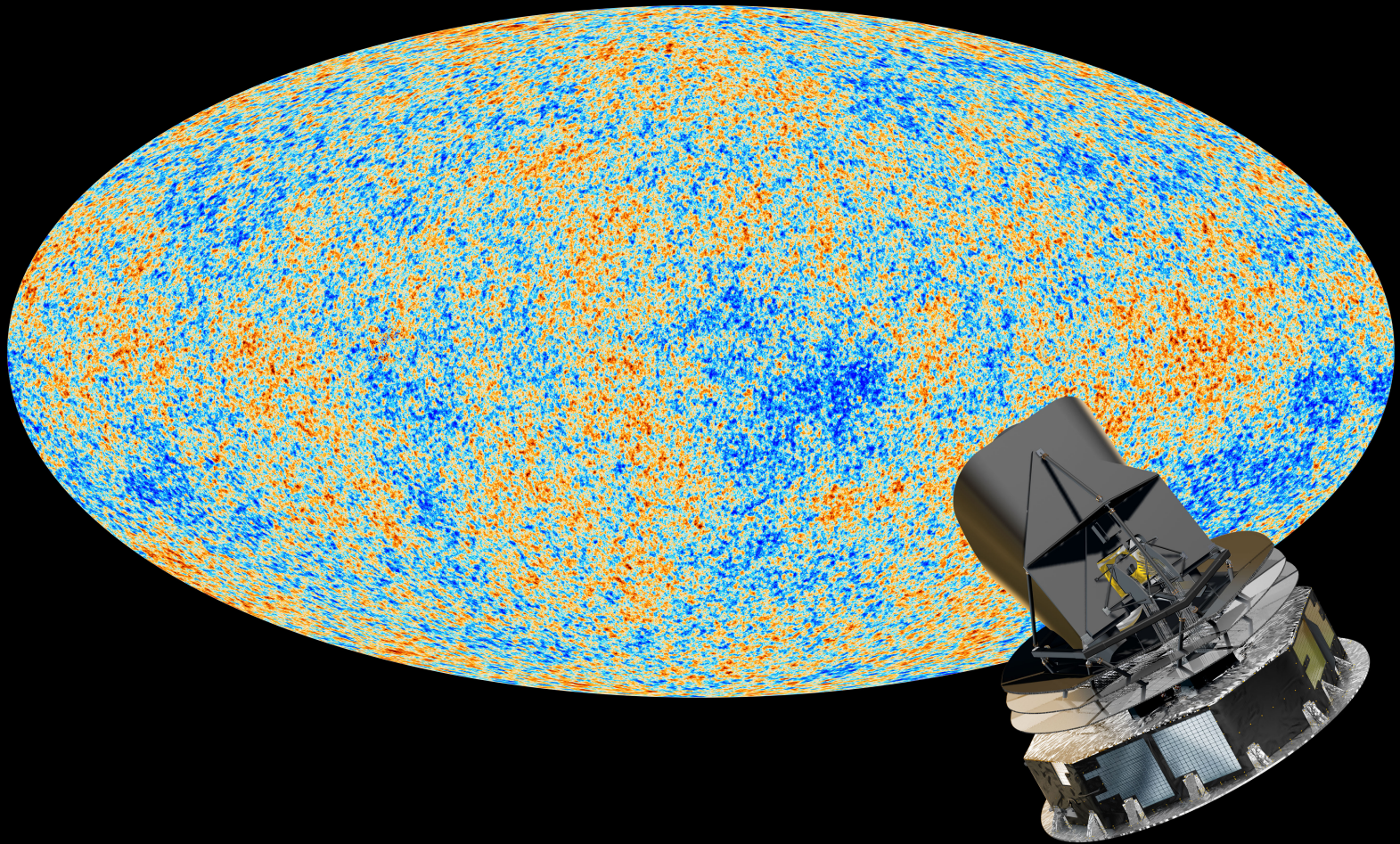
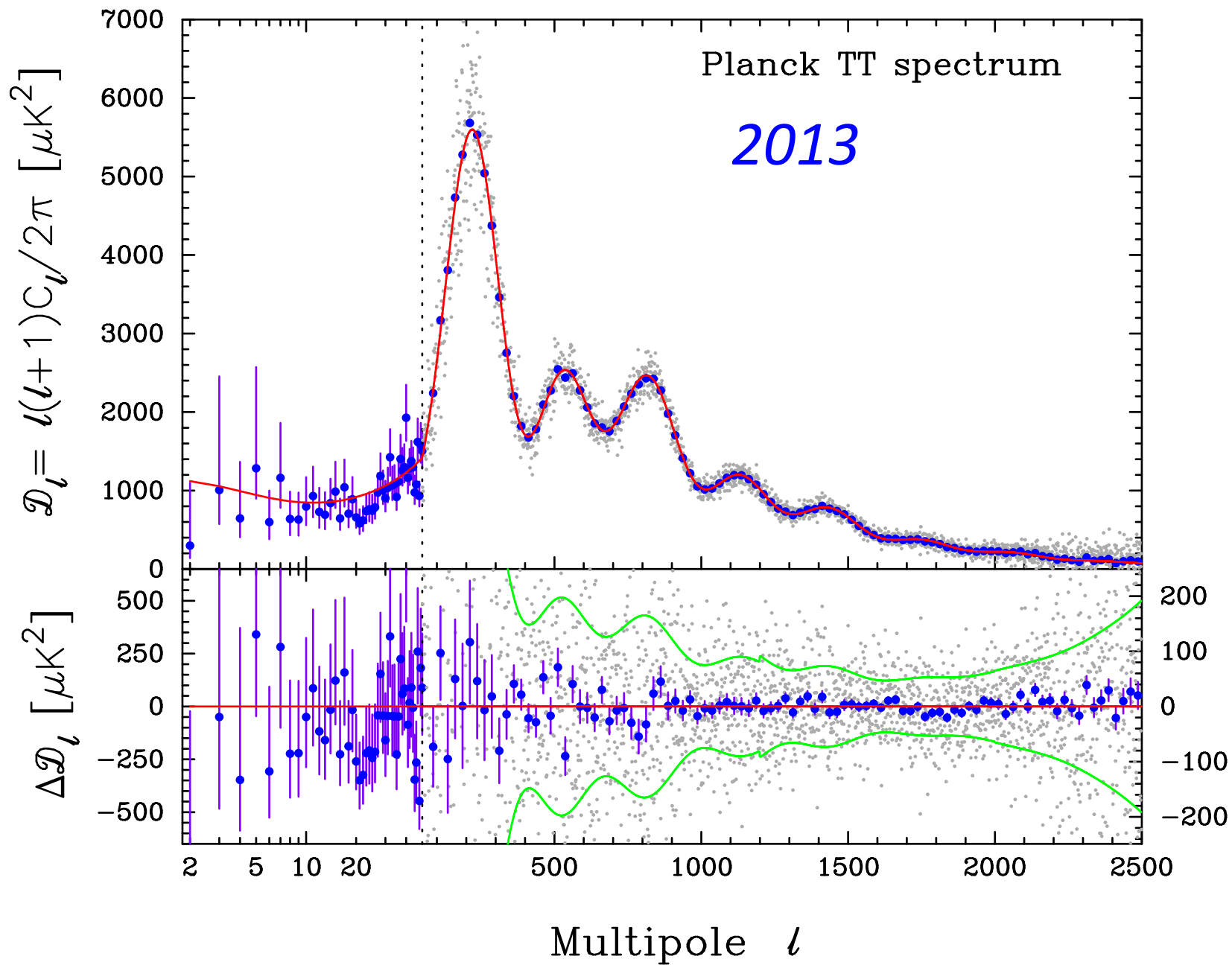


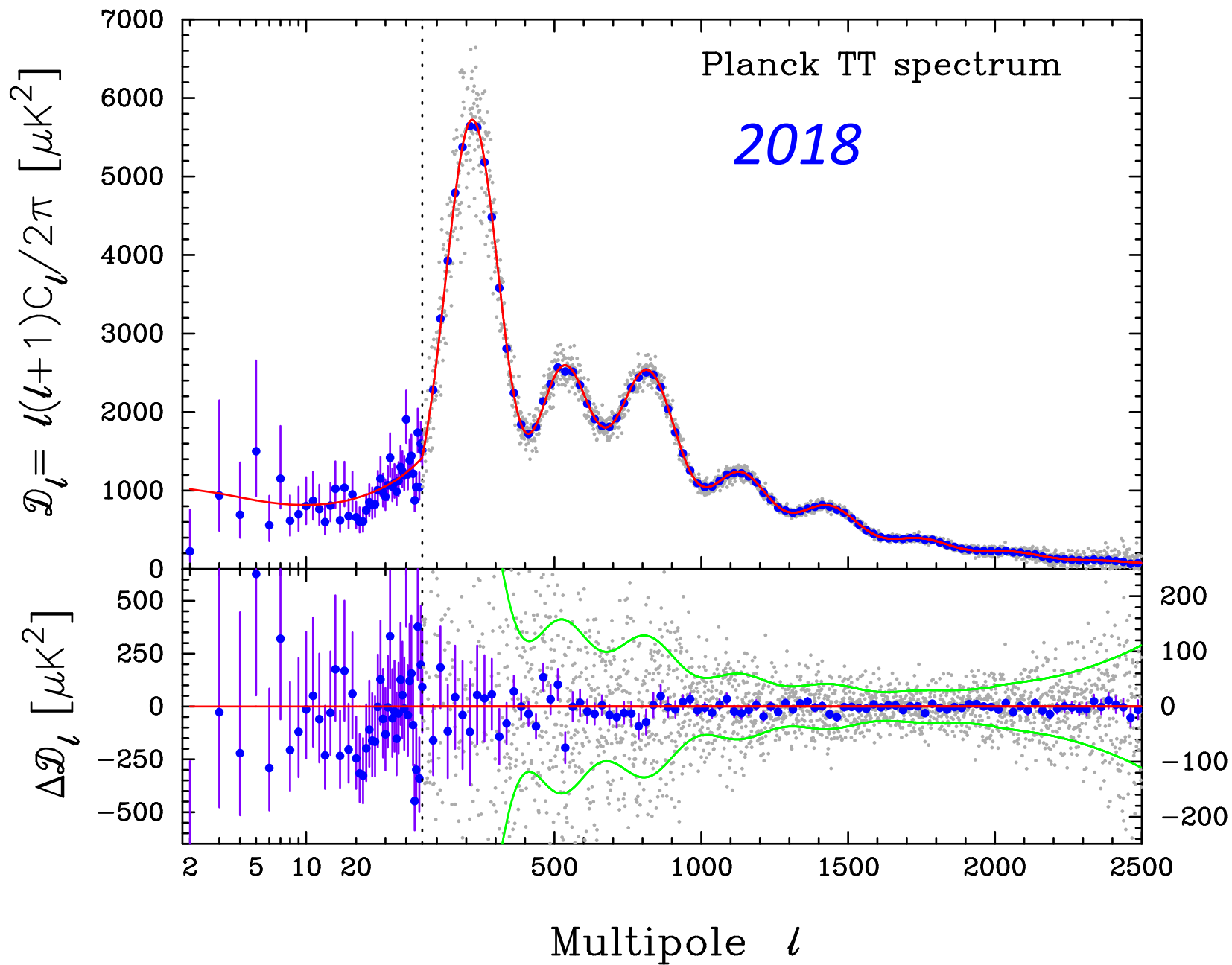
Planck, Λ CDM & H_0

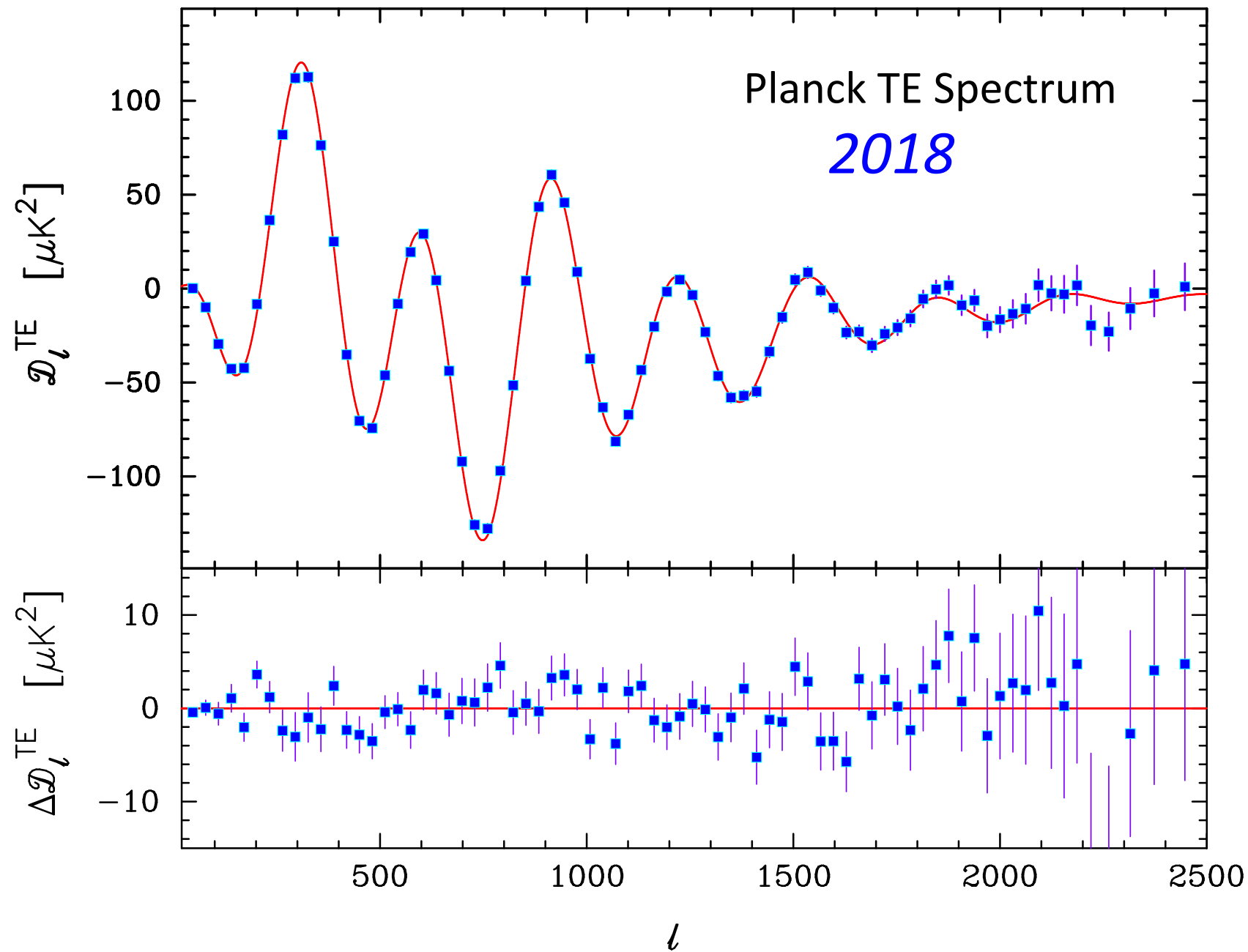
George Efstathiou KICC

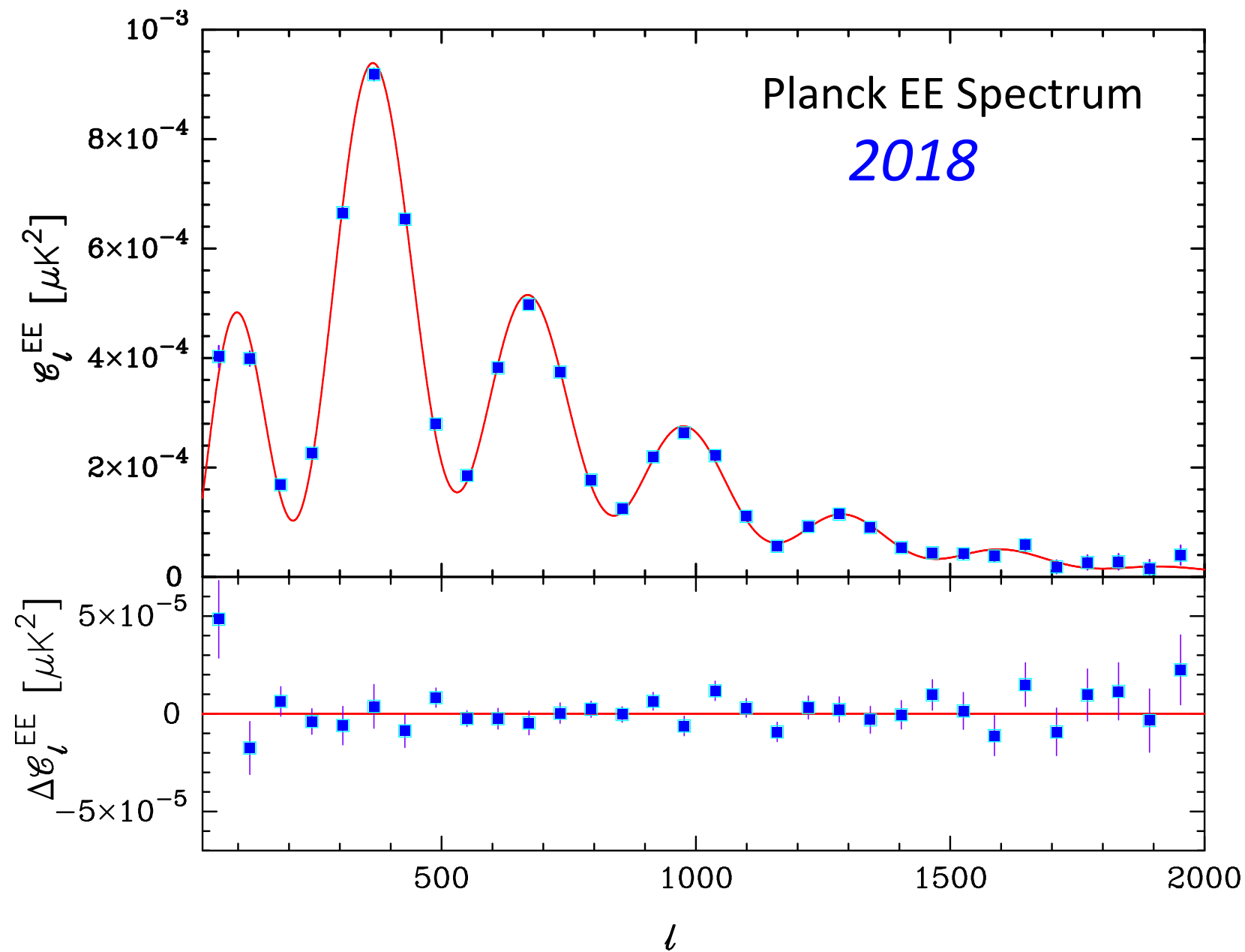


30th March 2021





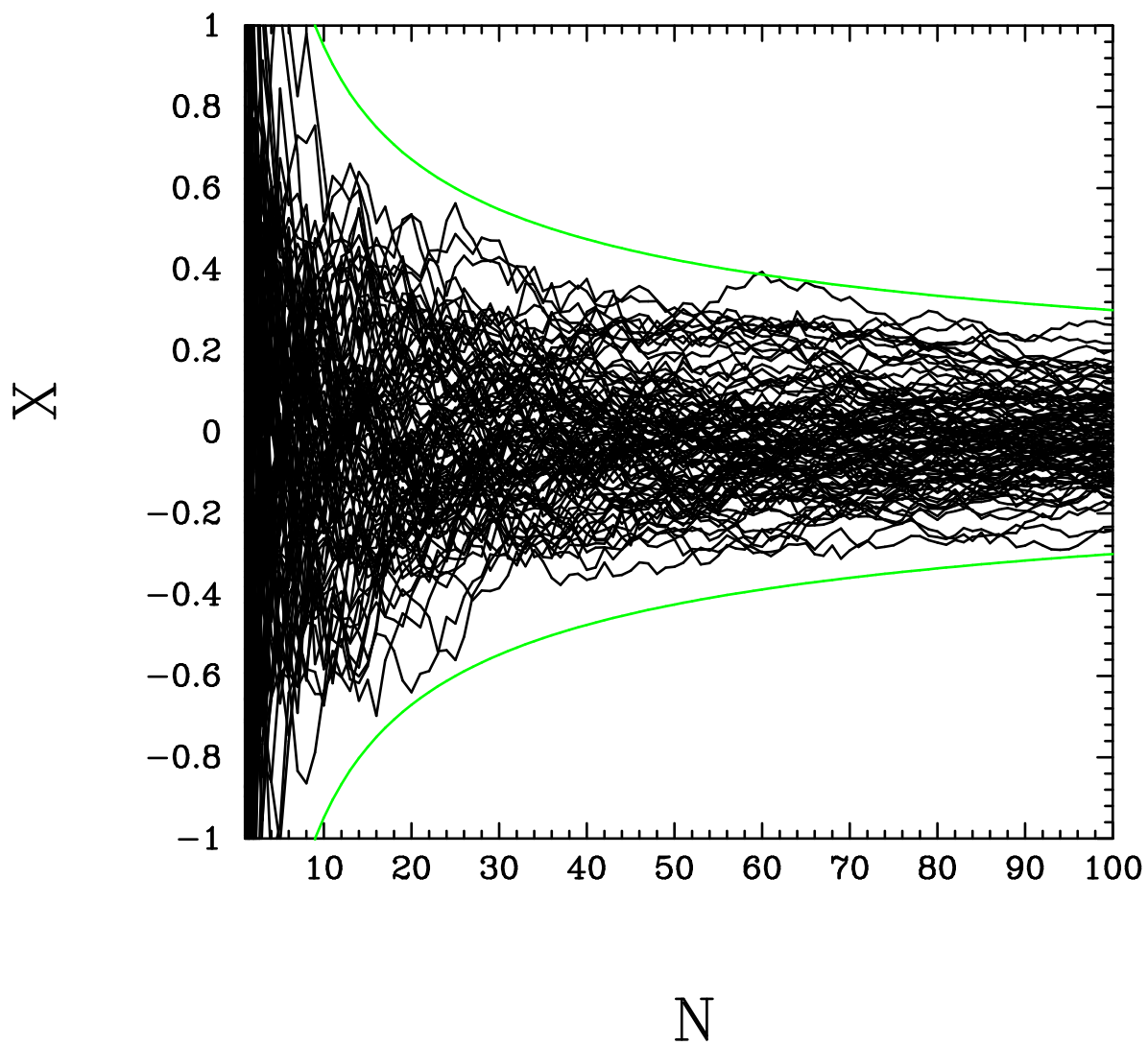




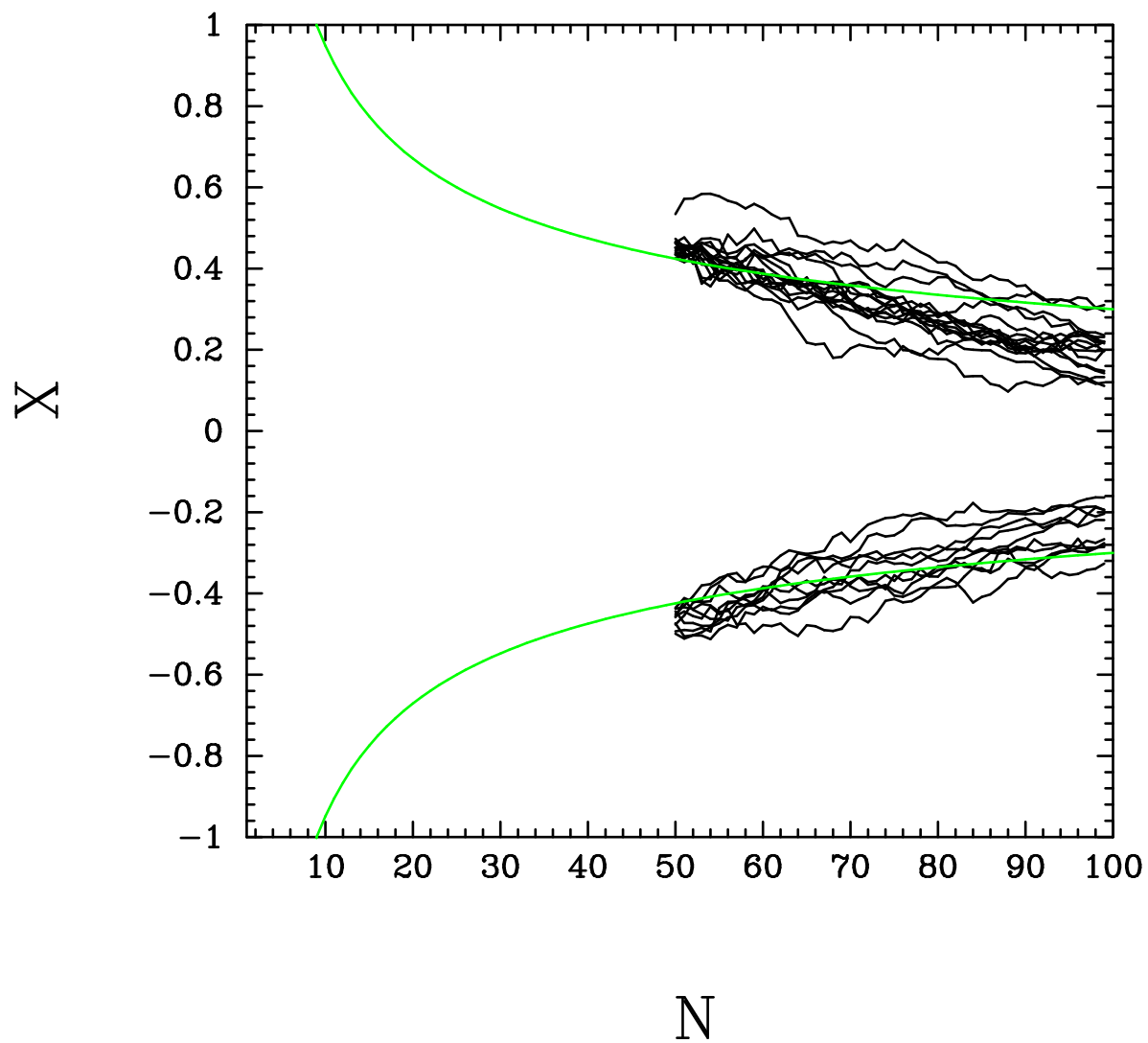
'Odd' aspects of Planck spectra?

- ❑ Planck temperature spectra want more lensing ($A_L > 1$) .
- ❑ Planck data favour closed universes.
- ❑ High multipoles ($\ell > 800$) give different parameters to lower multipoles (e.g. Addison et al 2016, ApJ, 818, 132).
- ❑ Outliers in TT spectrum and in TE spectrum (e.g. $\ell \sim 165$ in TE , Obied et al 2017, PRD, 083526).

$$X = \frac{1}{N} \sum_{i=1}^N x_i$$

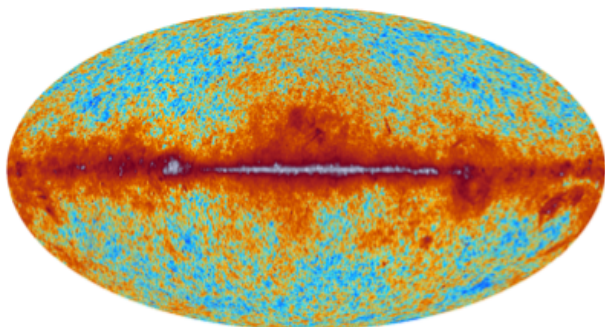


$$X = \frac{1}{N} \sum_{i=1}^N x_i$$

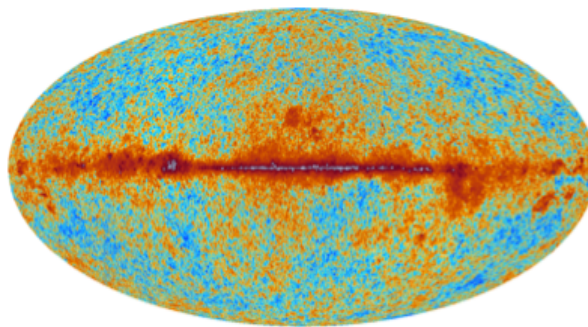


PLANCK FREQUENCYMAPS

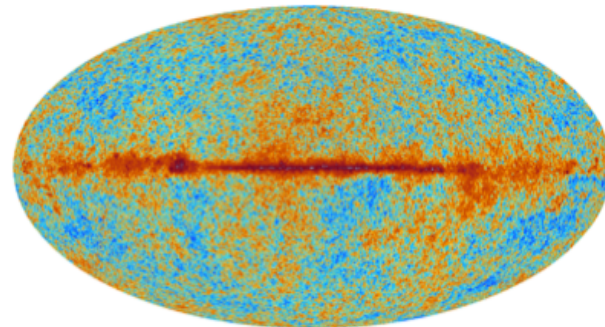
30 GHz



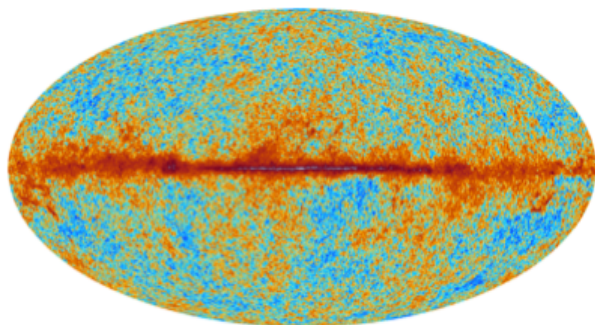
44 GHz



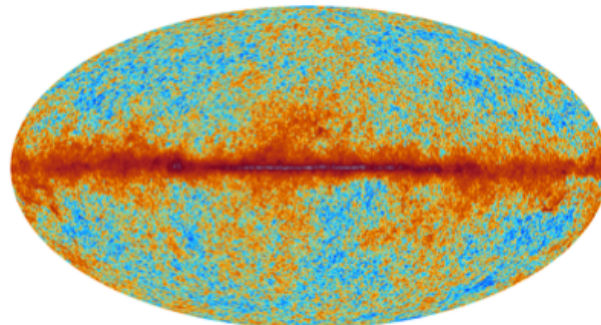
70 GHz



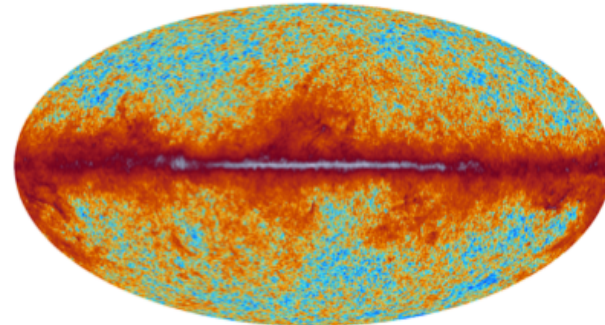
100 GHz



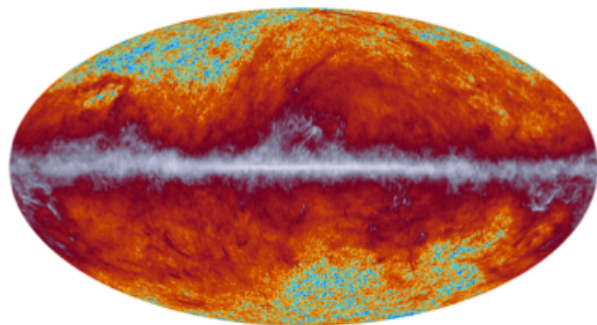
143 GHz



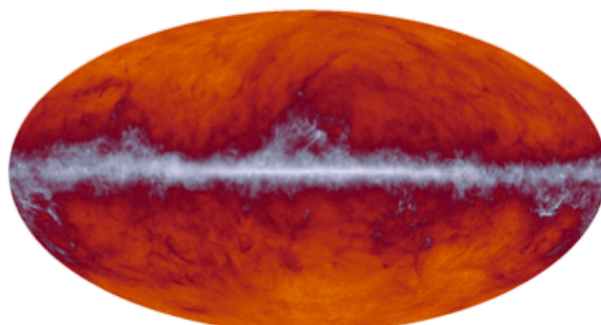
217 GHz



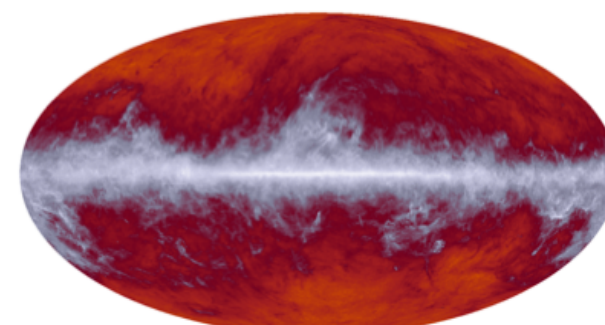
353 GHz



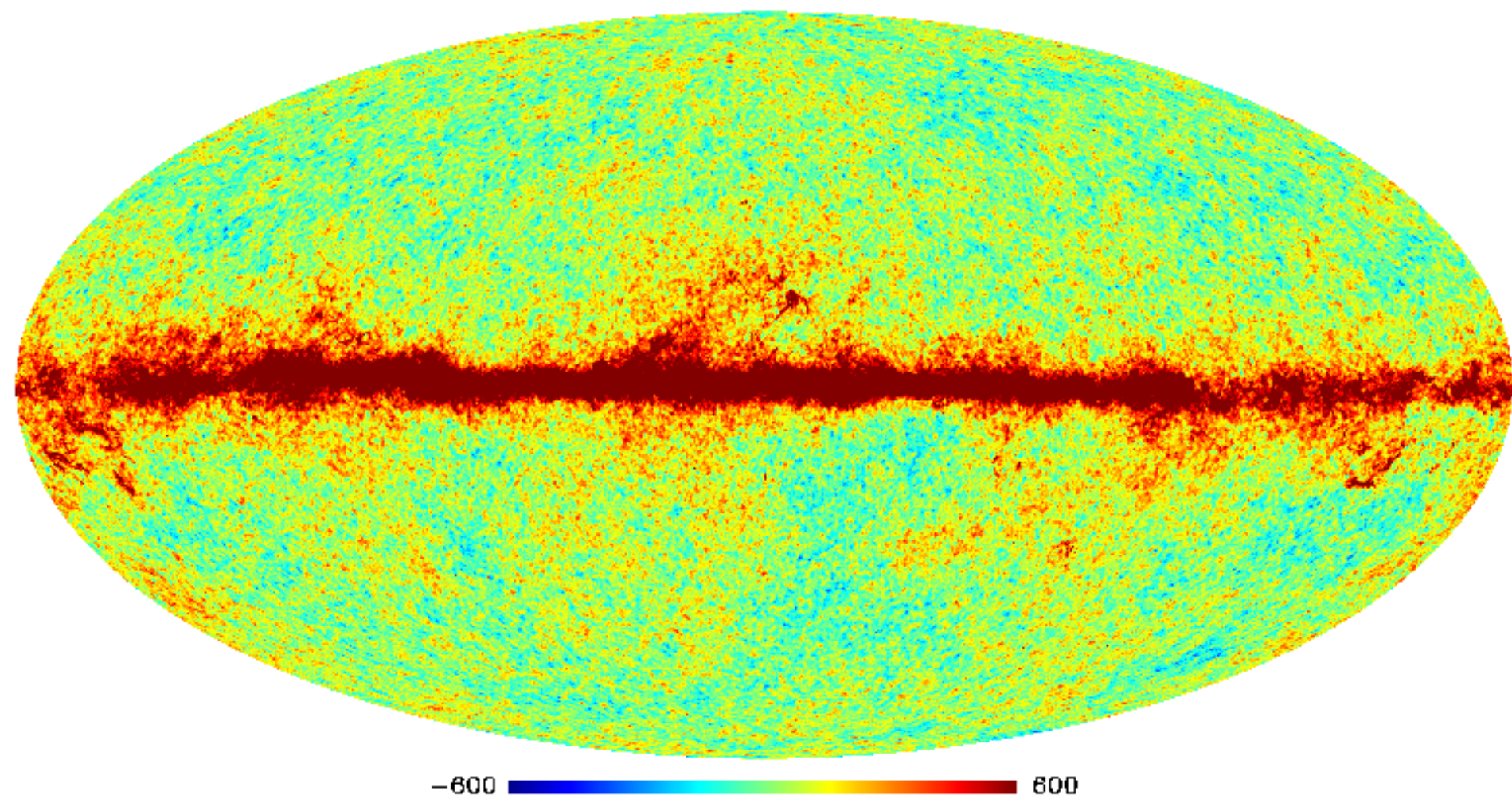
545 GHz



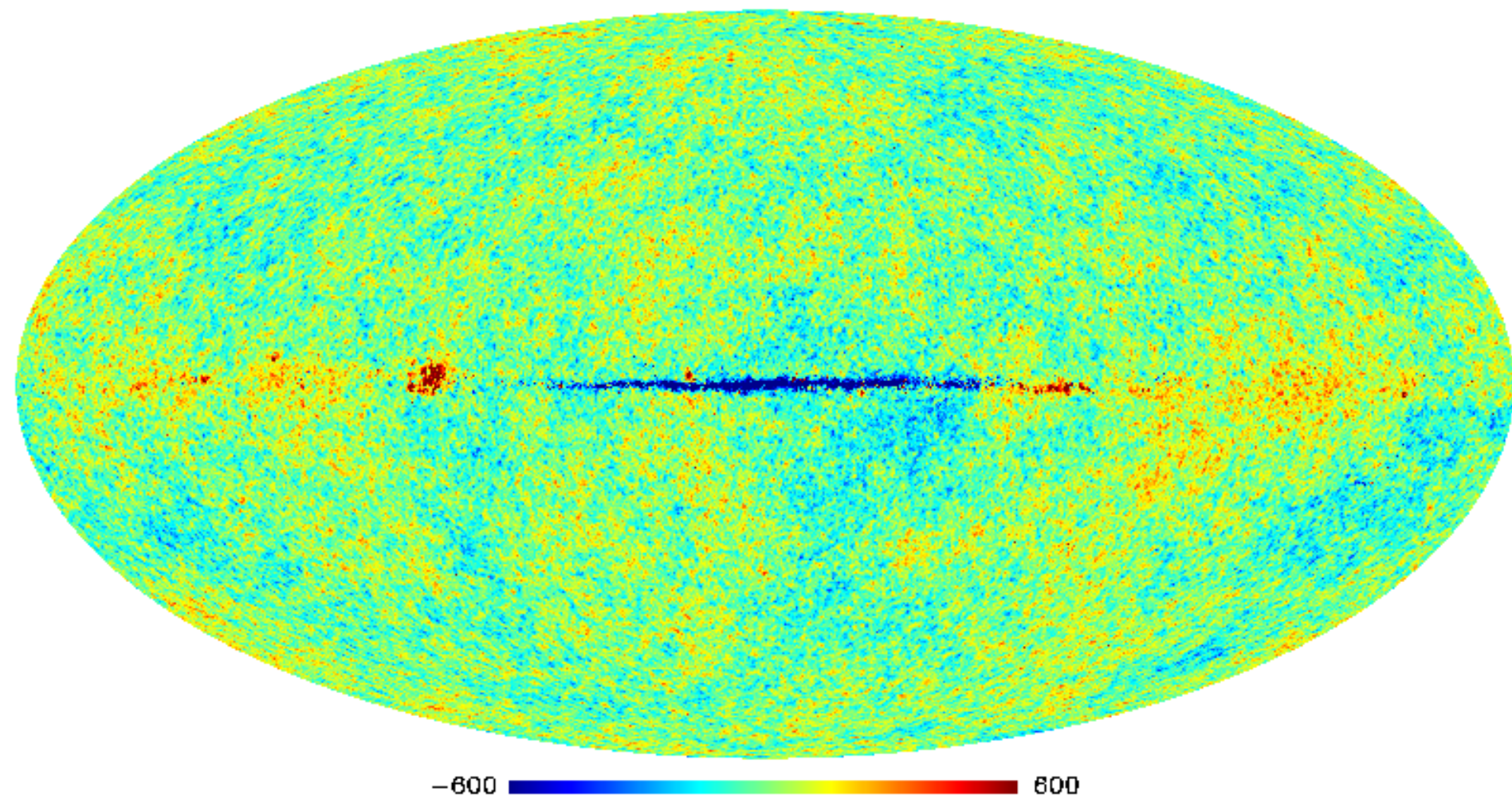
857 GHz



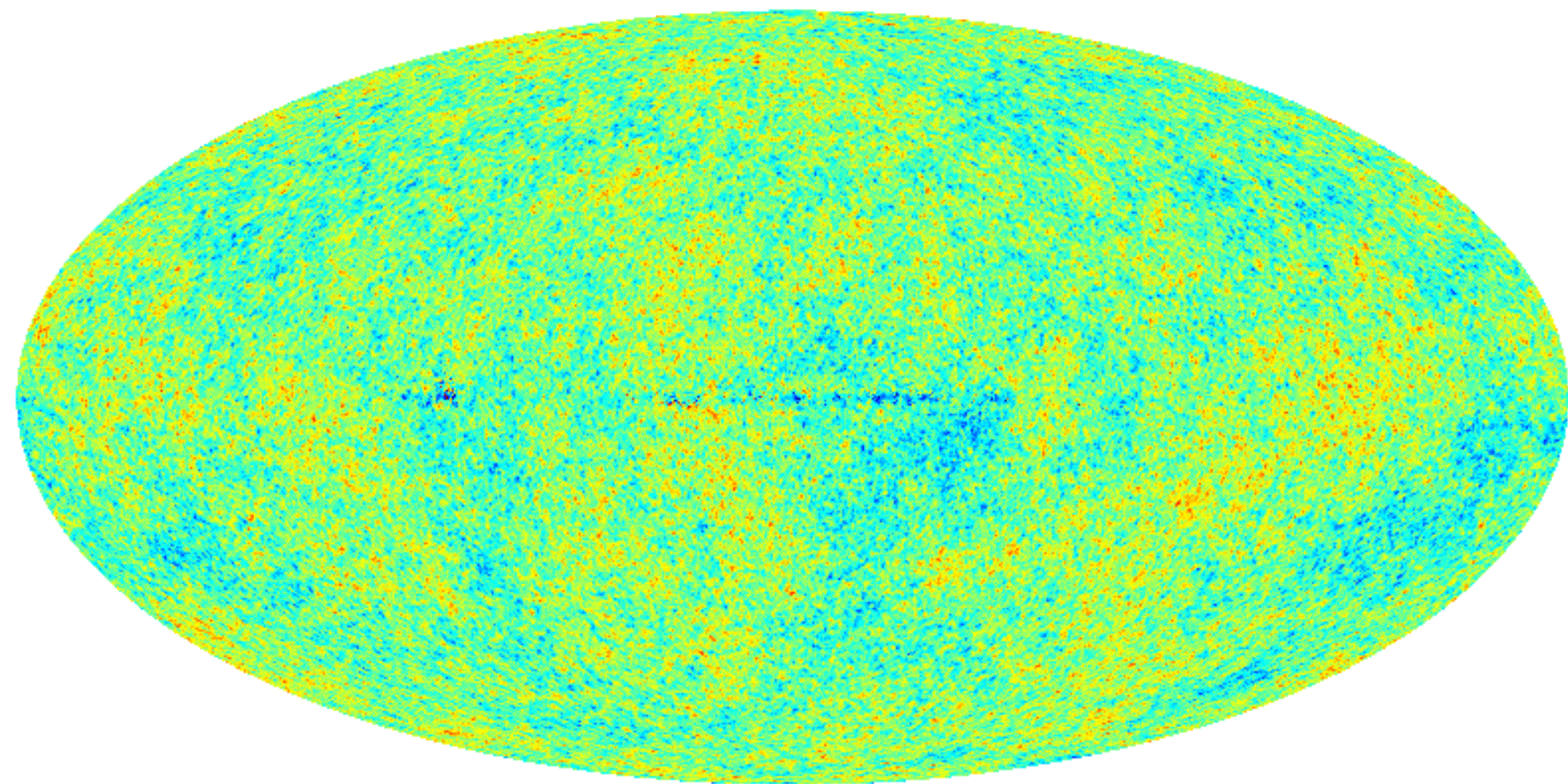
143



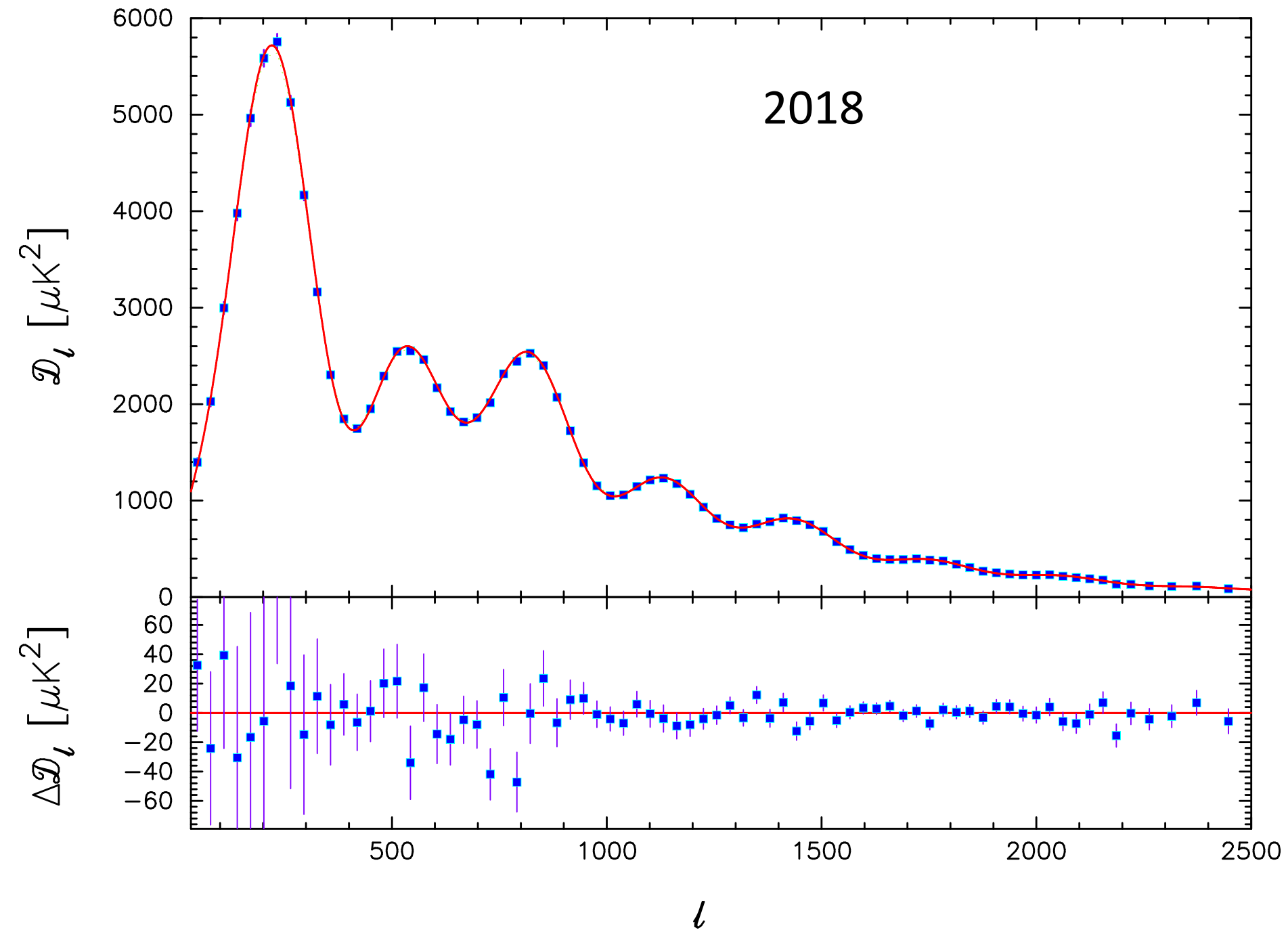
143 cleaned with 545

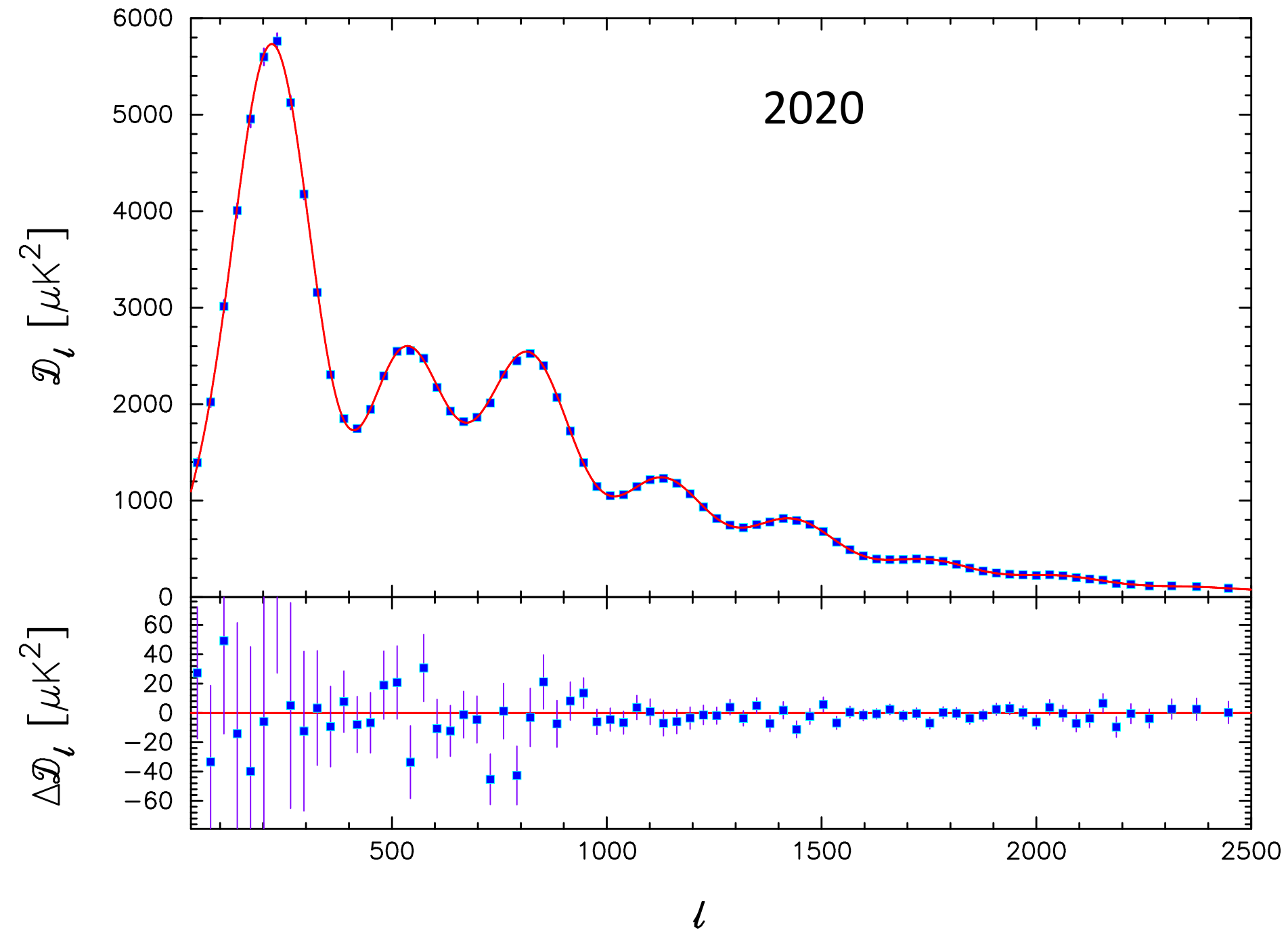


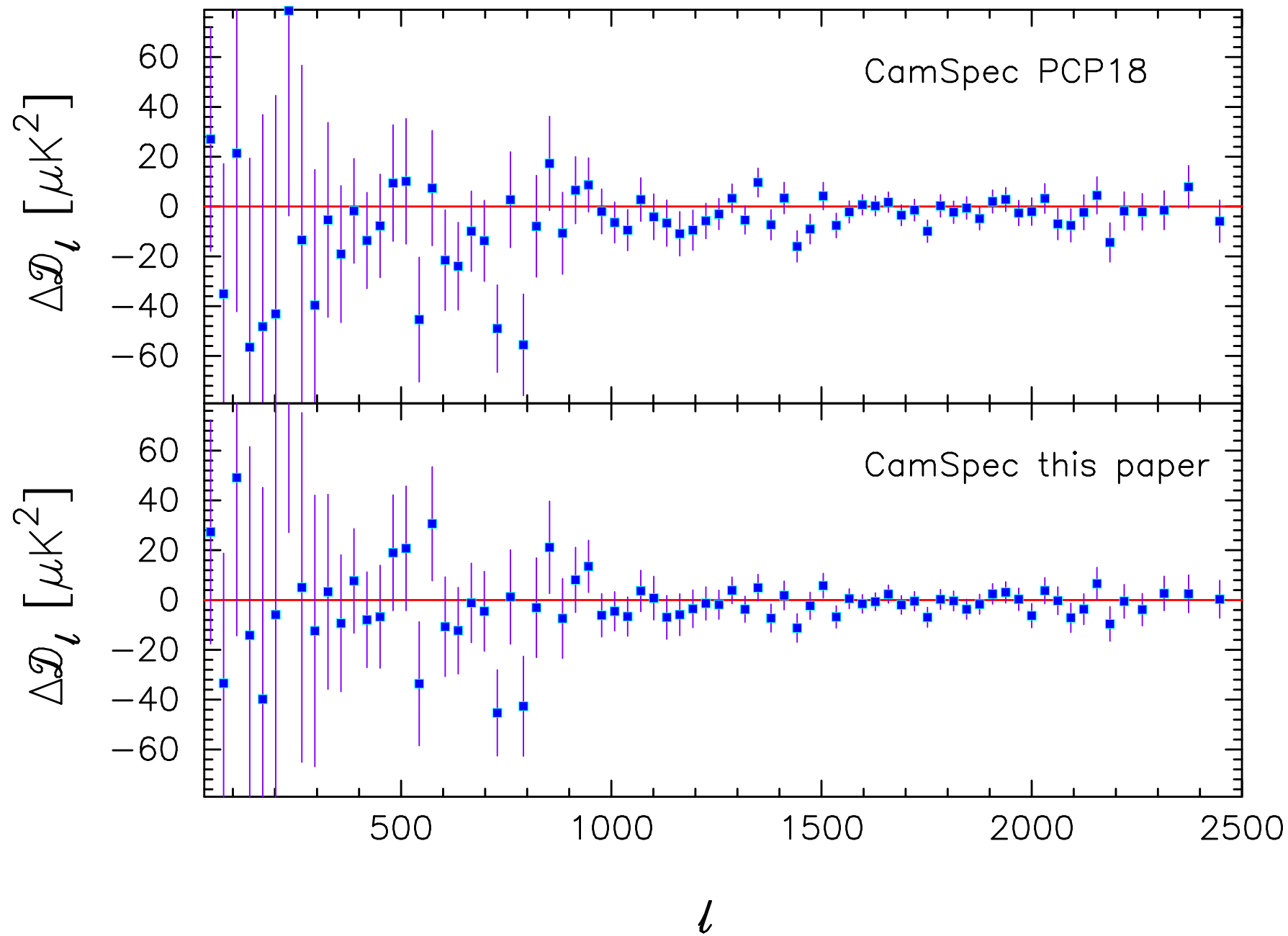
SMICA

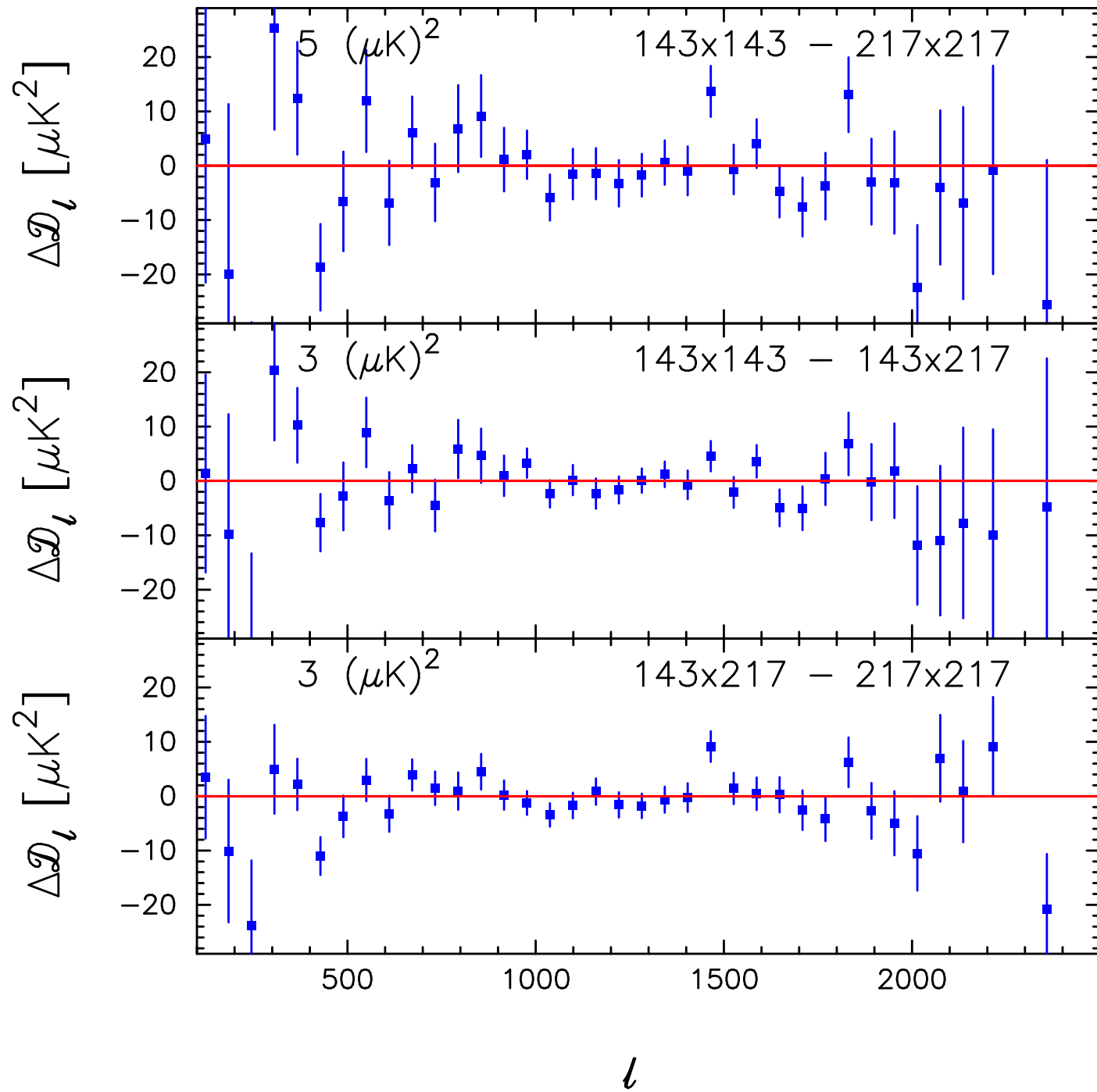


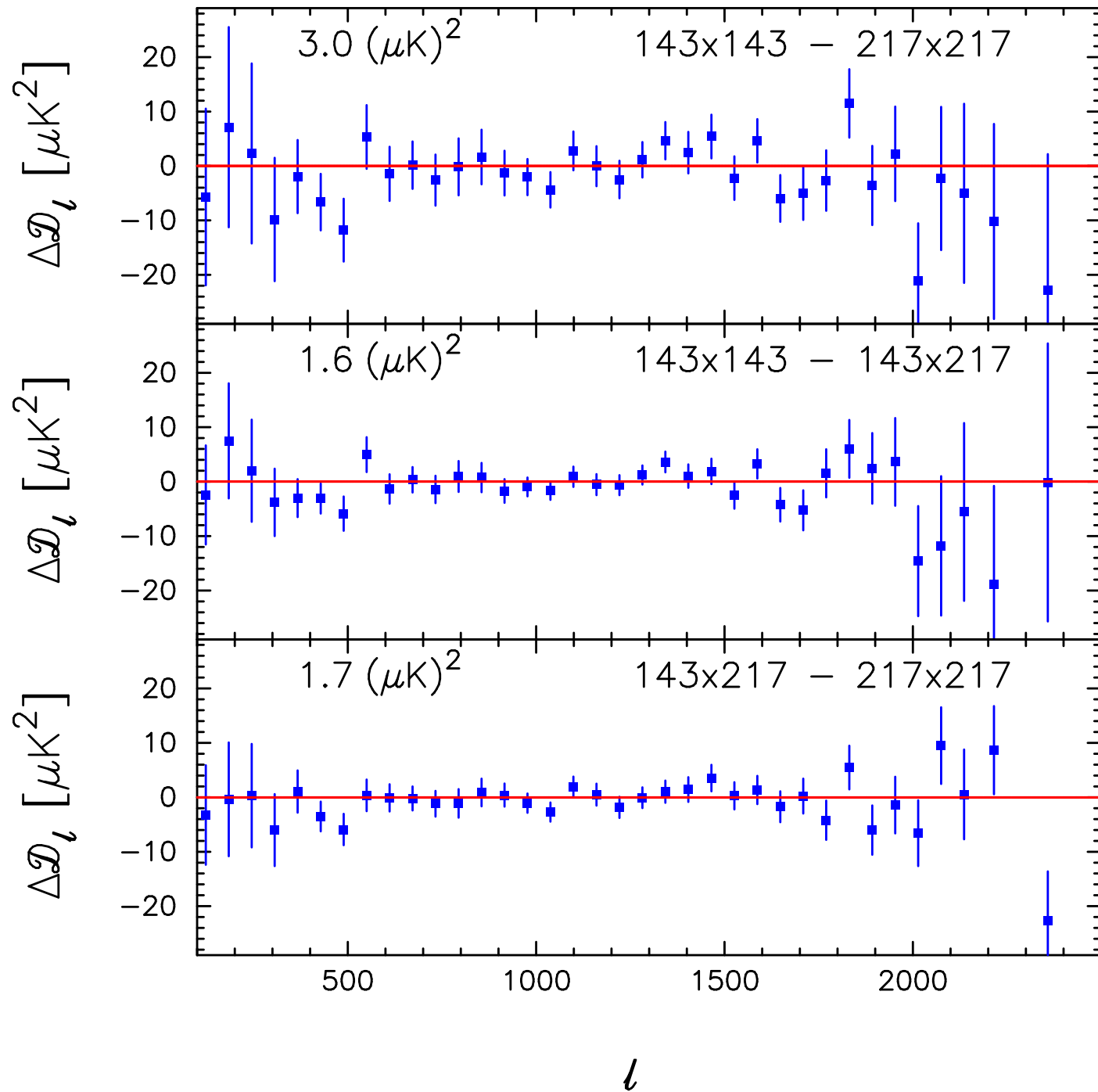
-600 600

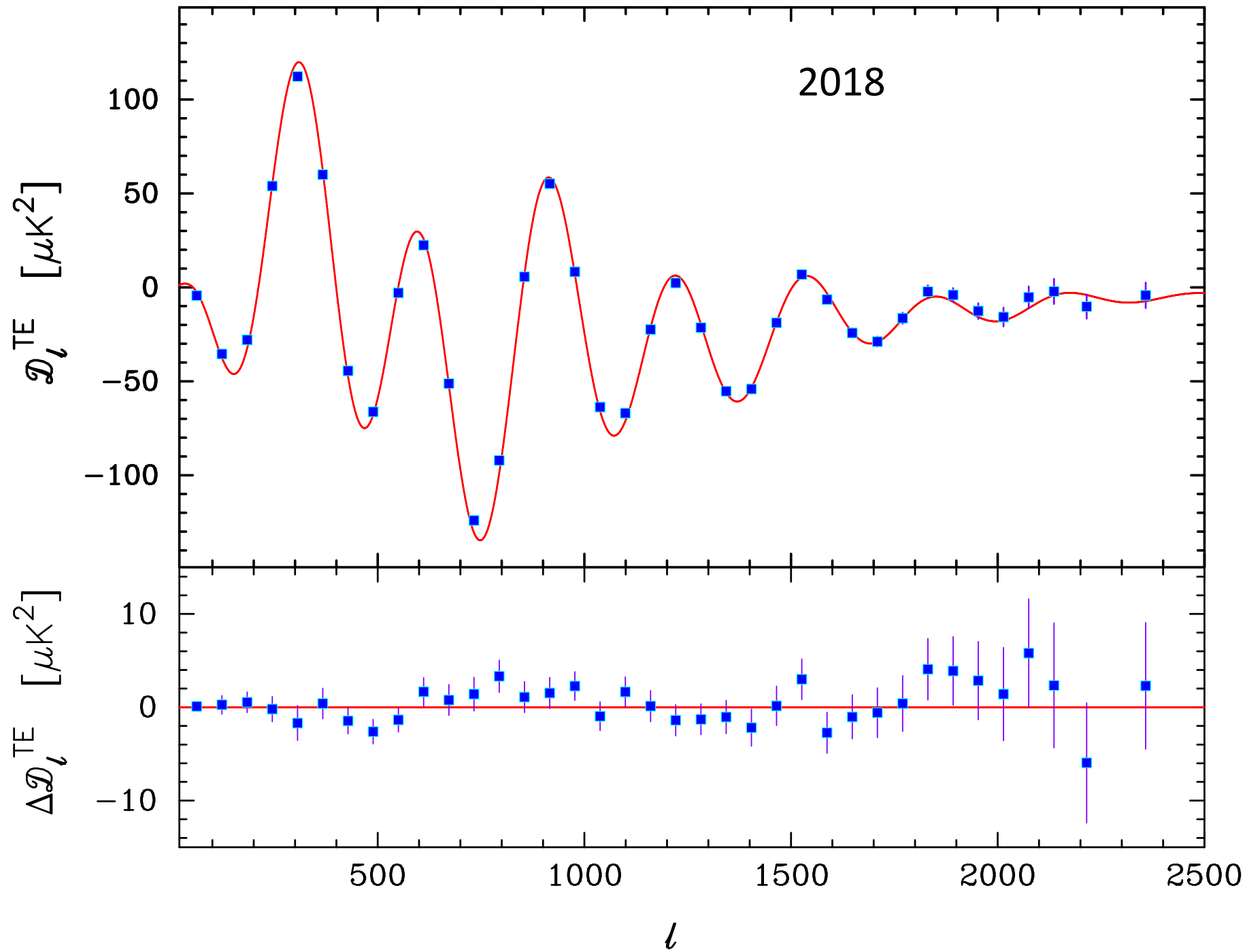


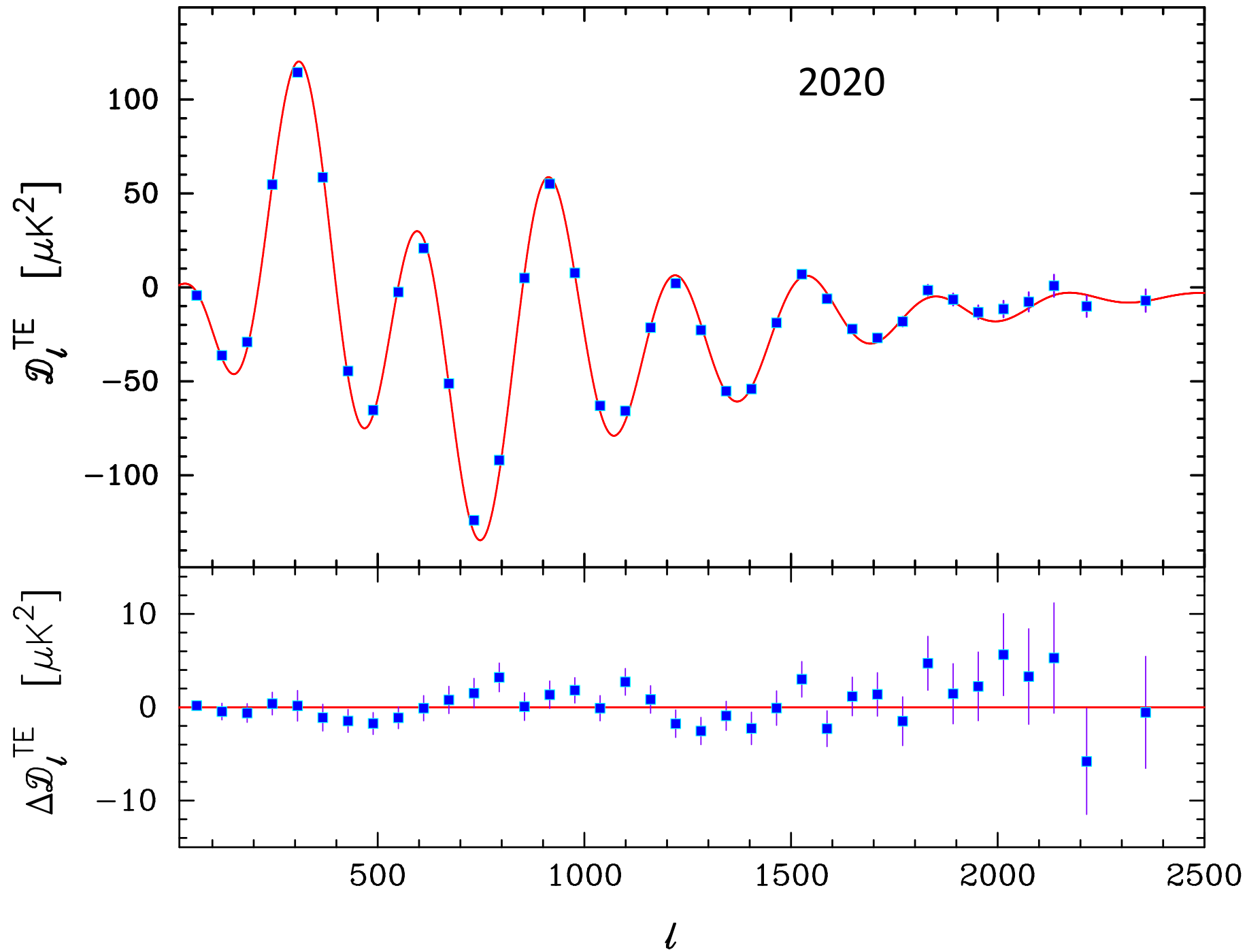


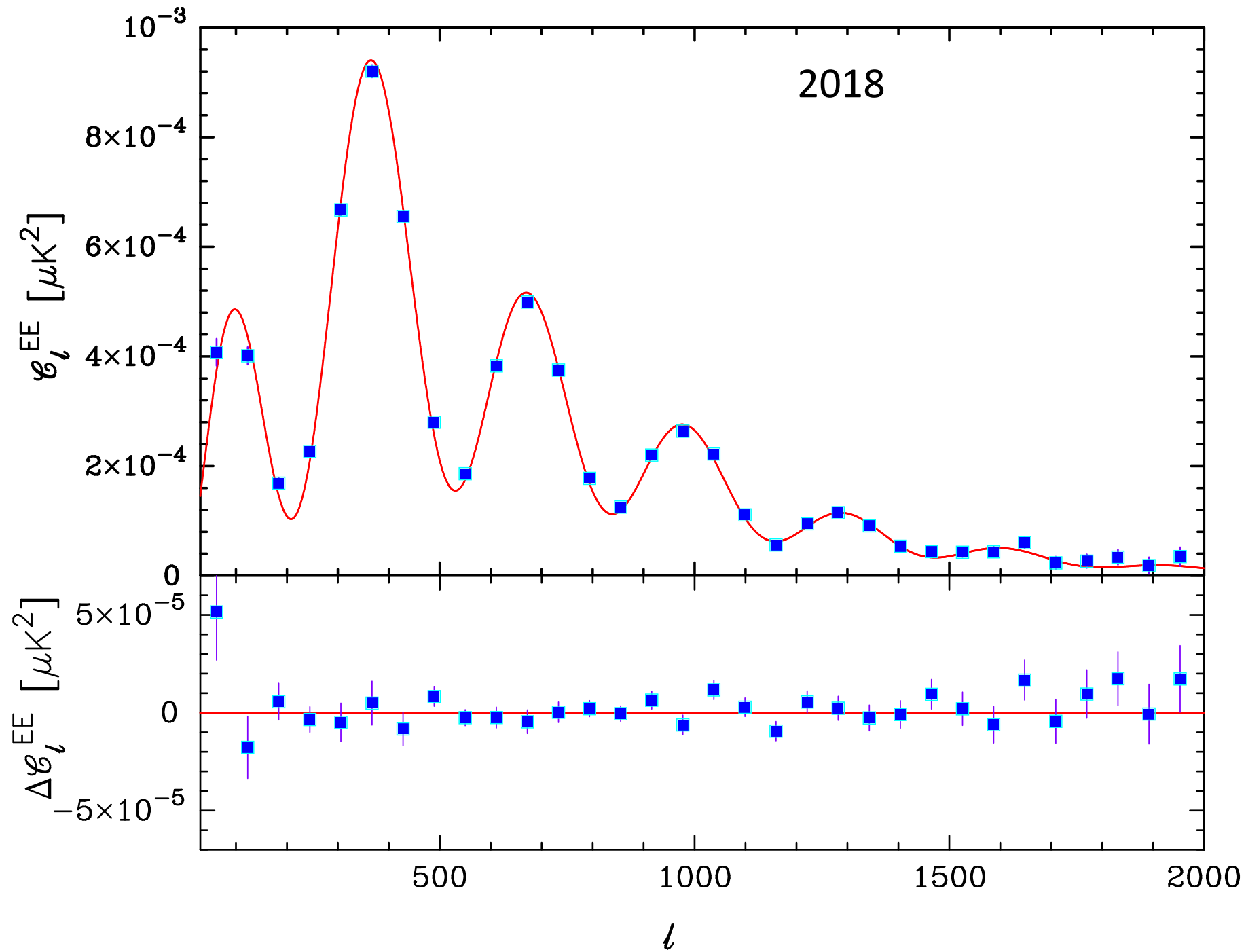


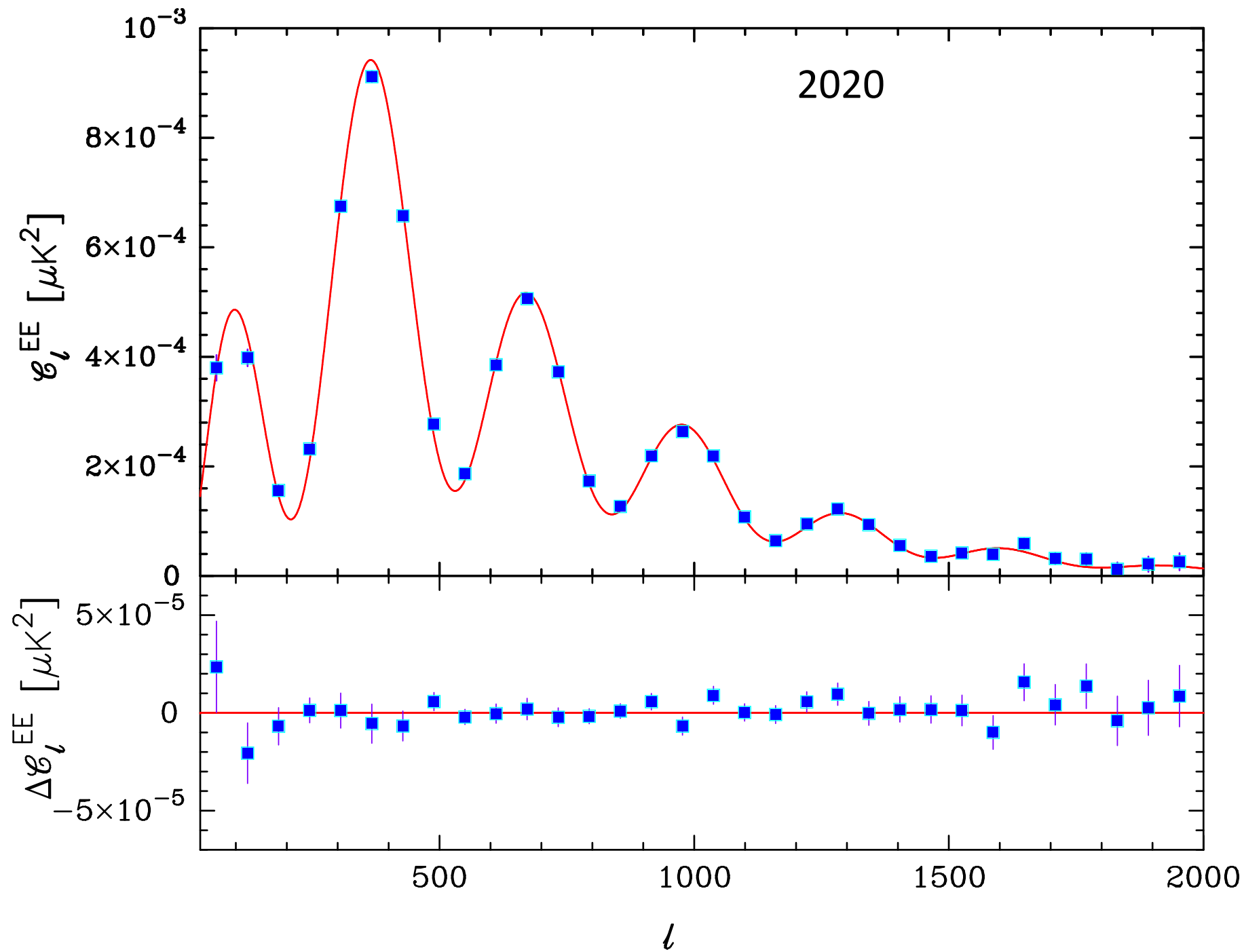




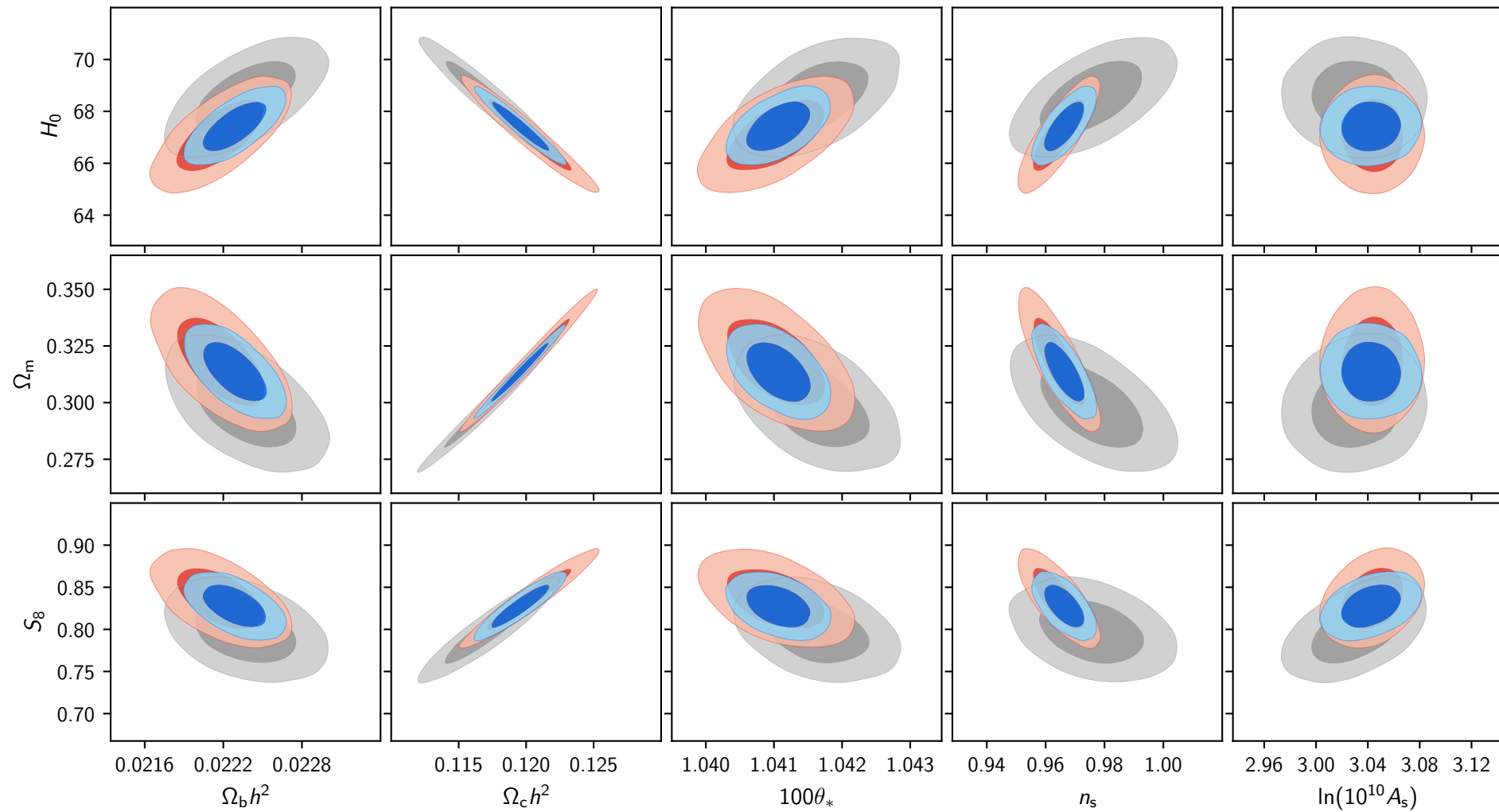




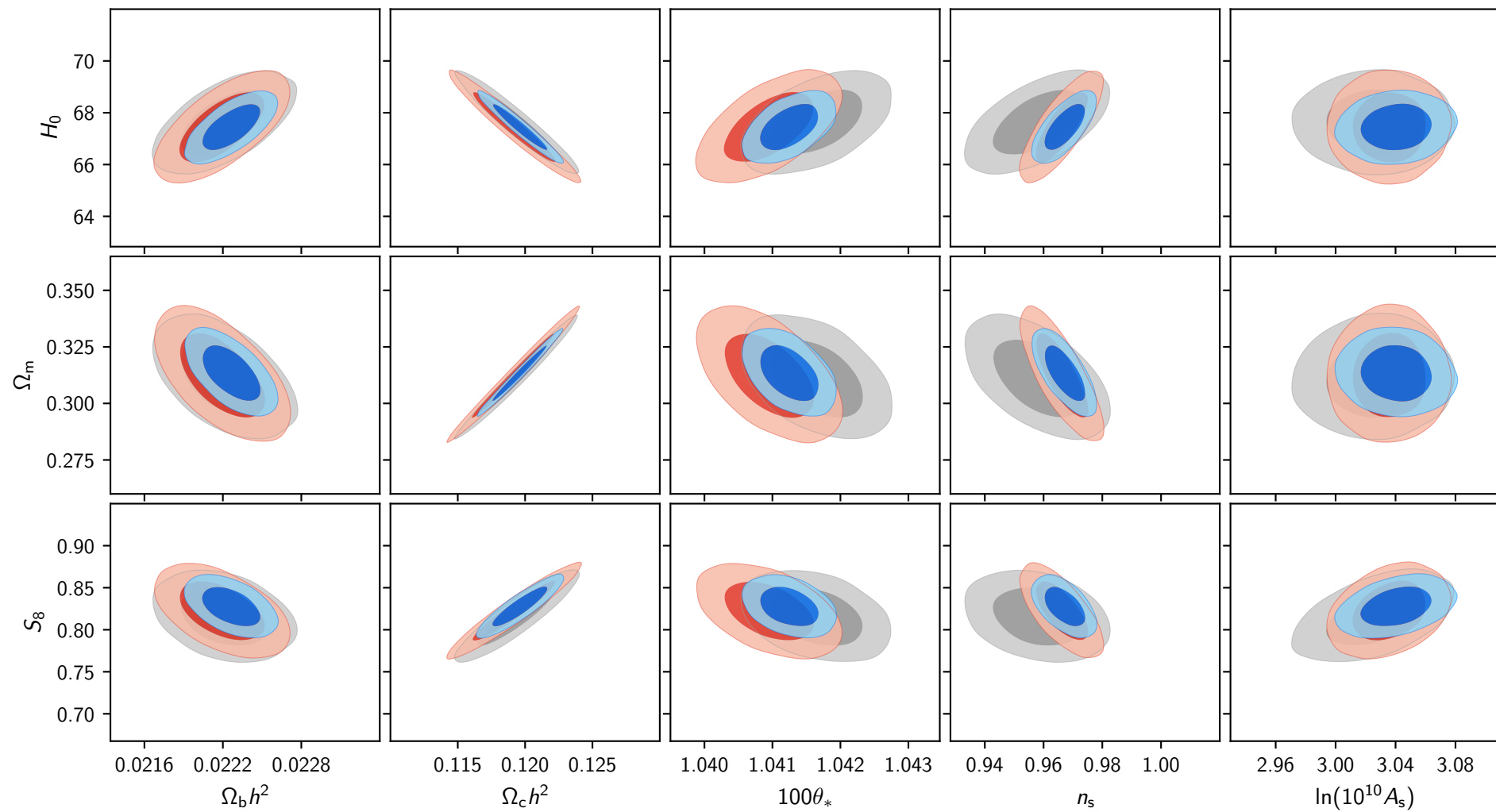




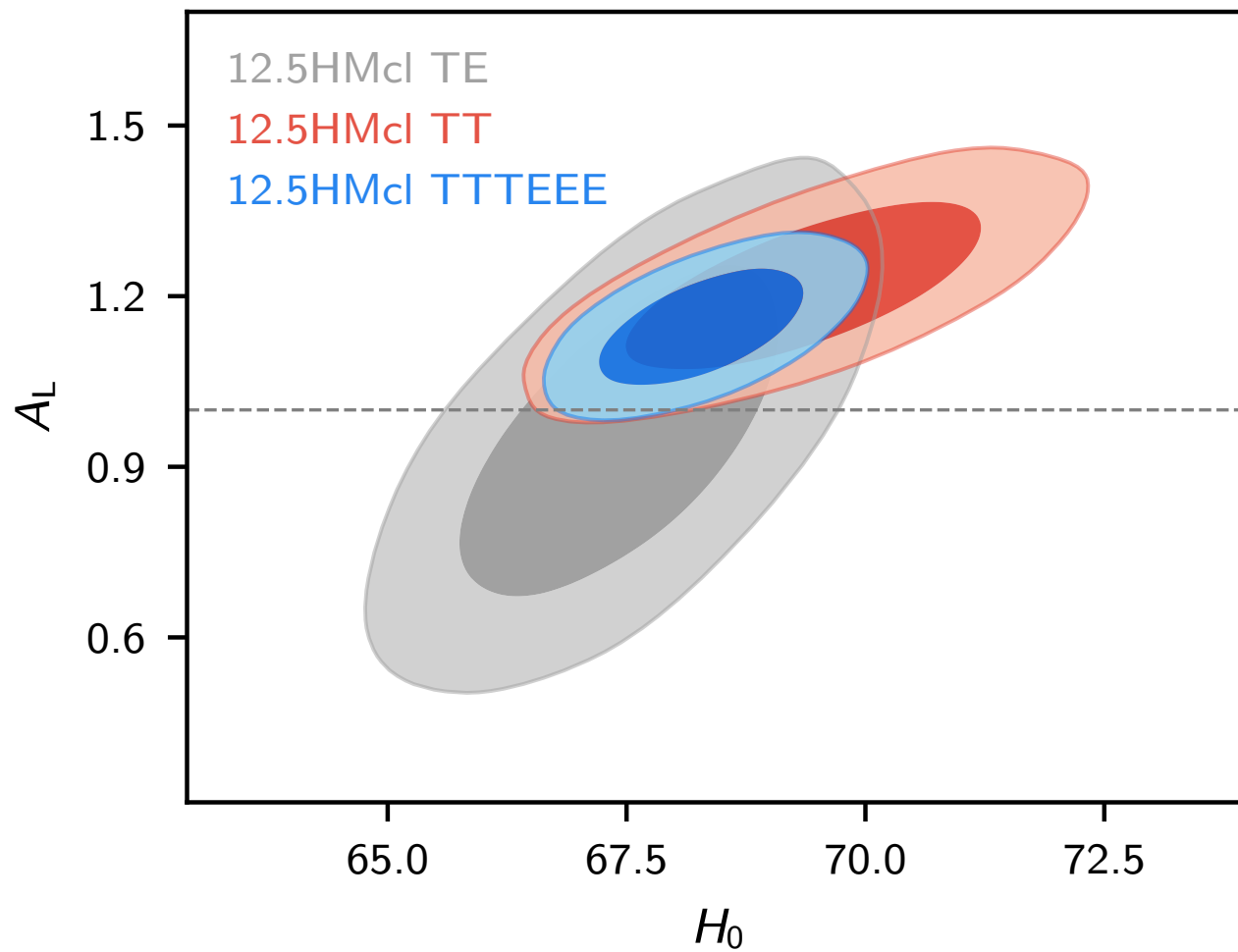
12.1HM TE 12.1HM TT 12.1HM TTTEEE



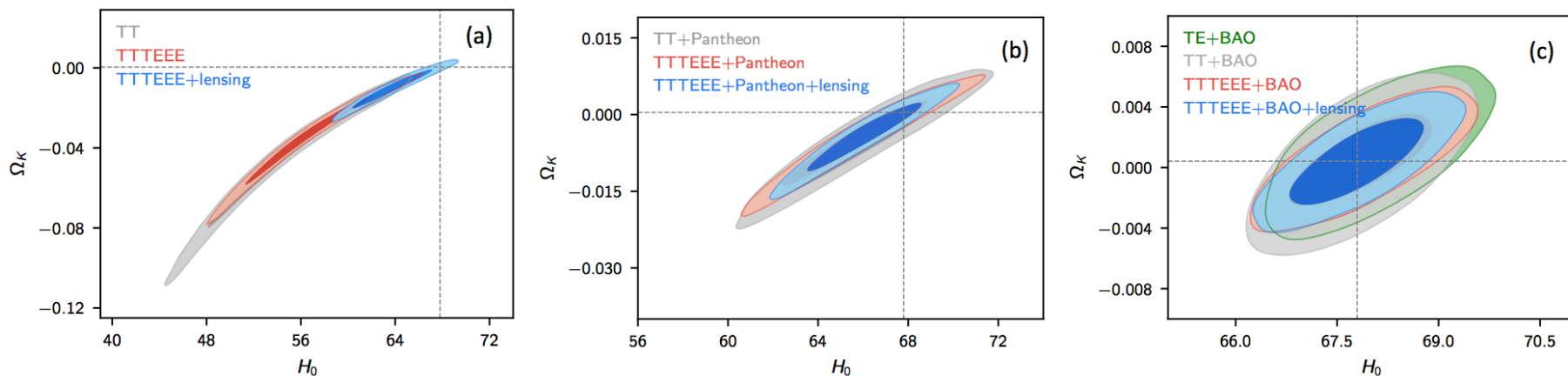
12.5HMcln TE 12.5HMlcn TT 12.5HMcln TTTEEE



Lensing parameter A_L

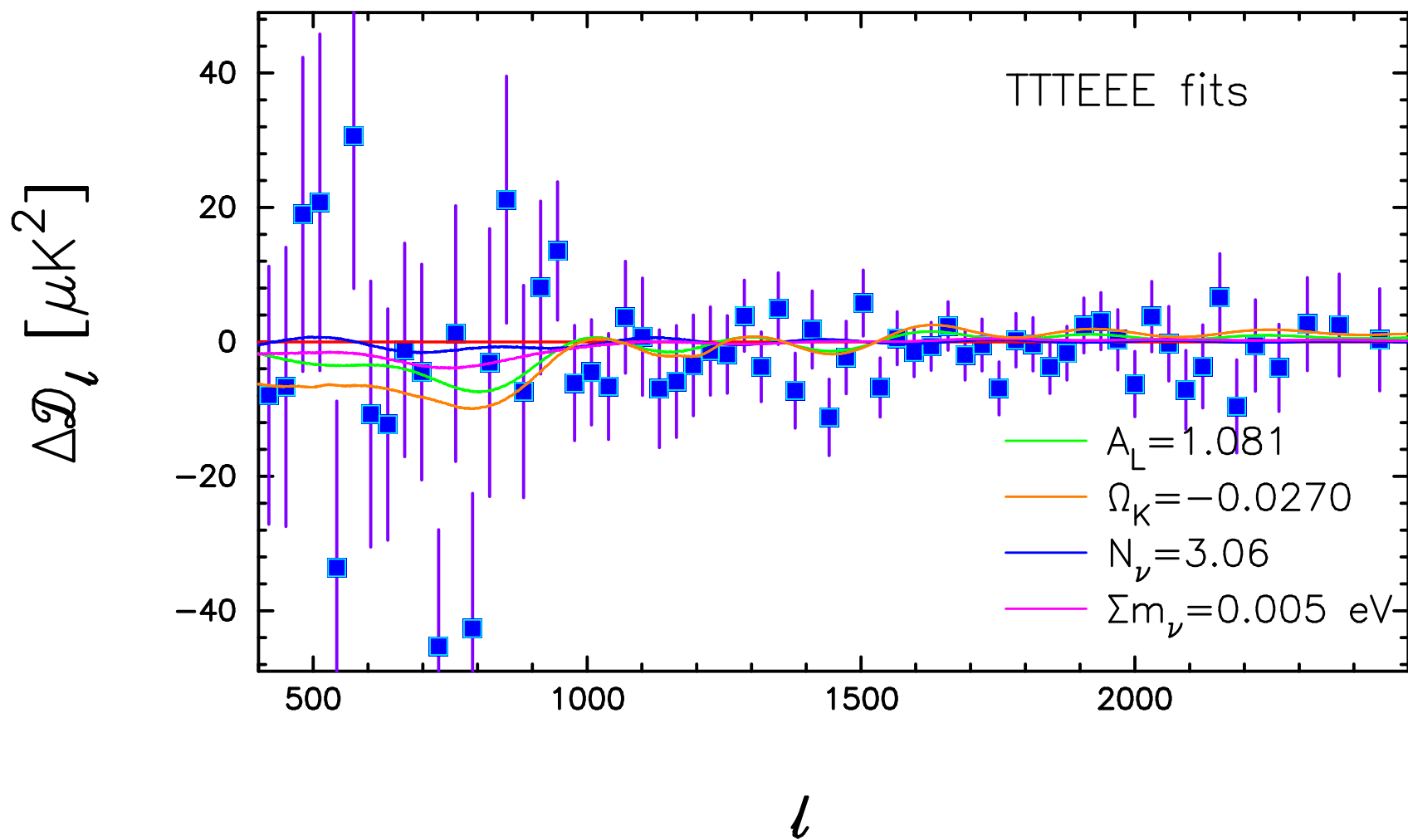


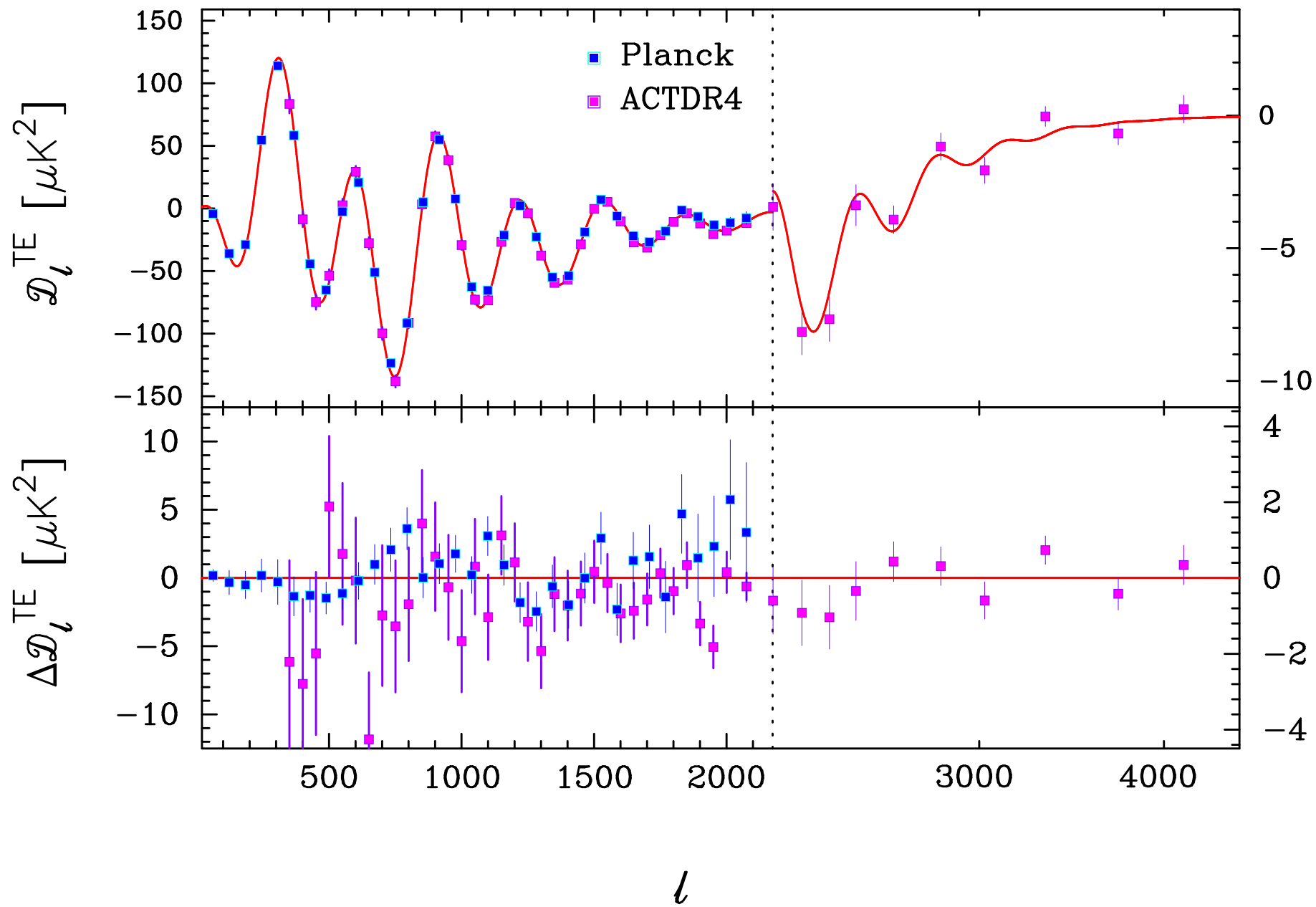
Curvature parameter

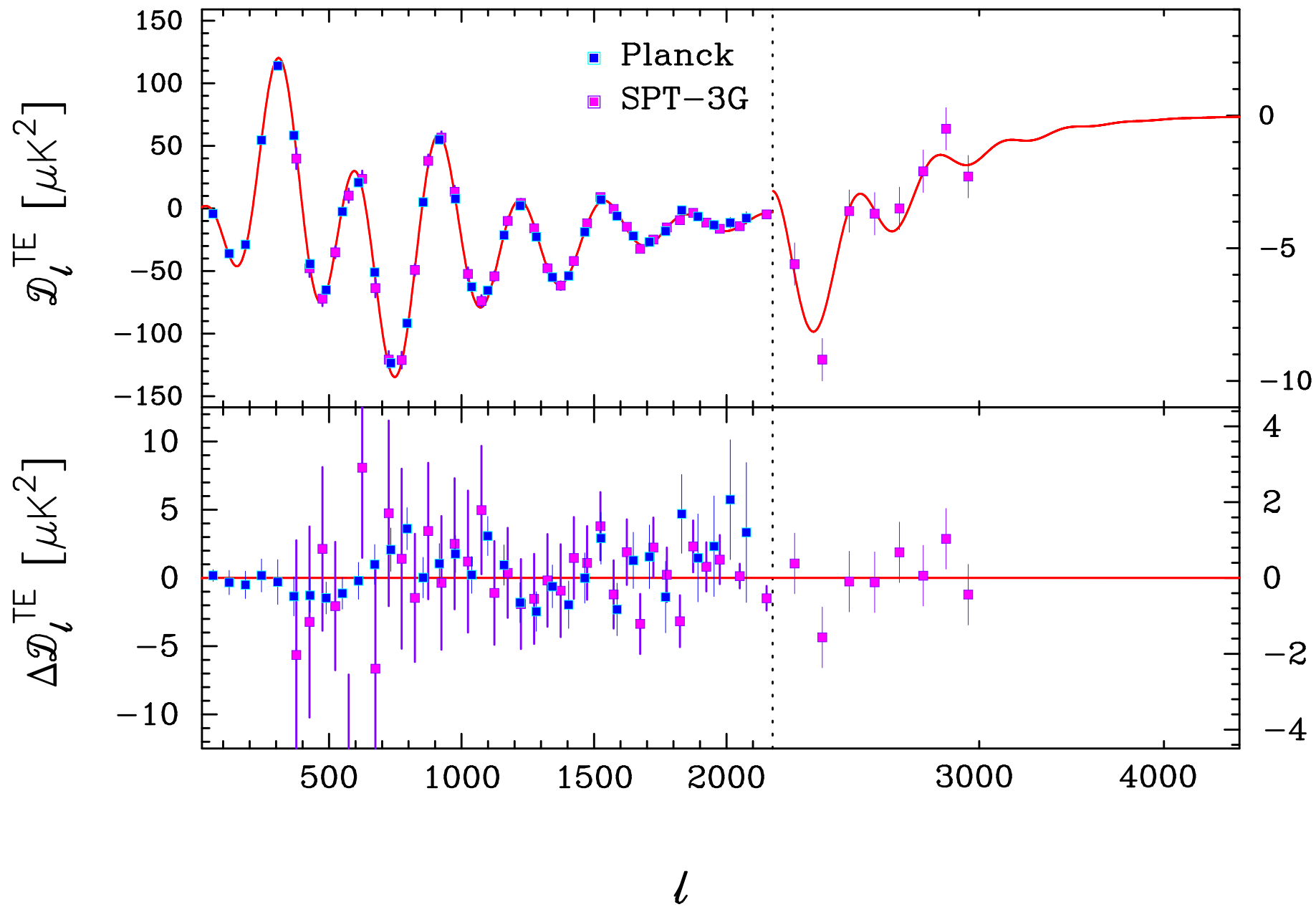


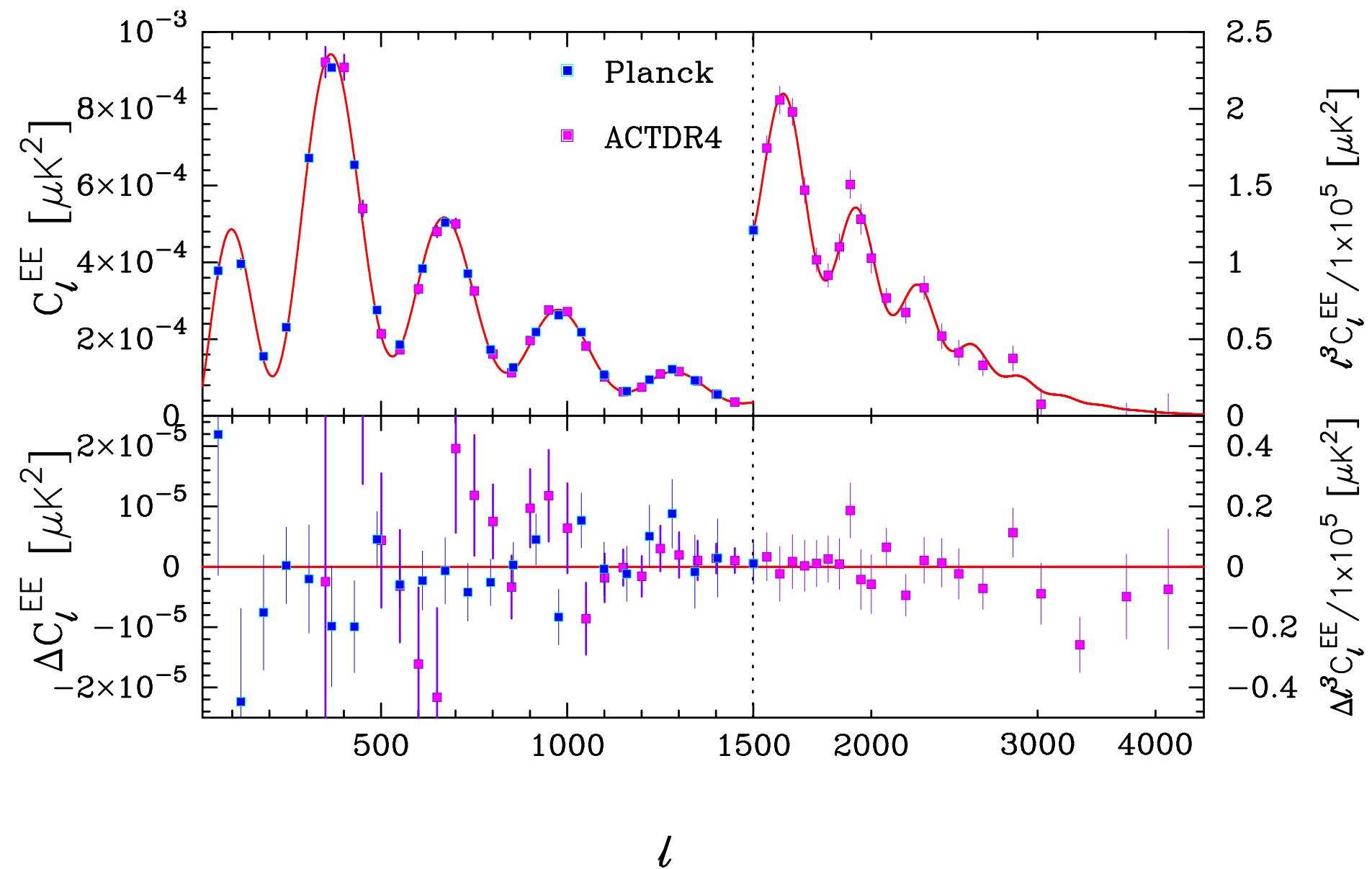
Planck TTTEEE+BAO+pantheon+lensing $\Omega_k = 0.0004 \pm 0.0018$

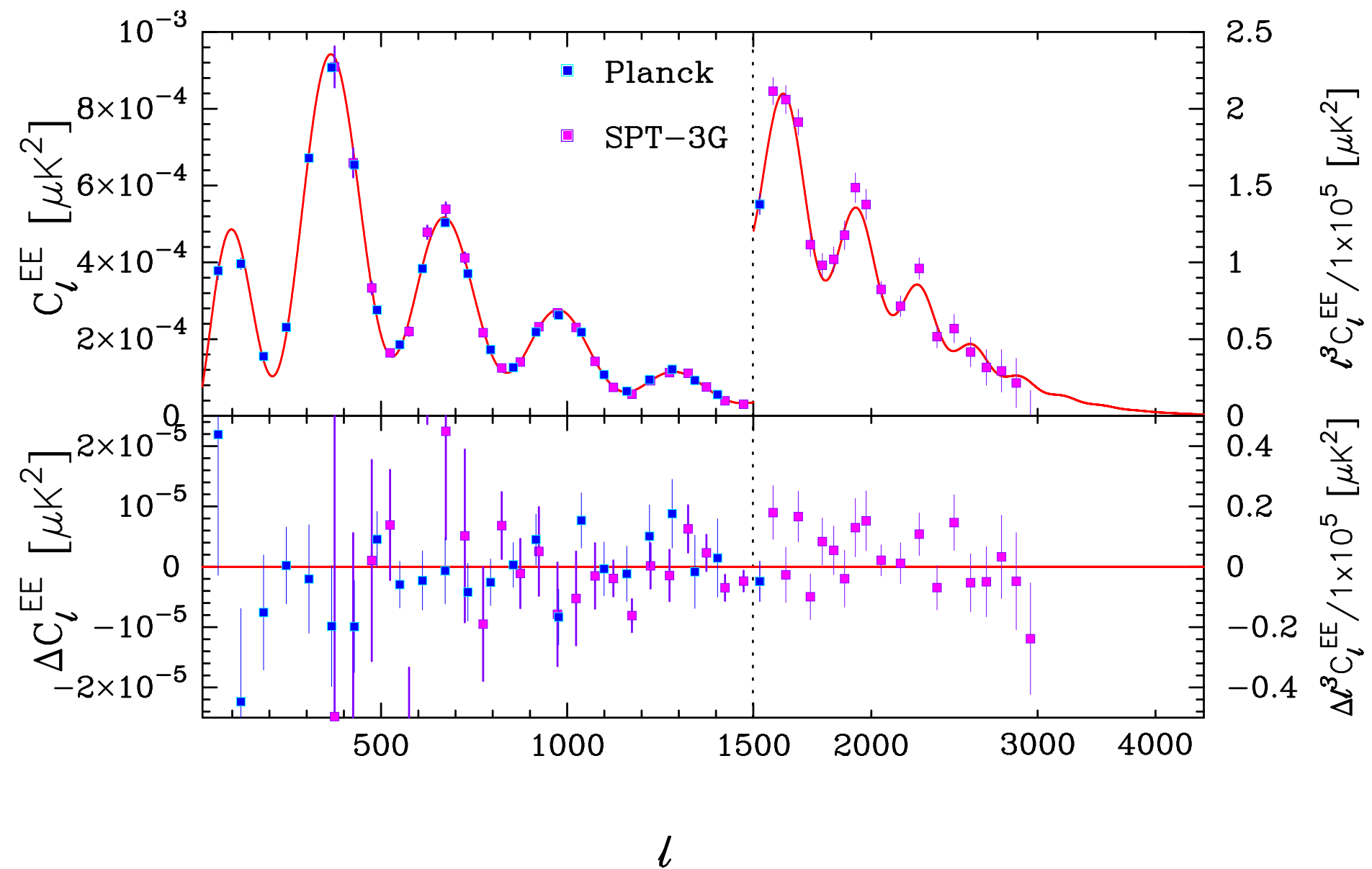
Addition of Ω_k as a parameter improves TTTEEE χ^2 by only -3.02



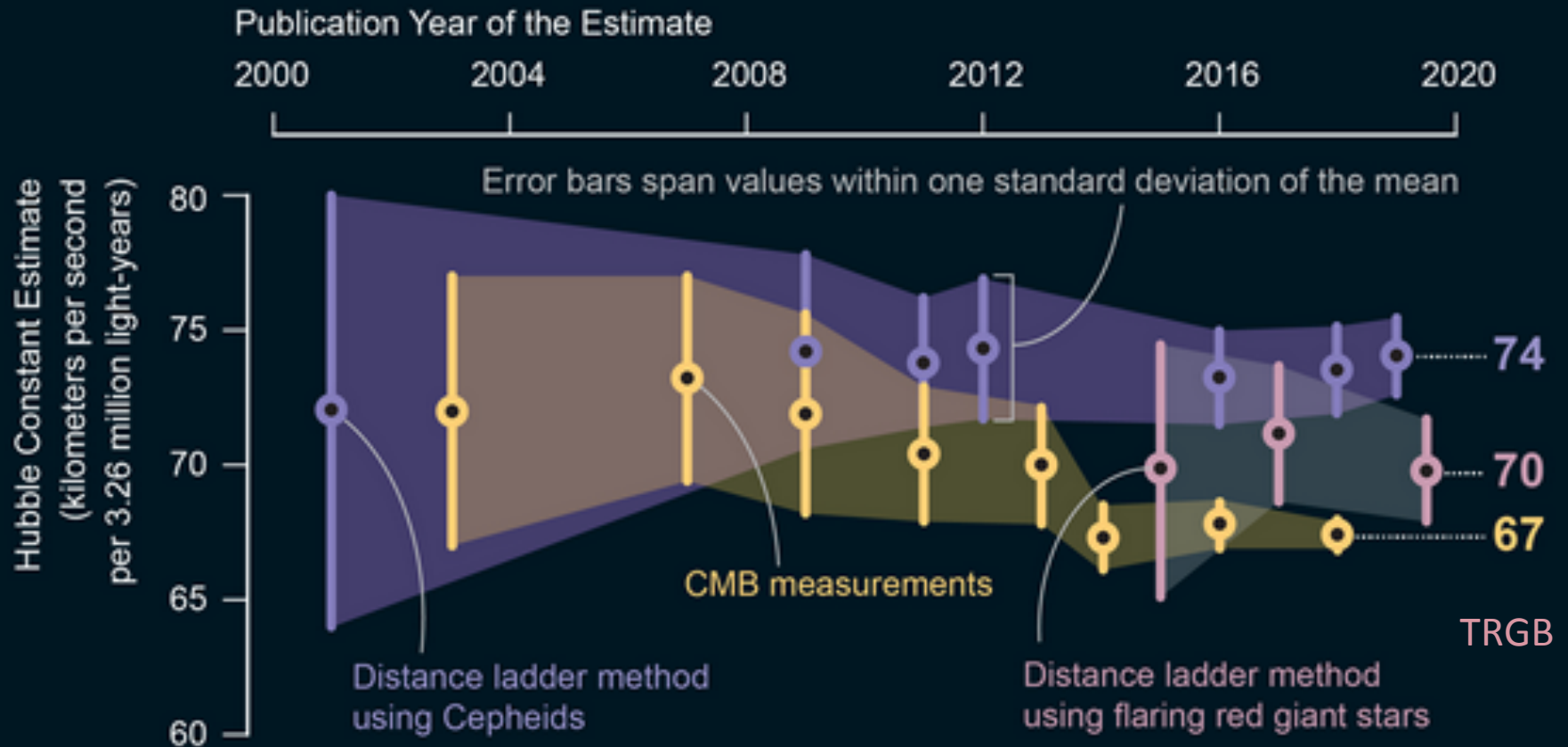








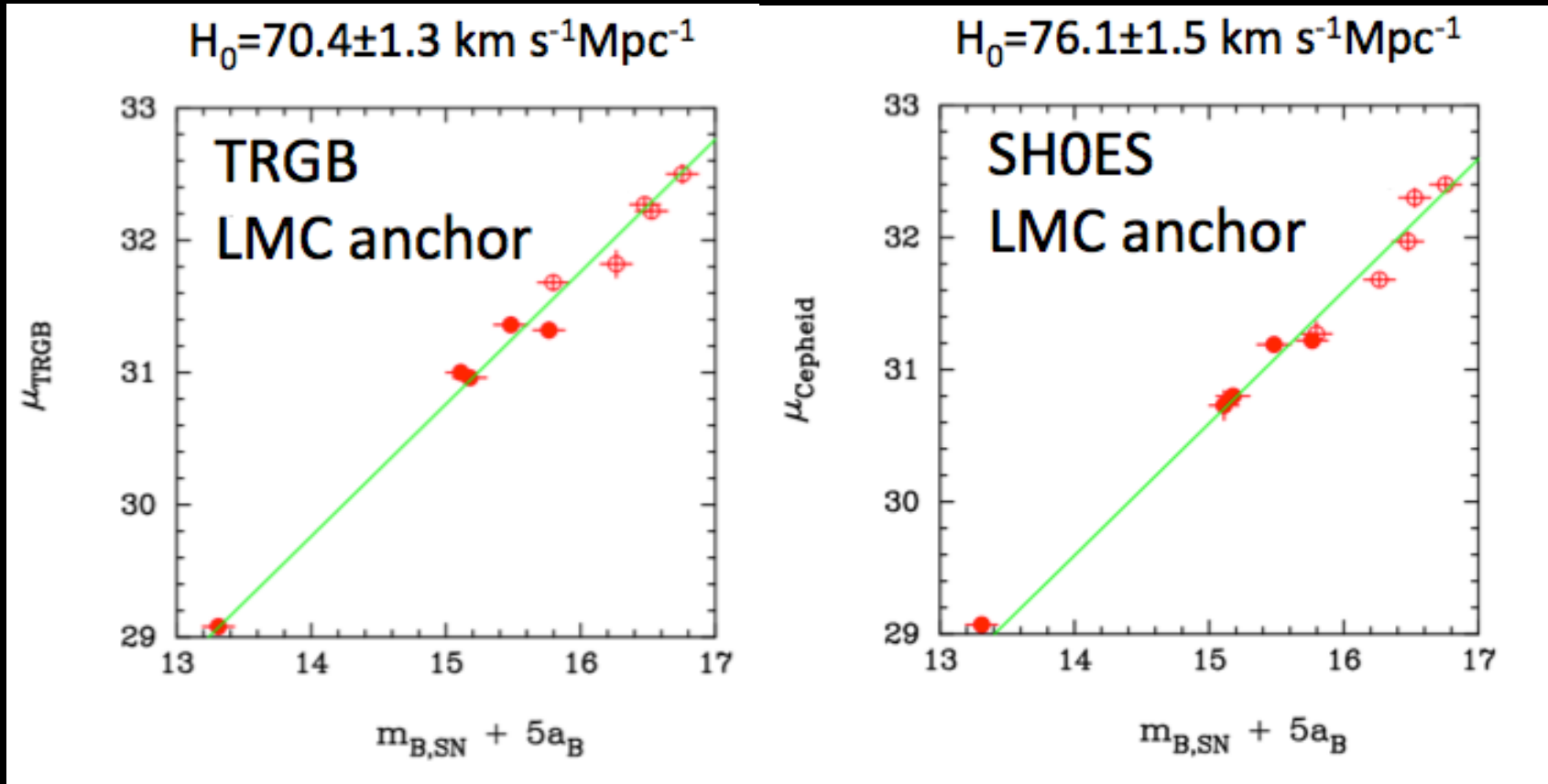
Three things that you may not know about the Hubble tension



Discrepancy of $6.4 \text{ km s}^{-1} \text{ Mpc}^{-1}$ -- a magnitude error of 0.1 mag gives a shift in H_0 of $3.2 \text{ km s}^{-1} \text{ Mpc}^{-1}$

Point 1 TRGB and Cepheid distance scales are highly discrepant.

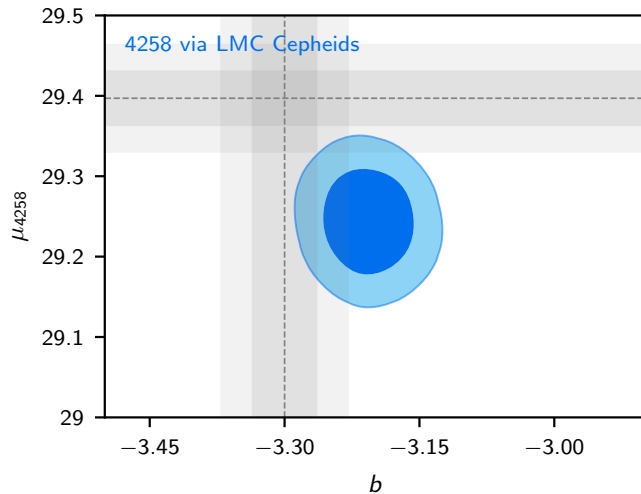
For 10 galaxies in common using the same anchor and Pantheon SN magnitudes:



$$\mu_{TRGB} - \mu_{Cepheid} = 0.14 \pm 0.02$$

Point 2 The SH0ES analysis is not internally consistent

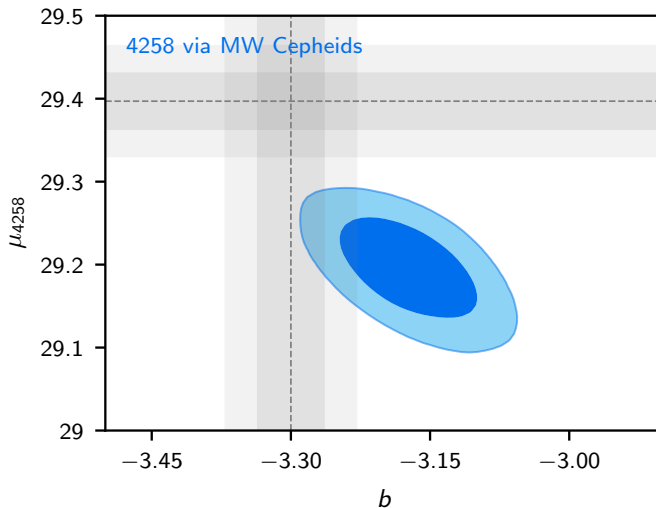
$$m = a + b \log_{10} P + Z \log_{10} (O/H)$$



$\mu_{N4258} = 29.397 \pm 0.033$ mag VLBI maser
(Reid et al 2019)

$\mu_{LMC} = 18.477 \pm 0.026$ mag Detached eclipsing
binaries (Pietrzynski et al 2019)

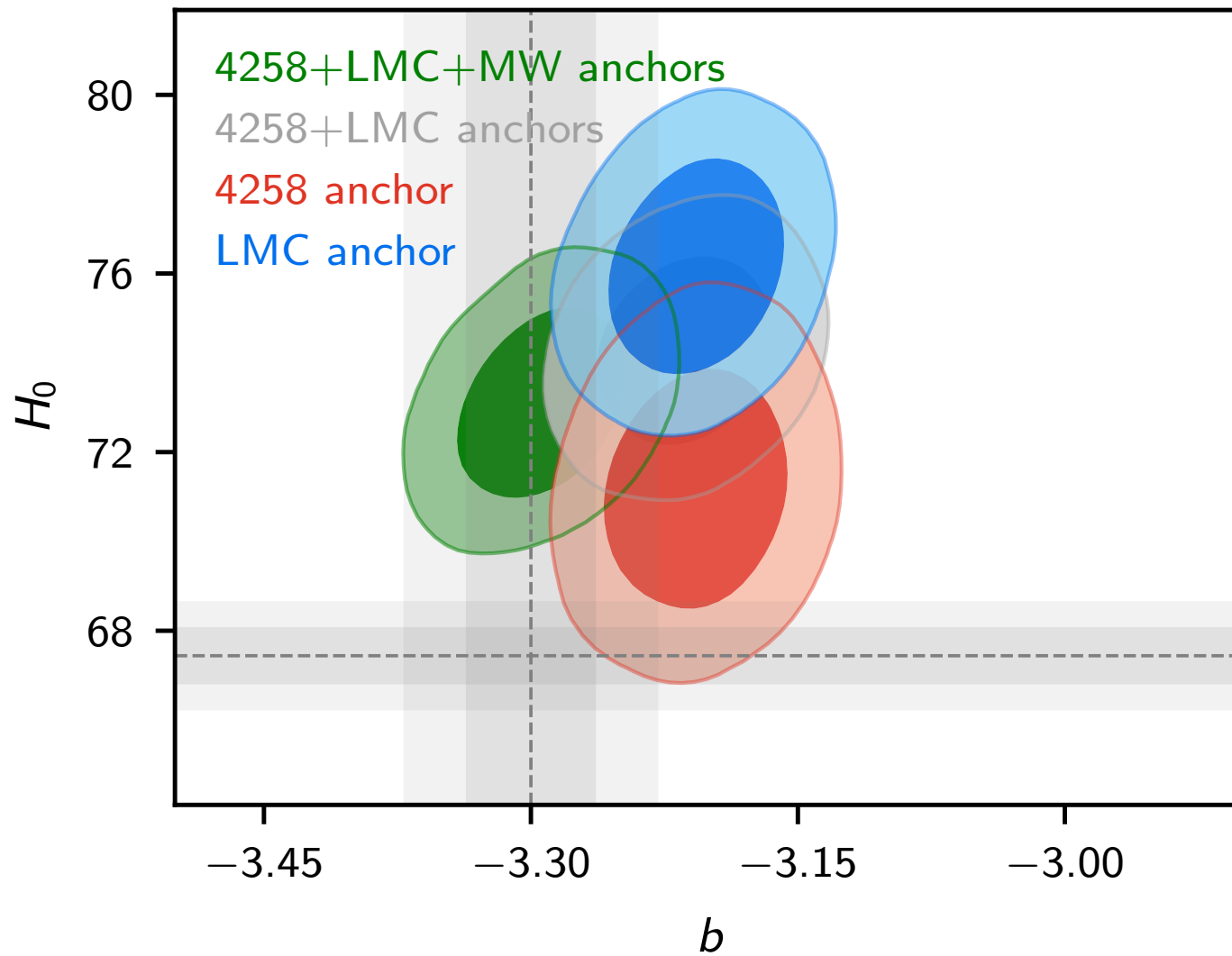
via SH0ES Cepheids $\mu_{N4258} = 29.220 \pm 0.029$ mag



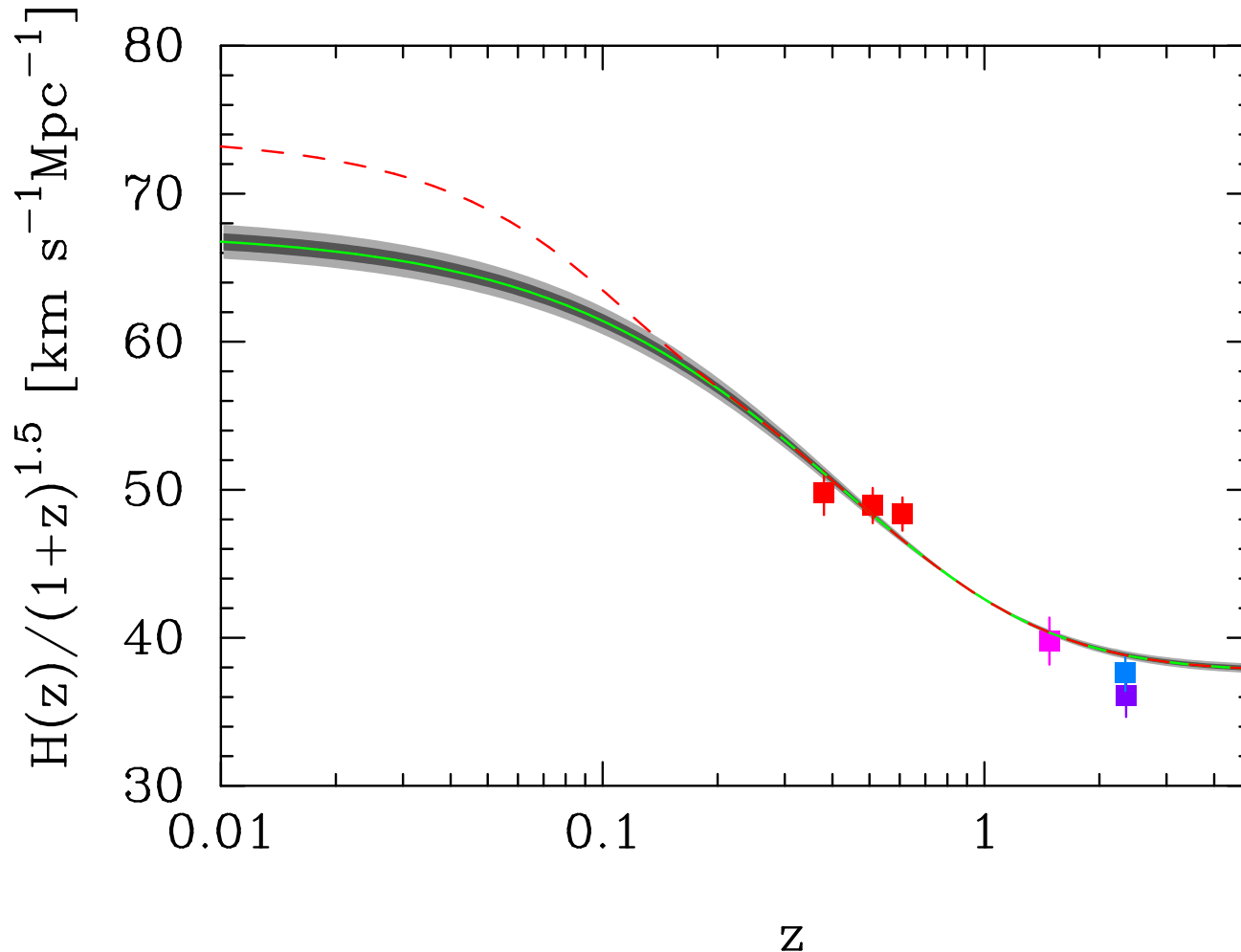
with GAIA DR3 parallaxes and SH0ES
Cepheids

$\mu_{N4258} = 29.211 \pm 0.038$ mag (3.7 σ discrepancy)

so beware of



Point 3: You cannot solve the Hubble tension by modifying late time physics (arXiv:2103.08723)

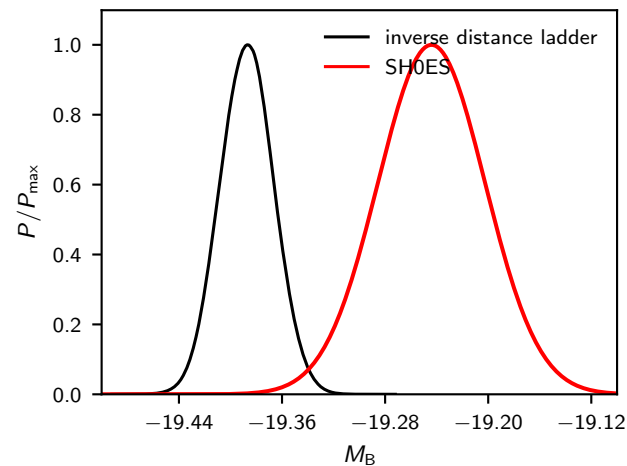
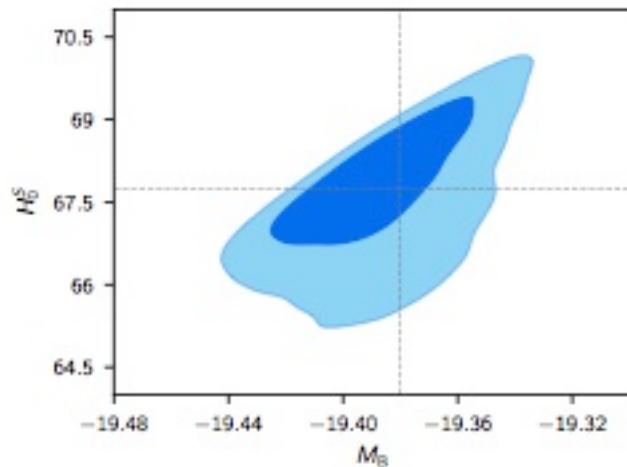


$$H(z) = H_0^f [\Omega_m (1+z)^3 + (1 - \Omega_m)(1 + \Delta \exp(-(z/z_c)^\beta))]^{1/2}$$

$$\begin{aligned} m &= M + 25 + 5 \log_{10} D_L(z), \\ &= -5a + 5 \log_{10} c \hat{d}_L(z) \end{aligned}$$

$$a_B = \left(\sum_{ij} C_{ij}^{-1} (\log_{10} \hat{d}_L(z) - 0.2m_B(i)) \right) / \sum_{ij} C_{ij}^{-1}$$

$$H_0^S = 10^{0.2(M_B + 5a_B + 25)}$$



Inverse distance ladder tightly constrains the form of $H(z)$ independent of dynamics. No late time solution can work - no phantom DE, interacting DE/DM

.....

The Conclusions:

- ❑ Λ CDM fits the Planck data **perfectly** within acceptable statistical errors
- ❑ If your experiment (CMB, LSS, H_0 ) disagrees with Planck, then either you are **wrong**, or there is new physics beyond Λ CDM.
- ❑ Any new physics must produce temperature and polarization spectra that are degenerate with base Λ CDM over the multipole range $2 \leq \ell \leq 2500$. Any such evidence is strongly dependent on the **fidelity** of other data.