

Edge of darkness: The splashback radius as a physical halo boundary

Benedikt Diemer

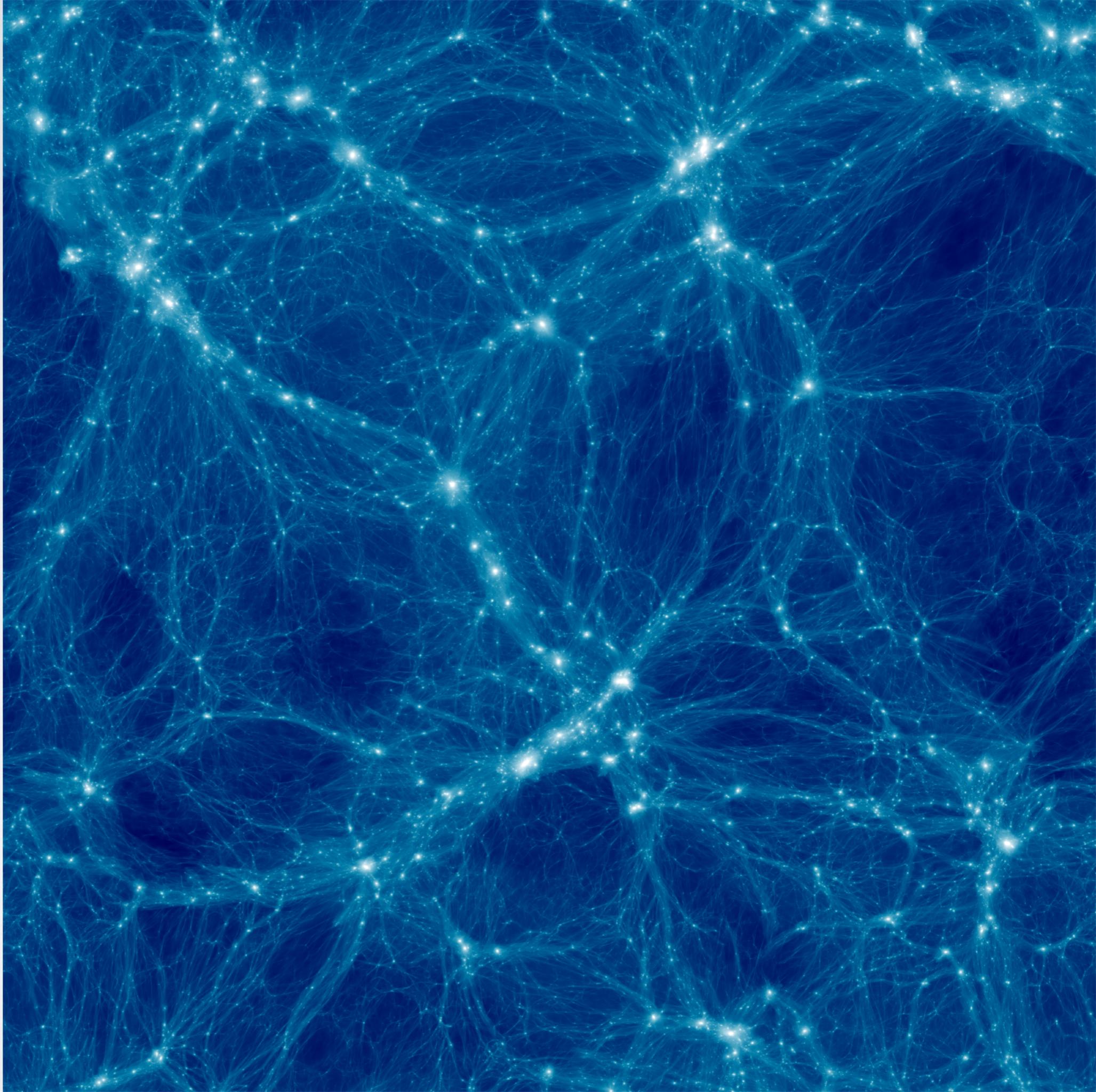
ITC Fellow, Harvard-Smithsonian Center for Astrophysics

In collaboration with:

Andrey Kravtsov, Surhud More, and Philip Mansfield

Cosmology Seminar • UC Berkeley • 04/17/2018

89 Mpc



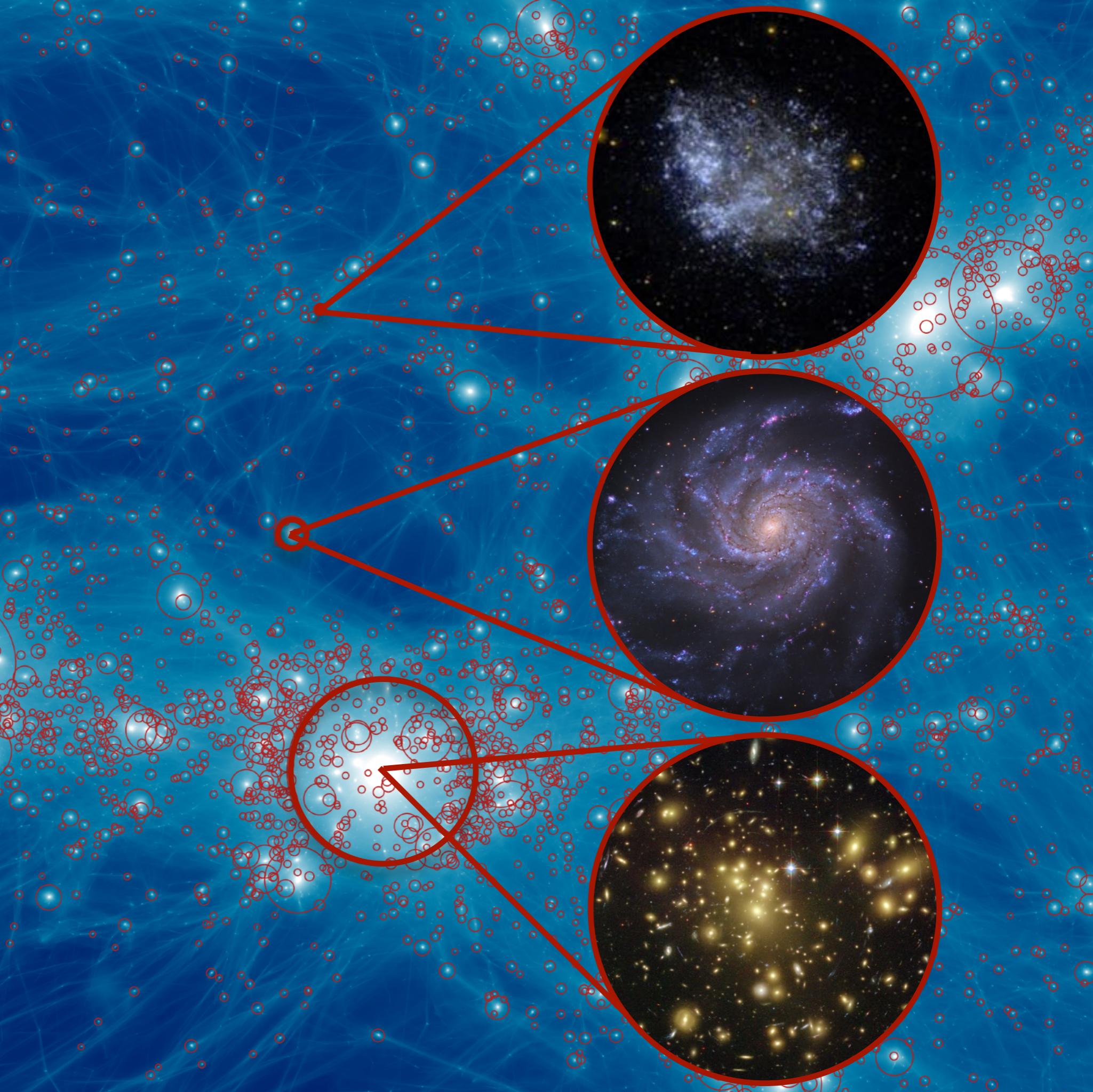
Visualization code:

Phil Mansfield

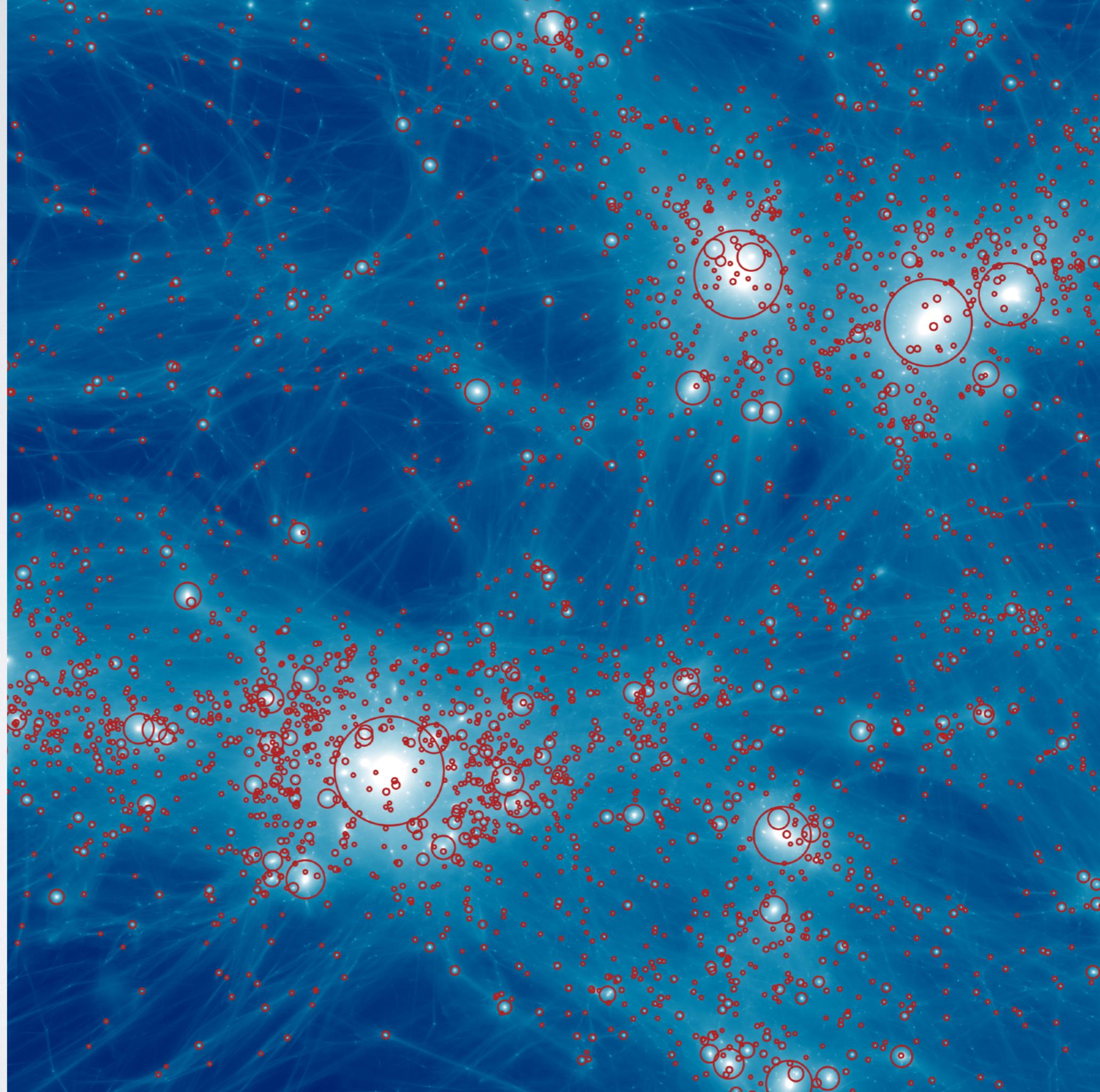
Algorithm:

Tom Abel, Oliver Hahn,
Ralf Kaehler

11 Mpc

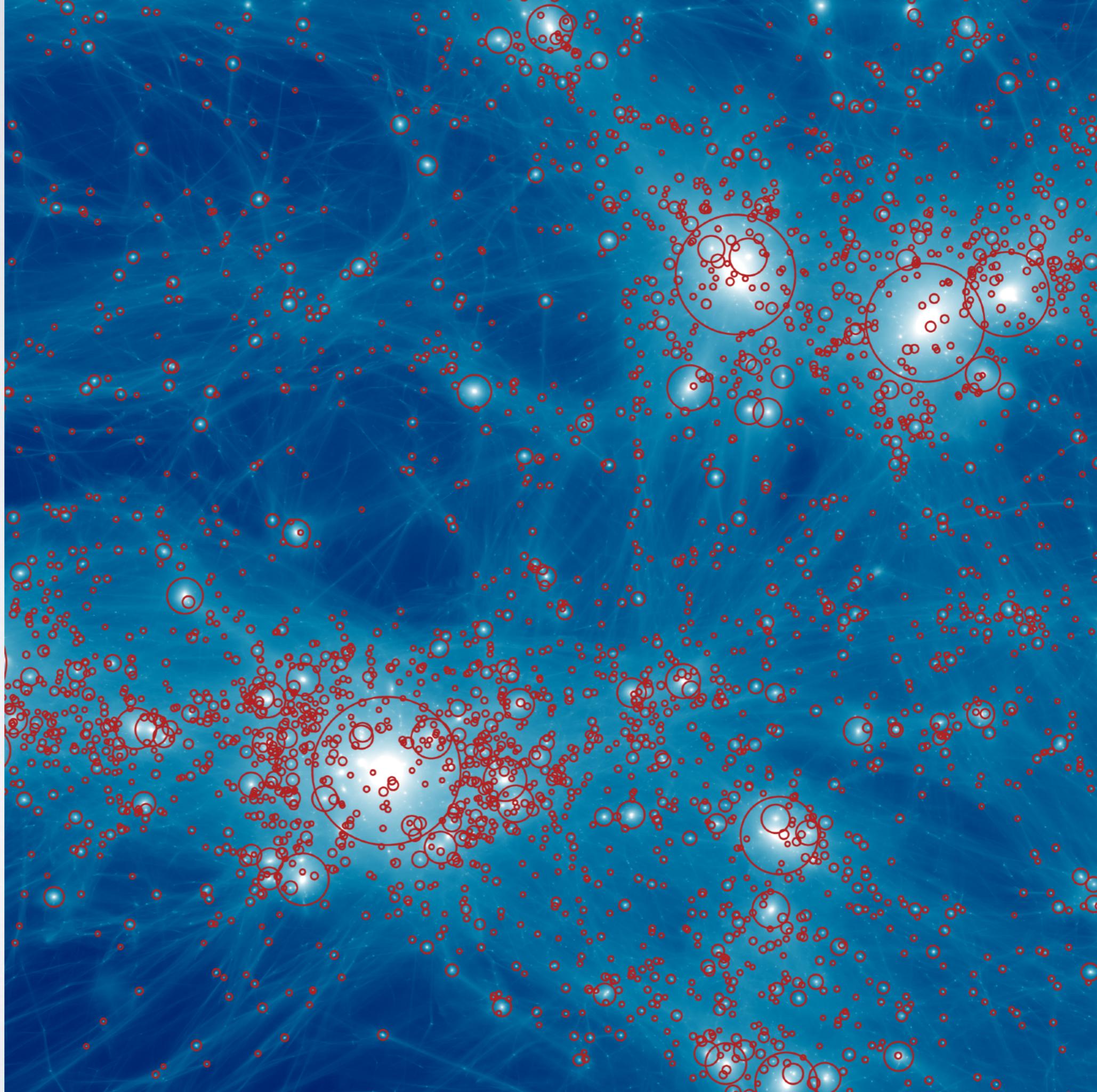


R_{500c}



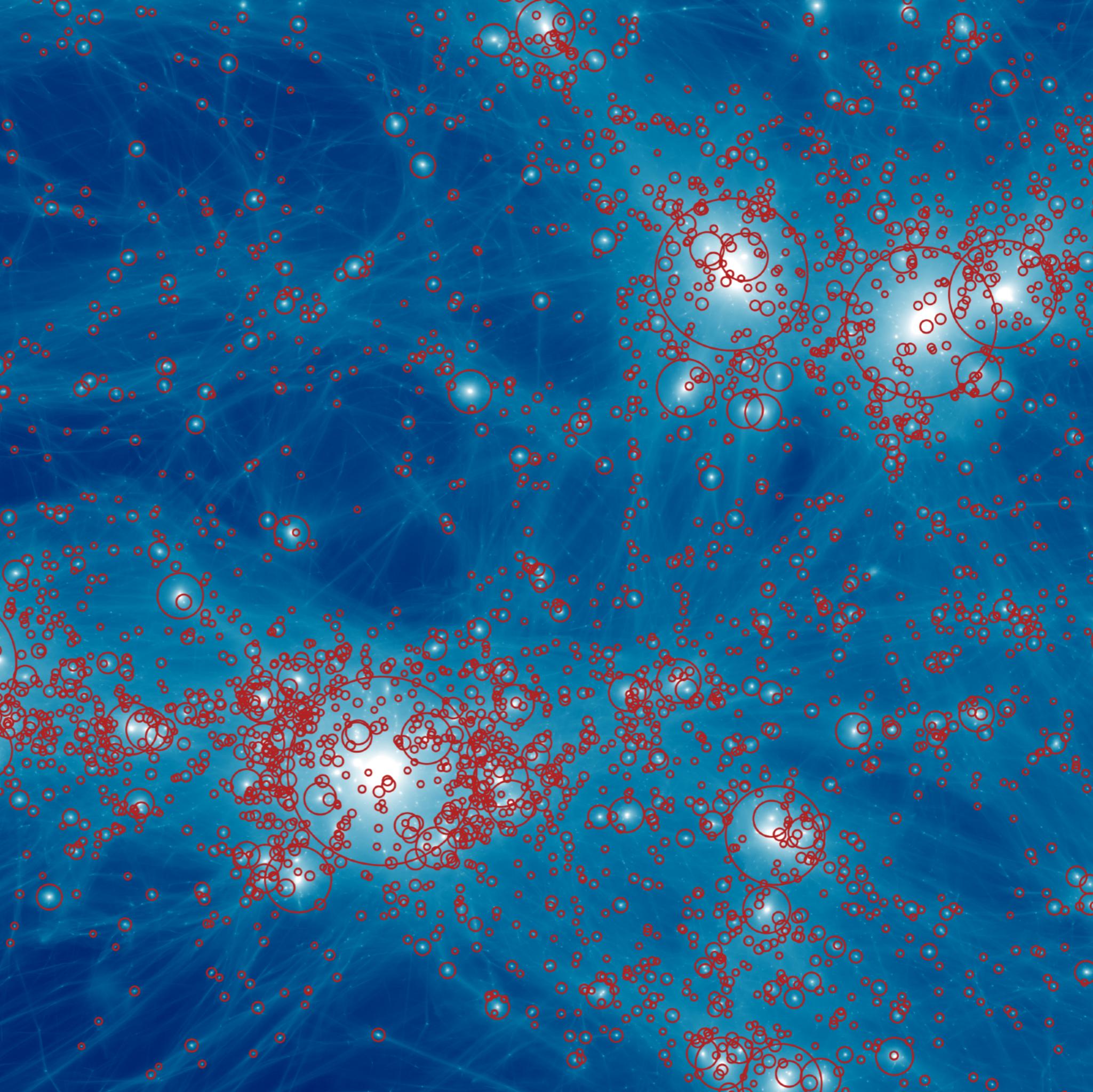
Halo finder: Rockstar
(Behroozi et al. 2013)

R_{200c}



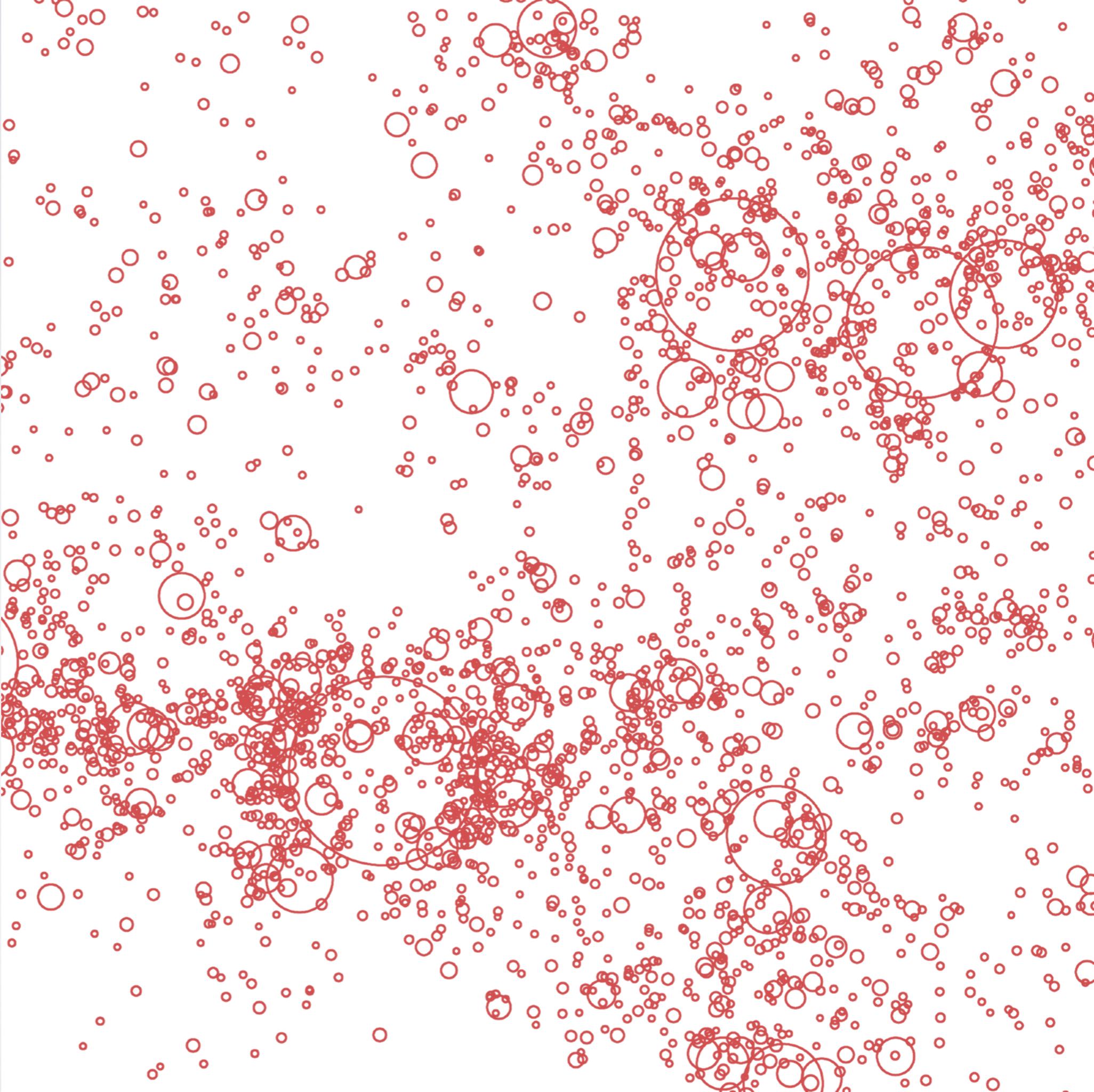
Halo finder: Rockstar
(Behroozi et al. 2013)

R_{vir}



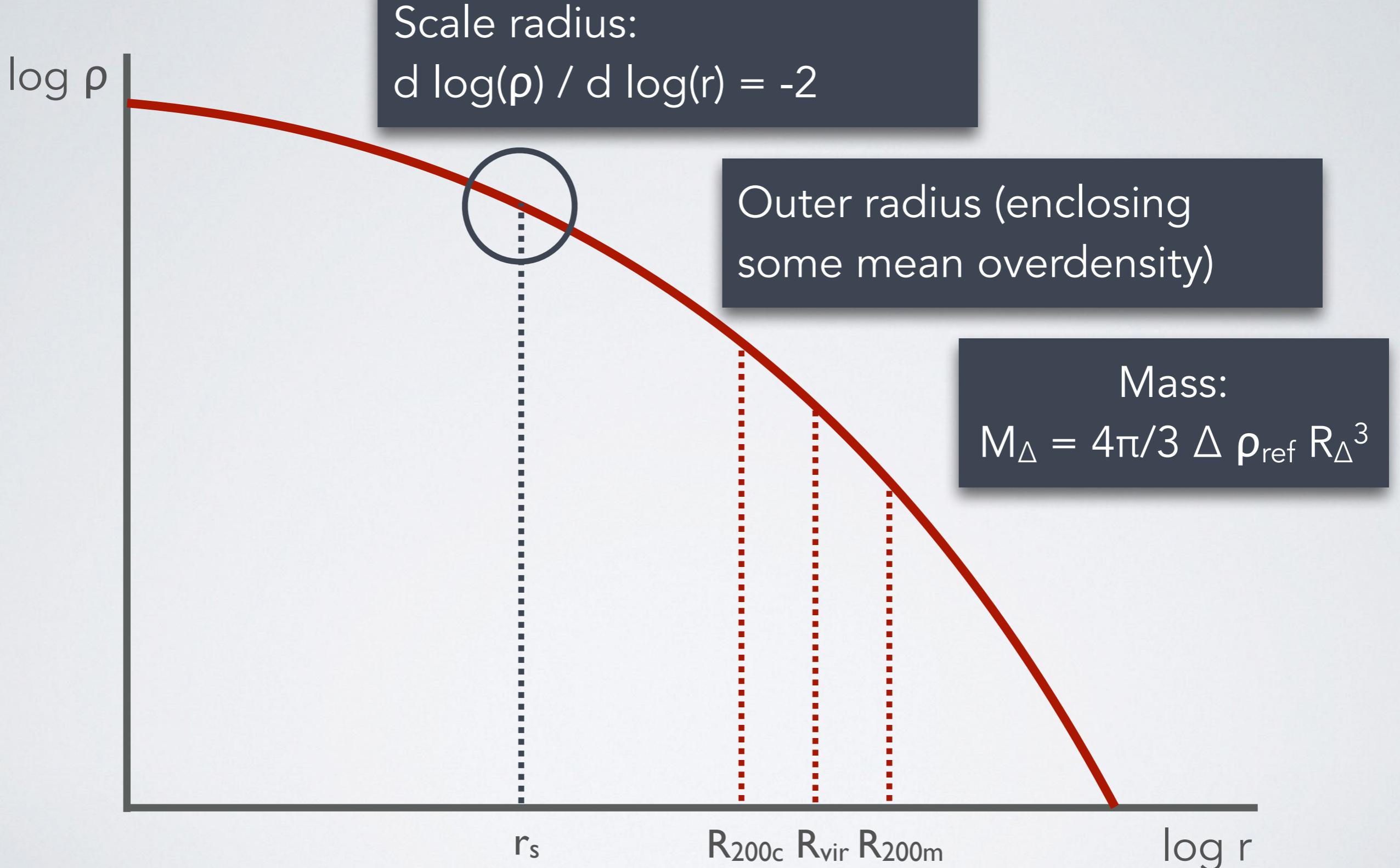
Halo finder: Rockstar
(Behroozi et al. 2013)

R_{vir}



Halo finder: Rockstar
(Behroozi et al. 2013)

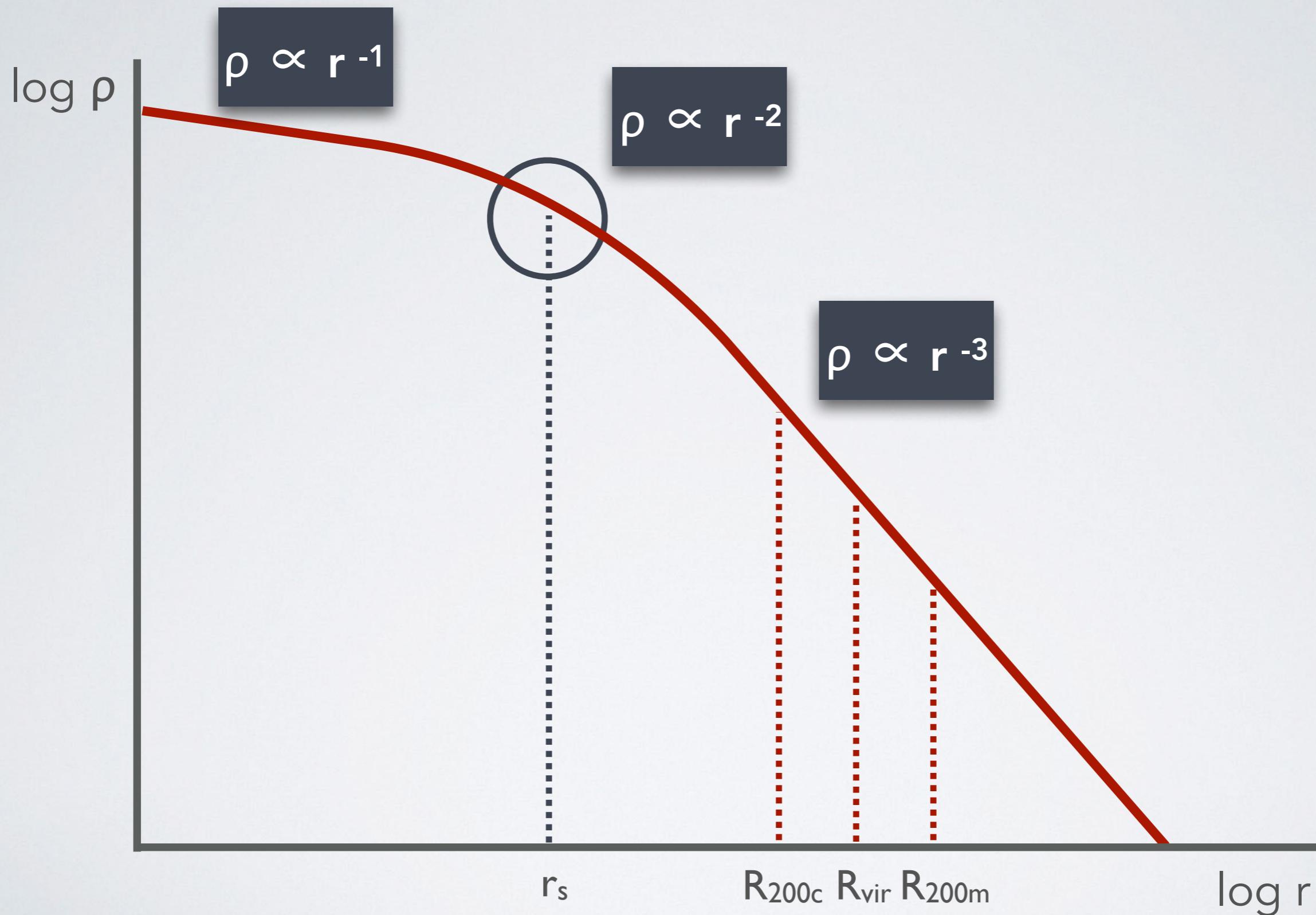
Density profile



Einasto 1965 • Frenk et al. 1988 • Hernquist 1990 • Dubinski & Carlberg 1991

Navarro et al. 1995/1996/1997/2004

Navarro-Frenk-White profile



The “virial” radius



The “virial” radius

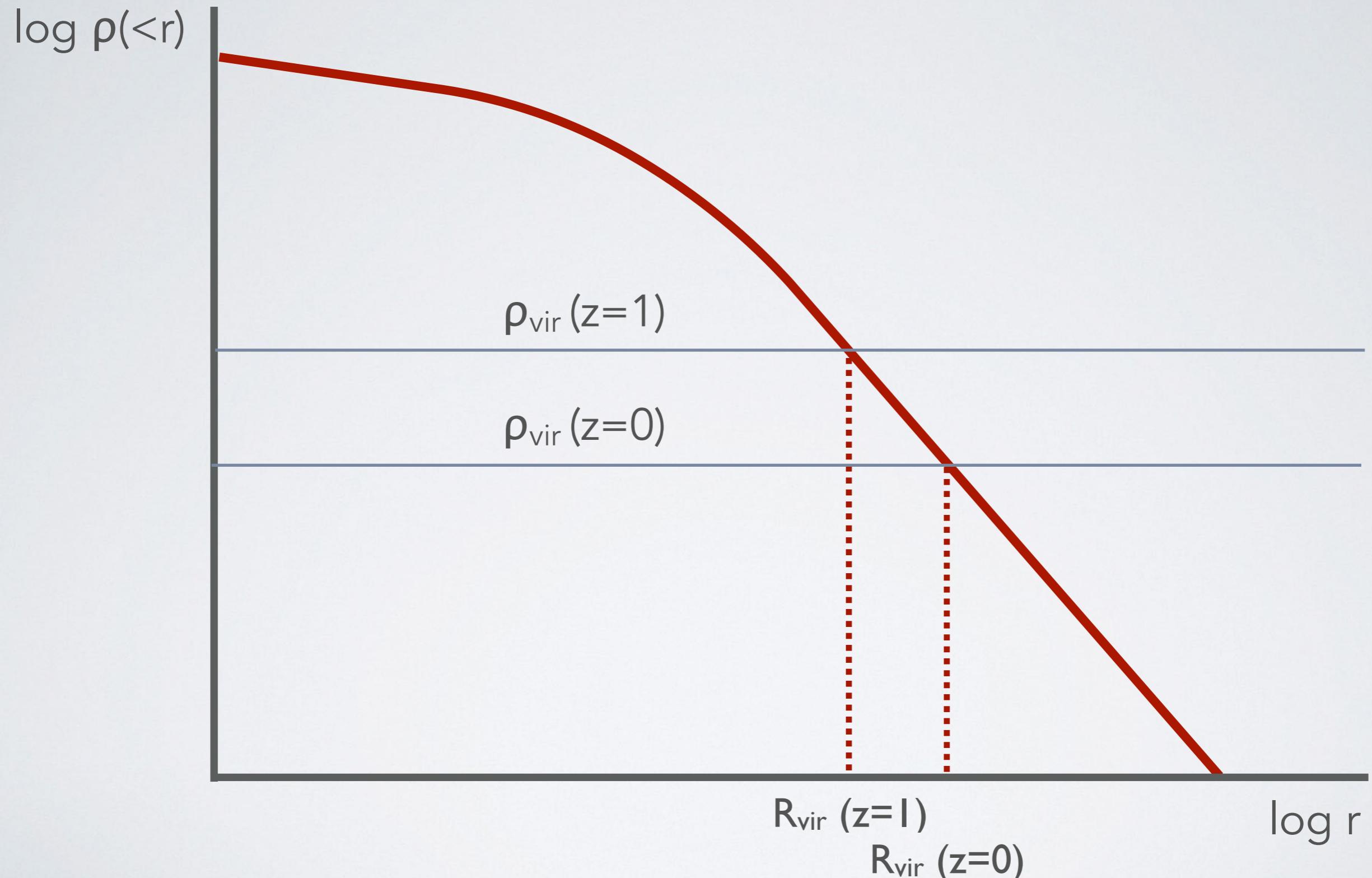


$$W = -2K$$

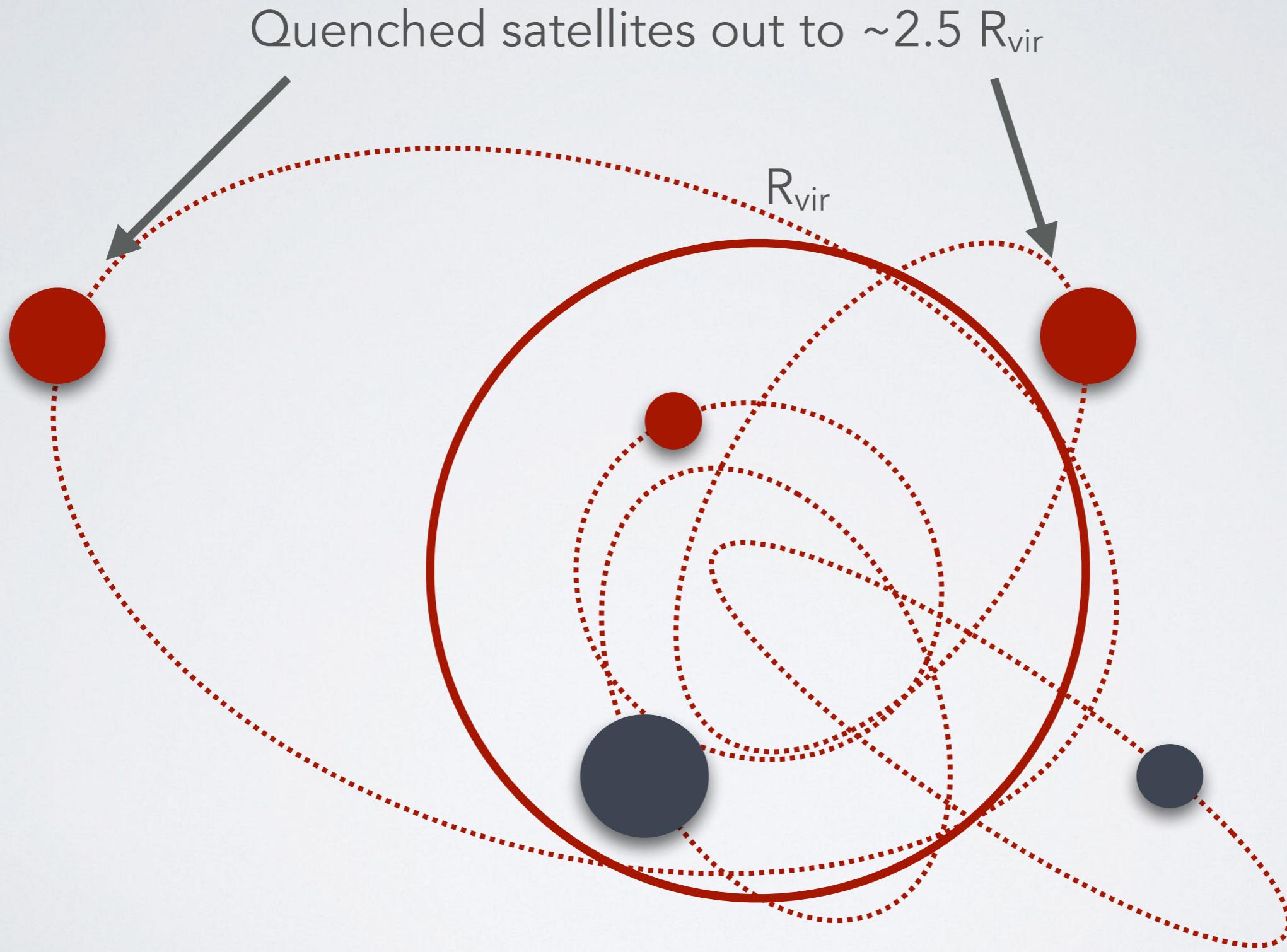
$$\rightarrow R_{\text{vir}} = 1/2 R_{\text{max}}$$

$$\rightarrow \Delta_{\text{vir}} = 18 \pi^2 = 178$$

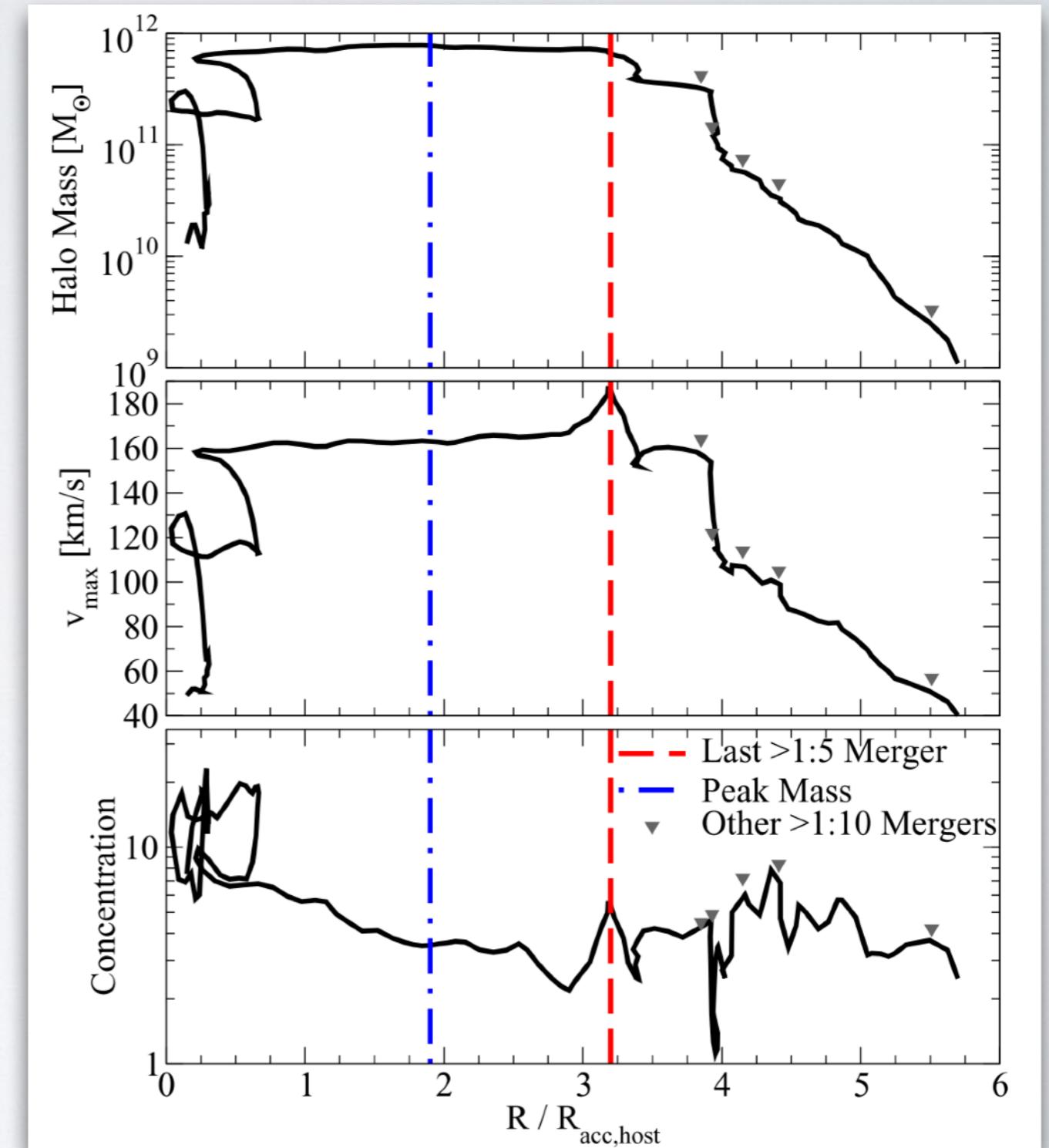
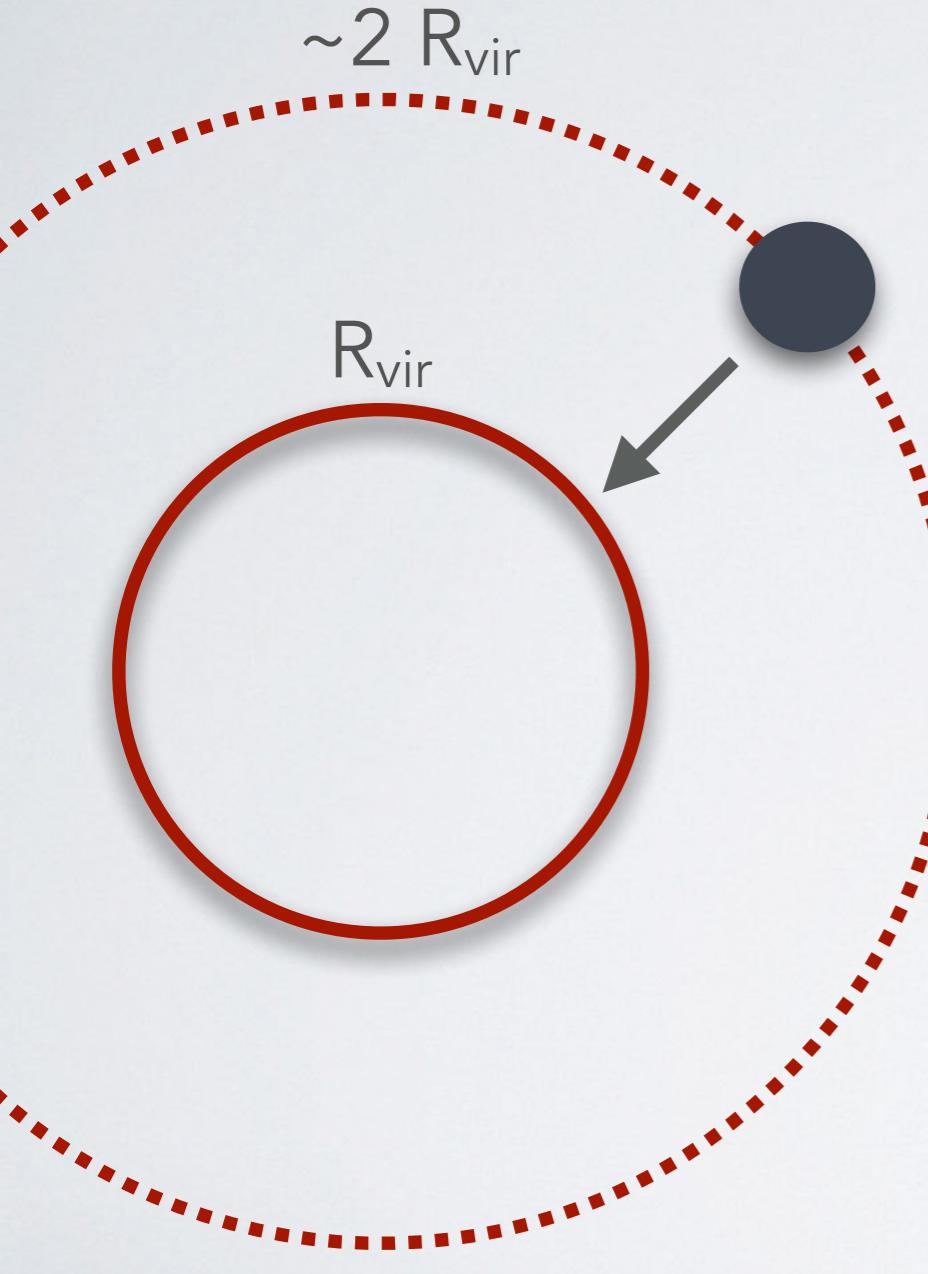
Pseudo-evolution



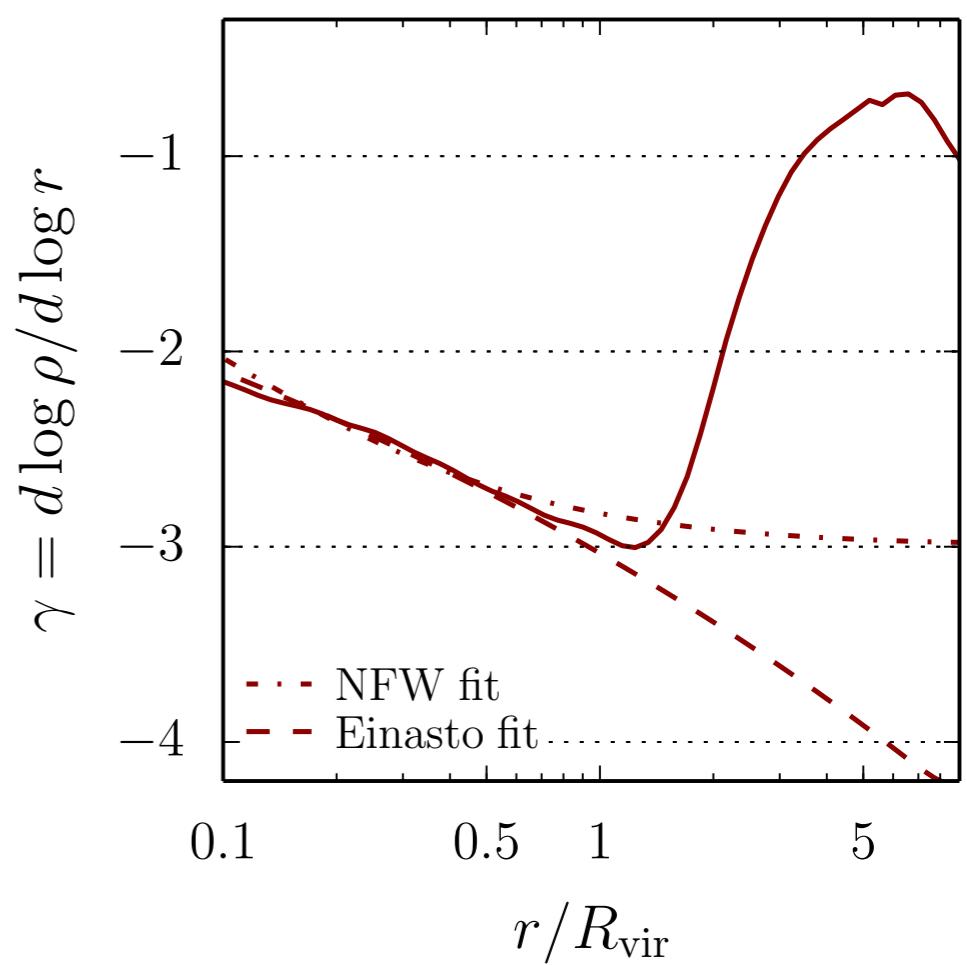
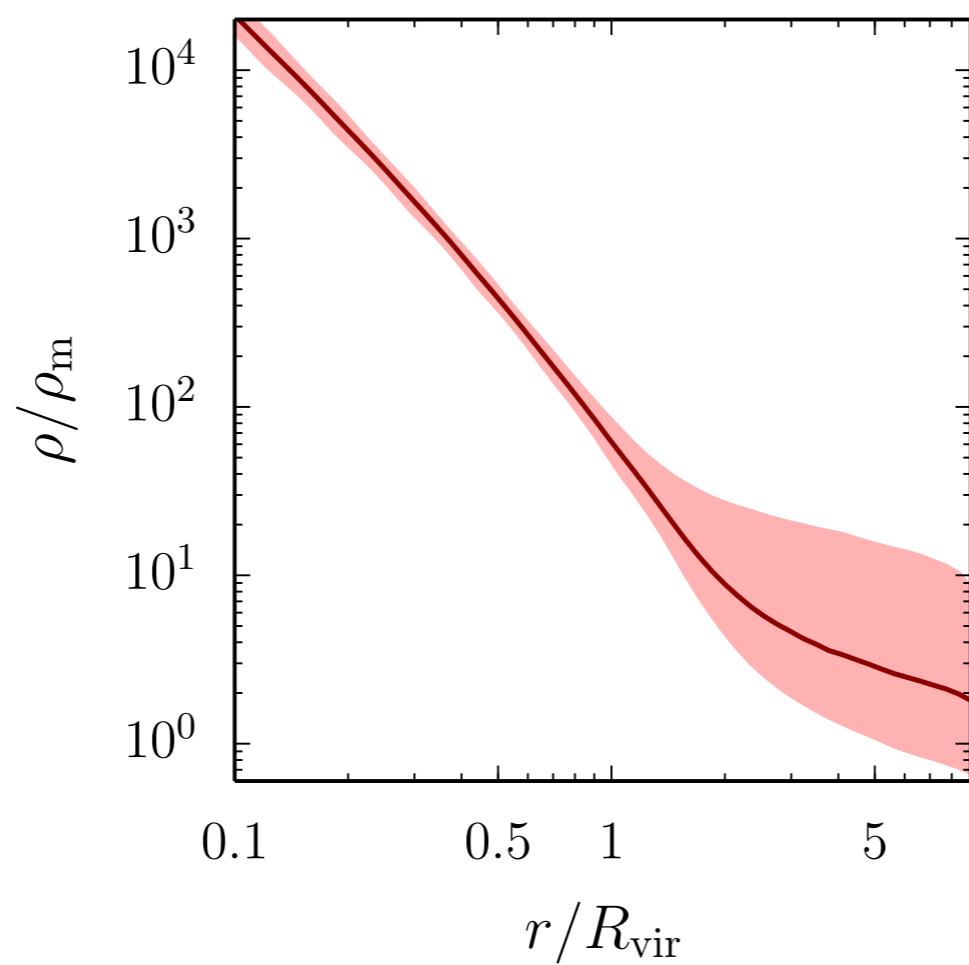
“Ejected” satellites



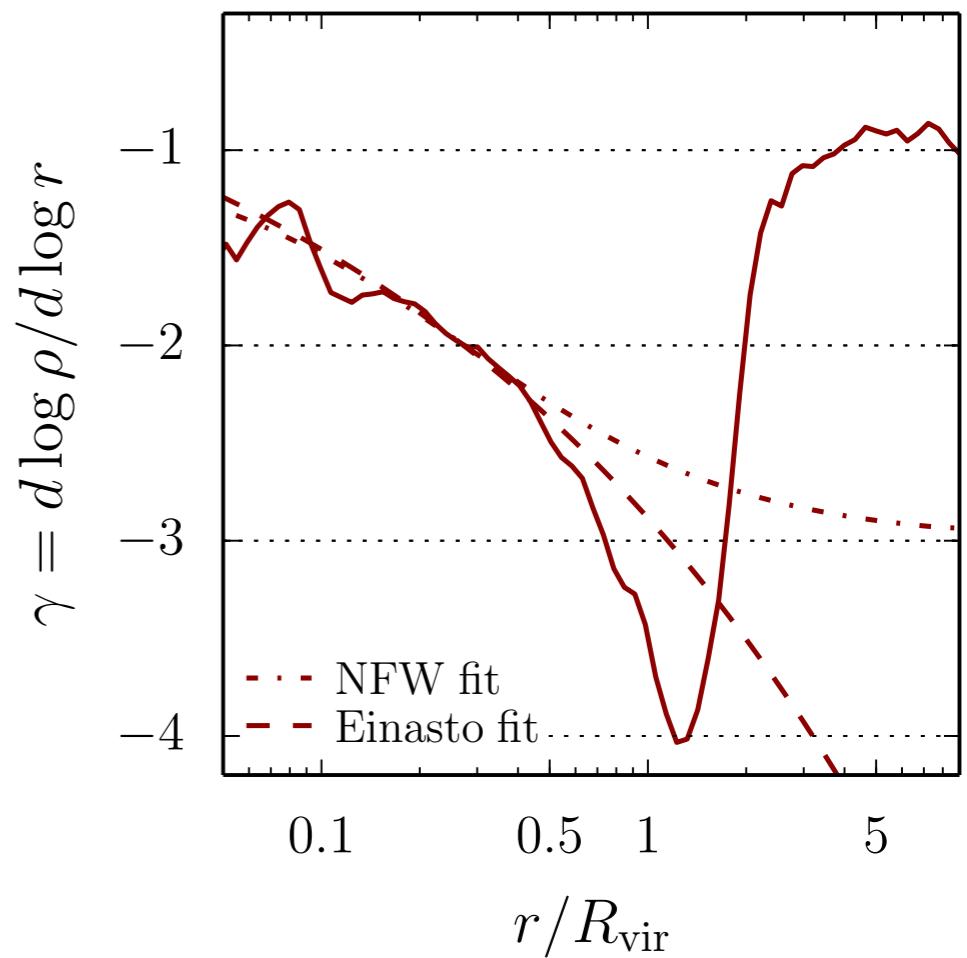
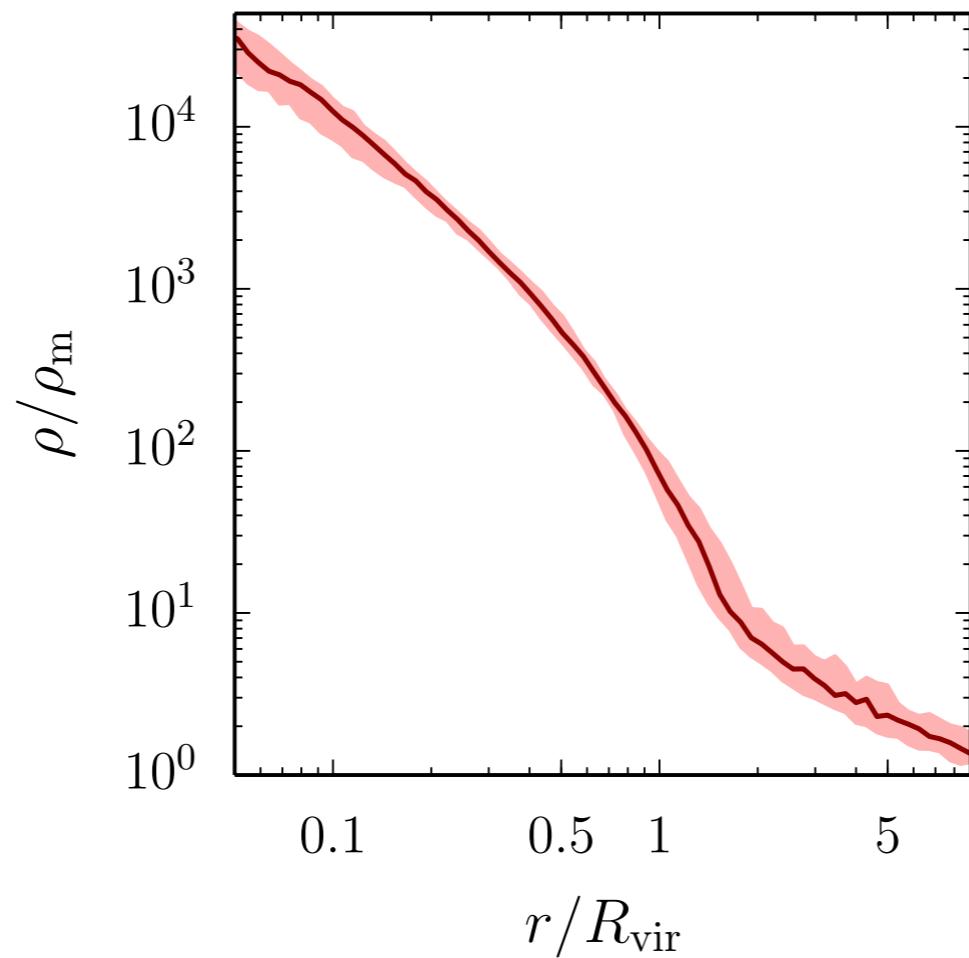
Sphere of influence



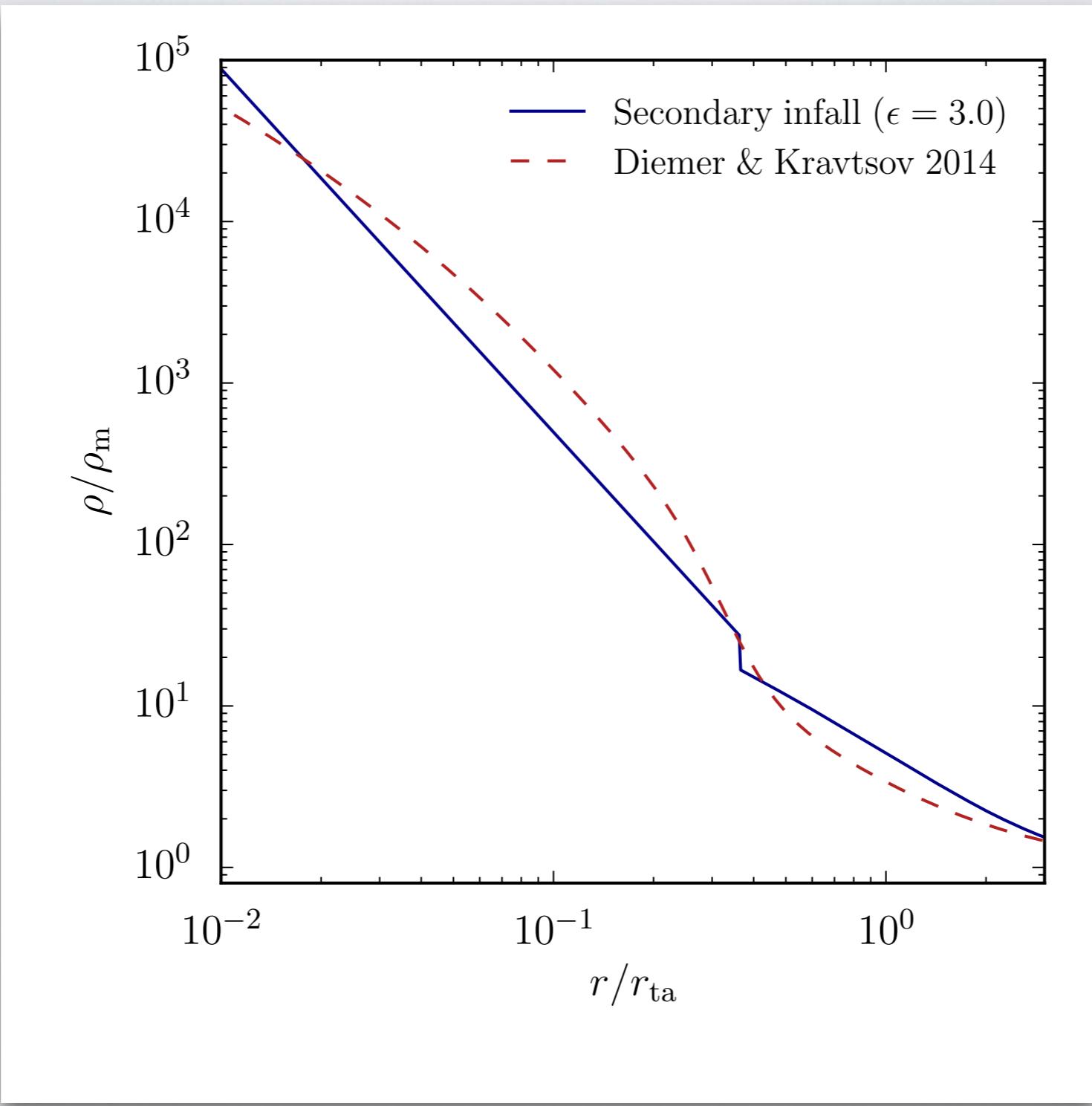
Small halos
 $(10^{10} < M < 3 \times 10^{11})$



Large halos
 $(M > 10^{15})$

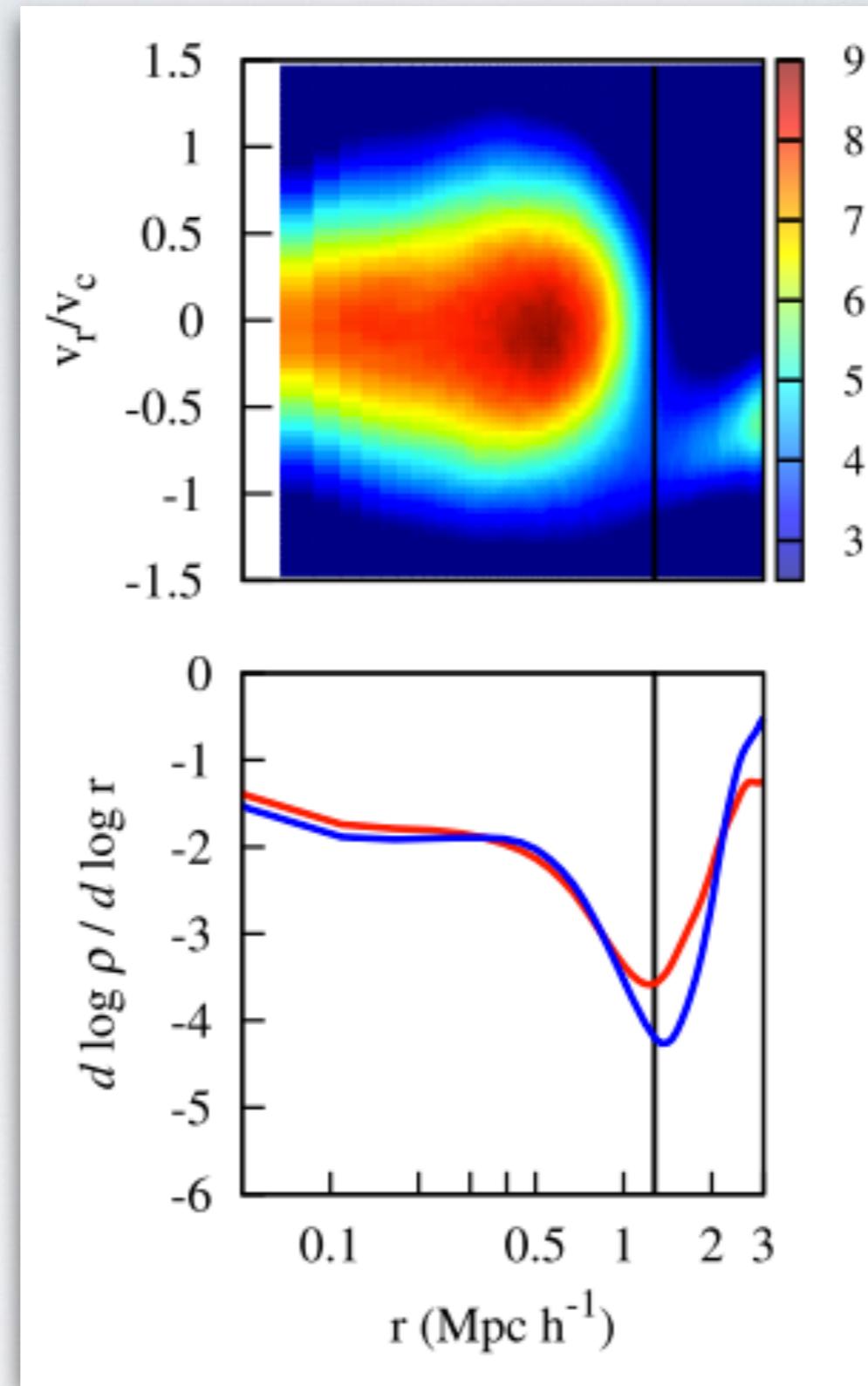


The Splashback Radius

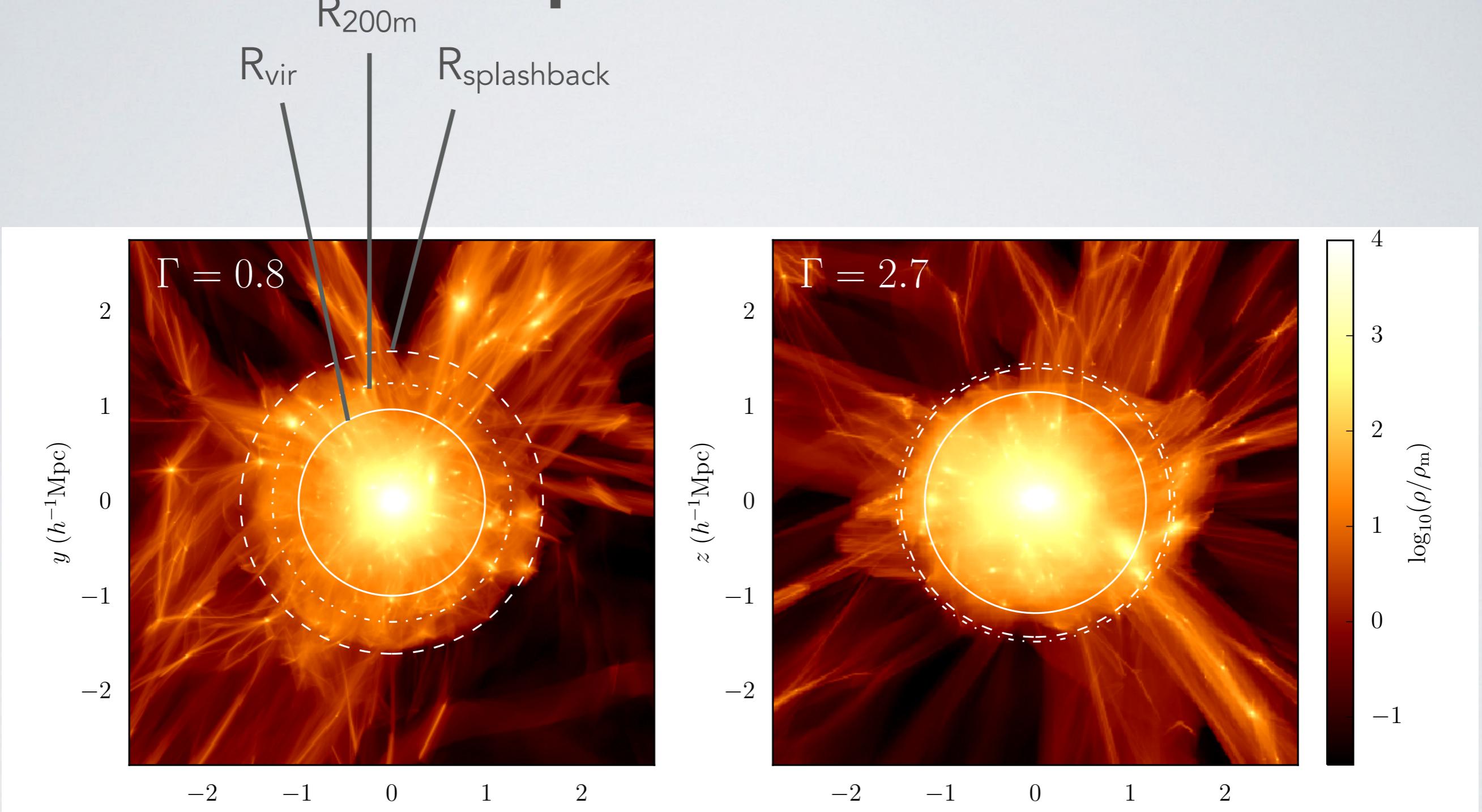


Fillmore & Goldreich 1984 • Bertschinger 1985 • Diemand & Kuhlen 2008
Vogelsberger et al. 2011 • Lithwick & Dalal 2011 • Adhikari et al. 2014 • Diemer & Kravtsov 2014

Phase space (stacked halos)



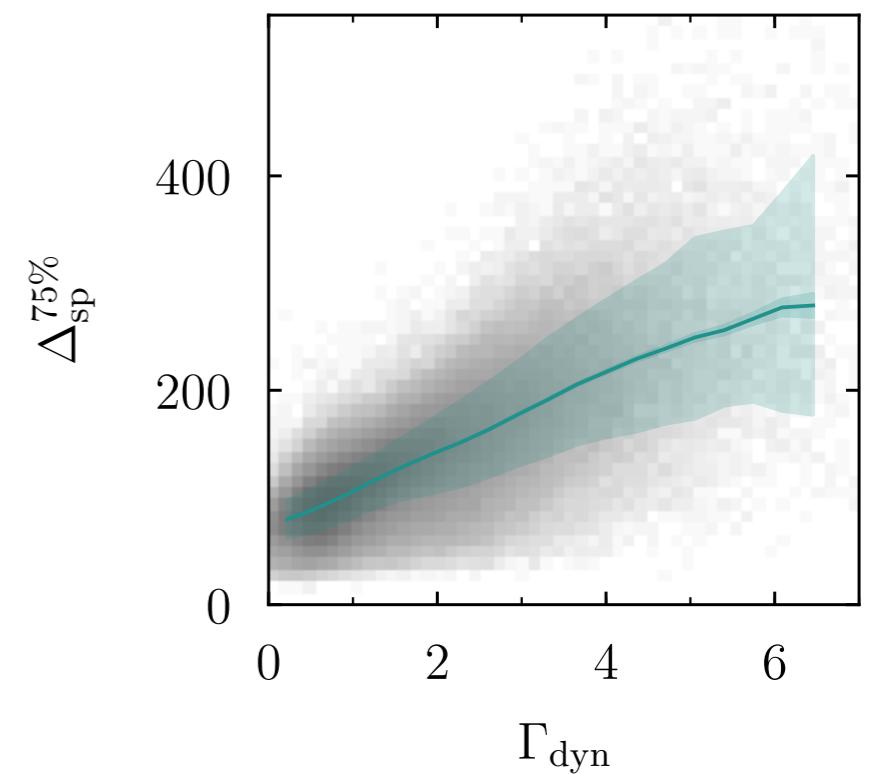
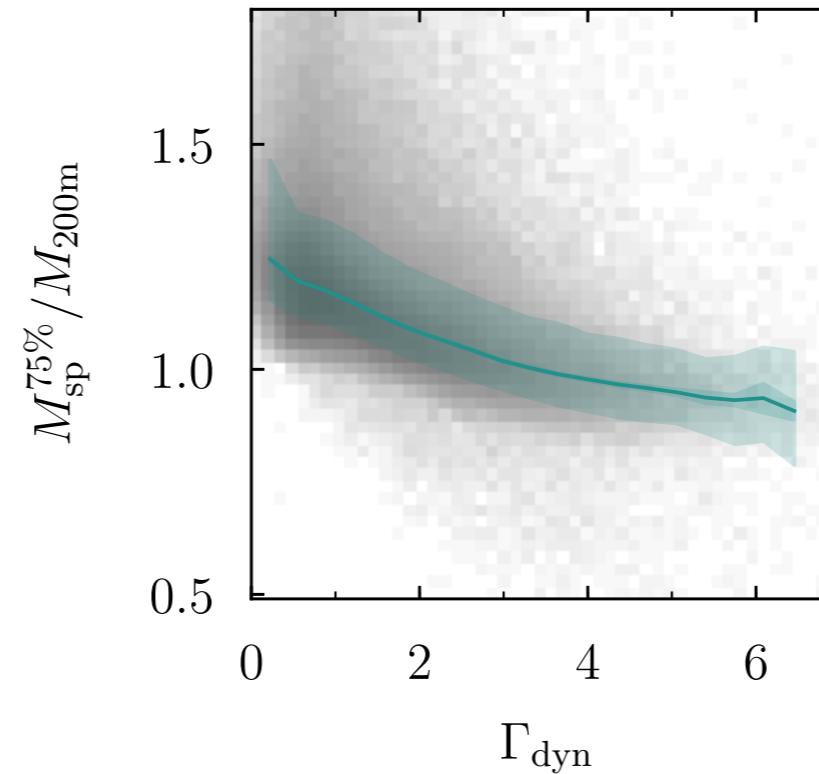
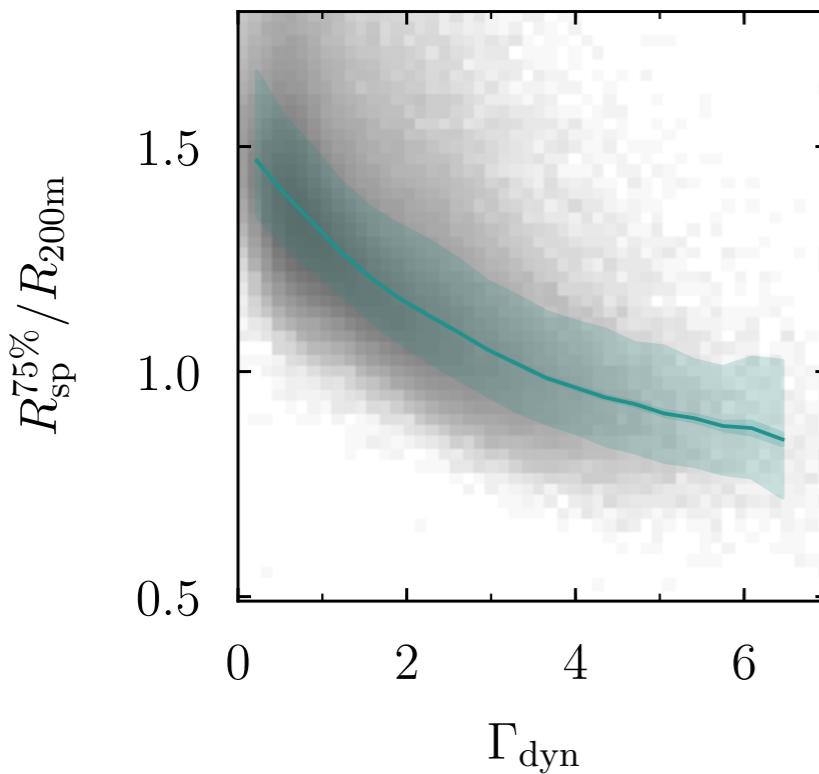
The Splashback Radius



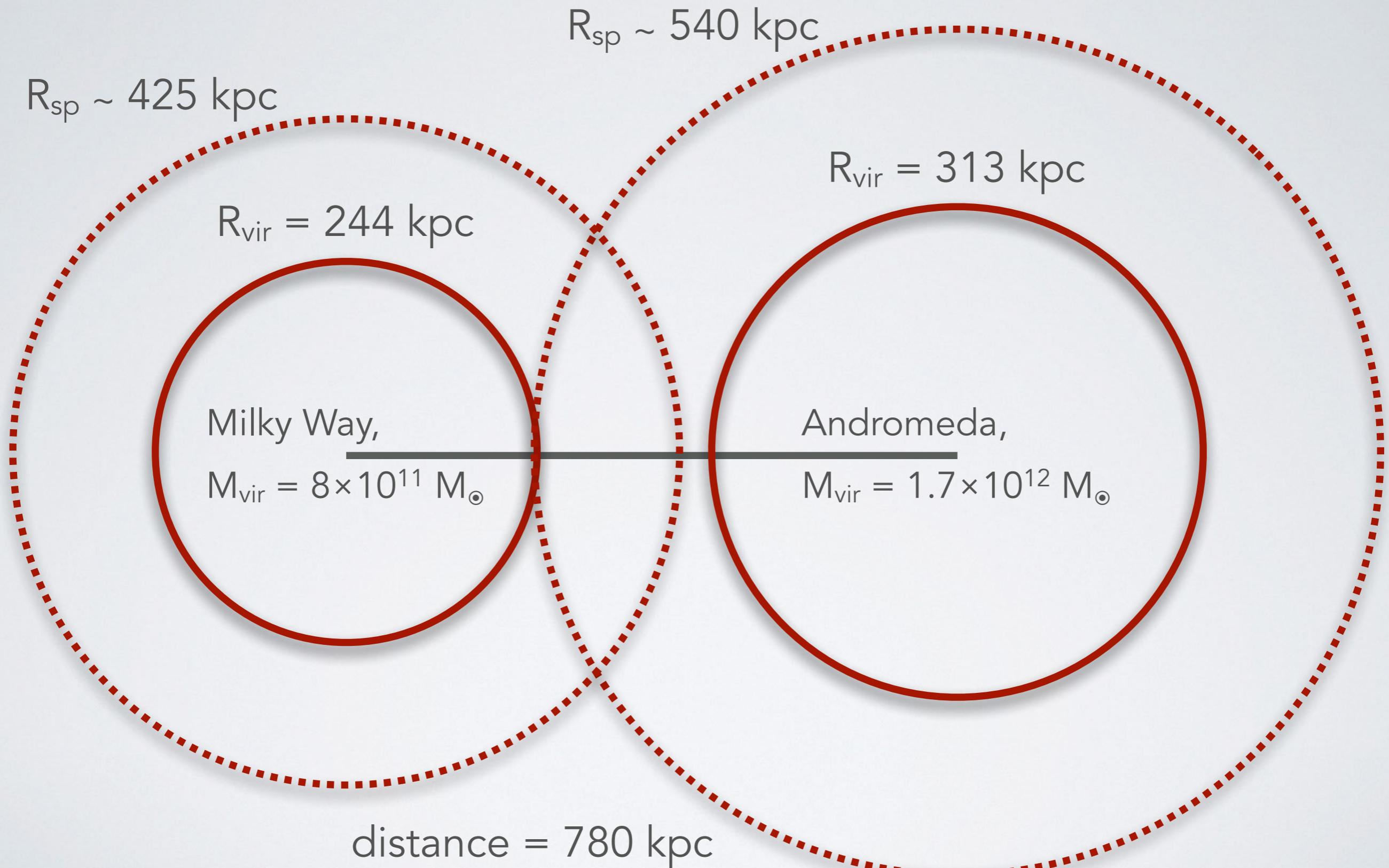
Low accretion rate

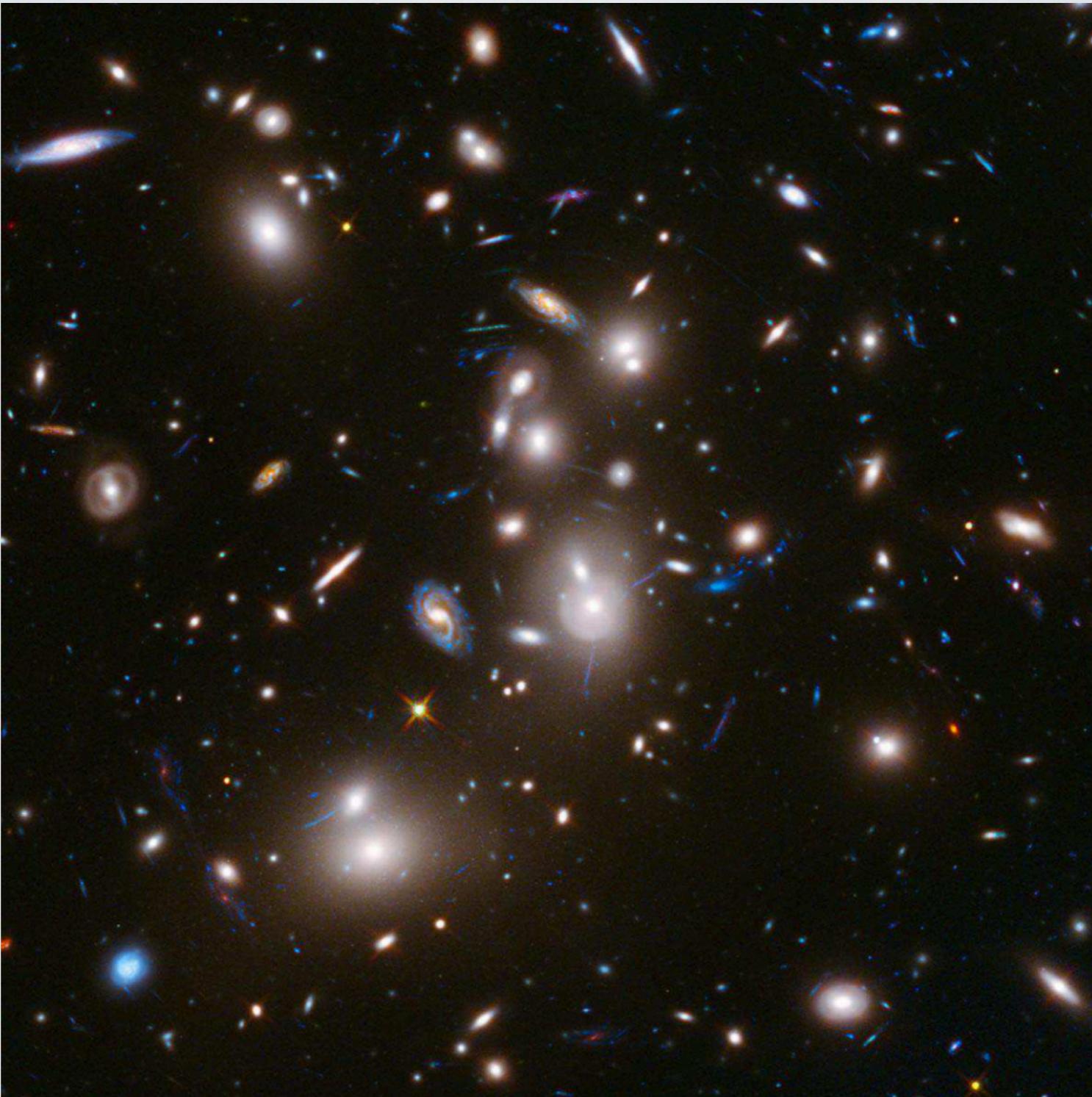
High accretion rate

The Γ - R_{sp} relation

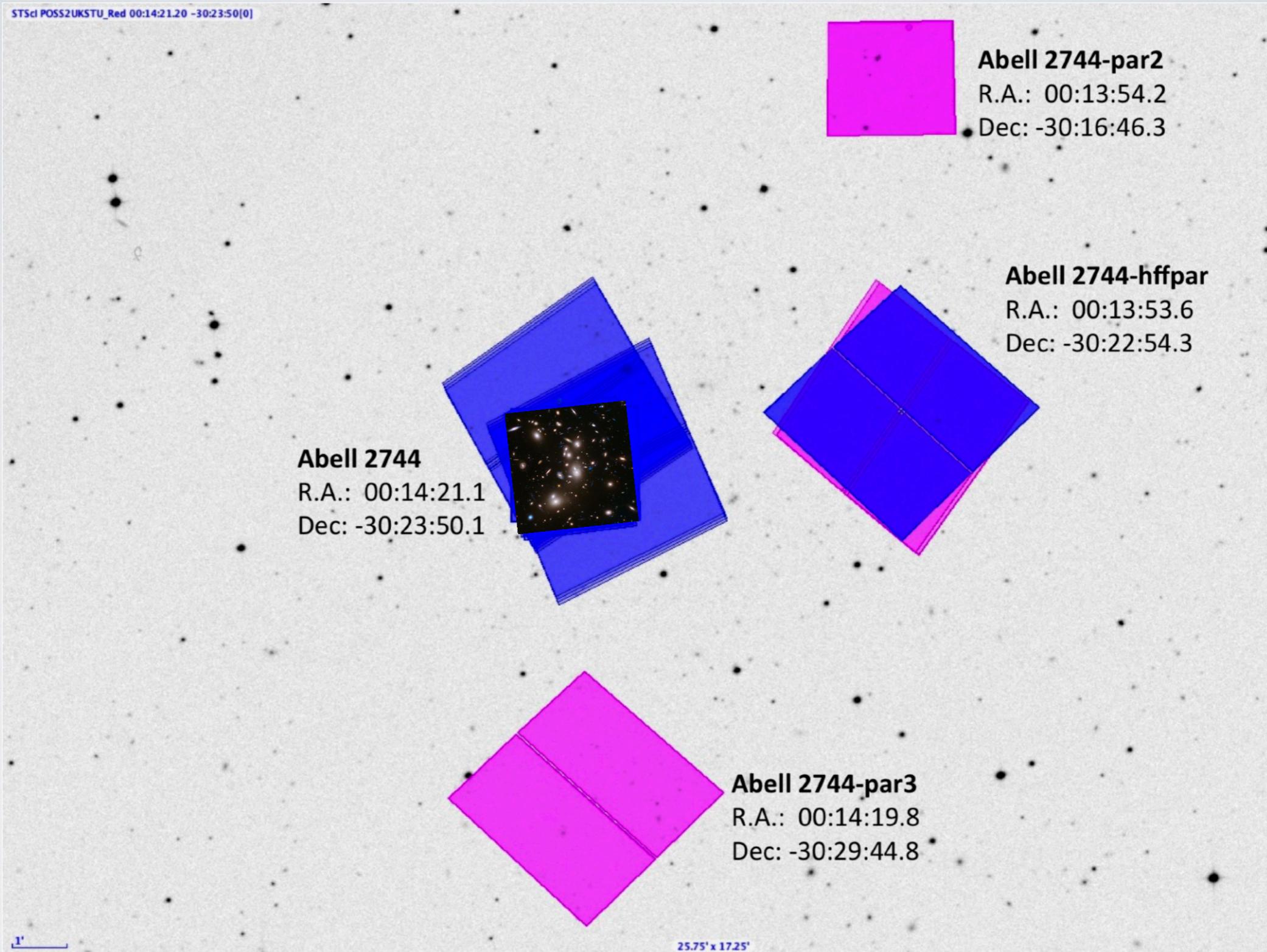


Do the Milky Way and Andromeda halos overlap?





55 kpc

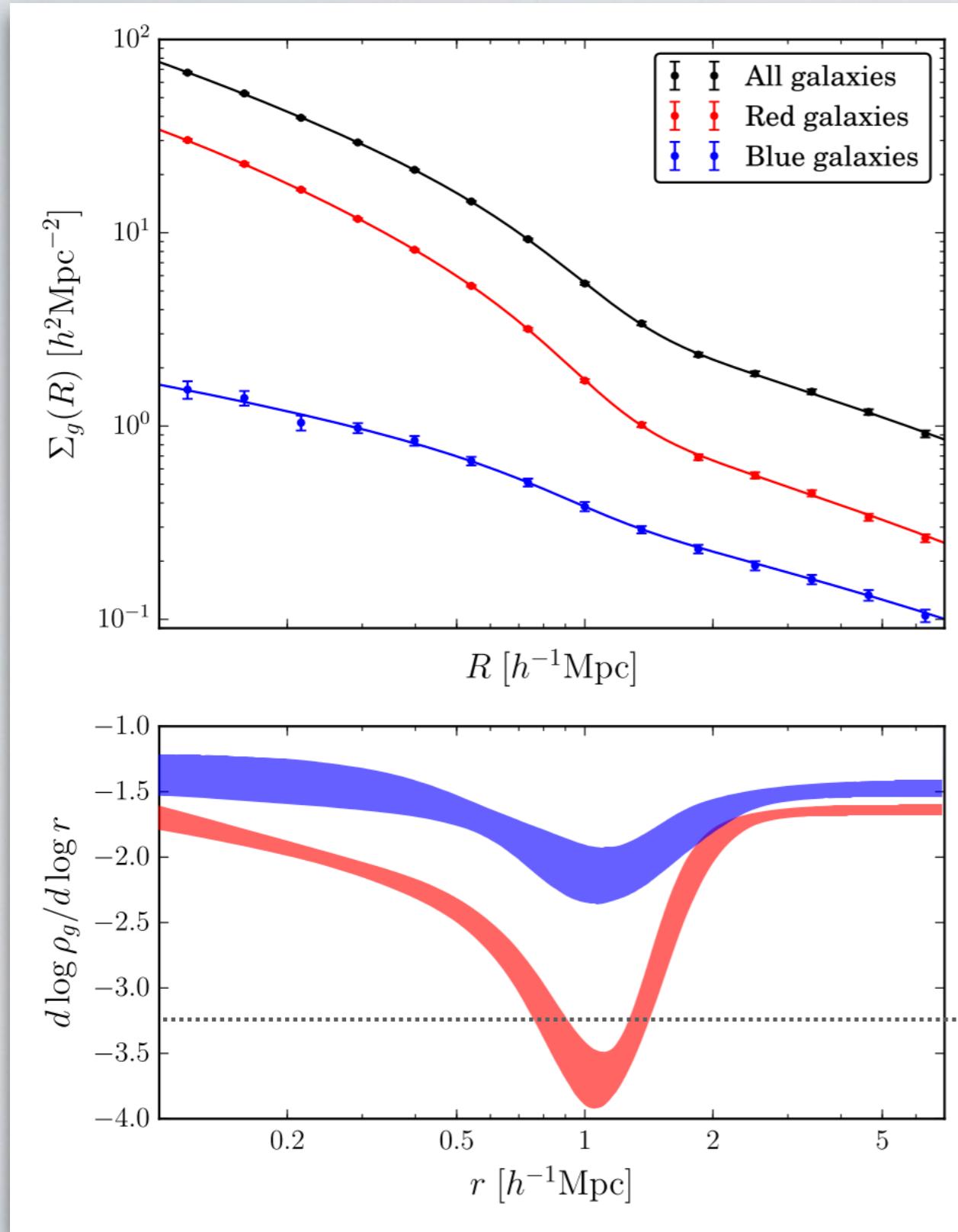


950 kpc



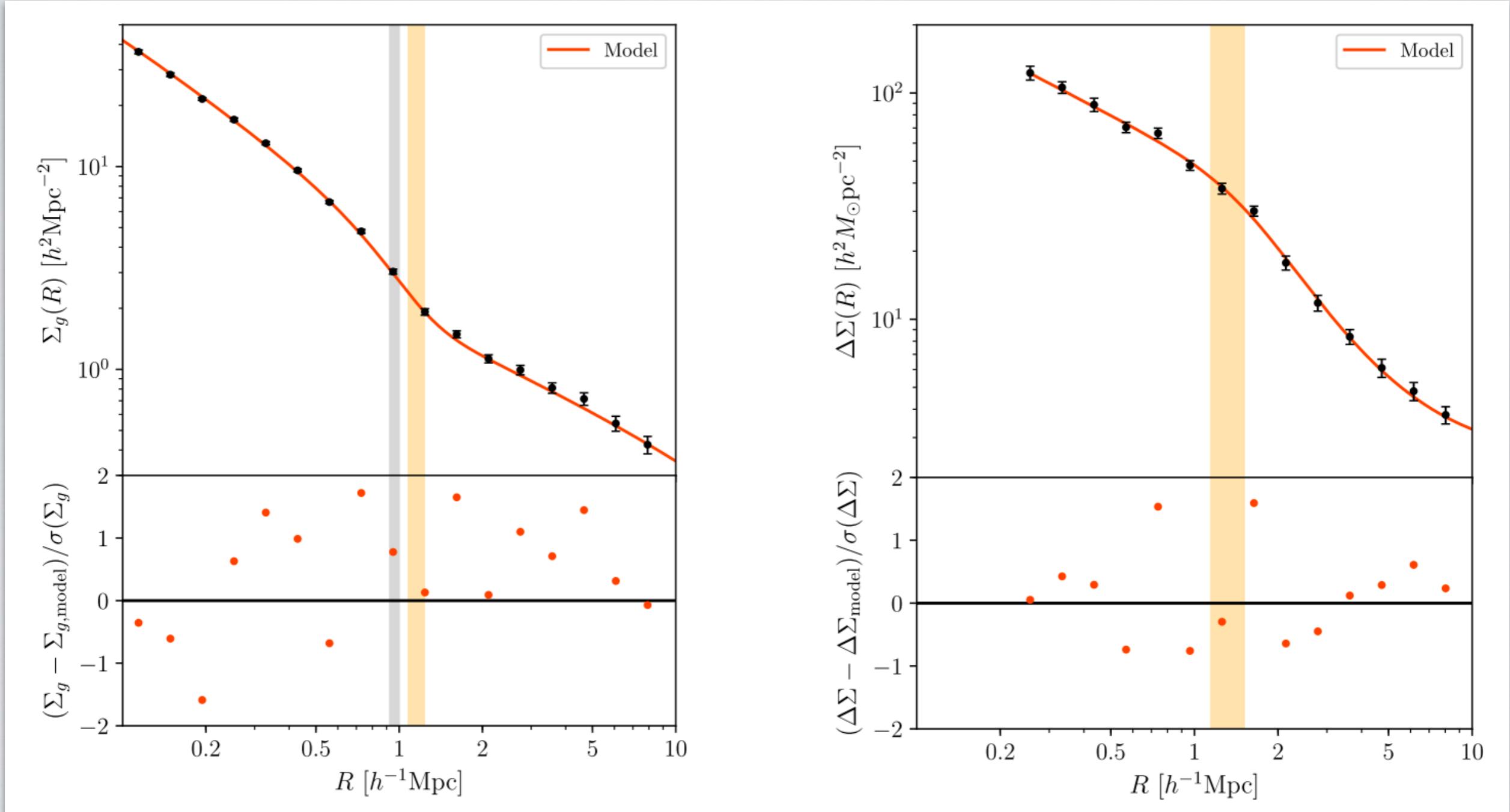
~3 Mpc

R_{sp} in cluster member profiles (SDSS)

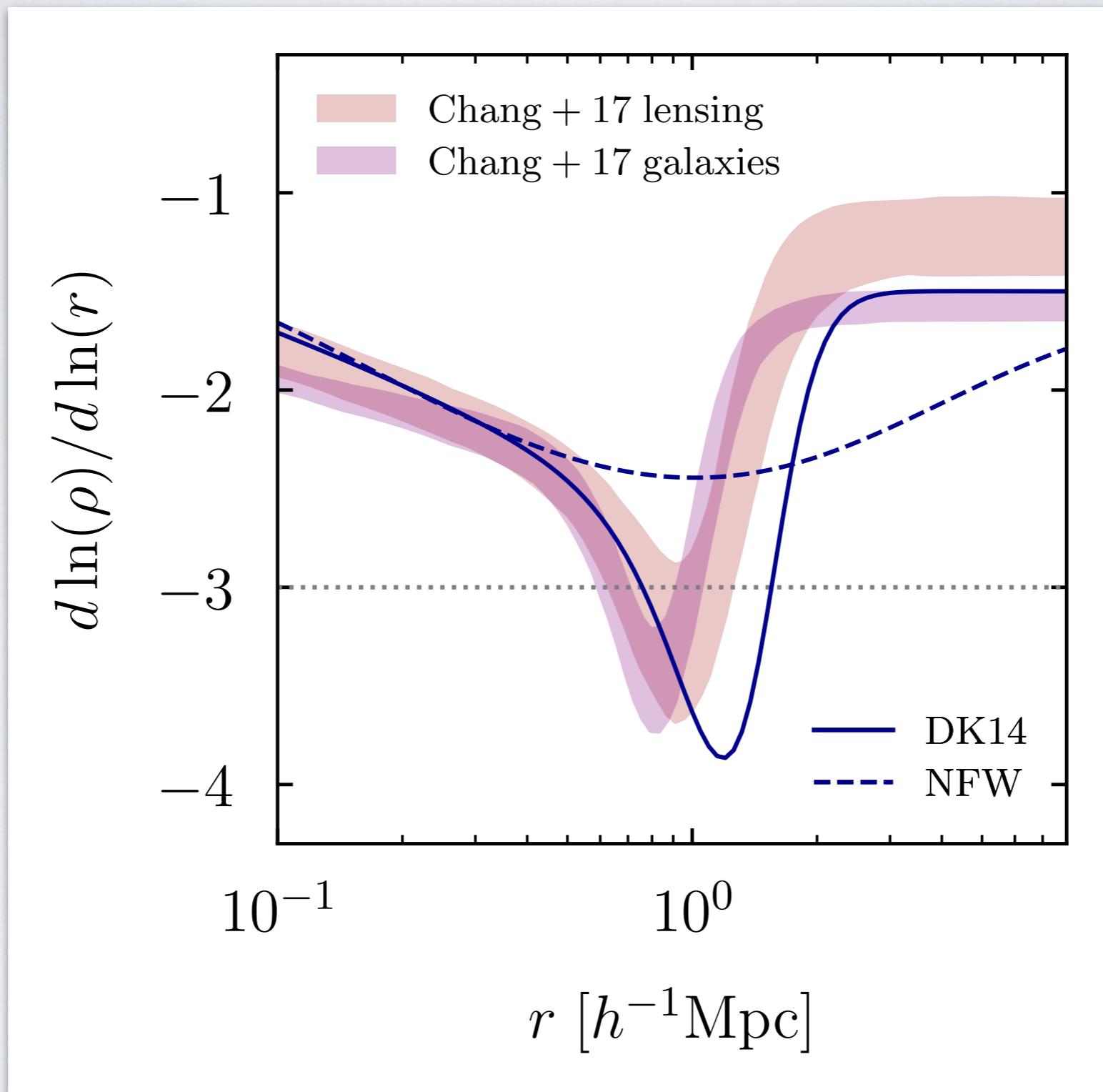


Assume $\Sigma_{\text{galaxies}} \sim \Sigma_{\text{subhalos}} \sim \Sigma_{\text{DM}}$

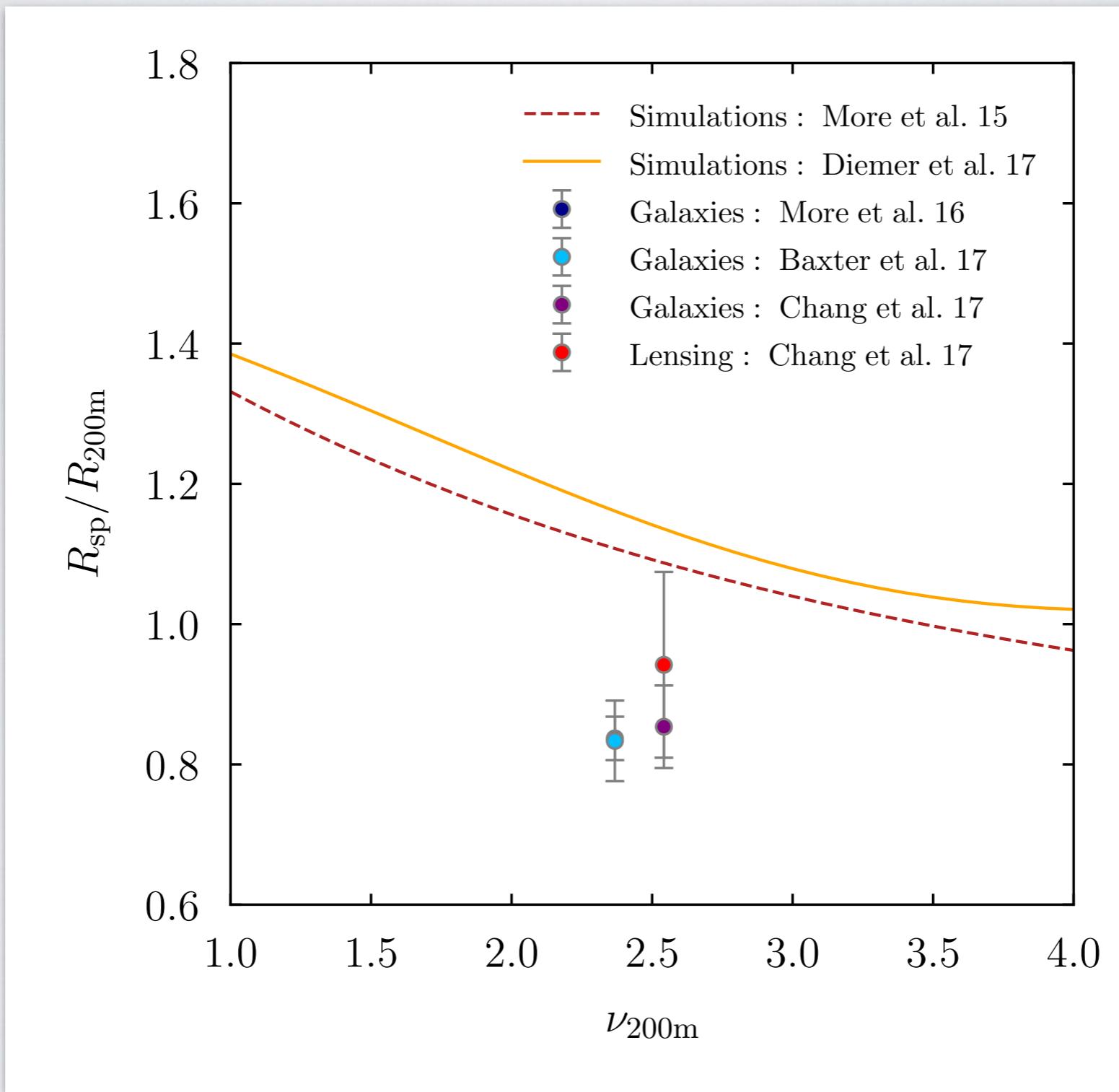
R_{sp} in cluster member profiles (DES)



R_{sp} in cluster member profiles (DES)



Summary of R_{sp} observations



More et al. 2015 • Diemer et al. 2017 • More et al. 2016 • Baxter et al. 2017 • Chang et al. 2017

Zu et al. 2017 • Busch & White 2017

R_{sp} of individual halos in simulations

1) Find drop in spherical density profiles

Too much sub-structure and noise

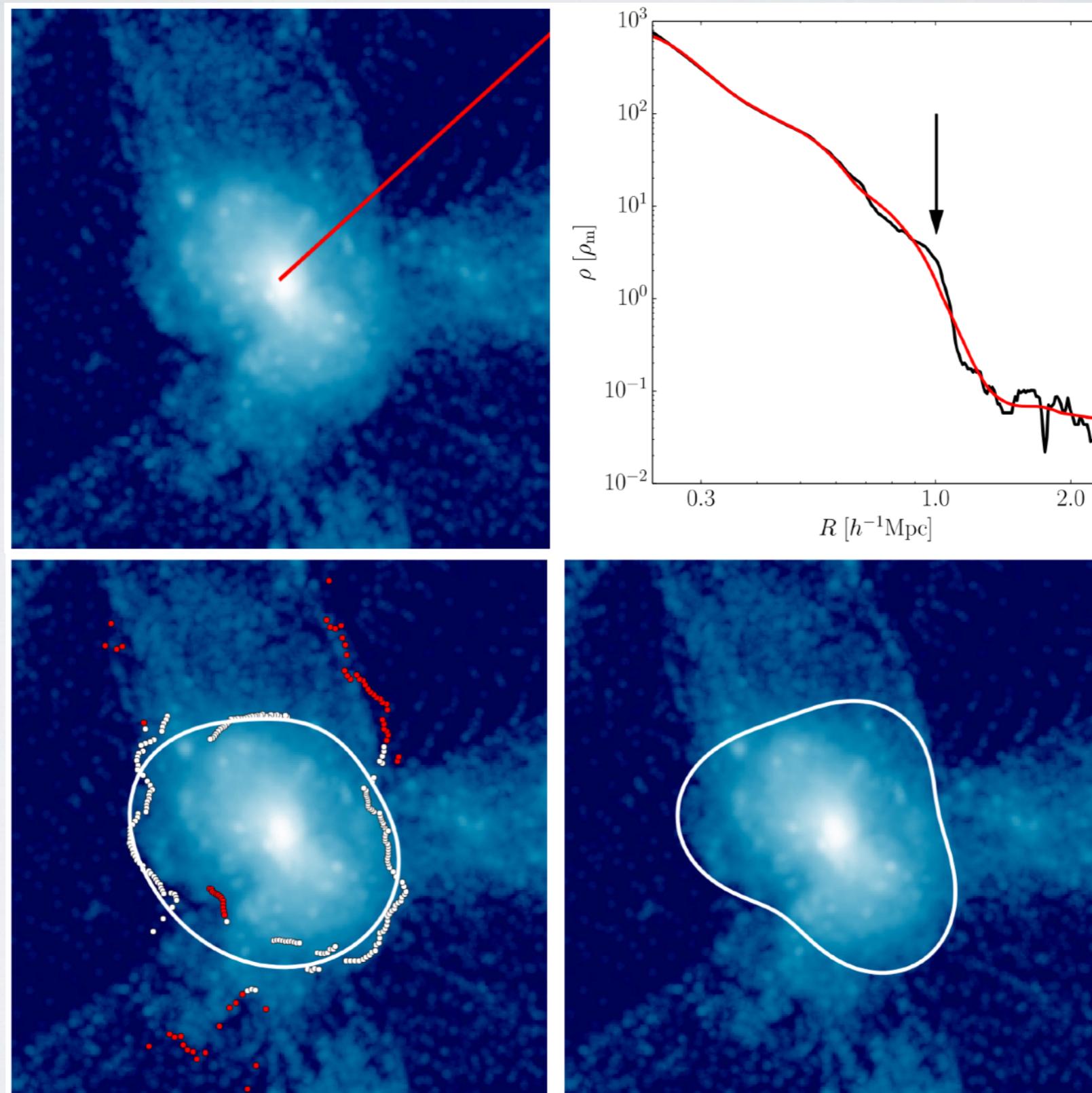
2) Consider the full 3D density information

Non-trivial, but can explore non-spherical effects

3) Measure the splashback in particle orbits

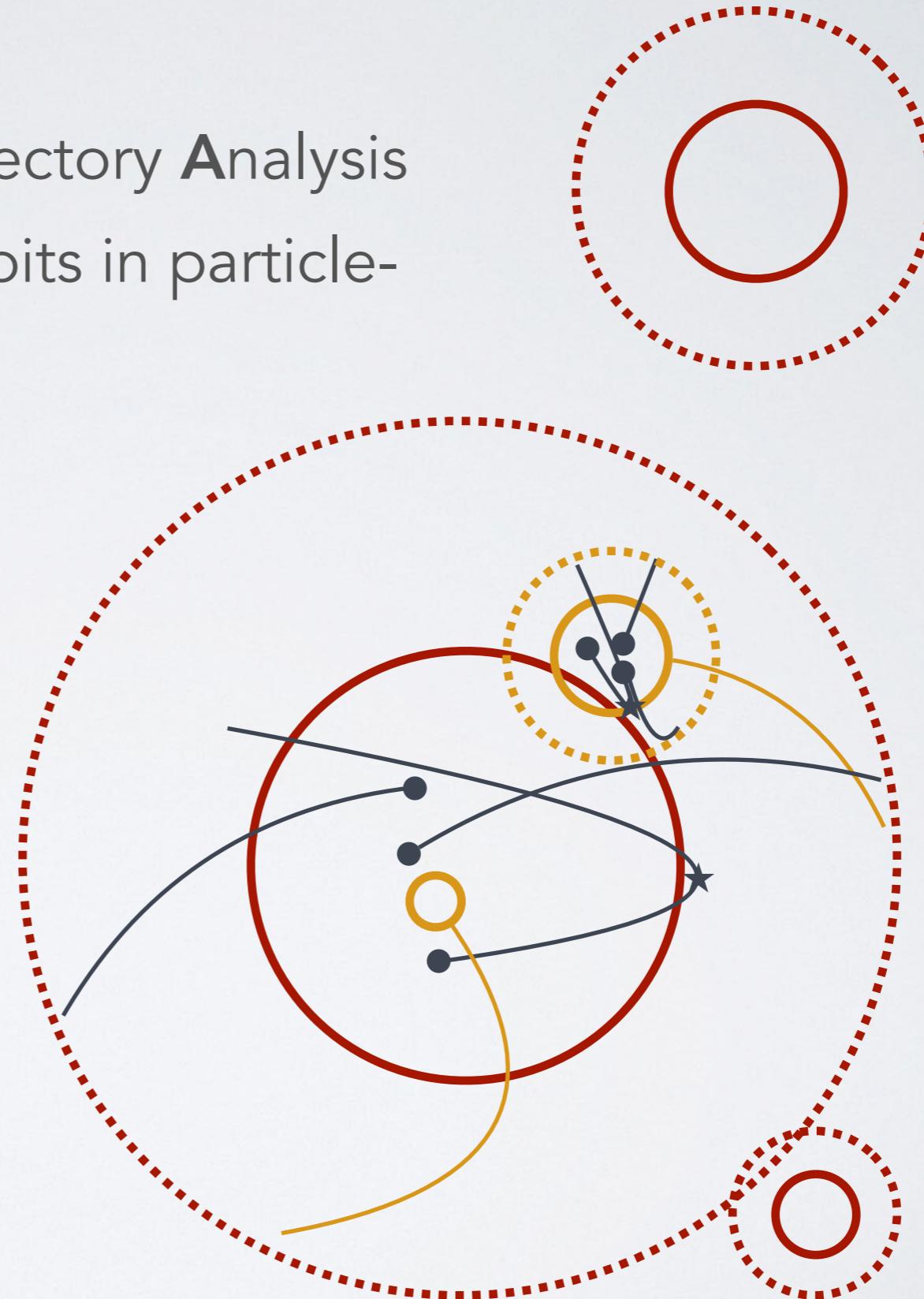
Sounds like a lot of work?

Option 2): Shell finding with SHELLFISH

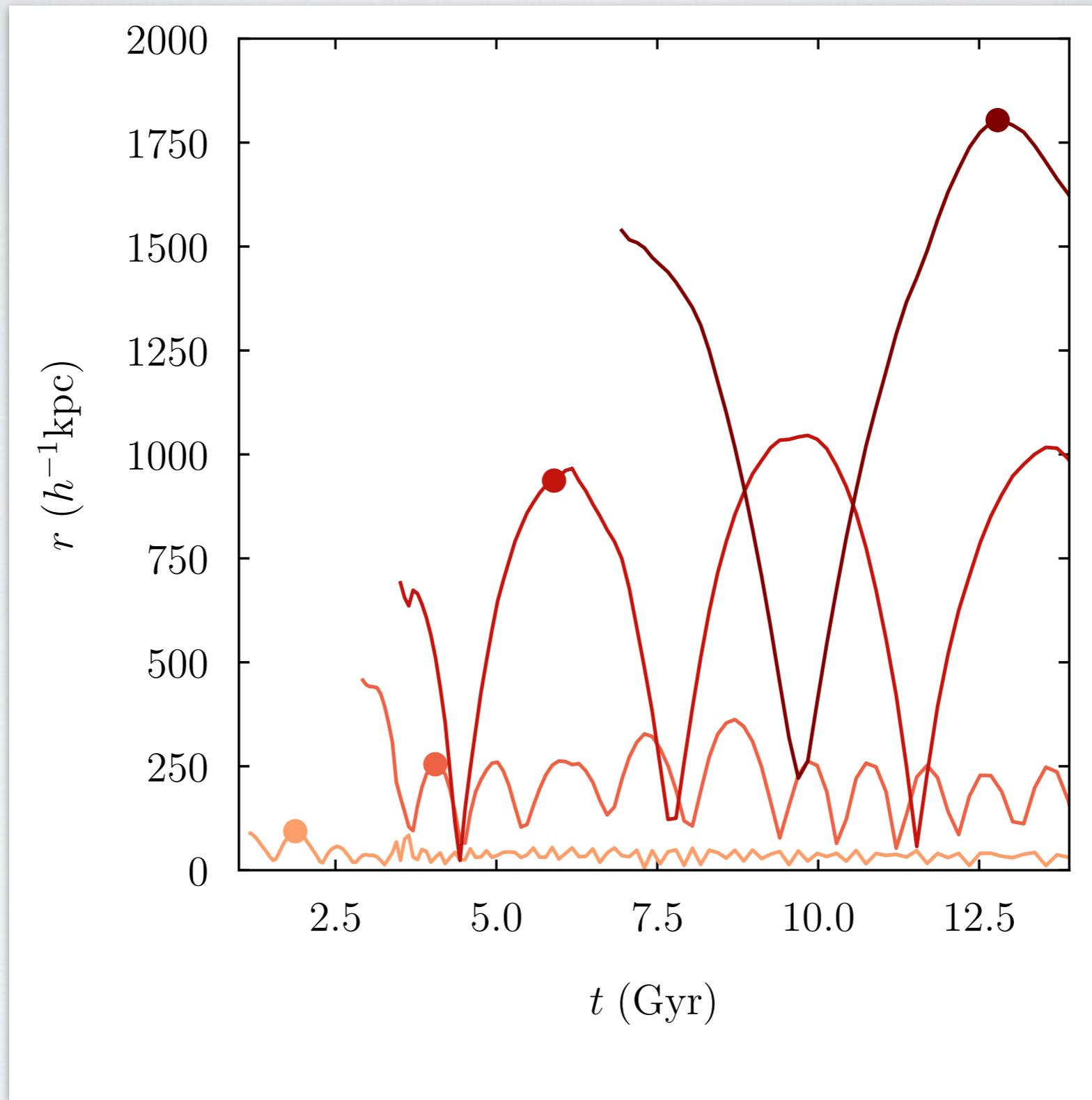


SPARTA

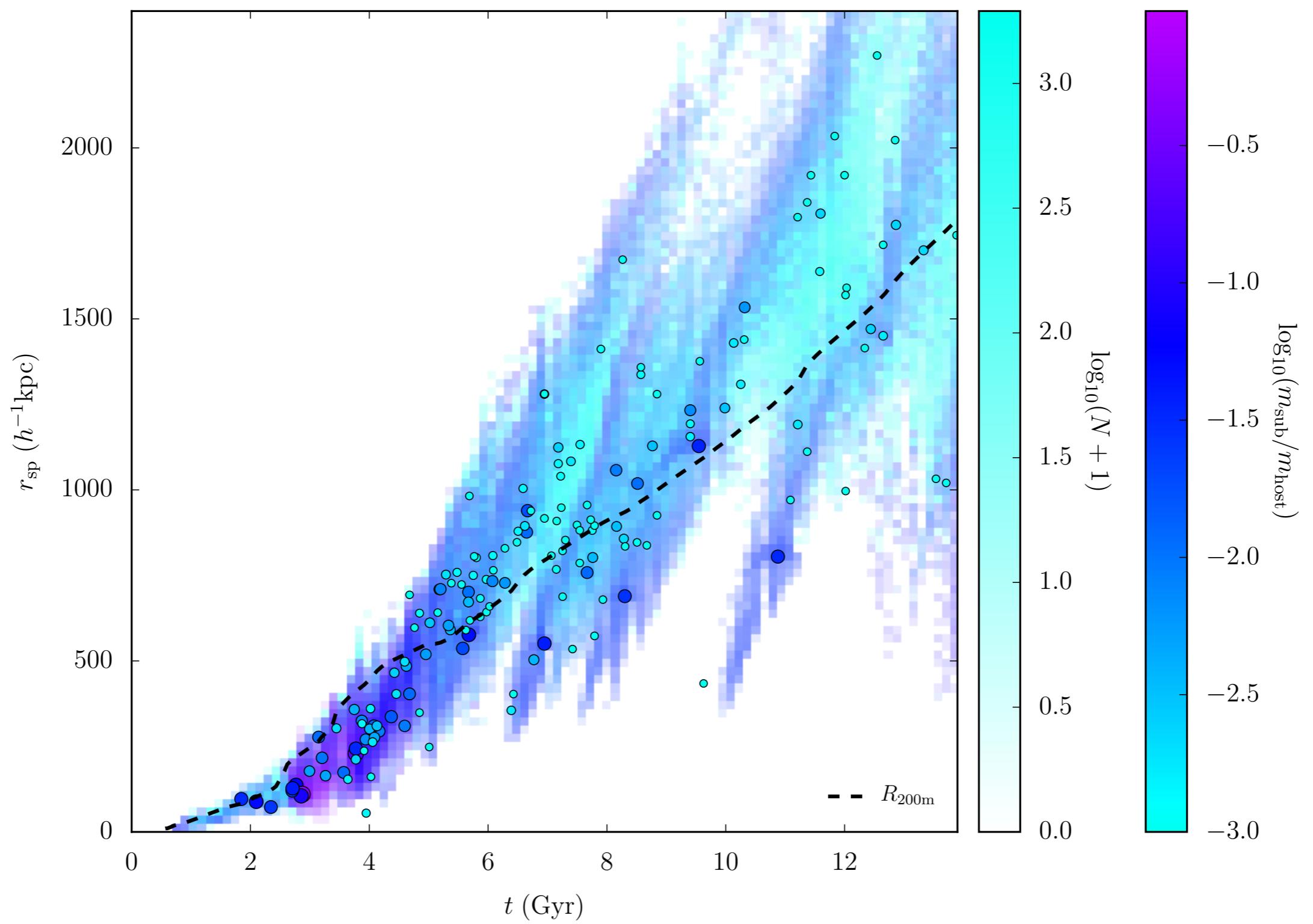
- Subhalo and PARticle Trajectory Analysis
- Framework for tracking orbits in particle-based simulations
- MPI-parallelized, pure C



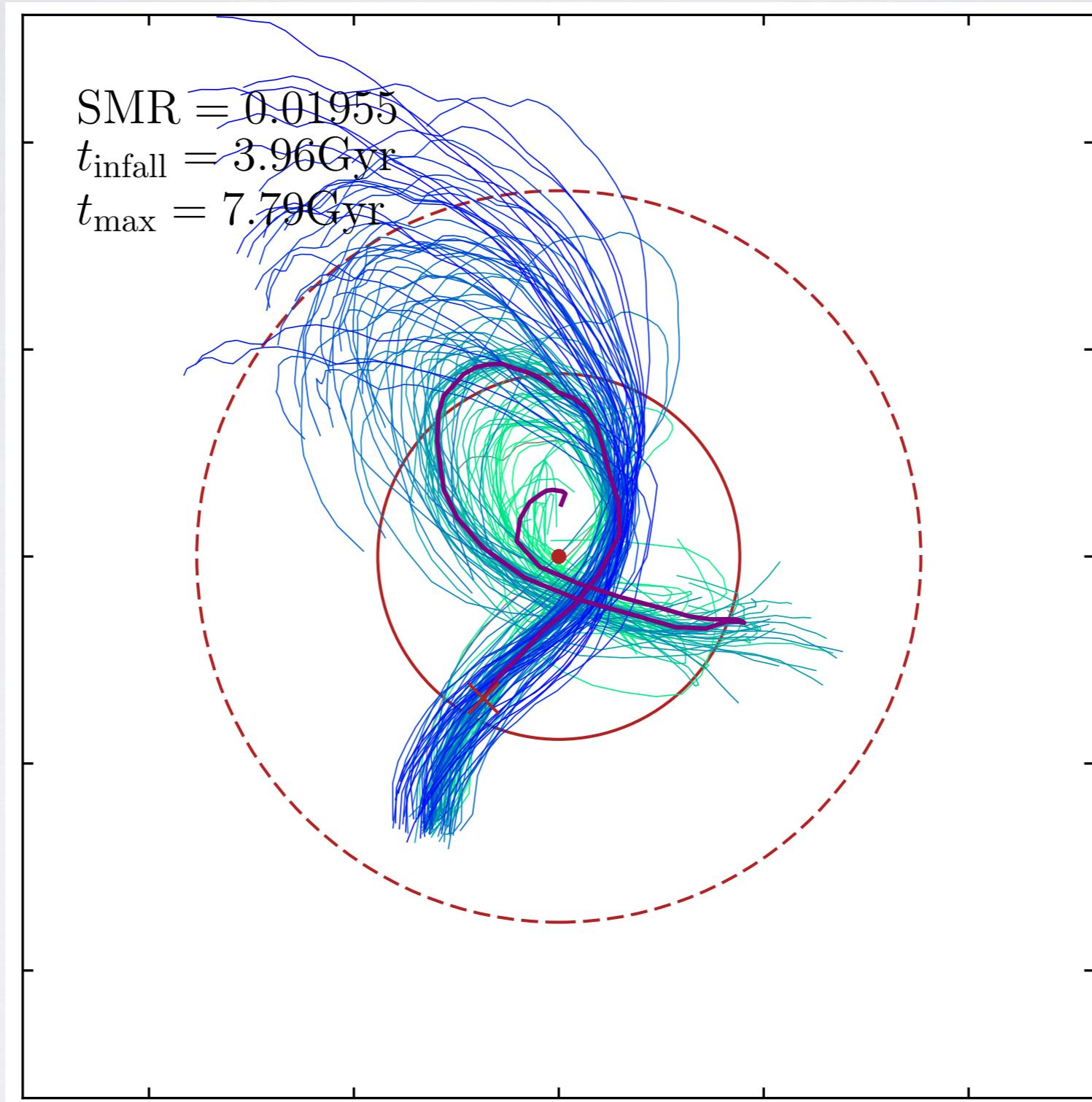
What do the orbits look like?



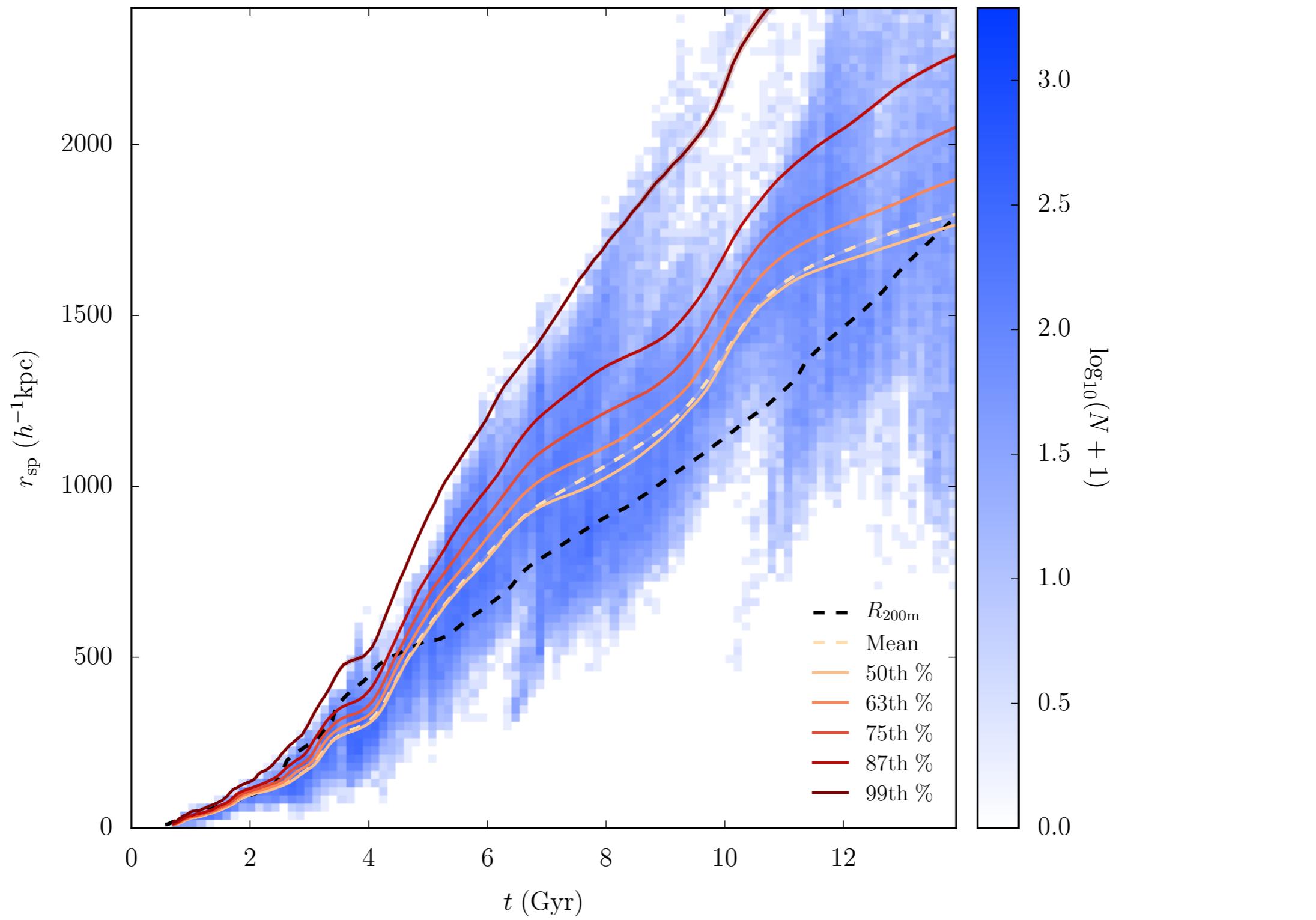
Splashback in an individual halo

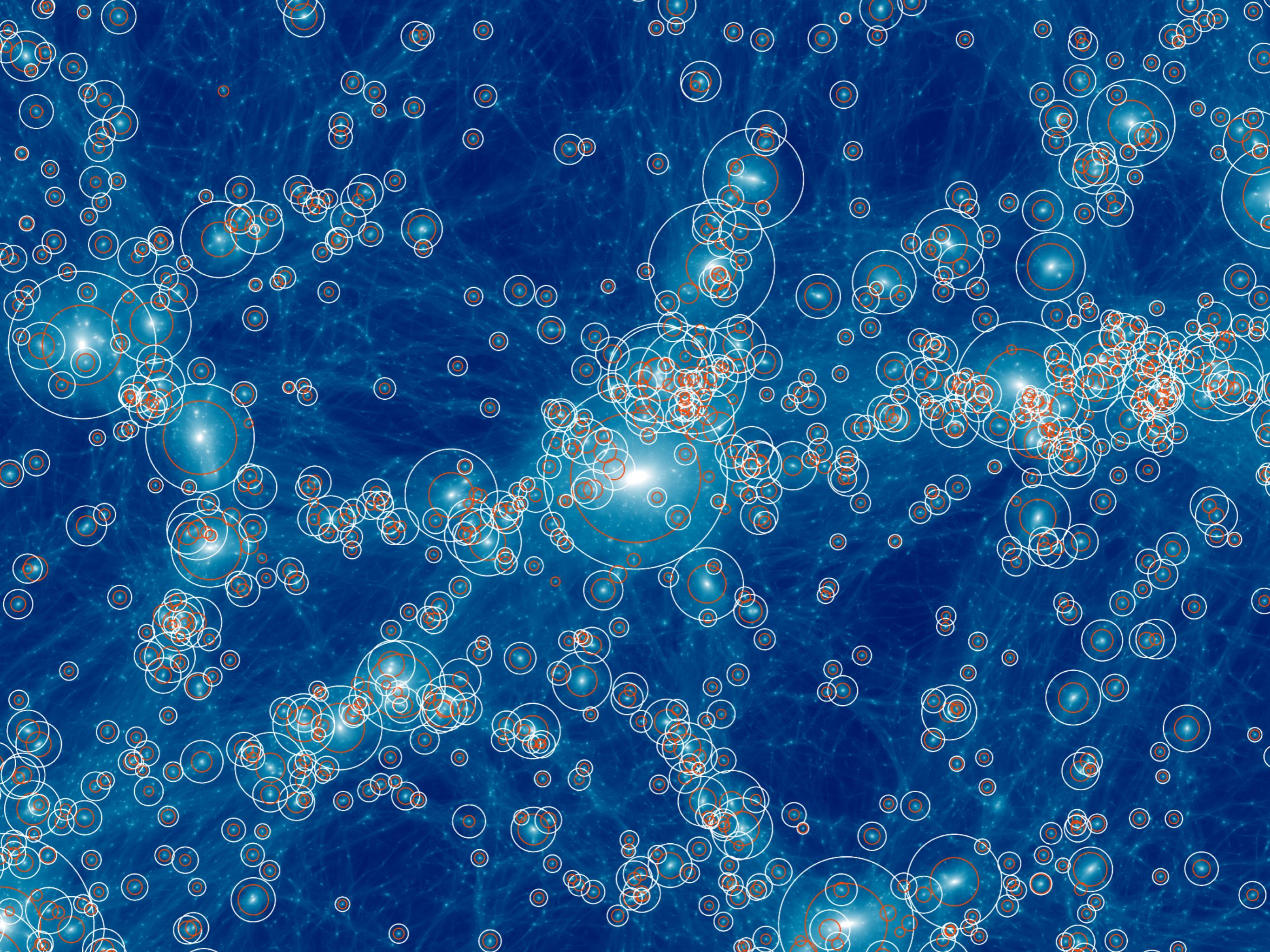


Subhalo disruption

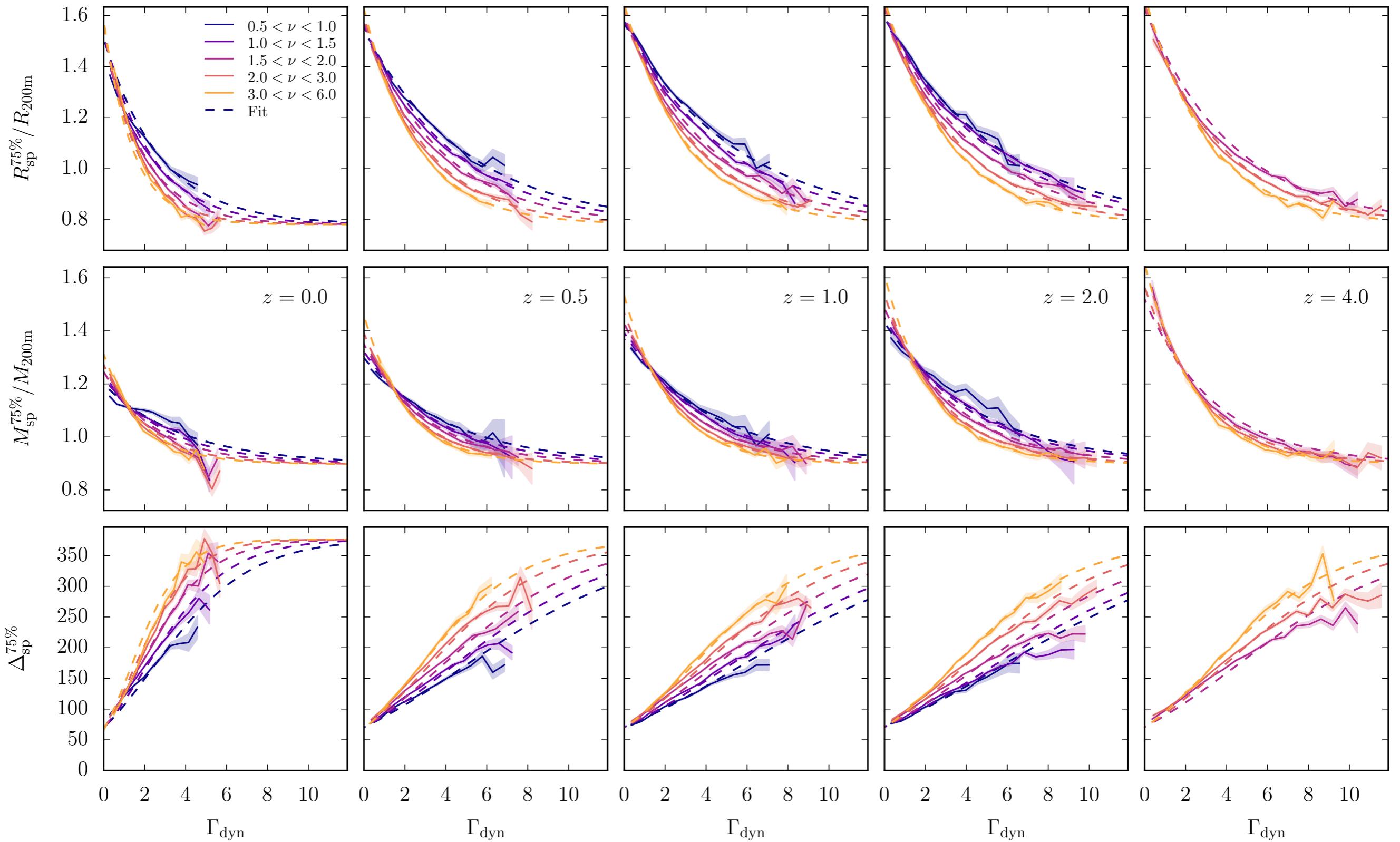


Splashback in an individual halo





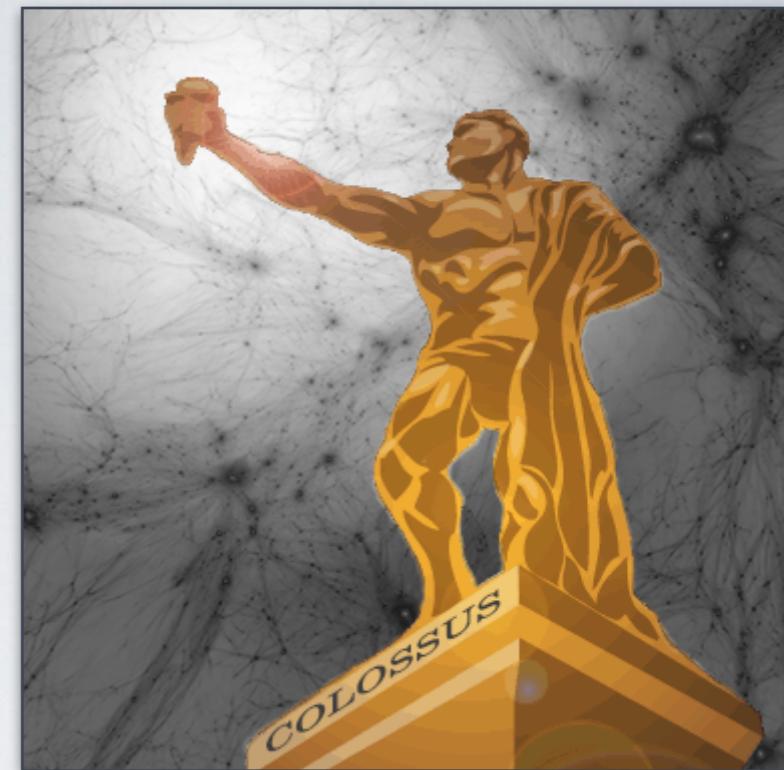
The Γ - R_{sp} relation



COLOSSUS

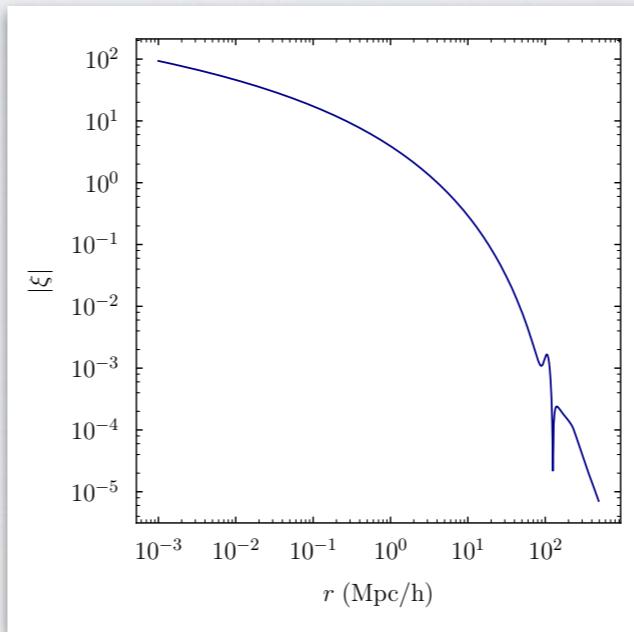
Cosmology, halos, and large-scale structure

benediktdiemer.com/code/colossus



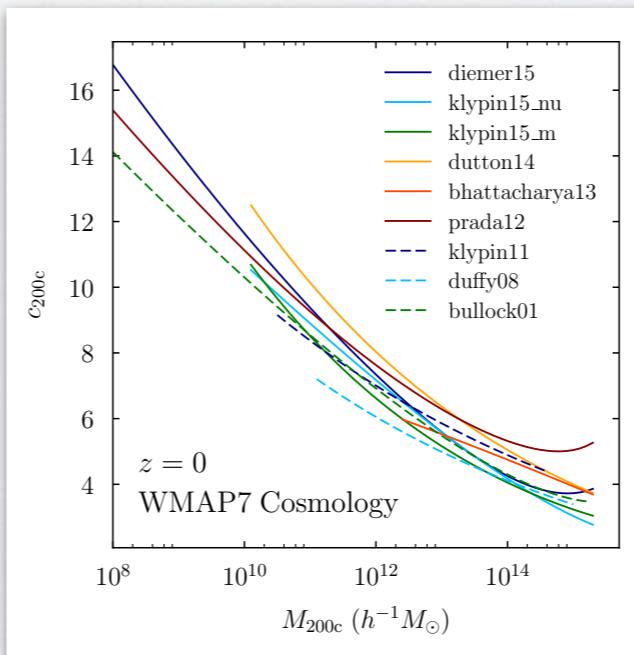
```
from colossus.cosmology import cosmology
```

```
cosmo = cosmology.setCosmology('WMAP9')
xi = cosmo.correlationFunction(10.0)
```



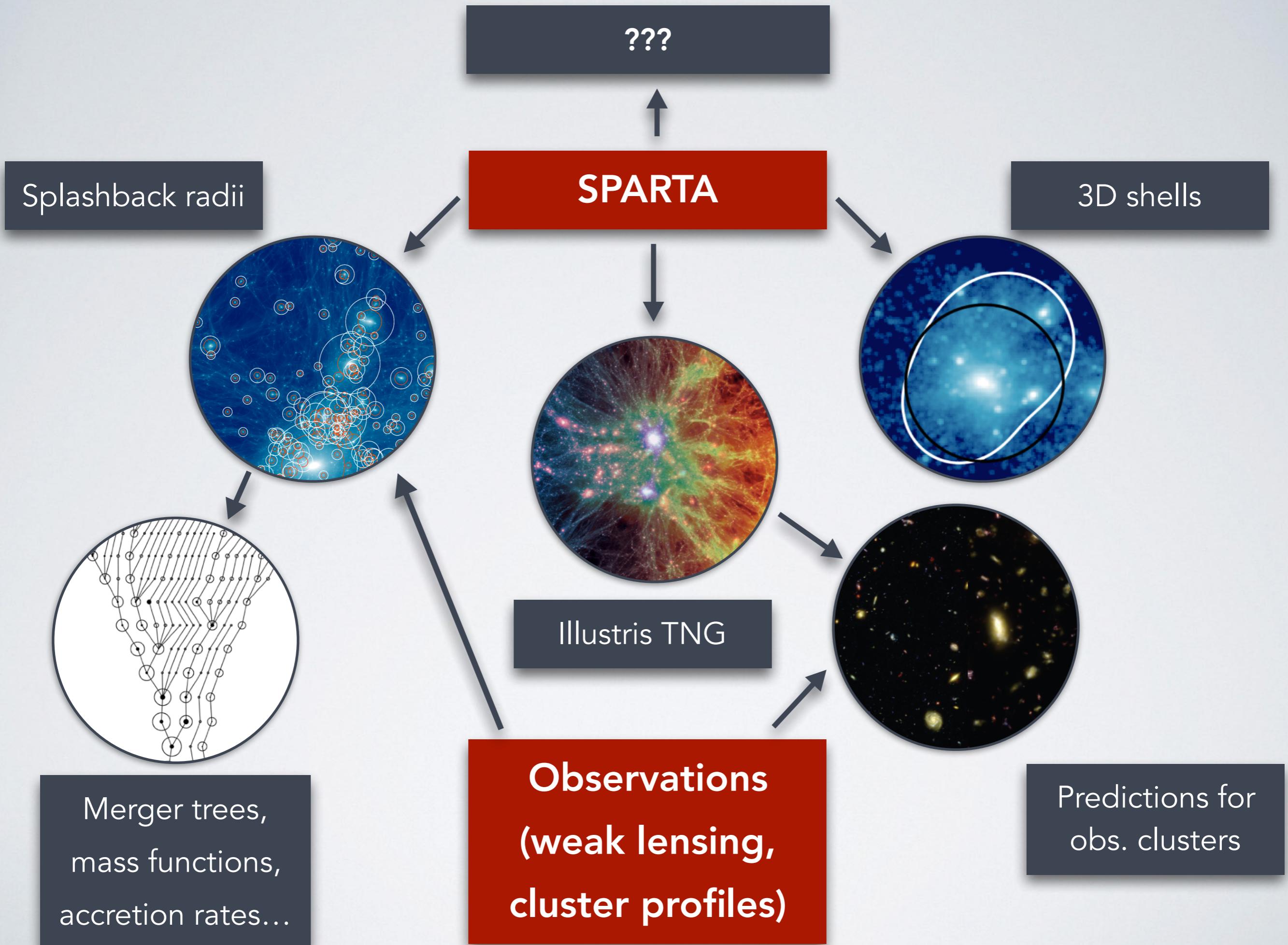
```
from colossus.halo import concentration as hc
```

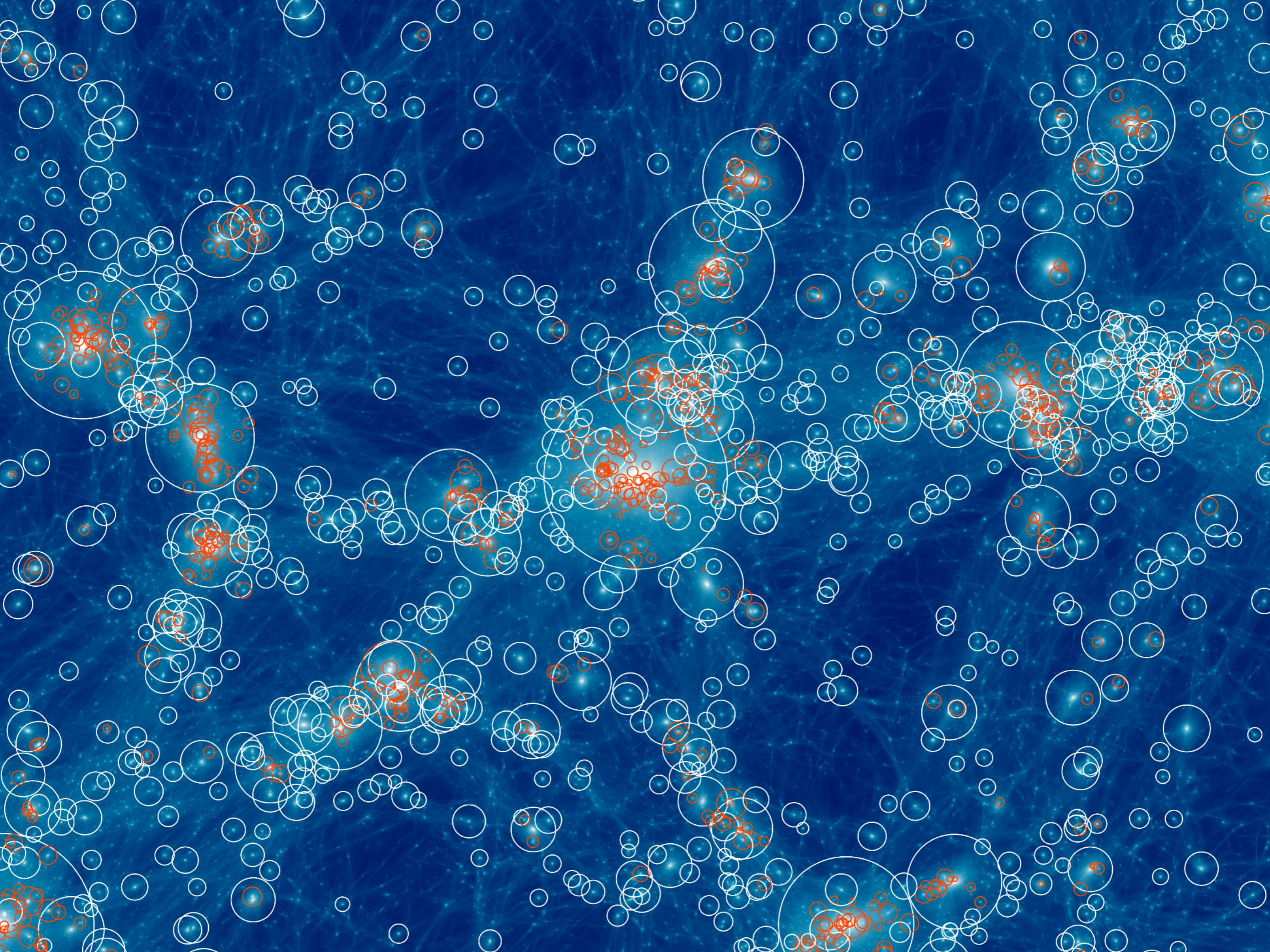
```
cosmology.setCosmology('bolshoi')
c = hc.concentration(1E12, 'vir', 0.0,
    model = 'bullock01')
```



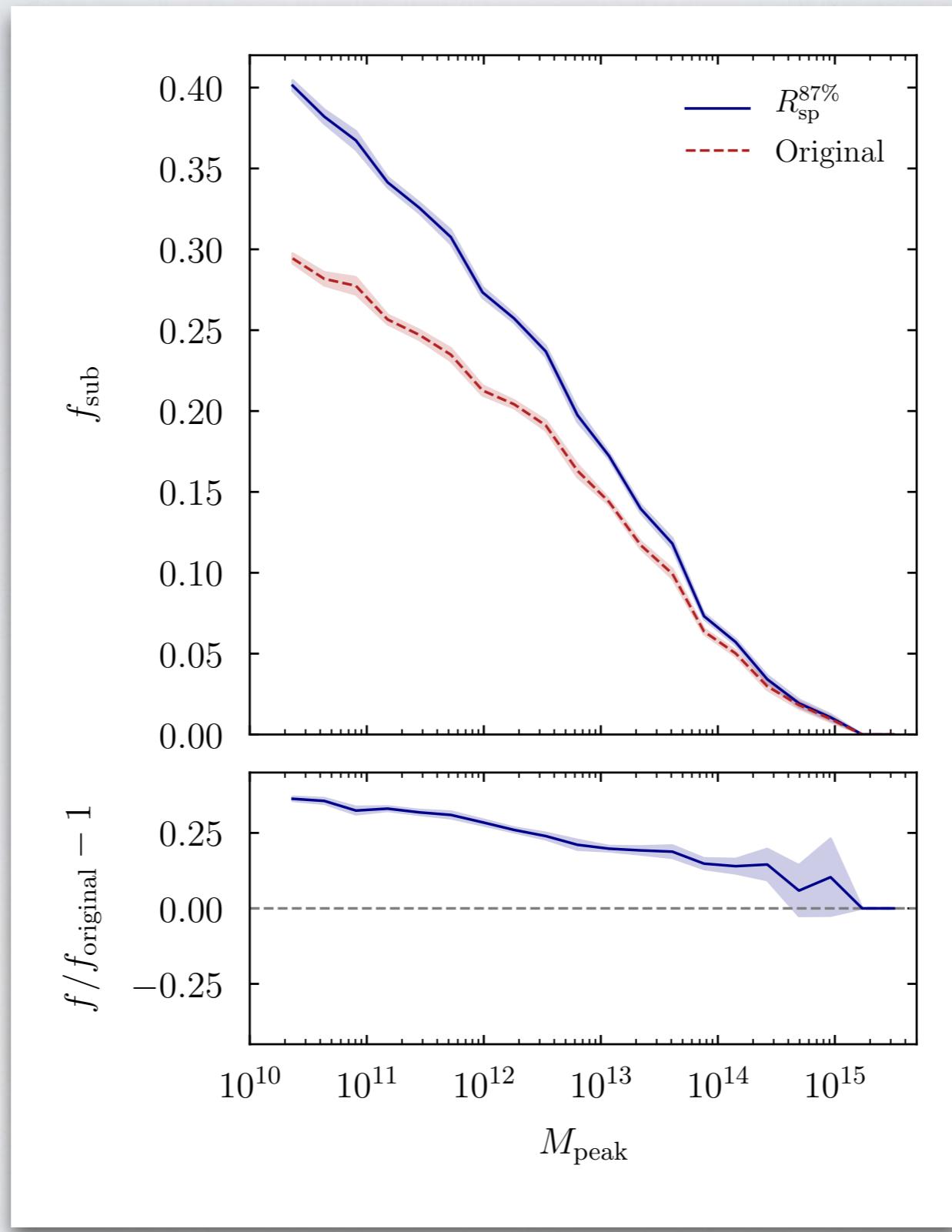
Modules:

- Cosmology
- Power spectrum
- Gaussian random peaks
- Halo mass function
- Bias
- Density profiles
- Halo mass definitions
- Concentration
- Splashback
- ...

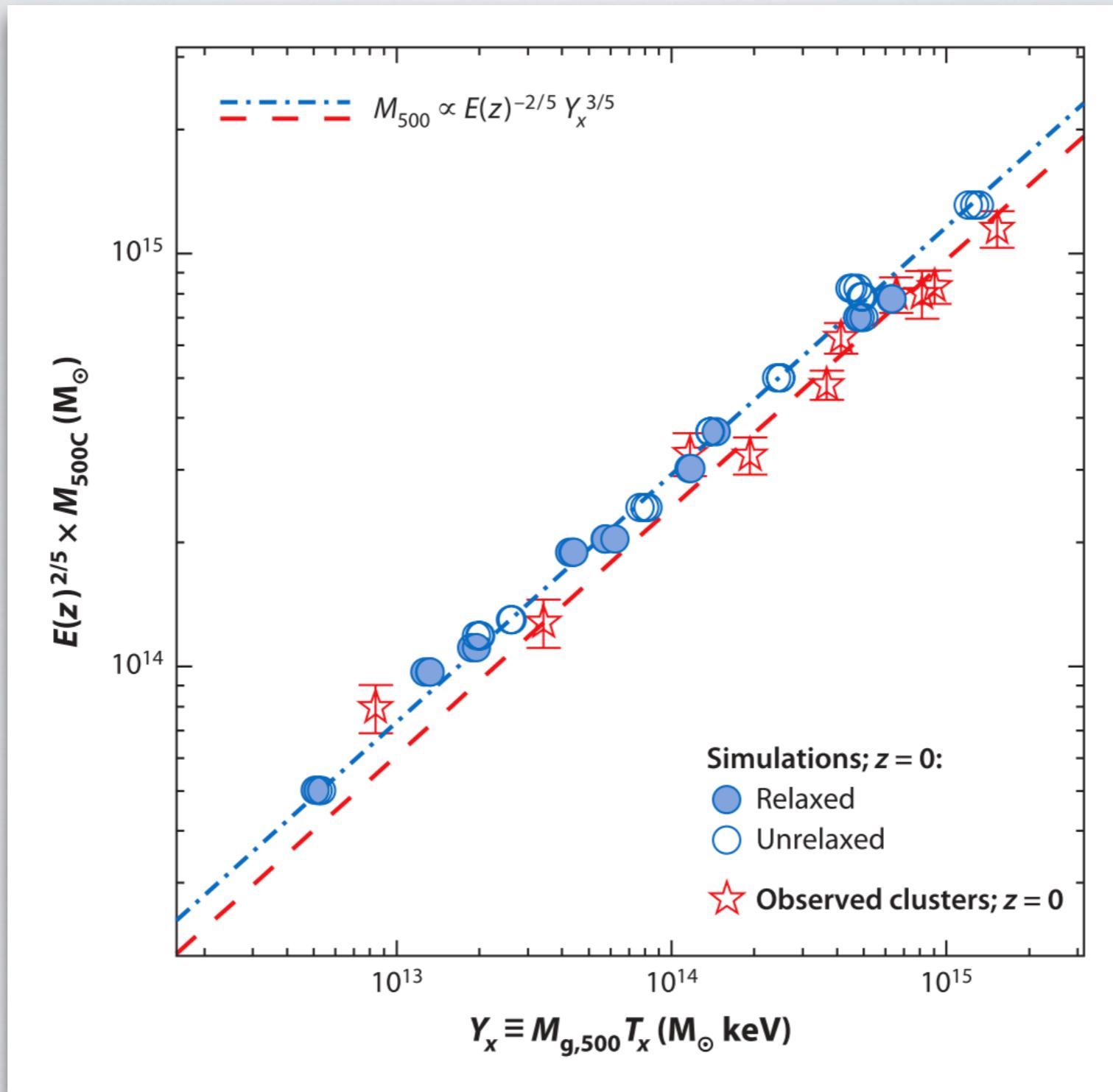




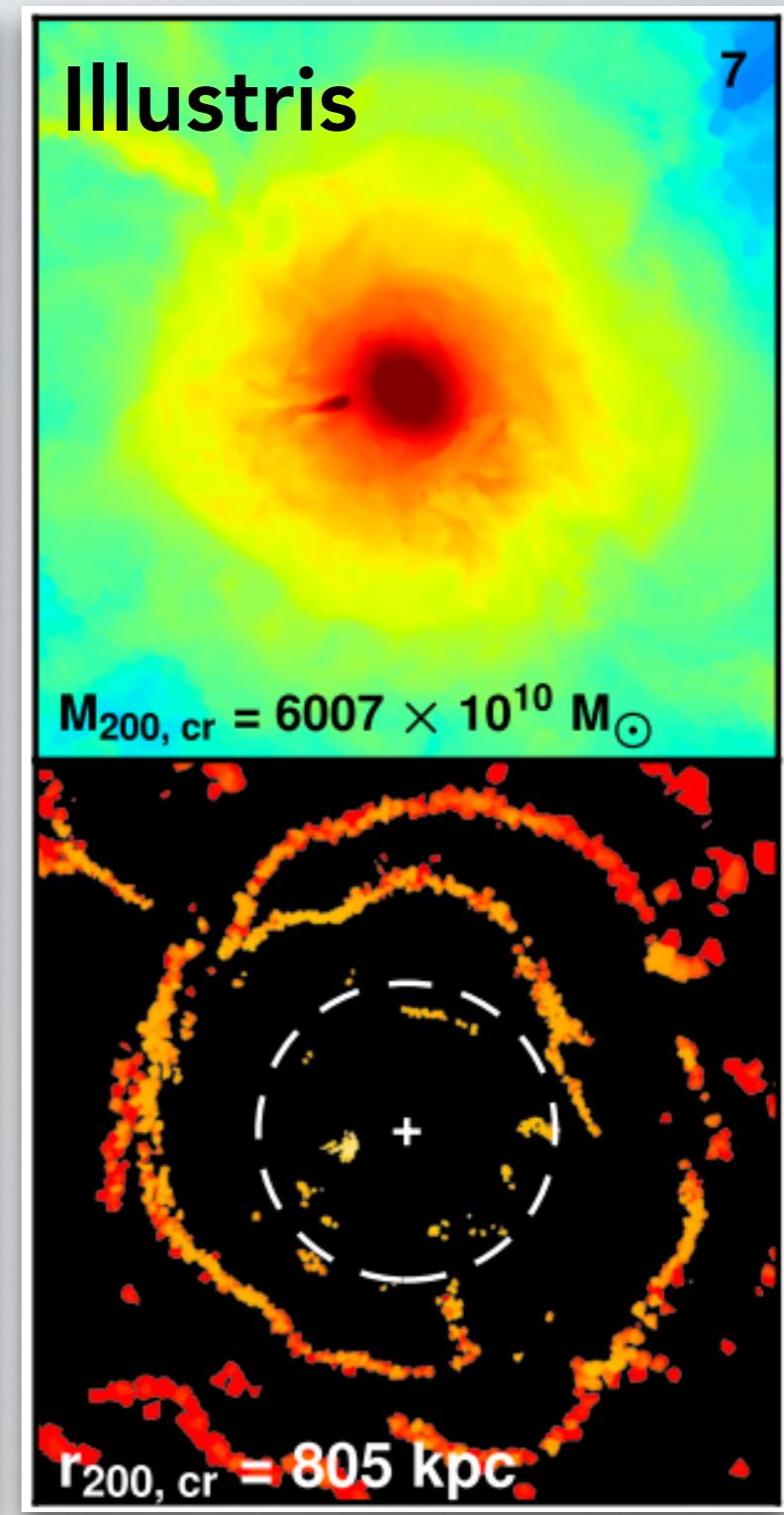
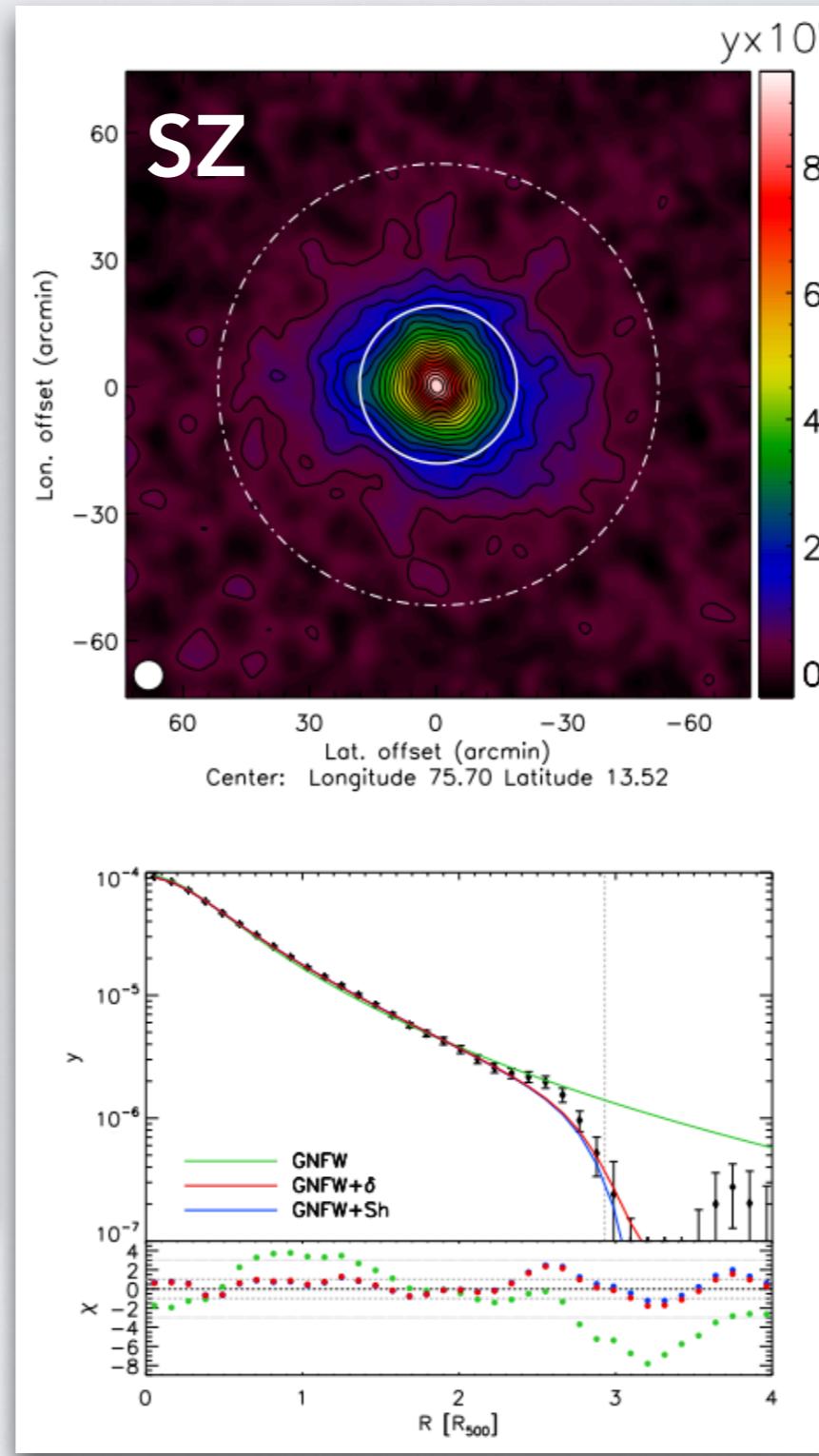
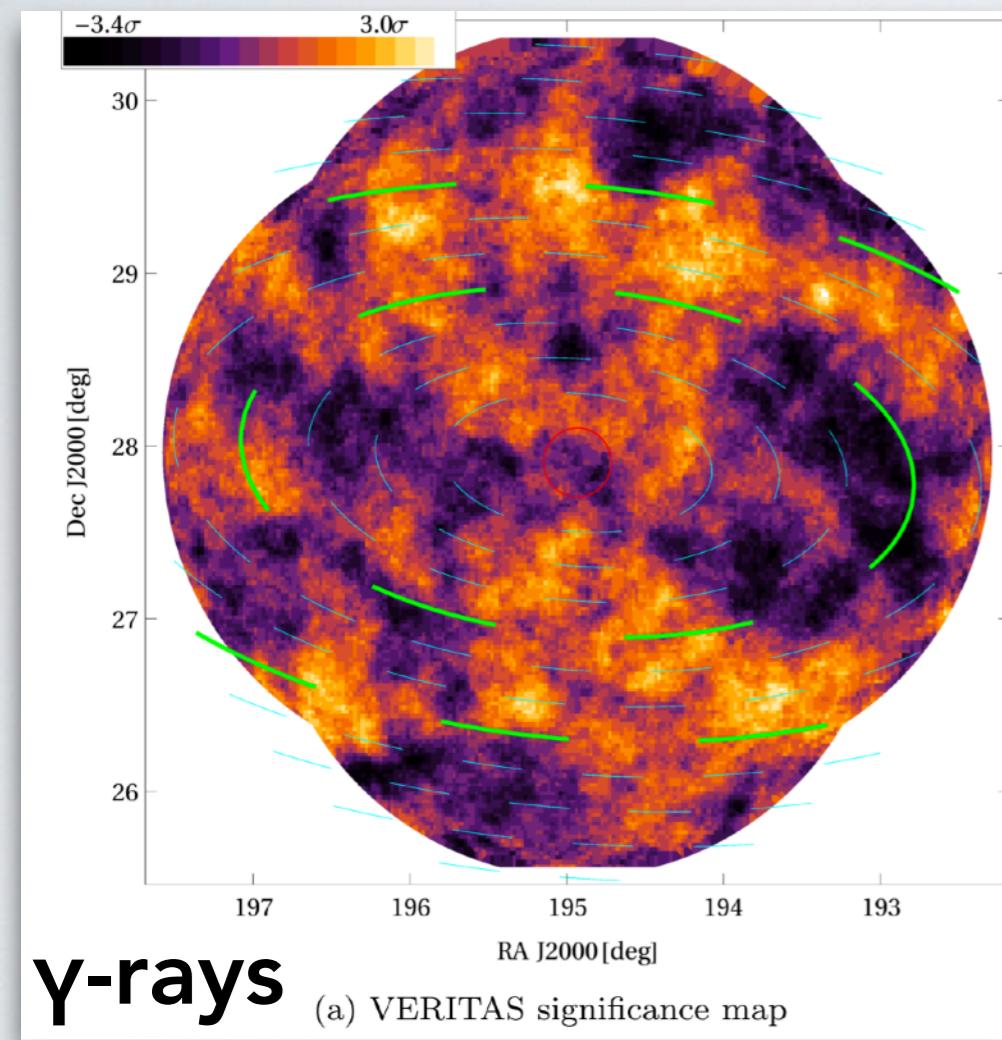
Subhalo fraction

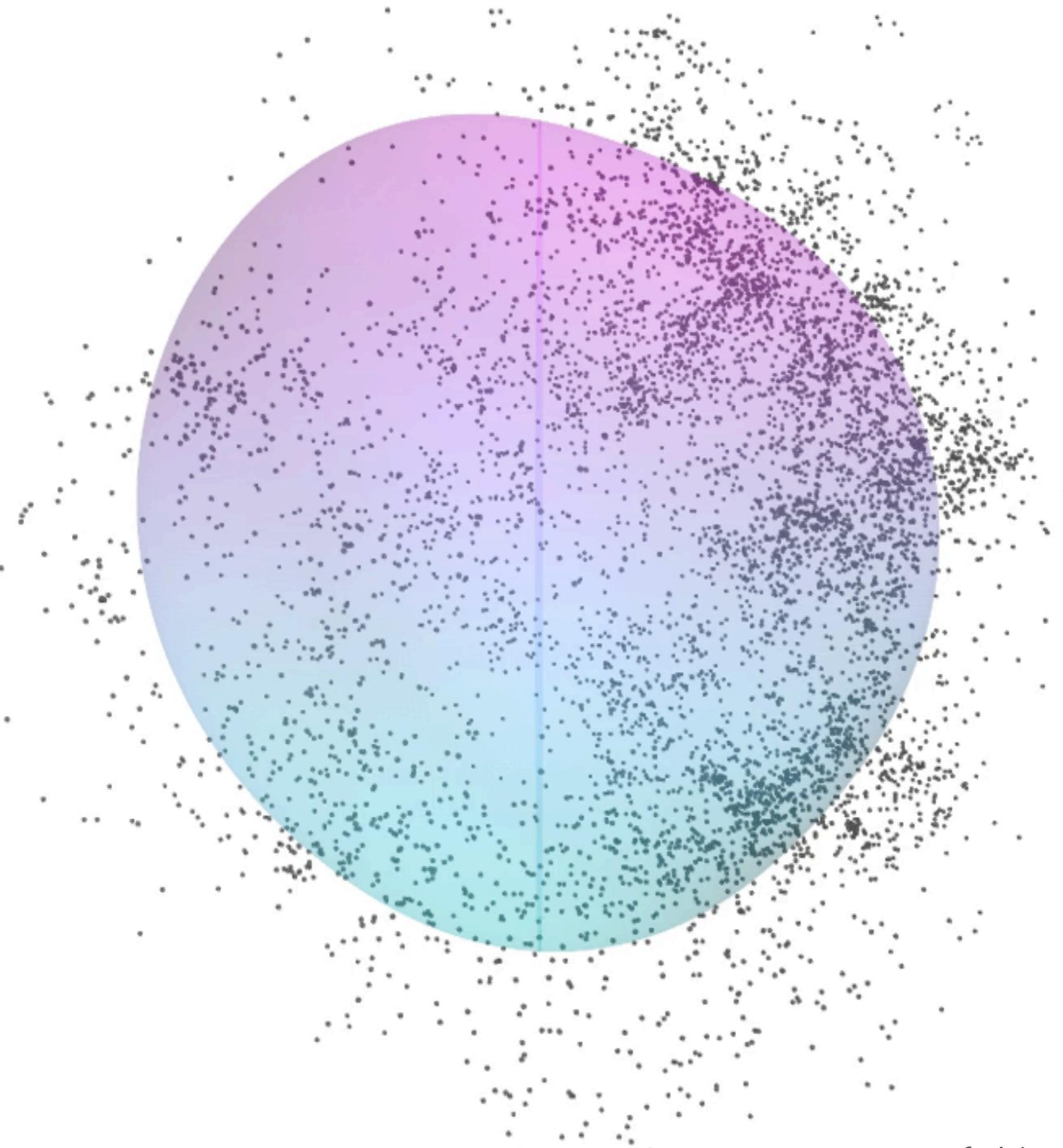


Scaling relations



A splashback-accretion shock relation?





Penna & Dines 2007 • Mansfield, Kravtsov & Diemer 2017



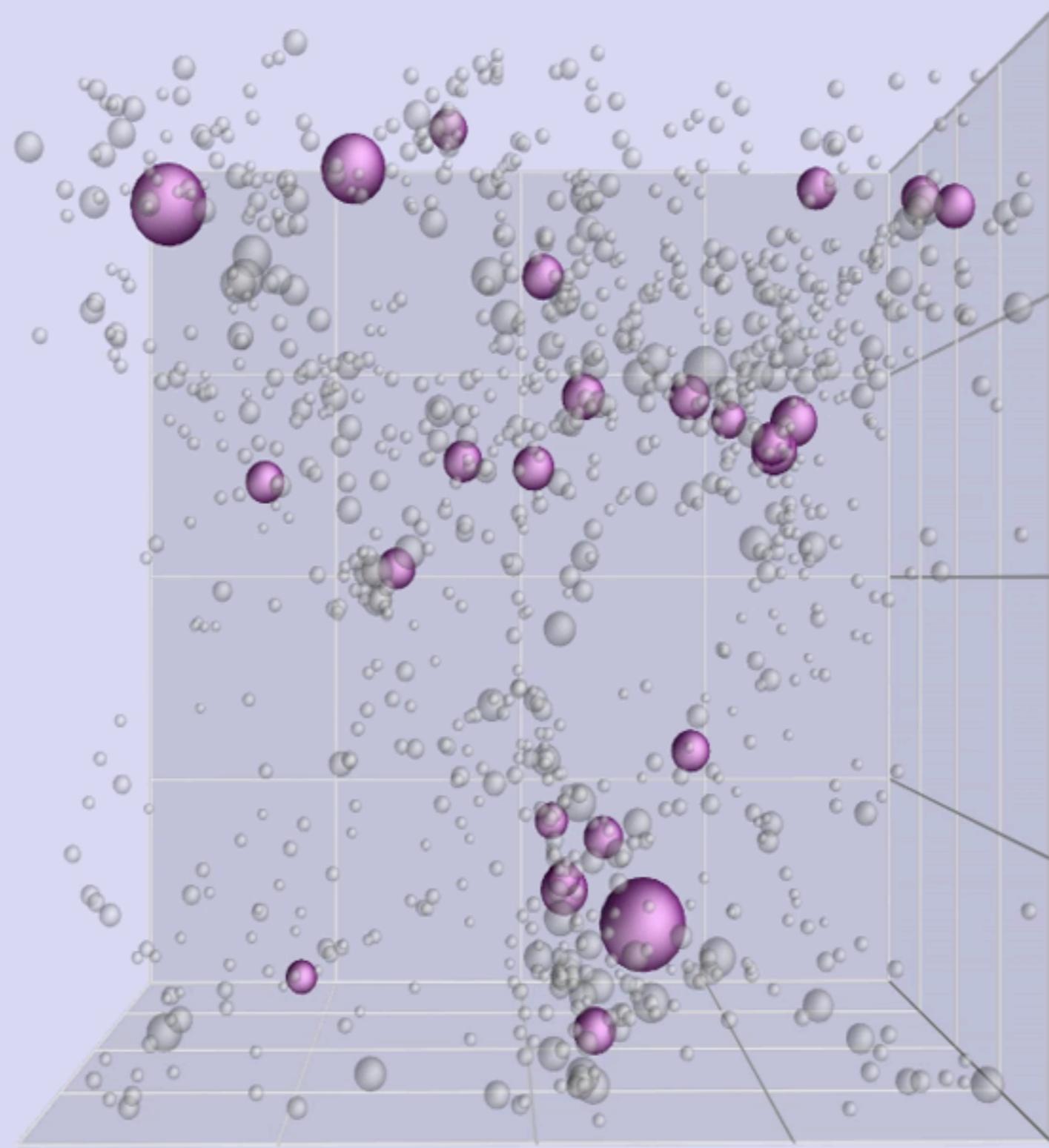
Diemer & Facio 2017 • The Fabric of the Universe



Diemer & Facio 2017 • The Fabric of the Universe

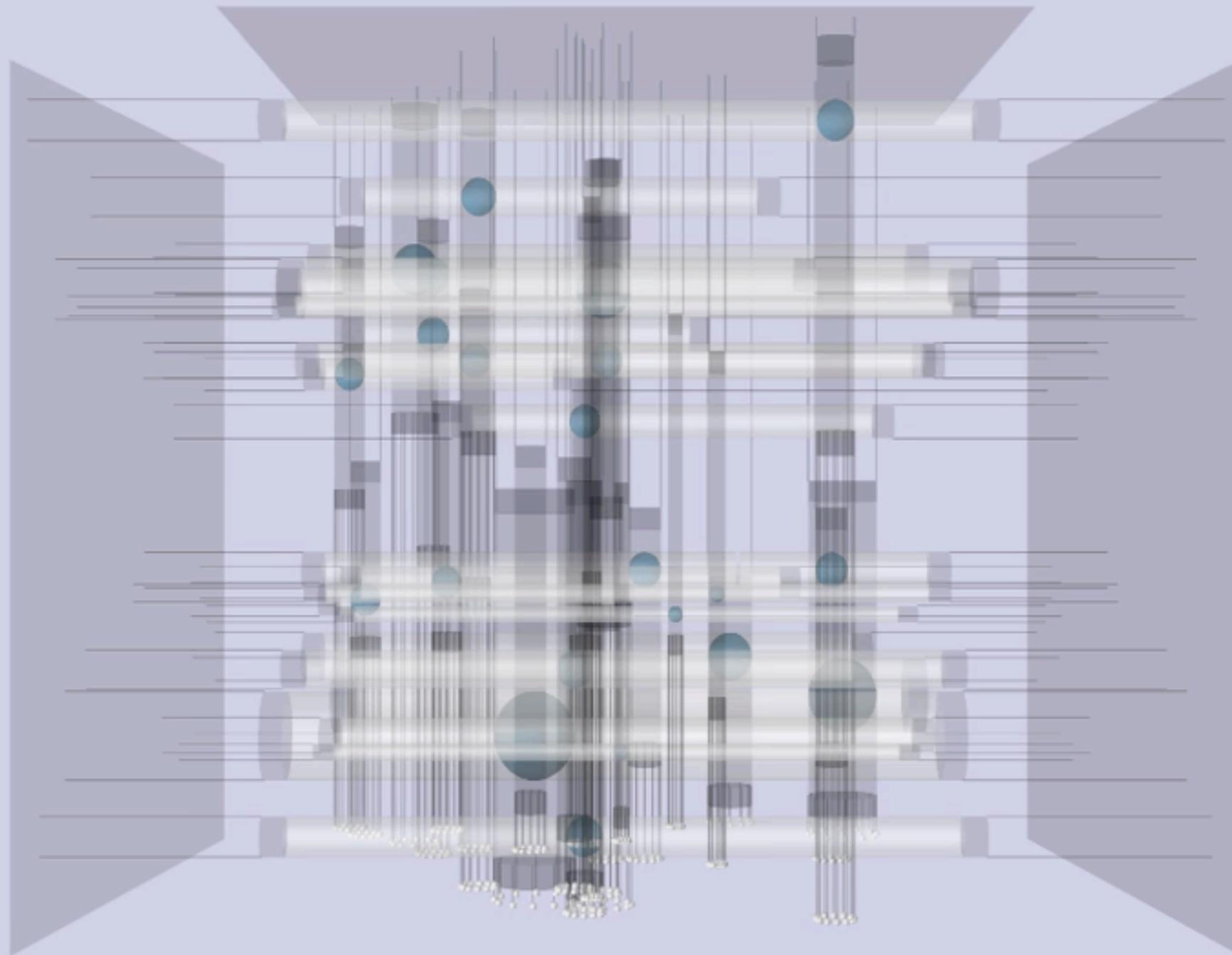


The Fabric of the Universe

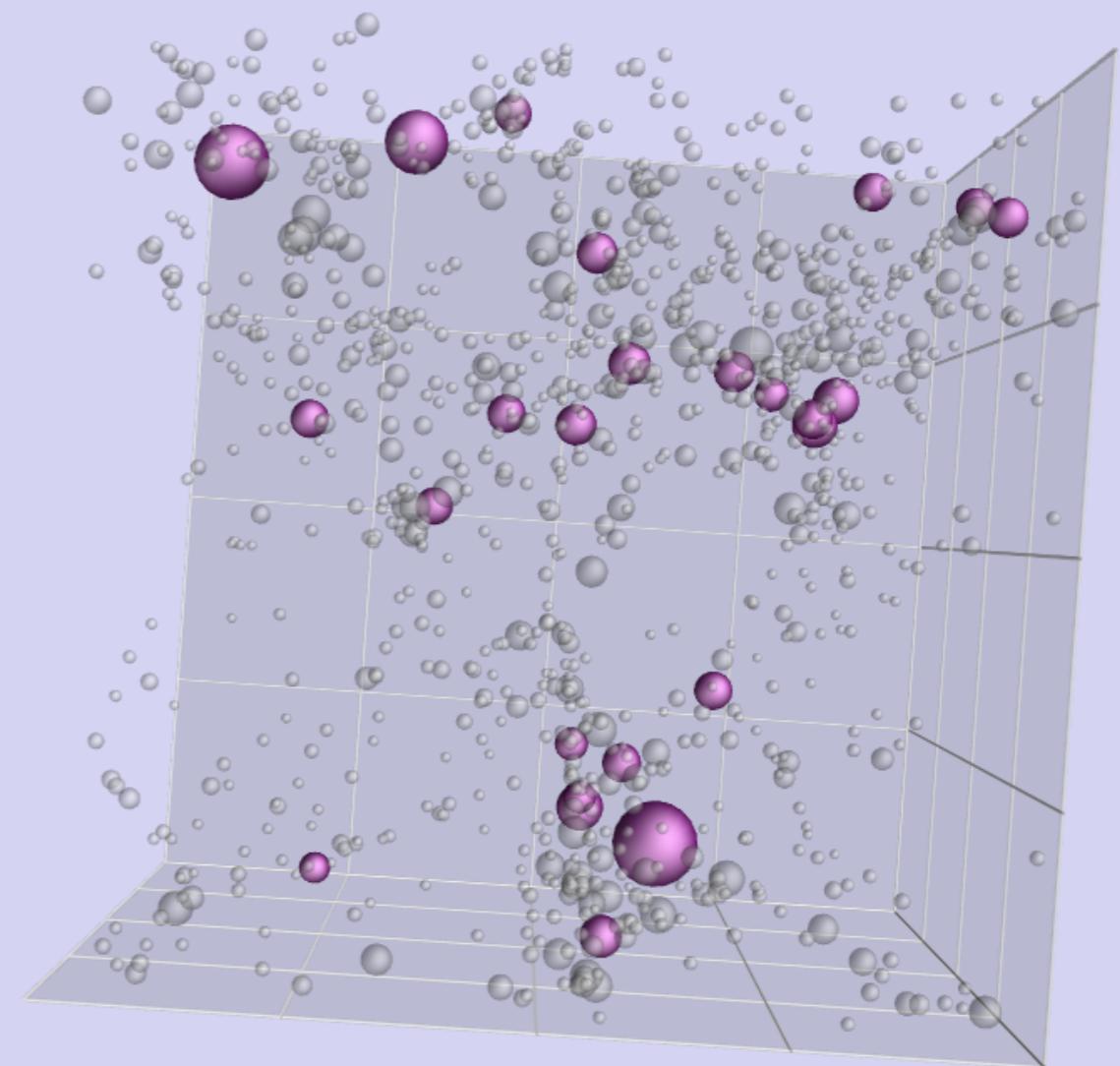




The Fabric of the Universe



The Fabric of the Universe



The Fabric of the Universe

Conclusions

- The **structure of CDM halos** is not a solved problem
- The **splashback radius** provides a physical halo boundary
- **SPARTA** and **SHELLFISH** provide entirely new ways of analyzing N-body simulations

Diemer & Kravtsov 2014 • ApJ 789, 1 • arXiv 1401.01216

More, Diemer & Kravtsov 2015 • ApJ 810, 36 • arXiv 1504.05591

Mansfield, Kravtsov & Diemer 2017 • arXiv 1612.01531

Diemer 2017 • arXiv 1703.09712

Diemer et al. 2017 • arXiv 1703.09716