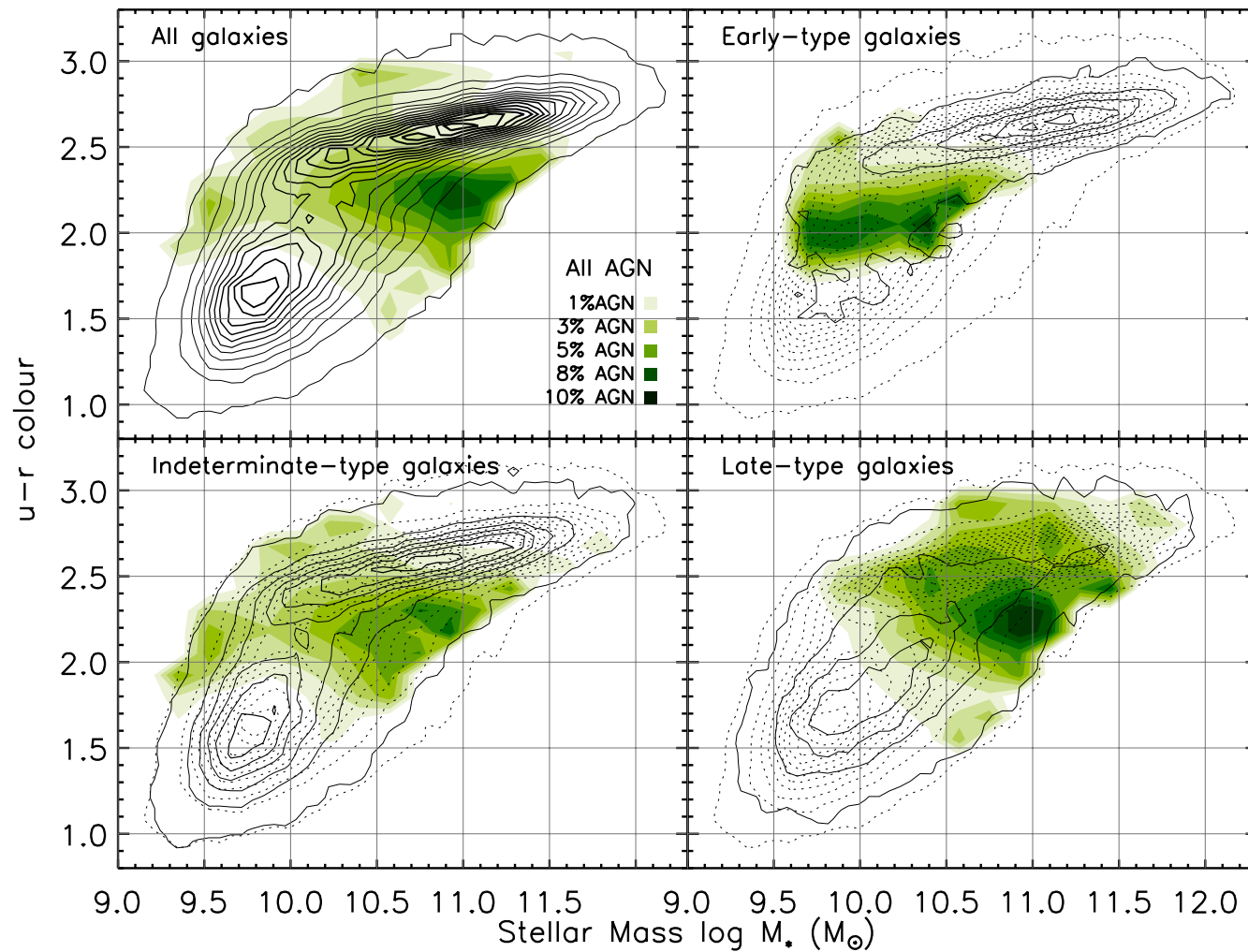


Schawinski+10a



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AGN duty cycle split by morphology



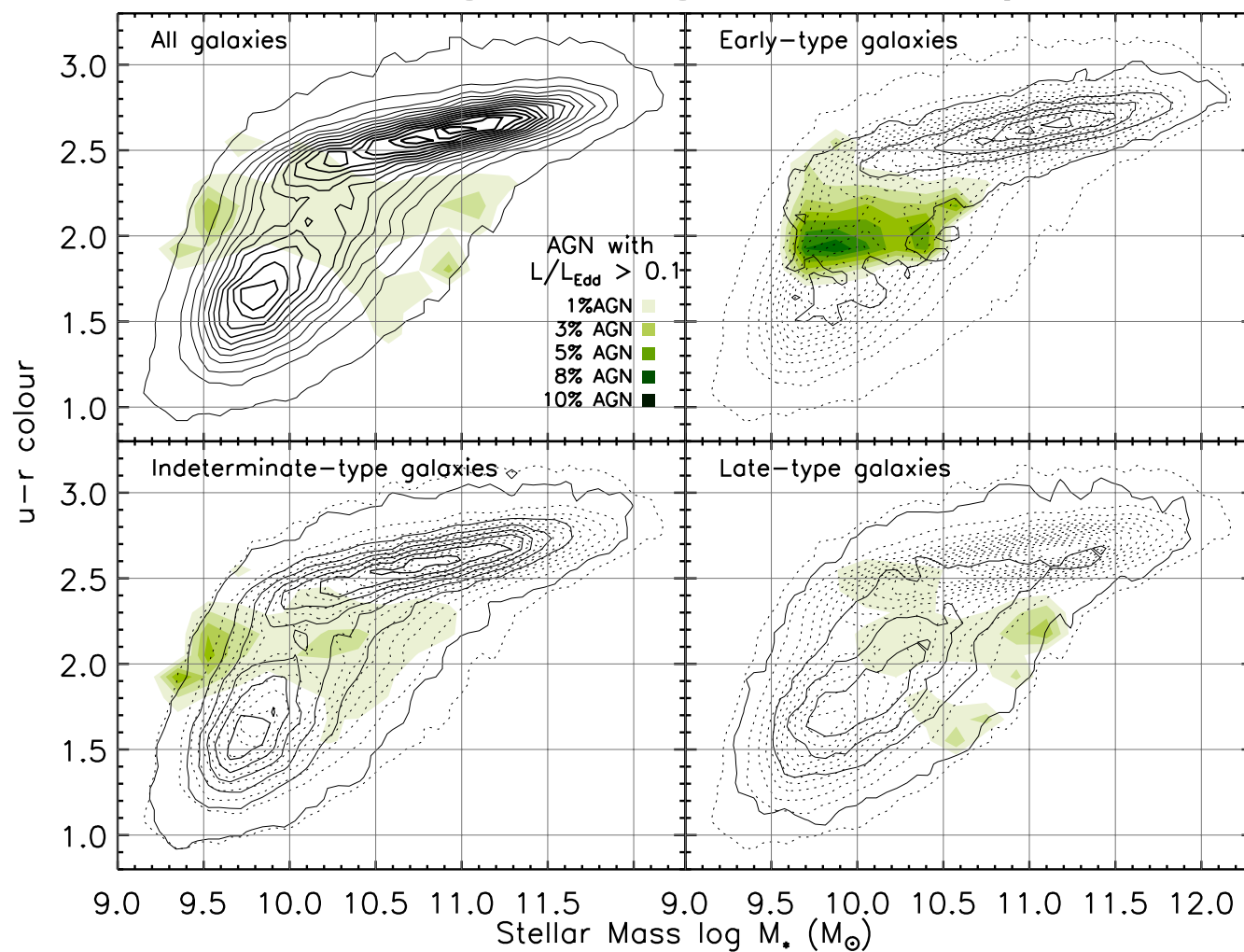
Schawinski+10a



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AGN duty cycle split by morphology

High Eddington ratio *only*

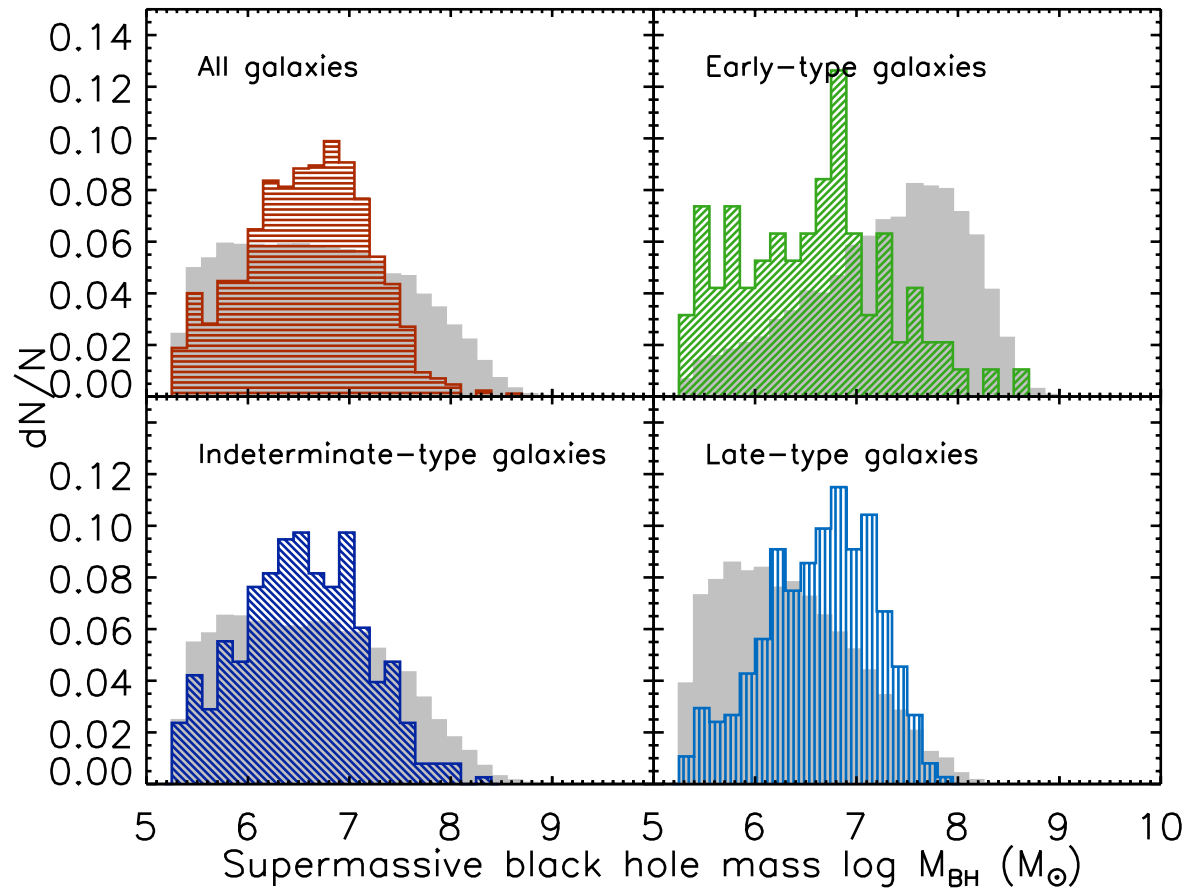


Schawinski+10a



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Which black holes are growing?

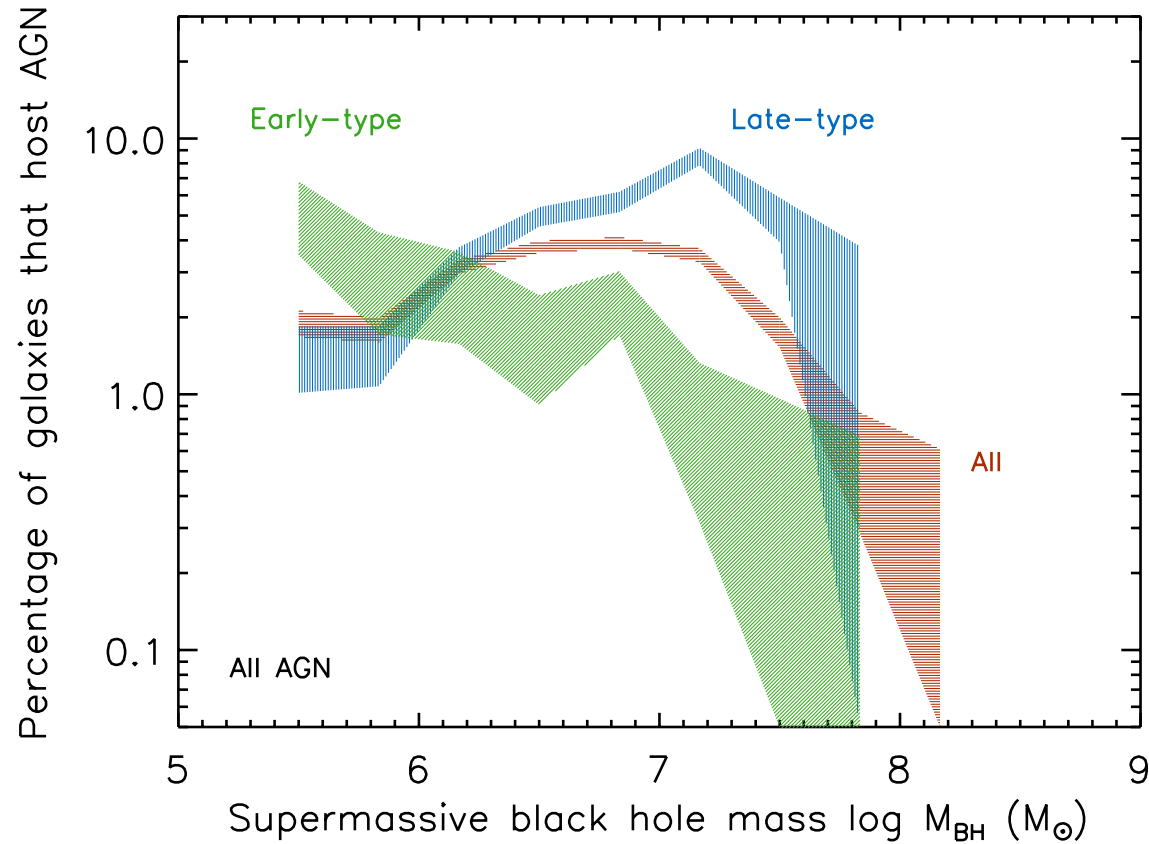


Schawinski+10a



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Which black holes are growing?

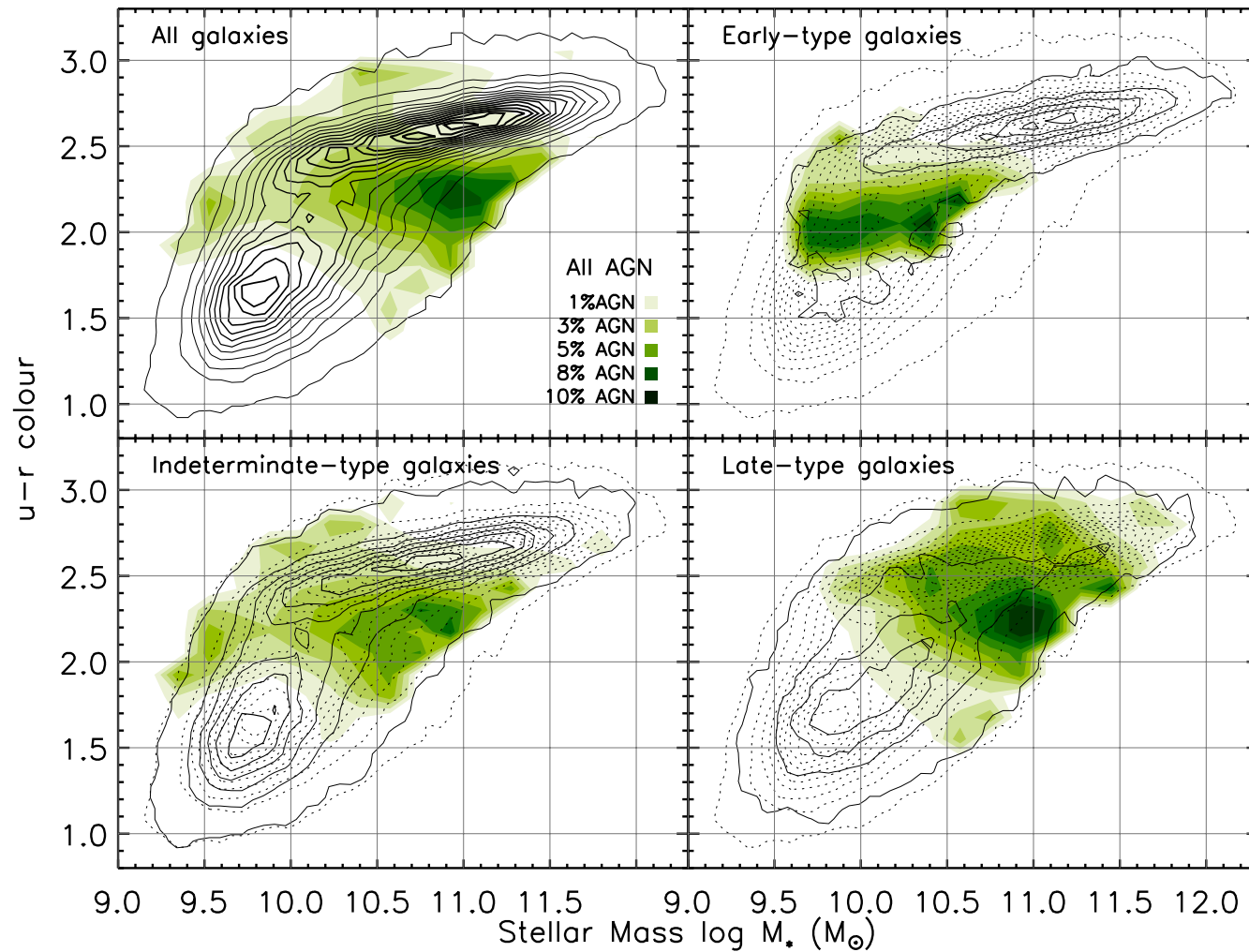


Schawinski+10a



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What does this mean for the role of the AGN for the host galaxy?

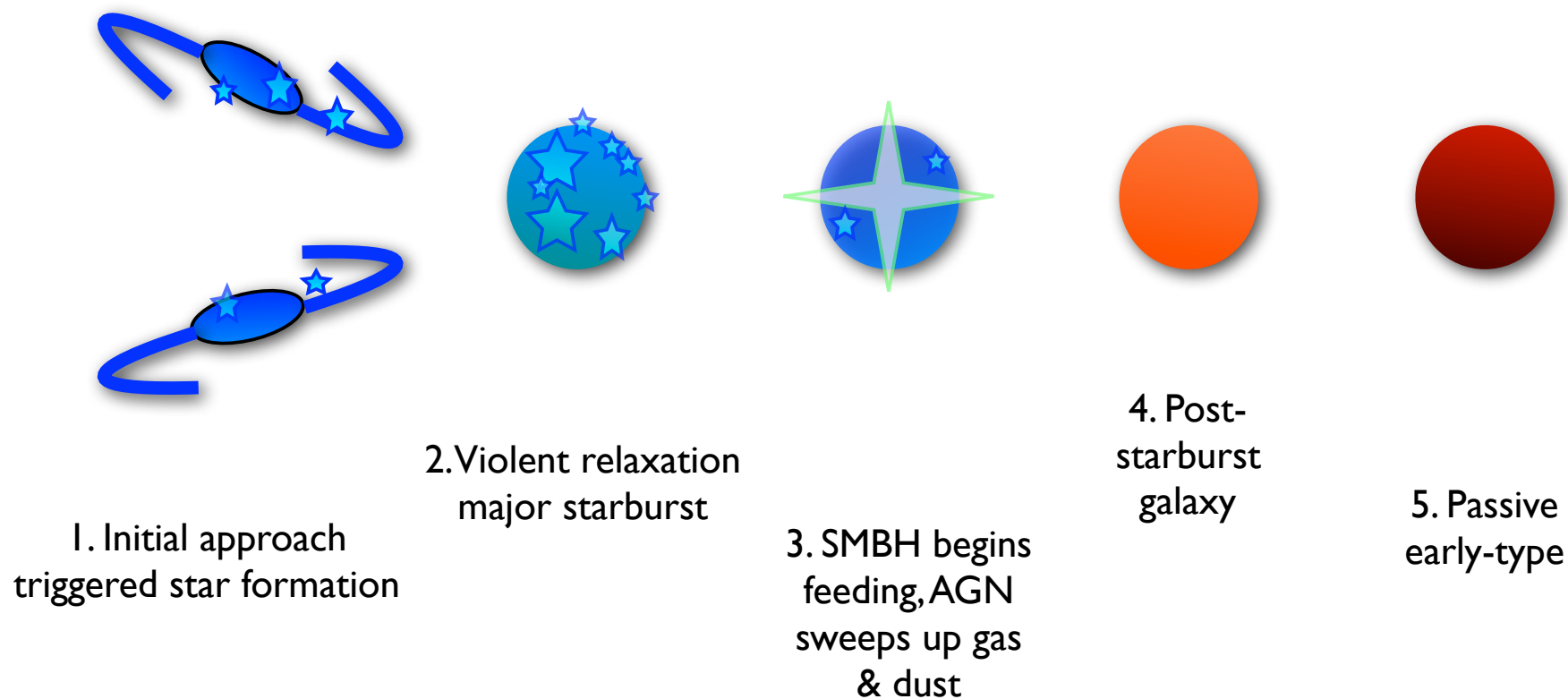


Schawinski+10a



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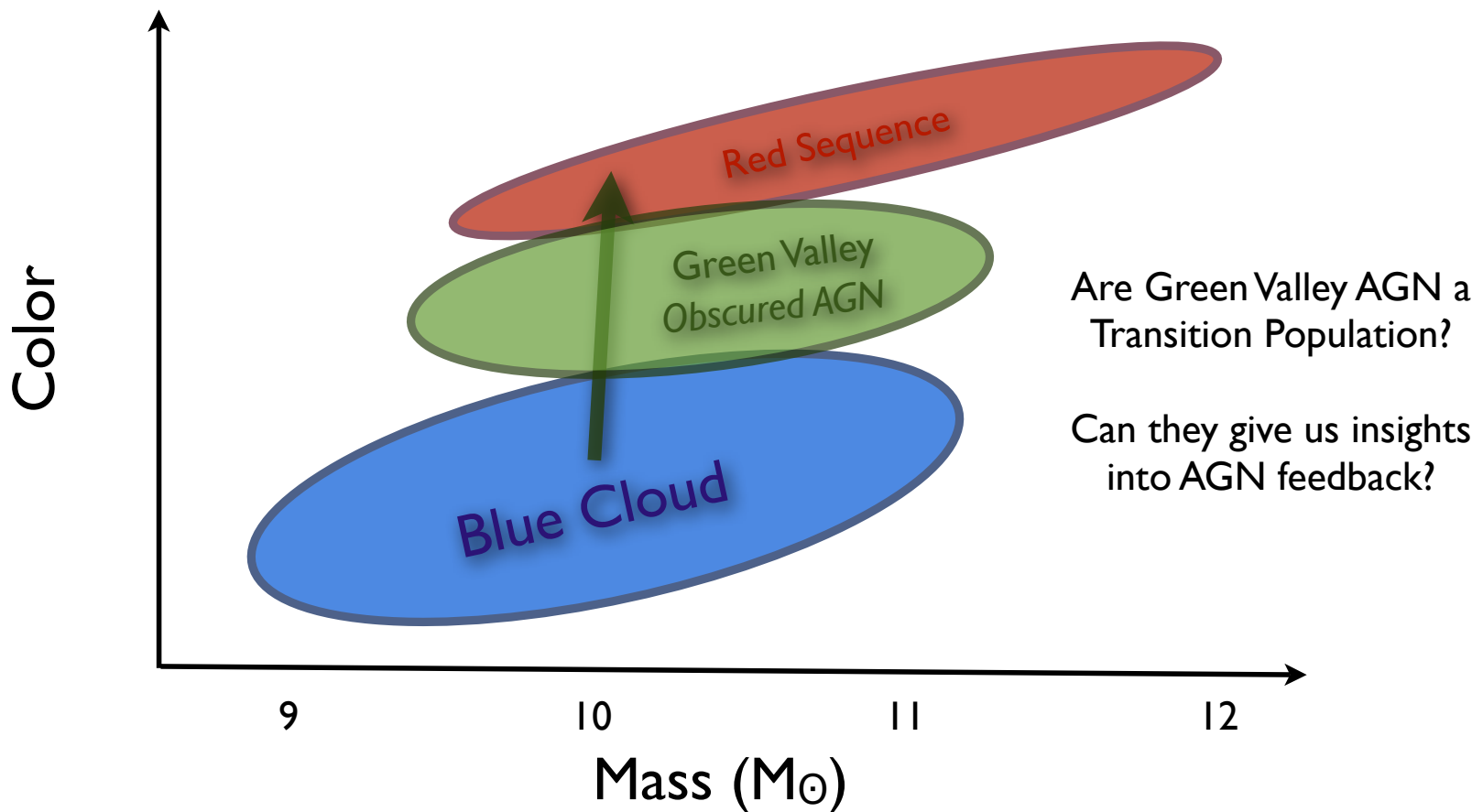
A Sequence of Events for AGN Feedback in the formation of early-type galaxies



Sanders et al. (1988), di Matteo et al. (2005), Springel et al. (2005), Hopkins et al. (2006), Johansson et al. (2009)

Observational Evidence for Migration from the blue cloud to the red sequence

...via an AGN phase in the green valley?



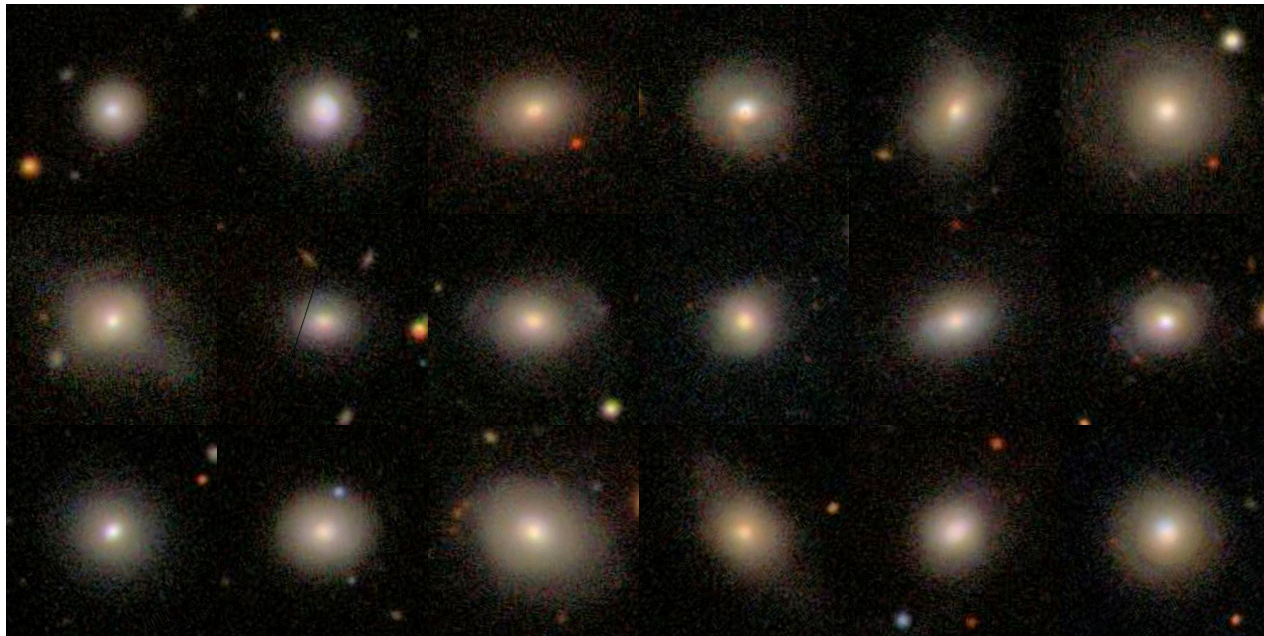
What does the Green Valley really mean?

Intermediate ('green') colours do not necessarily imply that star formation was recently shut down

but

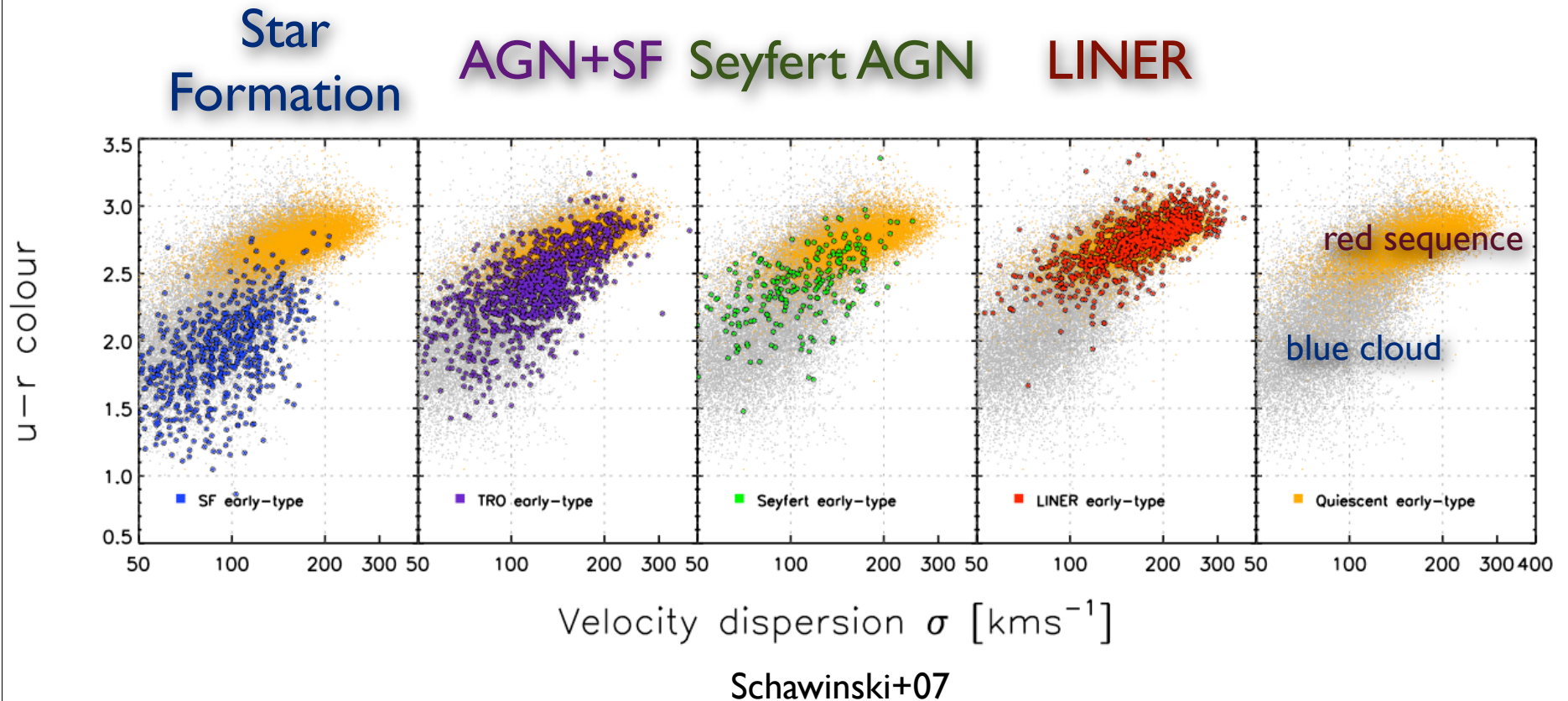
If star formation has recently been suppressed, then intermediate colours imply a time delay on the order of the lifetime of OB stars

Low-mass active early-types in SDSS



Schawinski et al. 2009, MNRAS

SDSS reveals an evolutionary sequence



Is this a genuine sequence? Are the AGN a part of it?

A Method for Recovering Star Formation Histories

Breaking degeneracies

1. Parameterisation

Describe the SFH as a two-component model
Derive age & mass-fraction of most recent burst.

- Old component – SSP (Maraston 2005)
- Young Component – CSP (exponential, $\tau=100$ Myr)
- Dust (Calzetti 2001)
- Metallicity from $-1.3 < [Z/H] < +0.67$ including asymmetric internal distribution

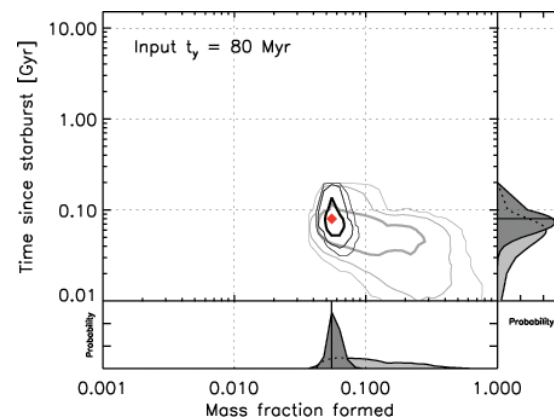
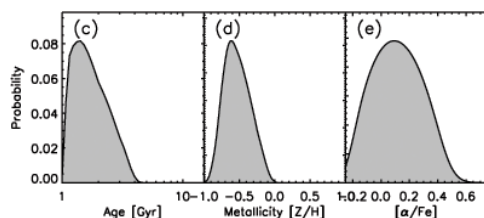
2a. Fitting Broad-band Photometry

Fit the photometry: FUV, NUV, u, g, r, i, z, J, H, K
Compute χ^2 for 5-d array. Total of ~5 million SFHs.

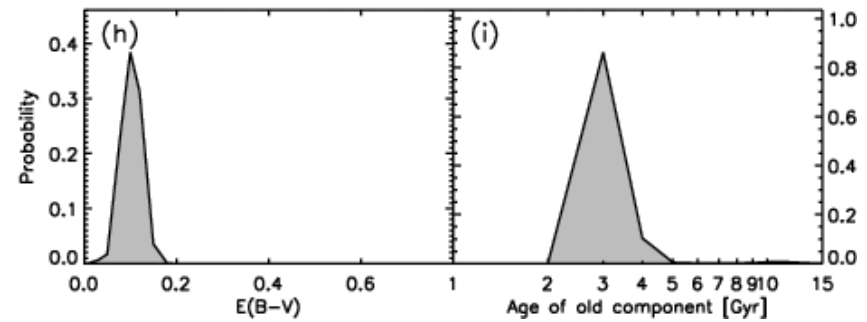
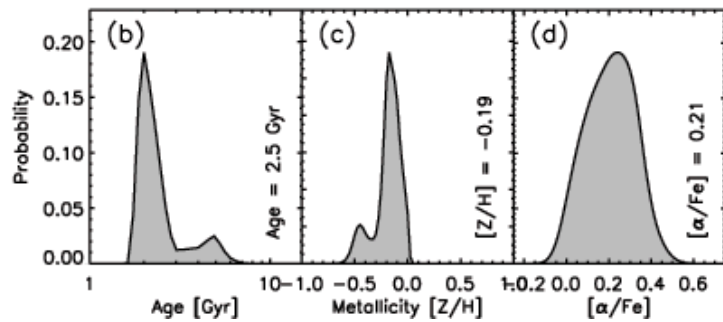
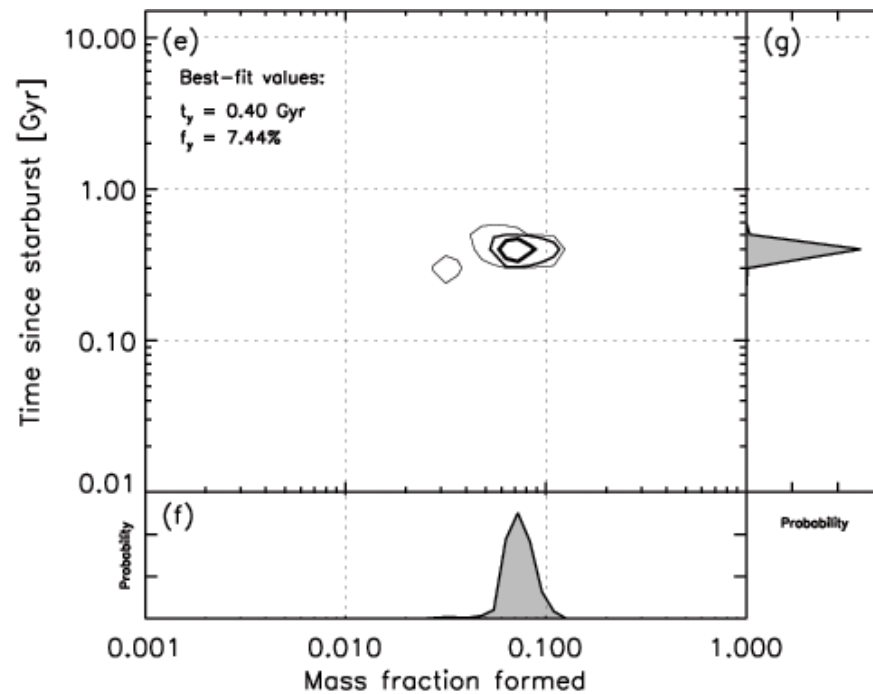
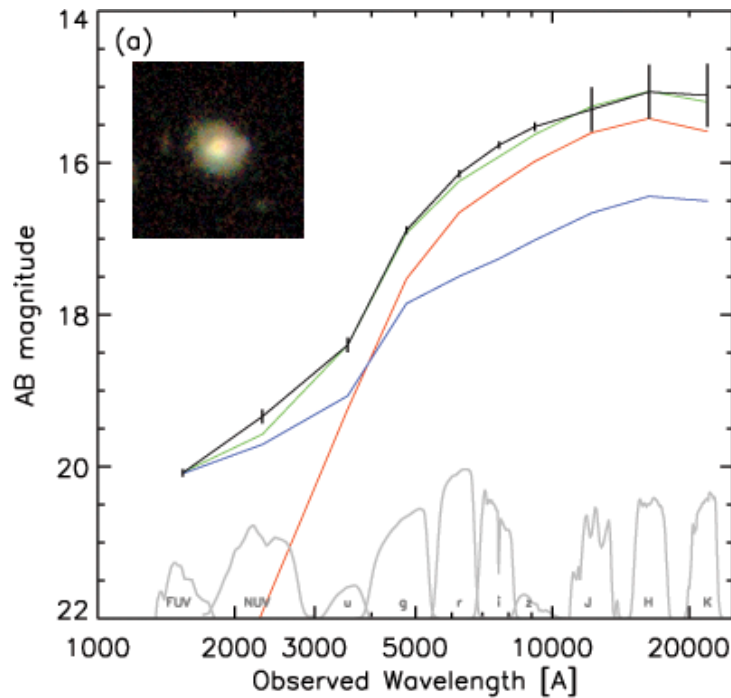
2b. Lick indices

Fit the 25 stellar absorption indices (Lick system) to models (Thomas, Maraston & Bender 2003) to get *spectroscopic* age & metallicity.

3. Combine the two Results

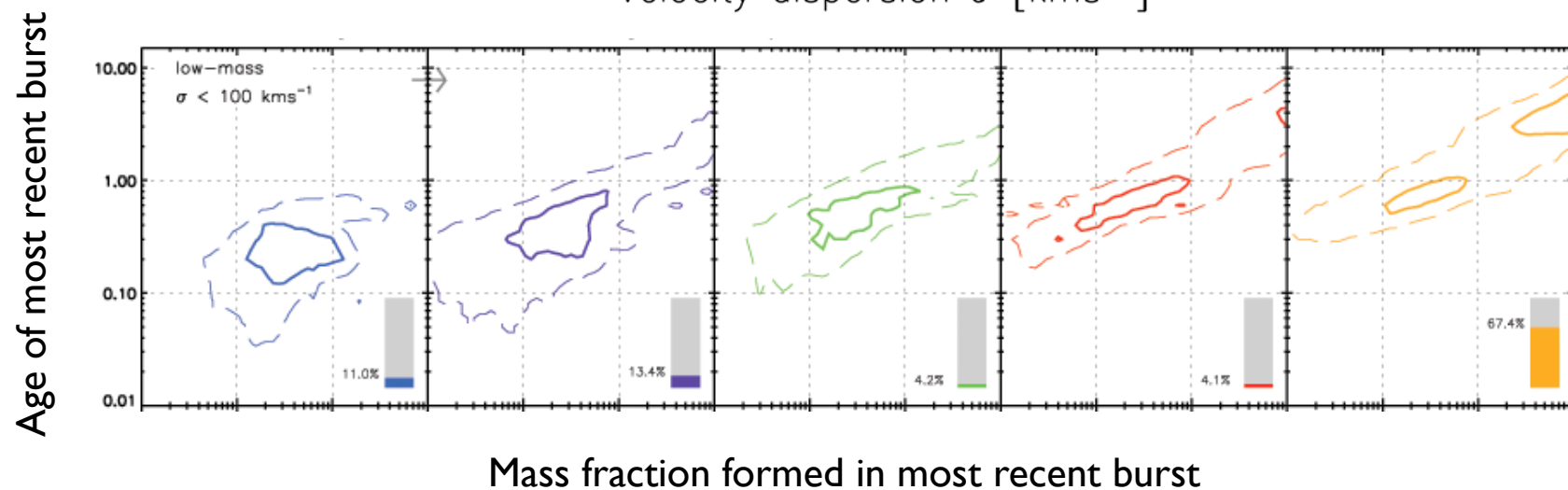
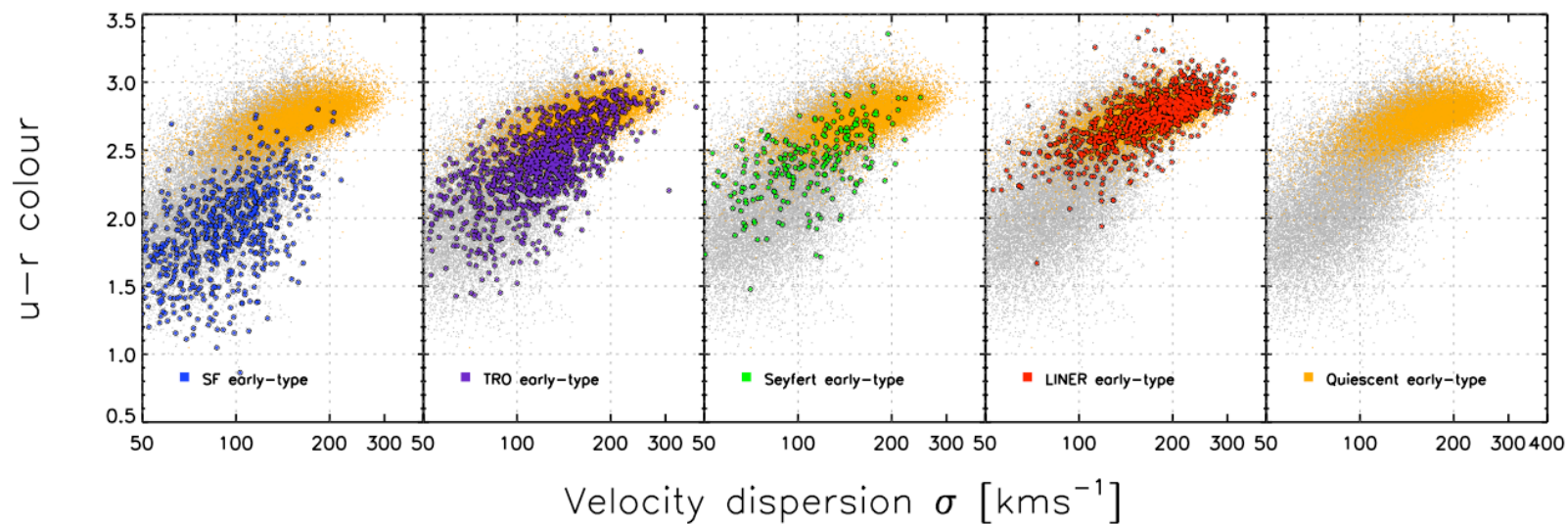


Application to Real Galaxies



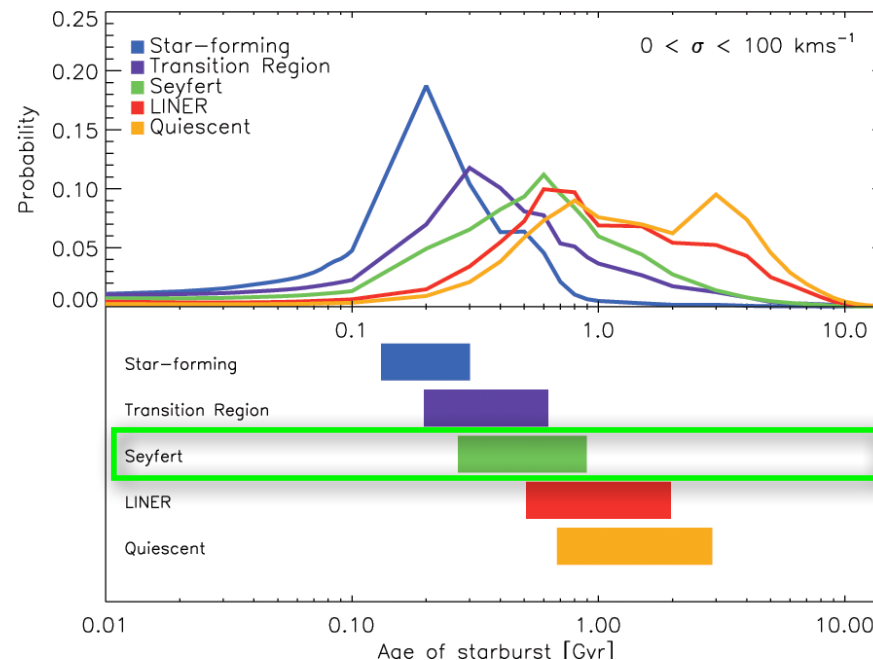
Schawinski+07

Recovered Star Formation Histories



Schawinski+07

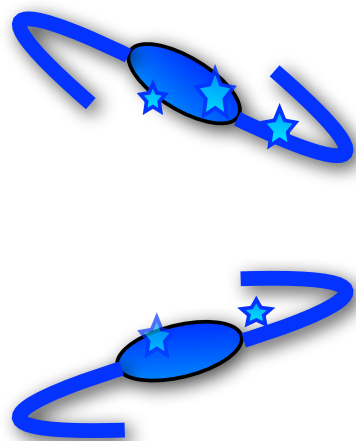
So, are we seeing AGN feedback suppressing star formation?



Transition time scales pose challenge!

Seyfert (moderate-luminosity obscured AGN) only become important ~ 0.5 Gyr after shutdown of star formation.

A Sequence of Events for AGN Feedback in the formation of early-type galaxies



1. Initial approach
triggered star formation

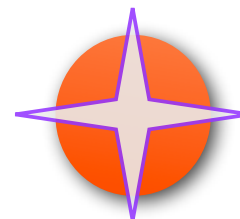


2. Violent relaxation
major starburst

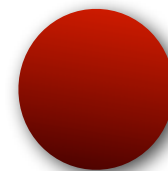


~~3. SMBH begins
feeding, AGN
sweeps up gas
& dust~~

SF shuts down

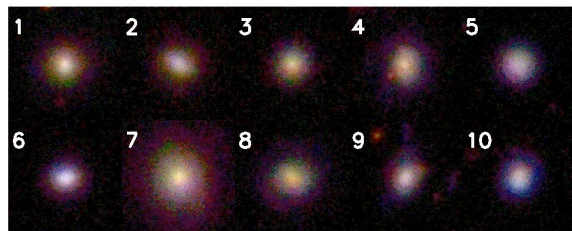
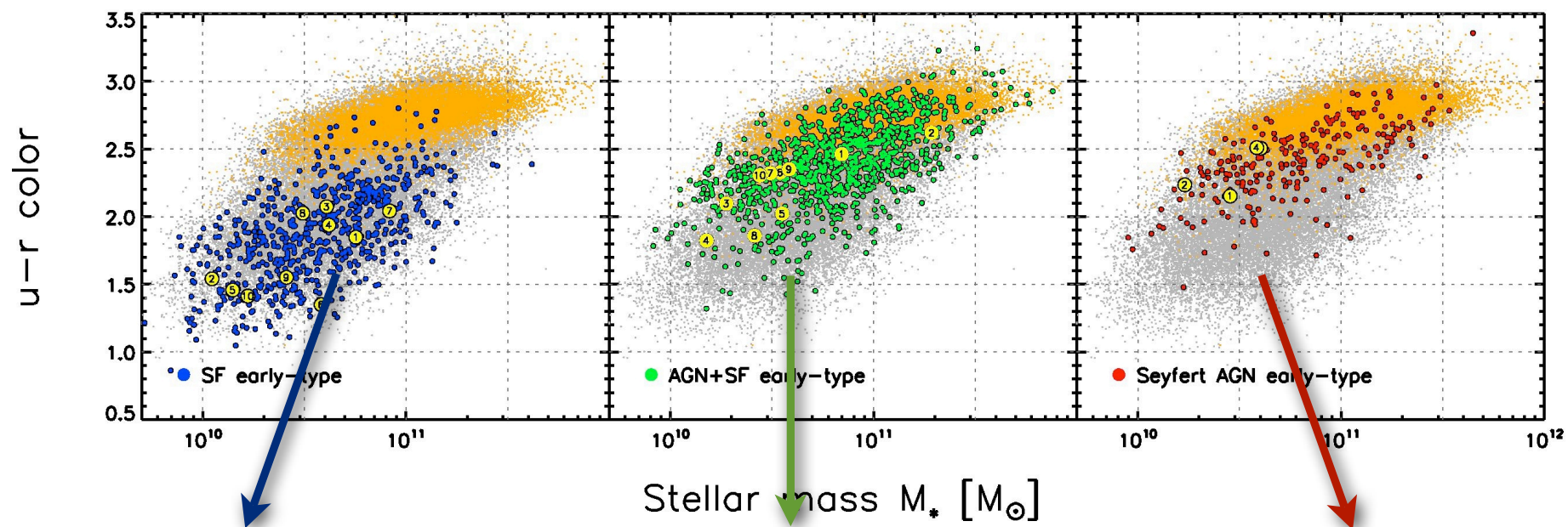


4. Post-
starburst
galaxy
+ AGN

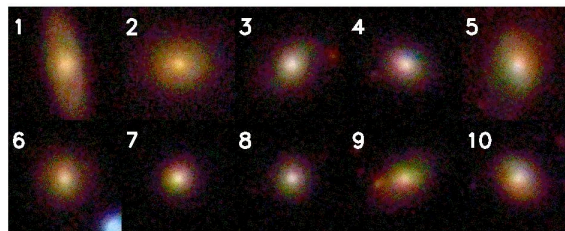


5. Passive
early-type

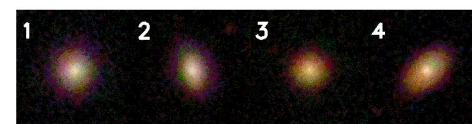
Tracing the fate of molecular gas



Starforming

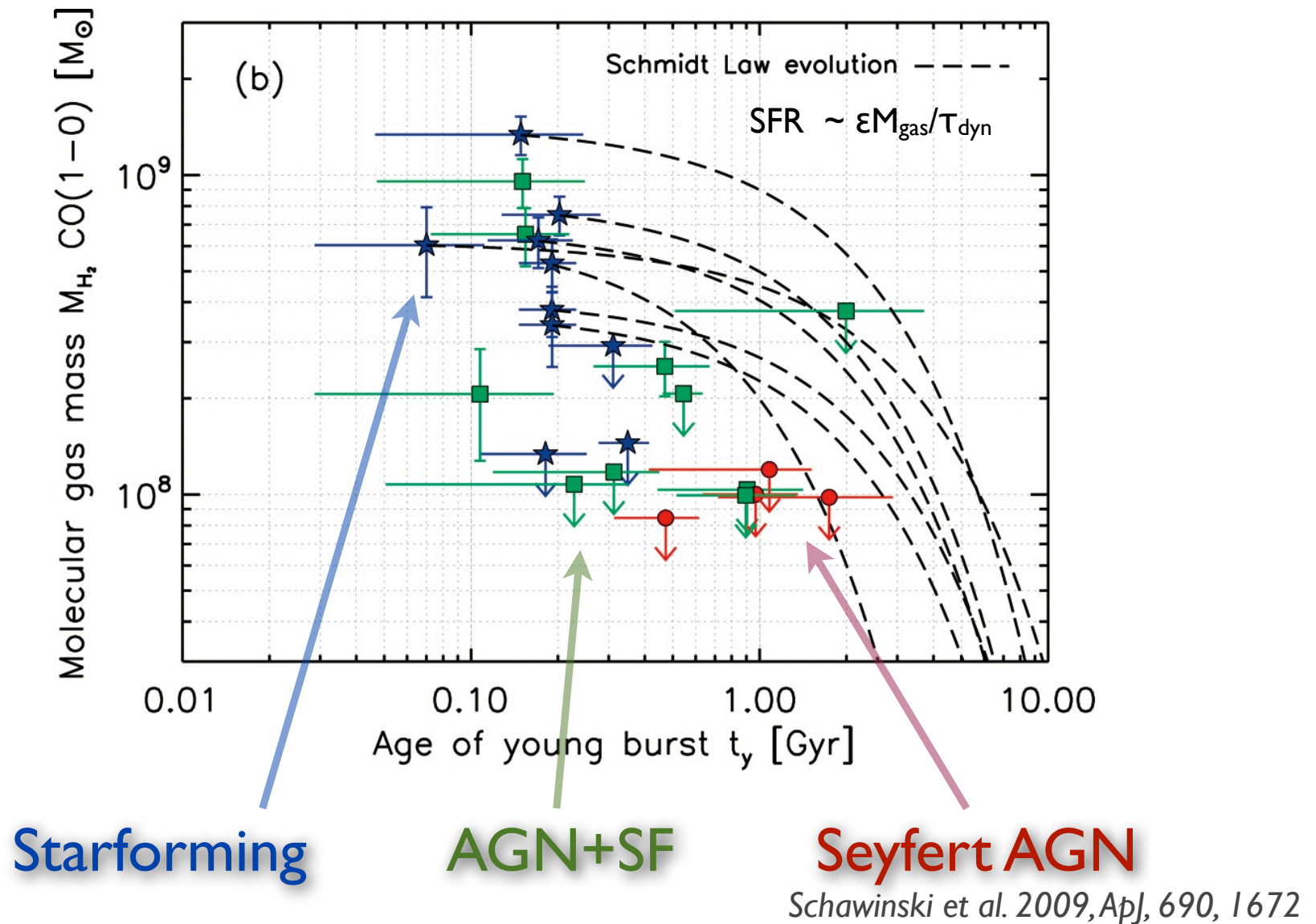


AGN+SF

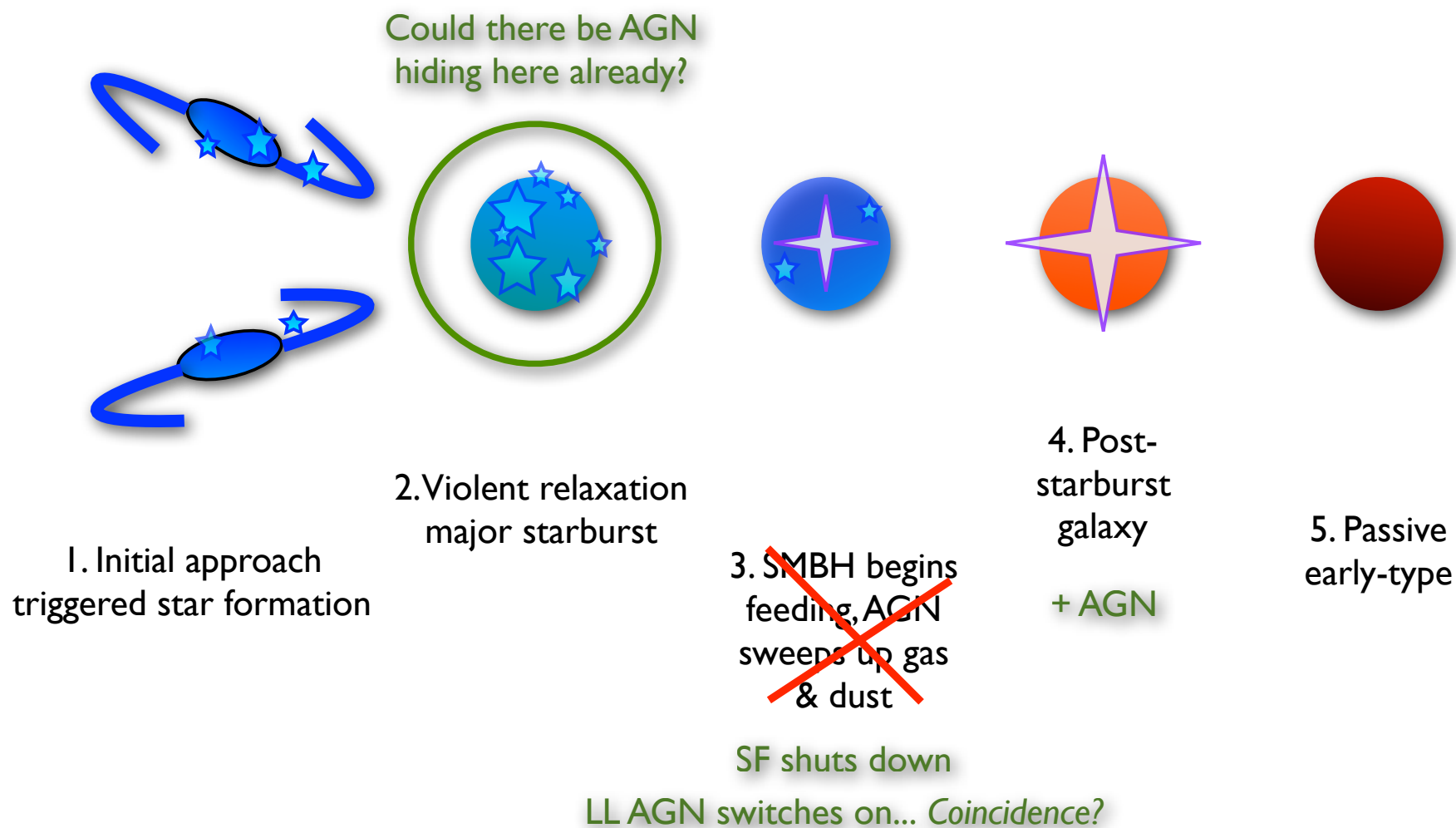


Seyfert AGN

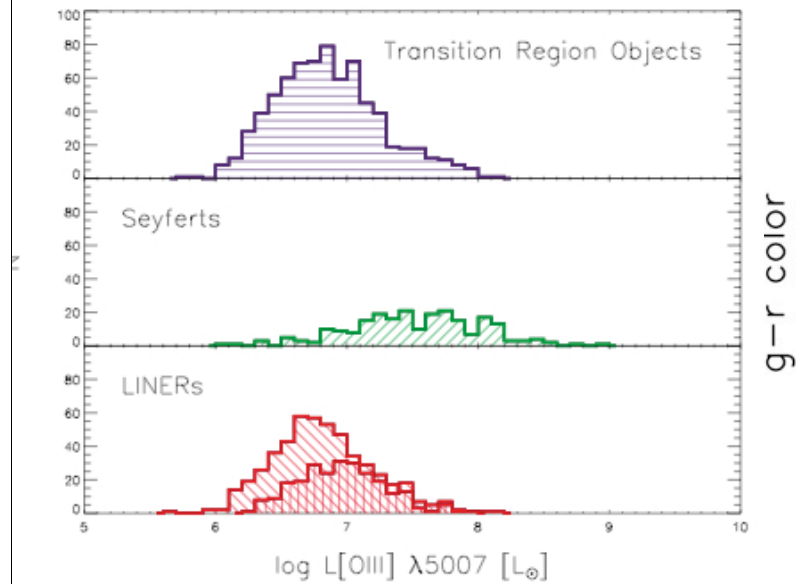
Can we rescue AGN feedback?



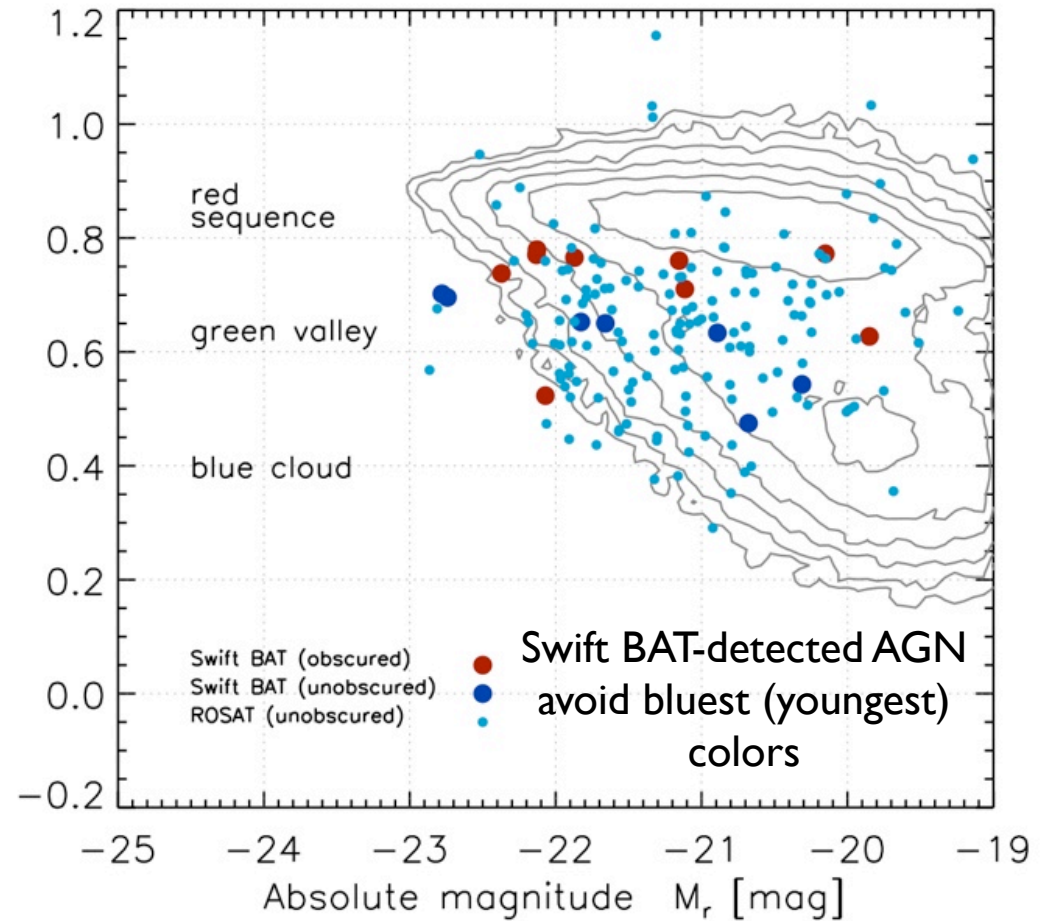
A Sequence of Events for AGN Feedback in the formation of early-type galaxies



Unlikely!



$L[\text{OIII}]$ increases w/ time

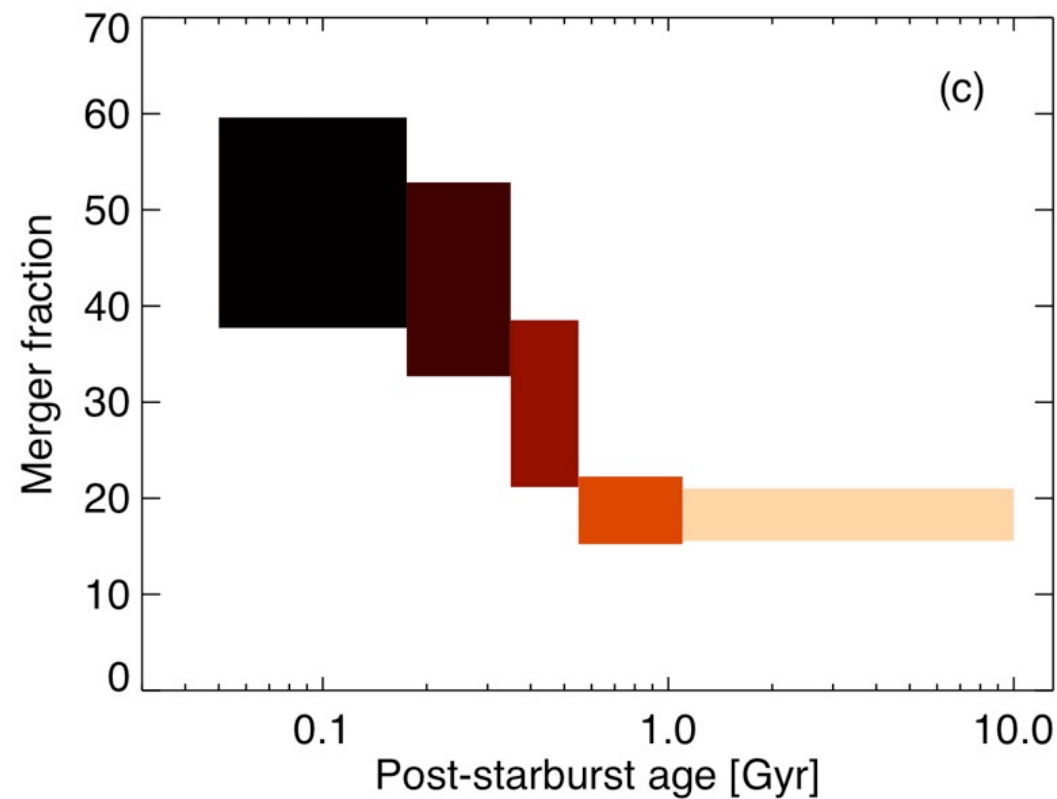


A Sequence of Events for AGN Feedback in the formation of early-type galaxies

What about the trigger?



Mergers trigger the migration from the blue cloud to the red sequence



Schawinski+10b

Summary

Early-type galaxies **continue to build** or return to the low-mass end of the red sequence.

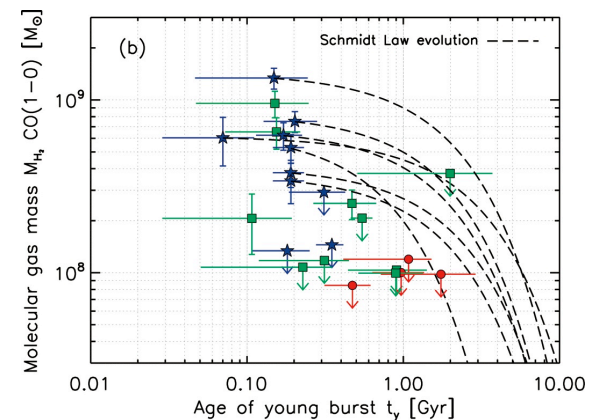
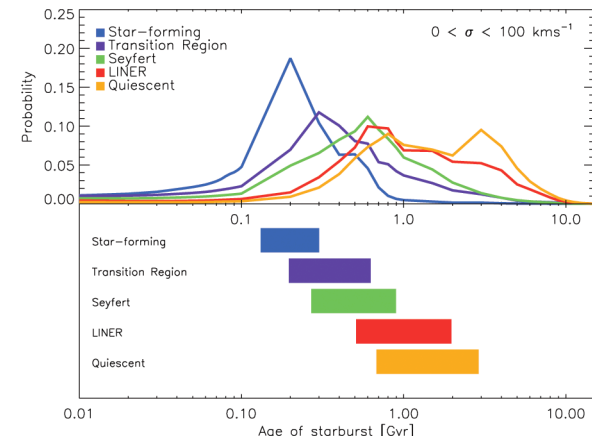
Time sequence of events includes a slow rise of AGN peaking in the green valley **~0.5 Gyr after shutdown**.

Molecular gas disappears **extremely rapidly** ~200 Myr after start of current starburst and **coincides with low-L AGN phase** -> kinetic feedback?

High-L AGN phase **too late** for suppression and cannot be responsible for shutdown of SF.

Up to 90% of local black hole growth is **not associated** with this movement of spheroidal galaxies to the red sequence - what do AGN do in late-type galaxies?

Time scales are a major puzzle!



A Sequence of Events for AGN Feedback in the formation of early-type galaxies



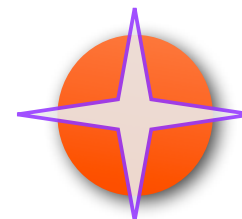
1. Initial approach
triggered star formation



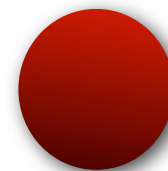
2. Violent relaxation
major starburst



3. LL AGN
switches on
-
at the same time
SF is shut down
as H_2 reservoir
is destroyed



4. Post-
starburst
galaxy hosting
Seyfert AGN



5. Passive
early-type

Thank you!